# Analytical solution for the spectrum of two ultracold atoms in a completely anisotropic confinement

Yue Chen,<sup>1</sup> Da-Wu Xiao,<sup>2</sup> Ren Zhang<sup>0</sup>,<sup>3</sup> and Peng Zhang<sup>1,2,\*</sup>

<sup>1</sup>Department of Physics, Renmin University of China, Beijing 100872, China <sup>2</sup>Beijing Computational Science Research Center, Beijing 100193, China

<sup>3</sup>School of Science, Xi'an Jiaotong University, Xi'an 710049, China

(Received 17 February 2020; accepted 16 April 2020; published 14 May 2020)

We study the system of two ultracold atoms in a three-dimensional (3D) or two-dimensional (2D) completely anisotropic harmonic trap. We derive the transcendental equation  $J_{3D}(E) = 1/a_{3D}$  ( $J_{2D}(E) = \ln a_{2D}$ ) for the eigenenergy *E* of this system in the 3D (2D) case, with  $a_{3D}$  and  $a_{2D}$  being the corresponding *s*-wave scattering lengths, and provide the analytical expressions of the functions  $J_{3D}(E)$  and  $J_{2D}(E)$ . In previous research, this type of equation was obtained for spherically or axially symmetric harmonic traps [T. Busch *et al.*, Found. Phys. **28**, 549 (1998); Z. Idziaszek and T. Calarco, Phys. Rev. A **74**, 022712 (2006)]. However, for our cases with a completely anisotropic trap, only the equation for the ground-state energy of some cases has been derived [J.-J. Liang and C. Zhang, Phys. Scr. **77**, 025302 (2008)]. Our results in this work are applicable for arbitrary eigenenergy of this system and can be used for the study of dynamics and thermal dynamics of interacting ultracold atoms in this trap, e.g., the calculation of the second virial coefficient or the evolution of two-body wave functions. In addition, our approach for the derivation of the above equations can also be used for other two-body problems of ultracold atoms.

DOI: 10.1103/PhysRevA.101.053624

## I. INTRODUCTION

The two-body problems of trapped interacting ultracold atoms are basic problems in cold atom physics [1-13]. They are of broad interest for the following reasons. First, the two-body systems are "minimum" interacting systems of trapped ultracold atoms, and one can obtain a primary understanding for the interaction physics of an ultracold gas from the analysis of such systems [14]. Second, the solutions to these problems can be directly used to calculate some important few- or many-body quantities [15–25], e.g., the second virial coefficient which determines the high-temperature properties of the ultracold gases. Third, these systems have been already realized in many experiments [14,26–38] where the trap of the two atoms can be created via an optical lattice site [26,27,30], optical tweezer [31–35], or nanostructure [36–38].

In these experiments, by measuring or controlling the energy spectrum or dynamics of these two atoms, one can, e.g., create a cold molecule in an optical tweezer [32,33], derive the parameters of interatomic interaction potential [26,27,30,34], or study various dynamical effects such as the interactioninduced density oscillation [31]. Theoretical results for the corresponding two-body problems are very important for these experimental studies.

The most fundamental systems of two trapped ultracold atoms are the ones with a harmonic trap which has the same frequencies for each atom, and an *s*-wave short-range interatomic interaction. For these systems, the center of mass (c.m.) and relative motion of the two atoms can be decoupled with each other, and the c.m. motion is just the same as a harmonic oscillator. Nevertheless, the dynamics of the relative motion of these two atoms is nontrivial. For the threedimensional (3D) systems, Busch *et al*. [1] and Idziaszek and Calarco [2,3] studied the cases with a spherically and axially symmetric harmonic trap, respectively. They showed that the eigenenergy E of the relative motion satisfies a transcendental equation with the form

$$J_{\rm 3D}(E) = \frac{1}{a_{\rm 3D}},$$
 (1)

with  $a_{3D}$  being the 3D *s*-wave scattering length, and provided the analytical expressions of the function  $J_{3D}(E)$  for these two cases [1–3], as shown in Table I, respectively. Moreover, for the two-dimensional (2D) systems with an isotropic harmonic trap, Busch et al. obtained the equation

$$J_{2D}(E) = \ln a_{2D},$$
 (2)

for the relative-motion eigenenergy E, where  $a_{2D}$  is the 2D *s*-wave scattering length, and derived the expression of the function  $J_{2D}(E)$  [1]. With these results, one can easily obtain the complete energy spectrum of these two-atom systems, as well as the corresponding eigenstates. Thus, these results have been widely used in the research of cold atom physics [5,14,26,27,30,34,39–43].

However, for two atoms trapped in a 3D completely anisotropic harmonic confinement, the transcendental equation for the *arbitrary* eigenenergy of the relative motion has not been obtained so far. Only the equation of the *ground-state* energy of a system with positive scattering length has been obtained (i.e., the eigenenergy E which satisfies  $E < E_0$ , with  $E_0$  being the relative ground-state energy for the noninteracting case) in Ref. [4] by Liang and Zhang. This equation has the form of Eq. (1) and the corresponding function  $J_{3D}(E)$  is

<sup>&</sup>lt;sup>\*</sup>pengzhang@ruc.edu.cn

TABLE I. Expressions of the function  $J_{3D}(E)$  of Eq. (1) for various 3D traps. Here,  $\omega_{\alpha}$  ( $\alpha = x, y, z$ ) is the trapping frequency in the  $\alpha$  direction, and we use the natural unit  $\hbar = 2\mu = \omega_z = 1$  with  $\mu$  being the reduced mass of the two atoms. The parameters  $\eta_x$  and  $\eta_y$  are defined as  $\eta_x = \omega_x/\omega_z$  and  $\eta_y = \omega_y/\omega_z$ , respectively,  $E_0 = (\eta_x + \eta_y + 1)/2$  is the ground-state energy of the relative motion in the noninteracting case, and  $\Gamma(z)$  and  $\zeta(z)$  are the Gamma and Riemann zeta function, respectively. The two expressions (39) and (40) of  $J_{3D}(E)$  given by this work are mathematically equivalent with each other. (\*): There are typographical errors in Eqs. (21) and (23) of Ref. [3].

Туре	Ε	$J_{3D}(E)$	Ref.
$\omega_x = \omega_y = \omega_z$	arbitrary E	$\sqrt{2}rac{\Gamma\left(-rac{E}{2}+rac{3}{4} ight)}{\Gamma\left(-rac{E}{2}+rac{1}{4} ight)}$	Ref. [1]
$\omega_x = \omega_y \neq \omega_z$	arbitrary E	$-\frac{\eta_x}{\sqrt{2}}\sum_{n=0}^{\infty}\left[\frac{\Gamma\left(-\frac{E-E_0}{2}+n\eta_x\right)}{\Gamma\left(\frac{1}{2}-\frac{E-E_0}{2}+n\eta_x\right)}-\frac{1}{\sqrt{\eta_x}\sqrt{n+1}}\right]-\sqrt{\frac{\eta_x}{2}}\zeta\left(\frac{1}{2}\right)$	Ref. [3] (*)
$\omega_x \neq \omega_y \neq \omega_z$	$E < E_0$	$-\frac{1}{2\sqrt{\pi}}\int_0^\infty dt \left\{ \frac{\sqrt{\eta_x \eta_y} \exp[(E-E_0)t/2]}{\sqrt{1-e^{-\eta_y t}}\sqrt{1-e^{-\eta_y t}}\sqrt{1-e^{-t}}} - \frac{1}{t^{3/2}} \right\}$	Ref. [4]
$\omega_x \neq \omega_y \neq \omega_z$	arbitrary E	Eq. (39) or Eq. (40)	this work

also shown in Table I. On the other hand, since these kinds of confinements are easy to experimentally prepare and are thus used in many experiments of ultracold gases [26–30,34,43], it would be helpful if we could derive a general equation satisfied by all of the eigenenergies.

In this work, we derive the equation for the arbitrary eigenenergy of two atoms in a completely anisotropic harmonic trap. We show that this equation also has the form of Eq. (1) and Eq. (2) for the 3D and 2D cases, respectively, and provide the corresponding analytical expressions of the functions  $J_{3D}(E)$  (Table I) and  $J_{2D}(E)$  [Eq. (60) or Eq. (62)]. As with the aforementioned results [1–3] for the spherically and axially symmetric traps, our equations are useful for the studies of various thermal-dynamical or dynamical properties of ultracold gases in completely anisotropic harmonic confinements. Furthermore, our calculation approach used in this work can also be generalized to other two-body problems of ultracold atoms in complicated confinements.

The remainder of this paper is organized as follows. In Secs. II and III, we derive the equations for eigenenergies of atoms in 3D and 2D completely anisotropic harmonic traps, respectively. In Sec. IV, we discuss how to generalize our approach to other problems. A brief summary and some discussions are given in Sec. V. Some details of our calculations are shown in the appendices.

## **II. 3D SYSTEMS**

We consider two ultracold atoms 1 and 2 in a 3D completely anisotropic harmonic trap which has the same frequencies for each atom. Here we denote  $\omega_{\alpha}$  ( $\alpha = x, y, z$ ) as the trapping frequency in the  $\alpha$  direction, which satisfies

$$\omega_x \neq \omega_y \neq \omega_z. \tag{3}$$

For convenience, in this work we use the natural unit

$$\hbar = 2\mu = \omega_z = 1, \tag{4}$$

with  $\mu$  being the reduced mass of the two atoms. We further introduce the aspect ratios

$$\eta_x = \frac{\omega_x}{\omega_z}, \quad \eta_y = \frac{\omega_y}{\omega_z}, \quad \eta_z = \frac{\omega_z}{\omega_z} = 1,$$
 (5)

where  $\eta_x$  and  $\eta_y$  describe the anisotropy of the trapping potential.

As mentioned in Sec. I, for this system we can separate out the c.m. degree of freedom of these two atoms and focus on the interatomic relative motion. The Hamiltonian operator of our problem is given by

$$\hat{H} = \hat{H}_0 + \hat{V}_I. \tag{6}$$

Here,  $\hat{H}_0$  is the free Hamiltonian for the relative motion and can be expressed as

$$\hat{H}_0 + = \hat{\mathbf{p}}^2 + \frac{1}{4} \Big[ \eta_x^2 \hat{x}^2 + \eta_y^2 \hat{y}^2 + \hat{z}^2 \Big], \tag{7}$$

with  $\hat{\mathbf{p}}$  and  $\hat{\mathbf{r}} \equiv (\hat{x}, \hat{y}, \hat{z})$  being the relative momentum and coordinate operators, respectively. In Eq. (6),  $\hat{V}_I$  is the interatomic interaction operator, which is modeled as the *s*-wave Huang-Yang pseudopotential. Explicitly, for any state  $|\psi\rangle$  of the relative motion, we have

$$\langle \mathbf{r} | \hat{V}_I | \psi \rangle = 4\pi a_{3\mathrm{D}} \delta(\mathbf{r}) \frac{\partial}{\partial r} [r \cdot \langle \mathbf{r} | \psi \rangle], \qquad (8)$$

with  $|\mathbf{r}\rangle$  being the eigenstate of the relative-coordinate operator  $\hat{\mathbf{r}}$  with eigenvalue  $\mathbf{r}$ ,  $r = |\mathbf{r}|$ , and  $a_{3D}$  being the 3D *s*-wave scattering length.

For our system, the parity with respect to the spatial inversion  $\mathbf{r} \rightarrow -\mathbf{r}$  is conserved and the contact pseudopotential  $\hat{V}_I$ only operates on the states with even parity. Therefore, in this work, we only consider the eigenenergies and eigenstates of  $\hat{H}$  in the even-parity subspace.

Now we deduce the transcendental equation for the eigenenergy E of the total Hamiltonian  $\hat{H}$ . We begin from the Schrödinger equation,

$$[\hat{H}_0 + \hat{V}_I]|\Psi\rangle = E|\Psi\rangle, \qquad (9)$$

satisfied by E and the corresponding eigenstate  $|\Psi\rangle$  of  $\hat{H}$ . This equation can be reexpressed as

$$|\Psi\rangle = \frac{1}{E - \hat{H}_0} \hat{V}_I |\Psi\rangle. \tag{10}$$

Using Eq. (8), we find that Eq. (10) yields

$$\langle \mathbf{r} | \Psi \rangle = 4\pi a_{3\mathrm{D}} G_0(E, \mathbf{r}) \left[ \left. \frac{\partial}{\partial r} [r \cdot \langle \mathbf{r} | \Psi \rangle] \right|_{\mathbf{r} = \mathbf{0}} \right], \qquad (11)$$

where  $G_0(E, \mathbf{r})$  is the Green's function of the free Hamiltonian  $\hat{H}_0$ , which is defined as

$$G_0(E, \mathbf{r}) = \langle \mathbf{r} | \frac{1}{E - \hat{H}_0} | \mathbf{0} \rangle.$$
(12)

We can derive the equation for the eigenenergy *E* by doing the the operation  $\frac{\partial}{\partial r}(r \cdot)|_{\mathbf{r}=0}$  on both sides of Eq. (11). In this operation, without loss of generality, we choose  $\mathbf{r} = z\mathbf{e}_z$ , with  $\mathbf{e}_z$  being the unit vector along the *z* direction. Then, we find that *E* satisfies

$$J_{3\rm D}(E) = \frac{1}{a_{3\rm D}},\tag{13}$$

which is just Eq. (1) of Sec. I, with the function  $J_{3D}(E)$  being defined as

$$J_{3\mathrm{D}}(E) = 4\pi \left\{ \left. \frac{\partial}{\partial |z|} \Big[ |z| \cdot G_0(E, z\mathbf{e}_z) \Big] \right|_{z \to 0} \right\}.$$
(14)

## A. Expression of $J_{3D}(E)$

Next we derive the expression of the function  $J_{3D}(E)$ . For our system, the ground-state energy of the relative motion of two noninteracting atoms is

$$E_0 = \frac{1}{2}(\eta_x + \eta_y + 1). \tag{15}$$

In the following, we first consider the case with  $E < E_0$ , which was also studied in Ref. [4], and then investigate the general case with arbitrary E.

## 1. Special case: $E < E_0$

When  $E < E_0$ , the Green's function  $G_0(E, z\mathbf{e}_z)$  (z > 0) can be expressed as the Laplace transform of the imaginary-time propagator [44–48], i.e.,

$$G_0(E, z\mathbf{e}_z) = -\int_0^{+\infty} K(z, E, \beta) d\beta, \qquad (16)$$

with the function  $K(z, E, \beta)$  defined as

$$K(z, E, \beta) = e^{\beta E} \langle z \mathbf{e}_z | e^{-\beta H_0} | \mathbf{0} \rangle$$
  
=  $\exp \left[ \beta E - \frac{z^2}{4} \coth \beta \right] \prod_{\alpha = x, y, z} \sqrt{\frac{\eta_\alpha}{4\pi \sinh(\eta_\alpha \beta)}}.$  (17)

Here we emphasize that when  $E < E_0$ , the function  $K(z, E, \beta)$  exponentially decays to zero in the limit  $\beta \rightarrow \infty$ , and thus the integration in Eq. (16) converges for any fixed nonzero *z*. Nevertheless, this integration diverges in the limit  $z \rightarrow 0$ . That is due to the behavior of the leading term  $e^{-z^2/(4\beta)}/(4\pi\beta)^{\frac{3}{2}}$  of the function  $K(z, E, \beta)$  in the limit  $\beta \rightarrow 0^+$  [44–48]. We can separate this divergence by reexpressing the integration as

$$G_{0}(E, z\mathbf{e}_{z}) = -\int_{0}^{+\infty} d\beta \frac{e^{-\frac{z^{2}}{4\beta}}}{(4\pi\beta)^{\frac{3}{2}}} - \int_{0}^{+\infty} d\beta \tilde{K}(r, E, \beta)$$
$$= -\frac{1}{4\pi|z|} - \int_{0}^{+\infty} d\beta \tilde{K}(z, E, \beta), \qquad (18)$$

where

$$\tilde{K}(z, E, \beta) = K(z, E, \beta) - \frac{e^{-\frac{z^2}{4\beta}}}{(4\pi\beta)^{\frac{3}{2}}}.$$
(19)

In Eq. (18), the integration  $\int_0^{+\infty} d\beta \tilde{K}(z, E, \beta)$  uniformly converges in the limit  $z \to 0$ . Using this result, we obtain the expansion of  $G_0(E, z\mathbf{e}_z)$  in this limit,

$$\lim_{z \to 0} G_0(E, z \mathbf{e}_z) = -\frac{1}{4\pi |z|} - \int_0^{+\infty} d\beta \tilde{K}(0; E, \beta) + O(z).$$
(20)

Substituting this result into Eq. (14) and using Eqs. (17) and (19), we finally obtain the expression

$$J_{3D}(E) = -\int_{0}^{+\infty} d\beta \left\{ \frac{e^{\beta E}}{2\sqrt{\pi}} \prod_{\alpha=x,y,z} \sqrt{\frac{\eta_{\alpha}}{\sinh(\eta_{\alpha}\beta)}} - \frac{1}{2\sqrt{\pi}} \frac{1}{\beta^{\frac{2}{3}}} \right\}$$
(for  $E < E_0$ ), (21)

where the integration converges for  $E < E_0$ , as the one in Eq. (16). This result was also derived by Liang and Zhang in Ref. [4].

#### 2. General case: Arbitrary E

In the general case with arbitrary energy E, we cannot directly use the above result in Eq. (21) because the integration in this equation diverges for  $E > E_0$ . For the systems with spherically or axially symmetric confinements, the authors of Refs. [1,2] successfully found the analytical continuation of this integration for all real E. However, for the current system with completely anisotropic traps, to our knowledge, so far such analytical continuation has not been found.

Now we introduce our approach to solve this problem. For convenience, we first define the eigenenergy of the free Hamiltonian  $\hat{H}_0$ , which is just a Hamiltonian of a 3D harmonic oscillator, as  $E_n$ . Here,

$$\mathbf{n} = (n_x, n_y, n_z), \tag{22}$$

with  $n_{\alpha} = 0, 1, 2, ...$  ( $\alpha = x, y, z$ ) being the quantum number of the  $\alpha$  direction. It is clear that we have

$$E_{\mathbf{n}} \equiv \epsilon_{n_x} + \epsilon_{n_y} + \epsilon_{n_z},\tag{23}$$

with

$$\epsilon_{n_{\alpha}} = \left(\frac{1}{2} + n_{\alpha}\right)\eta_{\alpha} \ (\alpha = x, y, z).$$
(24)

In addition, we further denote the eigenstate of  $\hat{H}_0$  corresponding to  $E_n$  as  $|\mathbf{n}\rangle$ .

Similar to above, the key step of our approach is to calculate the free Green's function  $G_0(E, z\mathbf{e}_z)$  defined in Eq. (12). Since the result in Eq. (16) cannot be used for our general case because the integration in this equation diverges for  $E > E_0$ , we need to find another expression for  $G_0(E, z\mathbf{e}_z)$ , which converges for any *E*. To this end, we separate all the eigenstates  $\{|\mathbf{n}\rangle\}$  of  $\hat{H}_0$  into two groups, i.e., the ones with  $\mathbf{n} \in L_E$  and  $\mathbf{n} \in U_E$ , respectively, with the sets  $L_E$  and  $U_E$  being defined as

$$L_E: \{ (n_x, n_y, n_z) | n_{x,y,z} = 0, 1, 2, \dots, \epsilon_{n_x} + \epsilon_{n_y} + \frac{1}{2} \leq E \}$$
(25)

and

$$U_E: \{ (n_x, n_y, n_z) | n_{x,y,z} = 0, 1, 2, \dots, \epsilon_{n_x} + \epsilon_{n_y} + \frac{1}{2} > E \},$$
(26)

respectively. It is clear that we have

$$E_{\mathbf{n}} > E$$
 for all states with  $\mathbf{n} \in U_E$ . (27)

Furthermore, we can reexpress the free Green's operator  $1/[E - \hat{H}_0]$  as

$$\frac{1}{E - \hat{H}_0} = \sum_{\mathbf{n} \in U_E} \frac{|\mathbf{n}\rangle \langle \mathbf{n}|}{E - E_{\mathbf{n}}} + \sum_{\mathbf{n} \in L_E} \frac{|\mathbf{n}\rangle \langle \mathbf{n}|}{E - E_{\mathbf{n}}}$$
(28)

$$= -\int_{0}^{+\infty} d\beta \left[ \sum_{\mathbf{n} \in U_{E}} |\mathbf{n}\rangle \langle \mathbf{n}| e^{-\beta(E_{\mathbf{n}}-E)} \right] + \sum_{\mathbf{n} \in L_{E}} \frac{|\mathbf{n}\rangle \langle \mathbf{n}|}{E - E_{\mathbf{n}}}.$$
(29)

Due to the fact (27), the integration of Eq. (29) *converges*. Moreover, the integrand in Eq. (29) can be rewritten as

$$\sum_{\mathbf{n}\in U_{E}} |\mathbf{n}\rangle \langle \mathbf{n}|e^{-\beta(E_{\mathbf{n}}-E)} = \sum_{\mathbf{n}} |\mathbf{n}\rangle \langle \mathbf{n}|e^{-\beta(E_{\mathbf{n}}-E)}$$
$$-\sum_{\mathbf{n}\in L_{E}} |\mathbf{n}\rangle \langle \mathbf{n}|e^{-\beta(E_{\mathbf{n}}-E)}$$
$$= e^{\beta E}e^{-\beta\hat{H}_{0}} - \sum_{\mathbf{n}\in L_{E}} |\mathbf{n}\rangle \langle \mathbf{n}|e^{-\beta(E_{\mathbf{n}}-E)},$$
(30)

where we have used  $e^{-\beta \hat{H}_0} = \sum_{\mathbf{n}} |\mathbf{n}\rangle \langle \mathbf{n}|e^{-\beta E_{\mathbf{n}}}$ . Substituting Eq. (30) into Eq. (29) and then into Eq. (12), we obtain

$$G_0(E, z\mathbf{e}_z) = -\int_0^{+\infty} d\beta [K(z; E, \beta) - F(z; E, \beta)] + Q(z; E),$$
(31)

where  $K(z; E, \beta)$  is defined in Eq. (17), and the functions  $F(z; E, \beta)$  and  $Q(z; E, \beta)$  are defined as

$$F(z; E, \beta) = \sum_{\mathbf{n} \in L_E} \langle z \mathbf{e}_z | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle e^{-\beta (E_{\mathbf{n}} - E)}, \qquad (32)$$

$$Q(z; E, \beta) = \sum_{\mathbf{n} \in L_E} \frac{\langle z \mathbf{e}_z | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle}{E - E_{\mathbf{n}}}.$$
 (33)

Due to the convergence of the integration in Eq. (29), the integration in Eq. (31) also converges for nonzero *z*, whether  $E > E_0$  or  $E < E_0$ . Therefore, Eq. (31) is the convergent expression of  $G_0(E, z\mathbf{e}_z)$  for arbitrary *E*.

Furthermore, similar to Sec. II A, the integration in Eq. (31) diverges in the limit  $|z| \rightarrow 0$  due to the leading term  $e^{-\frac{z^2}{4\beta}}/(4\pi\beta)^{\frac{3}{2}}$  of the integrand in the limit  $\beta \rightarrow 0^+$ . We can separate this divergence via the technique used in Eqs. (18)–(20), i.e., reexpress this integration as

$$-\int_{0}^{+\infty} d\beta [K(z; E, \beta) - F(z; E, \beta)]$$
  
=  $-\int_{0}^{+\infty} d\beta \frac{e^{-\frac{z^{2}}{4\beta}}}{(4\pi\beta)^{\frac{3}{2}}} - \int_{0}^{+\infty} d\beta [\tilde{K}(z; E, \beta) - F(z; E, \beta)]$   
=  $-\frac{1}{2} - \int_{0}^{+\infty} d\beta [\tilde{K}(z; E, \beta) - F(z; E, \beta)]$  (34)

$$= -\frac{1}{4\pi |z|} - \int_0 d\beta [\tilde{K}(z; E, \beta) - F(z; E, \beta)], \qquad (34)$$

with the function  $\tilde{K}(z; E, \beta)$  defined in Eq. (19). Using this technique, we obtain

$$\lim_{|z|\to 0} G^{(0)}(E, z\mathbf{e}_z) = -\frac{1}{4\pi |z|} + \left[ W_{3\mathrm{D}}(E) + \int_0^{+\infty} I_{3\mathrm{D}}(E, \beta) d\beta \right] + O(z),$$
(35)

where the *z*-independent functions  $W_{3D}(E)$  and  $I_{3D}(E, \beta)$  are defined as  $W_{3D}(E) \equiv Q(0, E, \beta)$  and  $I_{3D}(E, \beta) \equiv -\tilde{K}(0; E, \beta) + F(0; E, \beta)$ , respectively. The straightforward calculations (see Appendix A) show that they can be expressed as

$$W_{3D}(E) = -\frac{\pi}{2} \sqrt{\frac{\eta_x \eta_y}{2}} \sum_{(n_x, n_y) \in C_E^{(3D)}} \left\{ \frac{2^{n_x + n_y - 1} \Gamma\left[\frac{1}{4} - \frac{(E - \epsilon_{n_x} - \epsilon_{n_y})}{2}\right]}{\Gamma\left(\frac{1 - n_x}{2}\right)^2 \Gamma\left(\frac{1 - n_y}{2}\right)^2 \Gamma(n_x + 1) \Gamma(n_y + 1) \Gamma\left[\frac{3}{4} - \frac{(E - \epsilon_{n_x} - \epsilon_{n_y})}{2}\right]} \right\}$$
(36)

and

$$I_{3D}(E,\beta) = -e^{\beta E} \prod_{\alpha=x,y,z} \sqrt{\frac{\eta_{\alpha}}{4\pi \sinh(\eta_{\alpha}\beta)}} + \left(\frac{1}{4\pi\beta}\right)^{\frac{3}{2}} + \sqrt{\frac{\pi\eta_{x}\eta_{y}}{8\sinh\beta}} \sum_{(n_{x},n_{y})\in C_{E}^{(3D)}} \left\{ \frac{2^{n_{x}+n_{y}-\frac{1}{2}}}{\Gamma\left(\frac{1-n_{x}}{2}\right)^{2}\Gamma\left(\frac{1-n_{y}}{2}\right)^{2}\Gamma(n_{x}+1)\Gamma(n_{y}+1)} e^{\beta(E-\epsilon_{n_{x}}-\epsilon_{n_{y}})} \right\},$$
(37)

where  $\Gamma(z)$  is the Gamma function and  $C_E^{(3D)}$  is a set of *two*-dimensional number array  $(n_x, n_y)$ , which is defined as

$$C_E^{(3D)}: \{ (n_x, n_y) | n_{x,y} = 0, 2, 4, 6, \dots; \epsilon_{n_x} + \epsilon_{n_y} + \frac{1}{2} \leqslant E \}.$$
(38)

Finally, substituting Eq. (35) into Eq. (14), we obtain the expression of  $J_{3D}(E)$ ,

$$J_{3D}(E) = 4\pi \left[ W_{3D}(E) + \int_0^{+\infty} I_{3D}(E,\beta) d\beta \right], \quad (39)$$

with  $W_{3D}(E)$  and  $I_{3D}(E, \beta)$  given by Eqs. (36) and (37), respectively. Notice that the summations in Eqs. (36) and (37) are done for finite terms, and the integration in Eq. (39) converges. Therefore, using Eqs. (39), (36), and (37), one can efficiently calculate  $J_{3D}(E)$ . It is clear that Eq. (13) for the eigenenergy and the expression (39) of the function  $J_{3D}(E)$  are correct for not only the systems in completely anisotropic traps, but also the ones in spherically or axially symmetric traps. For the latter two cases, Eq. (39) is equivalent to the results derived by Refs. [1,2], which are shown in Table I.

### **B.** Techniques for fast calculation of $J_{3D}(E)$

There are some techniques which may speed up the calculation for the function  $J_{3D}(E)$ . First, as shown in Appendix B,  $J_{3D}(E)$  can be reexpressed as

$$J_{3D}(E) = 4\pi \left[ \int_0^{\Lambda} A_{3D}(E,\beta) d\beta + B_{3D}^{(1)}(E,\Lambda) + B_{3D}^{(2)}(E,\Lambda) + \left(\frac{1}{2\pi}\right)^{3/2} \frac{1}{\sqrt{2\Lambda}} \right],\tag{40}$$

where  $\Lambda$  is an arbitrary finite positive number, and the functions  $A_{3D}(E,\beta)$  and  $B_{3D}^{(1,2)}(E,\Lambda)$  are defined as

$$A_{3D}(E,\beta) = -e^{\beta E} \prod_{\alpha=x,y,z} \sqrt{\frac{\eta_{\alpha}}{4\pi \sinh(\eta_{\alpha}\beta)}} + \left(\frac{1}{4\pi\beta}\right)^{\frac{3}{2}},$$

$$(41)$$

$$B_{3D}^{(1)}(E,\Lambda) = (-1) \times \sum_{(n_x,n_y)\in C_E^{(3D)}} \left\{ \frac{2^{n_x+n_y-\frac{5}{2}}\sqrt{\pi\eta_x\eta_y} \quad \Gamma(\frac{1}{4} - \frac{E-\epsilon_{n_x}-\epsilon_{n_y}}{2})e^{(E-\epsilon_{n_x}-\epsilon_{n_y}-\frac{3}{2})\Lambda}}{\Gamma(\frac{1-n_x}{2})^2\Gamma(\frac{1-n_y}{2})^2\Gamma(n_x+1)\Gamma(n_y+1)\Gamma(\frac{5}{4} - \frac{E-\epsilon_{n_x}-\epsilon_{n_y}}{2})} \times \sqrt{e^{2\Lambda}-1} \times {}_2F^1\left[1,\frac{3}{4} - \frac{E-\epsilon_{n_x}-\epsilon_{n_y}}{2},\frac{5}{4} - \frac{E-\epsilon_{n_x}-\epsilon_{n_y}}{2},e^{-2\Lambda}\right] \right\},$$

$$(41)$$

and

$$B_{3D}^{(2)}(E,\Lambda) = \sqrt{\pi \eta_x \eta_y \operatorname{csch} \Lambda} \\ \times \sum_{(n_x,n_y) \notin C_E^{(3D)}} \frac{2^{n_x + n_y - 1} e^{(E - \epsilon_{n_x} - \epsilon_{n_y} - 2)\Lambda} (e^{2\Lambda} - 1) \times {}_2F^1 \Big[ 1, \frac{3}{4} - \frac{E - \epsilon_{n_x} - \epsilon_{n_y}}{2}, \frac{5}{4} - \frac{E - \epsilon_{n_x} - \epsilon_{n_y}}{2}, e^{-2\Lambda} \Big]}{\Gamma \Big( \frac{1 - n_x}{2} \Big)^2 \Gamma \Big( \frac{1 - n_y}{2} \Big)^2 \Gamma (n_x + 1) \Gamma (n_y + 1) [2(E - \epsilon_{n_x} - \epsilon_{n_y}) - 1]},$$
(43)

respectively, with  ${}_2F^1$  the hypergeometric function. In Appendix B, we prove that (40) is exactly equivalent to Eqs. (39) for any finite positive  $\Lambda$ . Thus, in the numerical calculation, for a specific problem one can choose the value of  $\Lambda$  by convenience.

Due to the following two facts, the calculation of  $J_{3D}(E)$  based on Eq. (40) may be faster than the one based on Eq. (39):

(a) An important difference between the expressions (39) and (40) is that the integrand  $I_{3D}(E, \beta)$  of Eq. (39) includes a summation  $\sum_{(n_x,n_y)\in C_E^{(3D)}}$ , while the integrand  $A_{3D}(E, \beta)$  of Eq. (40) does not include any summation. On the other hand, in the numerical calculations for these integrations for a fixed *E*, one needs to calculate the values of the integrands for many points in the  $\beta$  axis. Thus, to calculate the term  $\int_0^{+\infty} I_{3D}(E, \beta) d\beta$  of Eq. (39), one need to do the the summation  $\sum_{(n_x,n_y)\in C_E^{(3D)}}$  many times, while to calculate the term  $\int_0^{\Lambda} A_{3D}(E, \beta) d\beta$  of Eq. (40), one is not required to do that. This advantage is more significant for the large-*E* cases where this summation includes many terms.

(b) Furthermore, the terms in the summation  $\sum_{(n_x, n_y) \notin C_E^{(3D)}}$  of Eq. (43) decay to zero in the limits  $n_{x,y} \to \infty$  and the decaying is *faster than exponential* (see Appendix B). There-

fore, although this summation includes infinite terms, it can converges fast.

Another technique which may be helpful for the fast calculation of  $J_{3D}(E)$  is based on the direct conclusions of Eq. (39) or Eq. (40), i.e., for two energies  $E_1$  and  $E_2$ , we have

$$J_{3\mathrm{D}}(E_2) - J_{3\mathrm{D}}(E_1)$$

$$= 4\pi \left[ \delta W_{3\mathrm{D}} + \int_0^{+\infty} \delta I_{3\mathrm{D}}(\beta) d\beta \right]$$
(44)

$$=4\pi \left[\int_0^\Lambda \delta A_{3\mathrm{D}}(\beta)d\beta + \delta B_{3\mathrm{D}}^{(1)}(\Lambda) + \delta B_{3\mathrm{D}}^{(2)}(\Lambda)\right], \quad (45)$$

where  $\delta W_{3D} = W_{3D}(E_2) - W_{3D}(E_1)$ ,  $\delta I_{3D}(\beta) = I_{3D}(E_2, \beta) - I_{3D}(E_1, \beta)$ ,  $\delta A_{3D}(\beta) = A_{3D}(E_2, \beta) - A_{3D}(E_1, \beta)$ , and  $\delta B_{3D}^{(1,2)}(\Lambda) = B_{3D}^{(1,2)}(E_2, \Lambda) - B_{3D}^{(1,2)}(E_1, \Lambda)$ . Therefore, if the value of  $J_{3D}(E_1)$  is already derived, one can calculate  $J_{3D}(E_2)$  by either Eq. (39) [Eq. (40)] or Eq. (44) [Eq. (45)], and it is possible that the summations and integration in Eq. (44) [Eq. (45)] can converge faster.

### C. Energy spectrum

By solving Eq. (13) with expression (39) of  $J_{3D}(E)$ , we can derive the complete eigenenergy spectrum of the complete



FIG. 1. The energy spectrum of the relative motion of two atoms in a 3D harmonic trap with aspect ratios (a) ( $\eta_x = \eta_y = 1$ ), (b) ( $\eta_x = \eta_y = 1.1$ ), (c) ( $\eta_x = 1.2$ ,  $\eta_y = 1.1$ ), and (d) ( $\eta_x = 5$ ,  $\eta_y = 3$ ). We show the results given by Eq. (1) and the expression (39) of function  $J_{3D}(E)$  with solid lines. In (a) and (b), we also show the results given by Refs. [1] and [2] with blue dots, respectively. Here,  $d = \sqrt{\hbar/(2\mu\omega_z)}$  is the characteristic length of our natural unit.

Hamiltonian  $\hat{H}$  of the two-atom relative motion. In Figs. 1(a)–1(d), we show the energy spectrum of the relative motion of two atoms in 3D harmonic traps with various aspect ratios  $\eta_{x,y}$ , which is derived via Eqs. (13) and (39). In Figs. 1(a) and 1(b), we consider the traps with spherical symmetry ( $\eta_x = \eta_y = 1$ ) and axial symmetry ( $\eta_x = \eta_y = 1.1$ ), respectively, and show that the results given by our approach are the same as those from the methods in Refs. [1] and [2]. In Figs. 1(c) and 1(d), we show the results for 3D completely anisotropic traps whose frequencies in every direction are similar to ( $\eta_x = 1.2$ ,  $\eta_y = 1.1$ ) and quite different from ( $\eta_x = 5$ ,  $\eta_y = 3$ ) each other, respectively.

As an example of the application of our method, we make a simple investigation of the energy-level distribution for the systems with integer aspect ratios  $\eta_x$  and  $\eta_y$ . This property is crucial for many important problems such as thermalization and quantum chaotic behaviors. It is clear that for noninteracting cases (i.e., for  $a_{3D} = 0$ ), the energy levels of these systems are distributed with equal spacing. However, when  $a_{3D} \neq 0$ , the energy-level distributions become uneven, i.e., the energy spacings become unequal to each other. Nevertheless, for systems with different aspect ratios, the "degree of unevenness" of the energy-level distributions, or the fluctuation of the energy spacings, is quite different. In Ref. [2], it was found that for the pancake-shaped traps with  $\eta_x \gg 1$  and  $\eta_y = 1$ , the energies are almost distributed with equal spacing (i.e., the level spacings are almost equal), while for the cigar-shaped traps with  $\eta_x = \eta_y \gg 1$ , the eigenenergies are distributed very

unevenly (i.e., the energy spacings fluctuate significantly). Here we study the intermediate cases between the "pancake limit" and the "cigar limit," i.e., the systems with  $a_{3D} = \infty$ ,  $\eta_x = 5$ , and various values of  $\eta_y$  ( $\eta_y = 1, 2, 3, 4, 5$ ). We define  $\Delta_l$  ( $l = 1, 2, 3, ...; \Delta_l > 0$ ) as the spacing between the *l*th and (l + 1)-th excited-state energy of our systems, and calculate the "normalized" energy spacing,

$$\delta_{l} \equiv \frac{\Delta_{l}}{\frac{1}{20} \sum_{l=1}^{20} \Delta_{l}} \quad (l = 1, \dots, 20)$$
(46)

for the lowest 20 excited states for each system. As shown in Fig. 2, when the system is crossed from the pancakeshaped limit to the cigar shape, i.e.,  $\eta_y$  is increased from 1 to 5, the fluctuation of  $\delta_l$  monotonically increases, i.e., the energy-level distribution becomes more and more uneven. In future works, we will perform more systematic investigations for the energy-level distribution of two ultracold atoms in a completely anisotropic harmonic trap.

# **D.** Eigenstate of $\hat{H}$

Furthermore, Eqs. (11) and (12) yield that for a given eigenenergy E of the total Hamiltonian  $\hat{H}$  for the two-atom relative motion, the corresponding eigenstate  $|\Psi\rangle$  satisfies

$$\langle \mathbf{r} | \Psi \rangle \propto \langle \mathbf{r} | \frac{1}{E - \hat{H}_0} | \mathbf{0} \rangle = \sum_{\mathbf{n}} \frac{\langle \mathbf{r} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle}{E - E_{\mathbf{n}}}.$$
 (47)



FIG. 2. The normalized energy spacing  $\delta_l$  defined in Eq. (46) for the lowest 20 excited states of the systems with  $a_{3D} = \infty$ ,  $\eta_x = 5$ , and  $\eta_y$  ( $\eta_y = 1, 2, 3, 4, 5$ ). Here we also use the natural unit  $\hbar = 2\mu = \omega_z = 1$ .

Thus, considering the normalization condition  $\langle \Psi | \Psi \rangle = 1$ , we obtain the expression for the wave function of the eigenstate,

$$\langle \mathbf{r} | \Psi \rangle = \frac{1}{\sqrt{\sum_{\mathbf{n}} |c_{\mathbf{n}}|^2}} \sum_{\mathbf{n}} c_{\mathbf{n}} \phi_{n_x}(\eta_x, x) \phi_{n_y}(\eta_y, y) \phi_{n_z}(\eta_z, z),$$
(48)

with

$$c_{\mathbf{n}} = \frac{1}{(E - E_{\mathbf{n}})} \phi_{n_x}(\eta_x, 0) \phi_{n_y}(\eta_y, 0) \phi_{n_z}(\eta_z, 0).$$
(49)

Here, (x, y, z) are the components of **r**, and the function  $\phi_n(\eta, X)$  is defined in Eq. (A3). It is just the eigenwave function of a one-dimensional harmonic oscillator and satisfies  $\langle \mathbf{r} | \mathbf{n} \rangle = \phi_{n_x}(\eta_x, x)\phi_{n_y}(\eta_y, y)\phi_{n_z}(\eta_z, z)$ . Using Eqs. (48) and (49), one can easily derive the normalized eigenstate  $|\Psi\rangle$  of  $\hat{H}$  from the eigenenergy E [1,3,31].

#### **III. 2D SYSTEMS**

Now we consider two atoms confined in a 2D anisotropic harmonic trap in the *x*-*z* plane. As above, here we assume the trapping frequencies  $\omega_x$  ( $\omega_z$ ) in the *x* (*z*) direction are the same for each atom. Thus, by separating out the c.m. degree of freedom, we can obtain the free Hamiltonian operator for the relative motion. In our natural unit with  $\hbar = 2\mu = \omega_z = 1$ , this free Hamiltonian can be expressed as

$$\hat{H}_0^{(2D)} = \hat{\mathbf{p}}_{2D}^2 + \frac{1}{4}(\eta_x \hat{x}^2 + \hat{z}^2), \tag{50}$$

where  $\eta_x = \omega_x / \omega_z$  is the aspect ratio, as in Sec. II. In Eq. (50),  $\hat{\mathbf{p}}_{2D}$  and  $\hat{\boldsymbol{\rho}} \equiv (\hat{x}, \hat{z})$  are the relative momentum and coordinate operators, respectively. We further assume that the two atoms experience a 2D *s*-wave zero-range interaction with 2D scattering length  $a_{2D}$ , and model this potential with the 2D Bethe-Peierls boundary condition (BPC). Explicitly, in our calculation, the eigenenergy *E* and corresponding eigenstate  $|\Psi\rangle$  for these two interacting atoms should satisfy the equation

$$\langle \boldsymbol{\rho} | \hat{H}_0^{(2D)} | \Psi \rangle = E \langle \boldsymbol{\rho} | \Psi \rangle \quad \text{for } \rho > 0,$$
 (51)

and the BPC [49]

$$\lim_{\rho \to 0} \langle \boldsymbol{\rho} | \Psi \rangle \propto (\ln \rho - \ln a_{2\mathrm{D}}), \tag{52}$$

where  $|\rho\rangle$  is the eigenstate of the relative-coordinate operator  $\hat{\rho}$ , with the corresponding eigenvalue  $\rho$ , and  $\rho = |\rho|$ . Namely, the wave function  $\langle \rho | \Psi \rangle$  satisfies the eigenequation of the free Hamiltonian  $\hat{H}_0^{(2D)}$  in the region other than the origin (i.e., the region with  $\rho > 0$ ) and has the singular behavior (52) in the limit  $\rho \rightarrow 0$ , which describes the interaction effect. As in the 3D cases, this zero-range interaction only acts on the states in the subspace with even parity with respect to the spatial inversion  $\rho \rightarrow -\rho$ , and thus in this work we only consider the eigenenergies in this subspace. In addition, in the long-range limit  $\rho \rightarrow \infty$ , the wave function  $\langle \rho | \Psi \rangle$  should satisfy

$$\lim_{\rho \to \infty} \langle \boldsymbol{\rho} | \Psi \rangle = 0.$$
 (53)

### A. Expression of $J_{2D}(E)$

The solution to Eq. (51) and the long-range condition (53) is proportional to the 2D Green's function, i.e.,

$$\langle \boldsymbol{\rho} | \Psi \rangle \propto G_0^{(\text{2D})}(E; \boldsymbol{\rho}) \equiv \langle \boldsymbol{\rho} | \frac{1}{E - \hat{H}_0^{(\text{2D})}} | \boldsymbol{0} \rangle.$$
 (54)

Therefore, we can derive the algebraic equation for the eigenenergy *E* by matching  $G_0^{(2D)}(E; \rho)$  with the BPC (52). To this end, we need to expand  $G_0^{(2D)}(E; \rho)$  in the limit  $\rho \to 0$ . This expansion can be done with the similar approach as in Sec. II, and we show the detail in Appendix C. Finally, we obtain

$$\lim_{\rho \to 0} G_0^{(2D)}(E; \boldsymbol{\rho}) = \frac{1}{2\pi} \ln \rho - 2 \bigg[ W_{2D}(E) + \int_0^{+\infty} I_{2D}(E, \beta) d\beta \bigg],$$
(55)

where the functions  $W_{2D}(E)$  and  $I_{2D}(E, \beta)$  are defined as

$$W_{2D}(E) = \frac{\sqrt{\eta_x}}{2} \sum_{n_x \in C_E^{(2D)}} \left\{ \frac{\Gamma\left[\frac{1}{4} - \frac{(E - \epsilon_{n_x})}{2}\right] 2^{n_x - 2}}{\Gamma\left(\frac{1 - n_x}{2}\right)^2 \Gamma(n_x + 1) \Gamma\left[\frac{3}{4} - \frac{(E - \epsilon_{n_x})}{2}\right]} \right\} - \frac{\gamma}{4\pi} - \frac{1}{8\pi} \ln\left(\frac{\kappa}{4}\right)$$
(56)

and

$$I_{2D}(E,\beta) = \frac{e^{\beta E}}{2} \prod_{\alpha=x,z} \sqrt{\frac{\eta_{\alpha}}{4\pi \sinh(\eta_{\alpha}\beta)}} - \frac{1}{8\pi\beta} e^{-\kappa\beta} - \sqrt{\frac{\eta_{x}}{4\sinh\beta}} \sum_{n_{x}\in C_{E}^{(2D)}} \left\{ \frac{2^{n_{x}-\frac{3}{2}}}{\Gamma\left(\frac{1-n_{x}}{2}\right)^{2}\Gamma(n_{x}+1)} e^{\beta(E-\epsilon_{n_{x}})} \right\},$$
(57)

respectively, in our natural unit. Here,  $\gamma = 0.5772...$  is the Euler's constant,  $\eta_z = 1$ ,  $\epsilon_{n_x} = (n_x + 1/2)\eta_x$ , the parameter  $\kappa$  could be any positive number, and the result is independent of the value of  $\kappa$ . In Eq. (57),  $C_E^{(2D)}$  is a number set defined as

$$C_E^{(2D)}: \{ n_x | n_x = 0, 2, 4, 6, \dots; \epsilon_{n_x} + \frac{1}{2} \leqslant E \}.$$
 (58)

Clearly, for  $E < E_0$ , the set  $C_E^{(2D)}$  is empty and the summation  $\sum_{n_x \in C_E^{(2D)}}$  in the expressions (56), (57) of  $W_{2D}$  and  $I_{2D}$  becomes zero.

Combining Eqs. (55), (54), and (52), we obtain the equation for the eigenenergy E of our 2D system, which has the form of Eq. (2),

$$J_{2D}(E) = \ln a_{2D}.$$
 (59)

Here the function  $J_{2D}(E)$  is given by

$$J_{2D}(E) = 4\pi \left[ W_{2D}(E) + \int_0^{+\infty} I_{2D}(E,\beta) d\beta \right], \quad (60)$$

with  $W_{2D}(E)$  and  $I_{2D}(E, \beta)$  defined in Eqs. (56) and (57), respectively. One can derive the eigenenergies of the relative motion of these two atoms by solving Eq. (59) and further obtain the corresponding eigenstates.

As in the 3D cases, the summations in Eqs. (56) and (57) are done for finite terms, and the integration in Eq. (60) converges. Therefore, using Eqs. (60), (56), and (57), one

can efficiently calculate  $J_{2D}(E)$ . In addition, for the systems with a 2D isotropic trap ( $\omega_x = \omega_z$  or  $\eta_x = 1$ ), Eq. (59) for the eigenenergy and the expression (60) for the function  $J_{2D}(E)$  are equivalent to the results derived by Ref. [1]. Explicitly, we have [50]

$$J_{2D}(E) = \frac{d}{dz} \ln \left[ \Gamma(z) \right] \Big|_{z = \frac{1-E}{2}} - \gamma + \frac{1}{2} \ln 2 \quad \text{(for } \omega_x = \omega_z\text{)}.$$
(61)

#### **B.** Techniques for fast calculation of $J_{2D}(E)$

The techniques shown in Sec. II B for the fast calculation of  $J_{3D}(E)$  can also be directly generalized to the 2D case (see Appendix D). In particular, as proved in Appendix D,  $J_{2D}(E)$ given by Eq. (60) can be reexpressed as

$$\begin{aligned} I_{\rm 2D}(E) &= 4\pi \left[ \int_0^{\Lambda} A_{\rm 2D}(E,\beta) d\beta + B_{\rm 2D}^{(1)}(E,\Lambda) + B_{\rm 2D}^{(2)}(E,\Lambda) \\ &- \frac{\Gamma(0,\kappa\Lambda)}{8\pi} \right], \end{aligned}$$
(62)

where  $\Lambda$  and  $\kappa$  are arbitrary finite positive numbers, as in Eqs. (56), (57), and (40),  $\Gamma[a, z]$  is the incomplete Gamma function, and the functions  $A_{2D}(E, \beta)$  and  $B_{2D}^{(1,2)}(E, \Lambda)$  are defined as

$$A_{2D}(E,\beta) = \frac{e^{\beta E}}{2} \prod_{\alpha=x,z} \sqrt{\frac{\eta_{\alpha}}{4\pi \sinh(\eta_{\alpha}\beta)}} - \frac{1}{8\pi\beta} e^{-\kappa\beta},$$
(63)  
$$B_{2D}^{(1)}(E,\Lambda) = -\frac{\gamma}{4\pi} - \frac{1}{8\pi} \ln\left(\frac{\kappa}{4}\right) + \sum_{n_x \in C_E^{(2D)}} \left\{ \frac{\sqrt{\eta_x} 2^{n_x - 3} e^{(E-\epsilon_{n_x} - \frac{3}{2})\Lambda} \sqrt{e^{2\Lambda} - 1} \cdot \Gamma\left(\frac{1}{4} - \frac{E-\epsilon_{n_x}}{2}\right)}{\Gamma\left(\frac{1-n_x}{2}\right)^2 \Gamma(n_x + 1) \Gamma\left(\frac{5}{4} - \frac{E-\epsilon_{n_x}}{2}\right)} \right.$$
$$\times {}_2F^1 \left[ 1, \frac{3}{4} - \frac{E-\epsilon_{n_x}}{2}, \frac{5}{4} - \frac{E-\epsilon_{n_x}}{2}, e^{-2\Lambda} \right] \right\},$$
(64)

and

$$B_{2D}^{(2)}(E,\Lambda) = -\sum_{n_x \notin C_E^{(2D)}} \frac{2^{n_x - \frac{3}{2}} \sqrt{\eta_x \operatorname{csch} \Lambda}}{\Gamma\left(\frac{1 - n_x}{2}\right)^2 \Gamma(n_x + 1)} \frac{e^{(E - \epsilon_{n_x} - 2)\Lambda} (e^{2\Lambda} - 1) \times_2 F^1\left[1, \frac{3}{4} - \frac{E - \epsilon_{n_x}}{2}, \frac{5}{4} - \frac{E - \epsilon_{n_x}}{2}, e^{-2\Lambda}\right]}{2(E - \epsilon_{n_x}) - 1},$$
(65)

respectively, with  $_2F^1$  being the hypergeometric function. As in the 3D cases, expression (62) of  $J_{2D}(E)$  has the advantages (a) and (b) shown in Sec. II B. Therefore, the numerical calculations based on Eq. (62) are quite possibly faster than the one based on Eq. (60), especially for the high-energy cases with large *E*.

## C. Energy spectrum and eigenstates

In Fig. 3(a), we illustrate the energy spectrum for the cases with an isotropic 2D trap ( $\eta_x = 1$ ) and show that our results are the same as those from Ref. [1]. In Figs. 3(b) and 3(c), we show the results for anisotropic 2D traps which have similar ( $\eta_x = 1.1$ ) and quite different ( $\eta_x = 5$ ) frequencies in the *x* and *z* directions.

In addition, similar to Sec. II D, one can also derive the corresponding eigenstate  $|\Psi\rangle$  with eigenenergy *E*, which satisfies Eq. (51) and the boundary conditions (52) and (53), as well as the normalization condition  $\langle \Psi | \Psi \rangle = 1$ . With Eq. (54) and the similar approach as in Sec. II D, we obtain

$$\langle \boldsymbol{\rho} | \Psi \rangle = \frac{1}{\sqrt{\sum_{\mathbf{m}} |d_{\mathbf{m}}|^2}} \sum_{\mathbf{m}} d_{\mathbf{m}} \phi_{m_x}(\eta_x, x) \phi_{m_z}(1, z), \quad (66)$$

with

$$d_{\mathbf{m}} = \frac{1}{\left(E - E_{\mathbf{m}}^{(2D)}\right)} \phi_{m_x}(\eta_x, 0) \phi_{m_z}(1, 0).$$
(67)

Here, (x, z) are the components of  $\rho$ ,  $\mathbf{m} = (m_x, m_y)$ , with  $m_{x,y} = 0, 1, 2, \ldots, E_{\mathbf{m}}^{(2D)} = (m_x + 1/2)\eta_x + (m_z + 1/2)$ , and the function  $\phi_n(\eta, X)$  is defined in Eq. (A3).



FIG. 3. The energy spectrum of the relative motion of two atoms in a 2D harmonic trap with (a)  $\eta_x = 1$ , (b)  $\eta_x = 1.1$ , and (c)  $\eta_x = 5$ . We show the results given by Eq. (2) and the expression (60) of the function  $J_{2D}(E)$  with solid lines. In (a), we also show the results given by Ref. [1] with blue dots. Here,  $d = \sqrt{\hbar/(2\mu\omega_z)}$  is the characteristic length of our natural unit.

# IV. GENERALIZATION OF OUR APPROACH TO OTHER PROBLEMS

At the end of this work, we briefly summarize the main ideas of our approach used in the above calculations and discuss how to generalize these ideas. Here we take the 3D systems as an example.

The key step for solving the two-body problem with zerorange interaction is the calculation of the free Green's operator  $\hat{G}_0(E) \equiv 1/[E - \hat{H}_0]$ . When *E* is less than the ground-state energy  $E_0$  of  $\hat{H}_0$ ,  $\hat{G}_0(E)$  can be directly expressed as the Laplace transform of the imaginary-time evolution operator  $e^{-\beta\hat{H}_0}$ , i.e., we have  $\hat{G}_0(E) = -\int_0^{+\infty} e^{\beta(E-\hat{H}_0)} d\beta$ . However, when  $E > E_0$ , this integration diverges. To solve this problem, we separate  $\hat{G}_0(E)$  into two parts, i.e., the contributions from the high-energy states with  $\mathbf{n} \in U_E$  and the ones from the lowenergy states with  $\mathbf{n} \in L_E$ , as shown in Eq. (28). Furthermore, as shown in Eqs. (29) and (30), the first part can still be expressed as the Laplace transform of the operator  $e^{-\beta\hat{H}_0^{(3D)}} - \sum_{\mathbf{n}\in L_E} |\mathbf{n}\rangle\langle\mathbf{n}|e^{-\beta(E_n)}$ , which converges for any nonzero *z*, and the second part only includes finite terms.

It is pointed out that the definitions of the sets  $L_E$  and  $U_E$  are not unique. The only requirements are the following:

(i)  $L_E$  is the complement of  $U_E$ .

(ii)  $E_{\mathbf{n}} > E$  for  $\forall \mathbf{n} \in U_E$ . Namely, the set  $L_E$  includes (but is not limited to) all  $\mathbf{n}$  which satisfies  $E_{\mathbf{n}} < E$ .

For instance, an alternative definition of these two sets is  $U_E : {\mathbf{n} | E_{\mathbf{n}} > E}$  and  $L_E : {\mathbf{n} | E_{\mathbf{n}} \leq E}$ .

Using this method, we can derive the helpful expression (31) of Green's function  $G^{(0)}(E, z\mathbf{e}_z)$ , which is just the matrix element  $\langle z\mathbf{e}_z | \hat{G}_0(E) | \mathbf{0} \rangle$  of the free Green's operator  $\hat{G}_0(E)$ . In this expression, there is just a summation for finite terms and a one-dimensional integration  $\int_0^{+\infty} d\beta [K(z; E, \beta) - F(z; E, \beta)]$  which converges for any finite *z*.

To complete the calculation, we still must remove the divergent term of the above integration in the limit  $z \rightarrow 0$ . This divergent term is contributed by the leading term  $e^{-\frac{z^2}{4\beta}}/(4\pi\beta)^{\frac{3}{2}}$  of the integrand  $K(z; E, \beta) - F(z; E, \beta)$  in the limit  $\beta \rightarrow 0^+$ . Therefore, we can remove it via the technique used in Eqs. (18)–(20).

Our above approach for the calculation of the two-body free Green's function can be directly generalized to other fewbody problems of ultracold atoms, especially the ones where the analytical expressions of the eigenstates and imaginarytime propagator of the free Hamiltonian are known, e.g., the few-body problems in mixed-dimensional systems [51]. Here we emphasize that with the help of the Laplace transformation for the imaginary-time propagator, the free Green's operator can always be expressed as a one-dimensional integration, no matter how many degrees of freedoms are involved in the system. Thus, the free Green's function given by our method always includes a converged one-dimensional integration and summations for finite terms.

### V. SUMMARY

In this work, we derive the transcendental equations for the eigenenergies of two atoms in a 2D or 3D harmonic trap. Our results are applicable for general cases, whether the trap is completely anisotropic or has spherical or axial symmetry. Using our results, one can easily derive the complete energy spectrum, which can be used for further theoretical or experimental studies of dynamical or thermodynamical problems. Our approach can be used in other few-body problems of confined ultracold atoms.

*Note added.* Recently, we became aware of the related work [52]. The authors derived the expression of  $J_{2D}(E)$  for  $E < E_0$ , and a recurrence relation of  $J_{2D}(E)$  for arbitrary E. With this recurrence relation, they also obtained the complete energy spectrum of two atoms in a 2D anisotropic confinement, as well as the eigenstates.

## ACKNOWLEDGMENTS

This work is supported by the National Key R&D Program of China [Grants No. 2018YFA0306502 (P.Z.) and No. 2018YFA0307601 (R.Z.)], NSFC Grants No. 11804268 (R.Z.), No. 11434011 (P.Z.), and No. 11674393 (P.Z.), NSAF Grant No. U1930402, as well as the Research Funds of Renmin University of China under Grant No. 16XNLQ03 (P.Z.).

### APPENDIX A: PROOF OF EQS. (36) AND (37)

In this Appendix, we prove Eqs. (36) and (37) of Sec. II A. To this end, we first show some results on a 1D harmonic oscillator, which will be used in our calculation.

### 1. Some properties of 1D harmonic oscillator

Let us consider a 1D harmonic oscillator with frequency  $\eta$  and mass  $\mu$ . The Hamiltonian of this oscillator is ( $\hbar = 2\mu = 1$ )

$$\hat{H}_{\rm ho} = \hat{P}^2 + \frac{\eta^2 \hat{X}^2}{4},$$
 (A1)

with  $\hat{X}$  and  $\hat{P}$  being the coordinate and momentum operator, respectively. The eigenenergy of  $\hat{H}_{ho}$  is

$$E_n = \left(n + \frac{1}{2}\right)\eta \quad (n = 0, 1, 2, ...),$$
 (A2)

and the wave function of the eigenstate  $|n\rangle$  corresponding to  $E_n$  can be expressed as

$$\langle X|n\rangle \equiv \phi_n(\eta, X) = \left(\frac{\eta}{2\pi}\right)^{\frac{1}{4}} \frac{e^{-\frac{\eta X^2}{4}}}{\sqrt{2^n \Gamma(n+1)}} H_n\left(\sqrt{\frac{\eta}{2}}X\right),\tag{A3}$$

with  $|X\rangle$  being the eigenstate of the position operator  $\hat{X}$  with eigenvalue X, and  $H_n(X)$  and  $\Gamma(\alpha)$  being the Hermitian polynomial and the Gamma function, respectively. The wave function  $\phi_n(\eta, X)$  also satisfies

$$\begin{aligned} \langle X|e^{-\beta\hat{H}_{ho}}|0\rangle &= \sum_{n=0}^{+\infty} \phi_n(\eta, X)\phi_n^*(\eta, 0)e^{-\beta(n+1/2)\eta} \\ &= \sqrt{\frac{\eta}{4\pi\,\sinh(\eta\beta)}}\exp\left[-\frac{\eta X^2\coth\left(\eta\beta\right)}{4}\right], \end{aligned}$$
(A4)

for  $\beta > 0$ .

Now we consider the Green's function  $g(\xi; \eta; X)$  of the 1D harmonic oscillator, which is defined as

$$g(\xi;\eta;X) \equiv \langle X | \frac{1}{\xi - \hat{H}_{ho}} | 0 \rangle = \sum_{n=0}^{+\infty} \frac{\phi_n(\eta, X) \phi_n^*(\eta, 0)}{\xi - E_n}.$$
(A5)

This function satisfies the differential equation

$$\xi \cdot g(\xi;\eta;X) + \frac{d^2}{dX^2}g(\xi;\eta;X) - \frac{\eta^2 X^2}{4}g(\xi;\eta;X) = \delta(X)$$
(A6)

and the boundary condition

$$\lim_{|X| \to \infty} g(\xi; \eta; X) = 0.$$
 (A7)

To derive  $g(\xi; \eta; X)$ , we can first solve Eq. (A6) in the regions X > 0 and X < 0 with the boundary condition (A7), and then match the solution with the connection condition at X = 0, which is given by the term  $\delta(X)$  in Eq. (A6). With this approach, we obtain

$$g(\xi;\eta;X) = \sum_{n=0}^{+\infty} \frac{\phi_n(\eta,X)\phi_n^*(\eta,0)}{\xi - E_n}$$
$$= -\frac{\Gamma\left(\frac{1}{4} - \frac{\xi}{2\eta}\right)}{2^{\frac{\xi}{4} + \frac{\xi}{2\eta}}\sqrt{\pi\eta}} D_{\frac{\xi - \eta/2}{\eta}}(\sqrt{\eta}X), \qquad (A8)$$

where  $D_{\lambda}(\alpha)$  is the parabolic cylinder function. Equation (A8) and the property of the parabolic cylinder function further yields

$$g(\xi;\eta;0) = \sum_{n=0}^{+\infty} \frac{|\phi_n(\eta,0)|^2}{\xi - (n+1/2)\eta} = -\frac{\Gamma\left(\frac{1}{4} - \frac{\xi}{2\eta}\right)}{2\sqrt{2\eta}\Gamma\left(\frac{3}{4} - \frac{\xi}{2\eta}\right)}.$$
(A9)

#### 2. Proof of the two equations

Now we prove Eq. (36) and Eq. (37) of our main text, which are about the expressions of the functions  $W_{3D}(E)$  and  $I_{3D}(E, \beta)$ , respectively. As shown in Sec. II A, these two functions are defined as

$$W_{3D}(E) \equiv Q(0, E, \beta) \tag{A10}$$

and

$$I_{3D}(E,\beta) \equiv -\tilde{K}(0;E,\beta) + F(0;E,\beta), \qquad (A11)$$

with the functions  $F(z; E, \beta)$ ,  $Q(z; E, \beta)$ , and  $\tilde{K}(z; E, \beta)$  defined in Eq. (32), Eq. (33), and Eq. (19), respectively. Thus, we have

$$Q(0; E, \beta) = \sum_{\mathbf{n} \in L_E} \frac{\langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle}{E - E_{\mathbf{n}}}, \qquad (A12)$$

$$F(0; E, \beta) = \sum_{\mathbf{n} \in L_E} \langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle e^{-\beta(E_{\mathbf{n}} - E)}.$$
(A13)

It is clear that the eigenstate  $|\mathbf{n}\rangle$  of the 3D free Hamiltonian  $\hat{H}_0$ , which is defined in Sec. II A, satisfies  $\langle \mathbf{0} | \mathbf{n} \rangle = 0$  when  $n_x$  or  $n_y$  is odd. Using this fact and the definitions of the sets  $L_E$  and  $C_E^{(3D)}$ , which are given in Eqs. (25) and (38) of our main text, we obtain

$$Q(0; E, \beta) = \sum_{(n_x, n_y) \in C_E^{(3D)}} \sum_{n_z=0}^{+\infty} \frac{\langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle}{E - E_{\mathbf{n}}} = \sum_{(n_x, n_y) \in C_E^{(3D)}} \left\{ |\phi_{n_x}(\eta_x, 0)|^2 |\phi_{n_y}(\eta_y, 0)|^2 \left[ \sum_{n_z=0}^{+\infty} \frac{|\phi_{n_z}(\eta_z, 0)|^2}{(E - \epsilon_{n_x} - \epsilon_{n_y}) - \epsilon_{n_z}} \right] \right\}$$
(A14)

and

$$F(0; E, \beta) = \sum_{(n_x, n_y) \in C_E^{(3D)}} \sum_{n_z=0}^{+\infty} \langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle e^{-\beta(E_{\mathbf{n}}-E)} = \sum_{(n_x, n_y) \in C_E^{(3D)}} \left\{ |\phi_{n_x}(\eta_x, 0)|^2 |\phi_{n_y}(\eta_y, 0)|^2 e^{\beta(E-\epsilon_{n_x}-\epsilon_{n_y})} \left[ \sum_{n_z=0}^{+\infty} |\phi_{n_z}(\eta_z, 0)|^2 e^{-\beta\epsilon_{n_z}} \right] \right\},$$
(A15)

of the Hermitian polynomial, we can derive Eq. (36). More-

over, substituting Eq. (A4) into Eq. (A15) and then into

Eq. (A11), and using Eqs. (A16), (19), and (17), we can derive

where  $\eta_{x,y,z}$  and  $\epsilon_{n_{x,y,z}}$  are defined in Sec. II, and the function  $\phi_n(\eta, X)$  is defined by Eq. (A3).

Substituting Eq. (A9) into Eq. (A14) and then into Eq. (A10), and using the property

$$H_n(0) = \frac{2^n \sqrt{\pi}}{\Gamma(\frac{1-n}{2})} \tag{A16}$$

### **APPENDIX B: PROOF OF EQ. (40)**

Eq. (37).

In this Appendix, we prove Eq. (40) in Sec. II B. To this end, we separate the integration in Eq. (39) into two parts, i.e.,

$$\int_0^\infty I_{3\mathrm{D}}(E,\beta)d\beta = \int_0^\Lambda I_{3\mathrm{D}}(E,\beta)d\beta + \int_\Lambda^\infty I_{3\mathrm{D}}(E,\beta)d\beta, \tag{B1}$$

with  $\Lambda$  being an arbitrary finite positive number. In addition, using the definition (37) of  $I_{3D}(E, \beta)$ , we immediately obtain

$$\begin{split} \int_{0}^{\Lambda} I_{3D}(E,\beta) d\beta &= \int_{0}^{\Lambda} A_{3D}(E,\beta) d\beta + \sum_{(n_{x},n_{y})\in C_{E}^{(3D)}} \int_{0}^{\Lambda} \frac{2^{n_{x}+n_{y}-2}\sqrt{\pi \eta_{x}\eta_{y}} e^{\beta(E-\epsilon_{n_{x}}-\epsilon_{n_{y}})}}{\Gamma\left(\frac{1-n_{y}}{2}\right)^{2} \Gamma\left(\frac{1-n_{y}}{2}\right)^{2} \Gamma(n_{x}+1)\Gamma(n_{y}+1)\sqrt{\sinh\beta}} d\beta \\ &= \int_{0}^{\Lambda} A_{3D}(E,\beta) d\beta + \sum_{(n_{x},n_{y})\in C_{E}^{(3D)}} \frac{2^{n_{x}+n_{y}-\frac{5}{2}}\sqrt{\pi \eta_{x}\eta_{y}}}{\Gamma\left(\frac{1-n_{x}}{2}\right)^{2} \Gamma\left(\frac{1-n_{y}}{2}\right)^{2} \Gamma(n_{x}+1)\Gamma(n_{y}+1)} \left\{ \frac{\sqrt{\pi}\Gamma\left(\frac{1}{4}-\frac{E-\epsilon_{n_{x}}-\epsilon_{n_{y}}}{2}\right)}{\Gamma\left(\frac{3}{4}-\frac{E-\epsilon_{n_{x}}-\epsilon_{n_{y}}}{2}\right)} - \frac{\Gamma\left(\frac{1}{4}-\frac{E-\epsilon_{n_{x}}-\epsilon_{n_{y}}}{2}\right)}{\Gamma\left(\frac{5}{4}-\frac{E-\epsilon_{n_{x}}-\epsilon_{n_{y}}}{2}\right)} e^{(E-\epsilon_{n_{x}}-\epsilon_{n_{y}}-\frac{3}{2})\Lambda} \sqrt{e^{2\Lambda}-1} \times {}_{2}F^{1}\left[1,\frac{3}{4}-\frac{E-\epsilon_{n_{x}}-\epsilon_{n_{y}}}{2},\frac{5}{4}-\frac{E-\epsilon_{n_{x}}-\epsilon_{n_{y}}}{2}, e^{-2\Lambda}\right] \right\}, \end{split}$$
(B2)

with  $A_{3D}(E, \beta)$  defined in Eq. (41).

Now we calculate the term  $\int_{\Lambda}^{\infty} I_{3D}(E, \beta) d\beta$  in Eq. (B1). We first notice that according to Eqs. (A11), (A15), (19), and (17),  $I_{3D}(E, \beta)$  can be reexpressed as

$$I_{3D}(E,\beta) = \left(\frac{1}{4\pi\beta}\right)^{\frac{3}{2}} - \sum_{n_x,n_y,n_z} \langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle e^{-\beta(E_\mathbf{n}-E)} + \sum_{(n_x,n_y)\in C_E^{(3D)}} \sum_{n_z=0}^{+\infty} \langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle e^{-\beta(E_\mathbf{n}-E)}$$
$$= \left(\frac{1}{4\pi\beta}\right)^{\frac{3}{2}} - \sum_{(n_x,n_y)\notin C_E^{(3D)}} \sum_{n_z=0}^{+\infty} \langle \mathbf{0} | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle e^{-\beta(E_\mathbf{n}-E)}$$
$$= \left(\frac{1}{4\pi\beta}\right)^{\frac{3}{2}} - \sum_{(n_x,n_y)\notin C_E^{(3D)}} \left\{ |\phi_{n_x}(\eta_x,0)|^2 |\phi_{n_y}(\eta_y,0)|^2 e^{\beta(E-\epsilon_{n_x}-\epsilon_{n_y})} \left[ \sum_{n_z=0}^{+\infty} |\phi_{n_z}(\eta_z,0)|^2 e^{-\beta\epsilon_{n_z}} \right] \right\},$$
(B3)

where  $\epsilon_{n_{x,y,z}}$  is defined in Sec. II, and the function  $\phi_n(\eta, X)$  is defined in Eq. (A3). Moreover, substituting Eq. (A4) and Eq. (A16) into Eq. (B3), we further obtain

$$I_{3D}(E,\beta) = \left(\frac{1}{4\pi\beta}\right)^{\frac{3}{2}} - \sqrt{\frac{\pi\eta_x\eta_y}{8\sinh\beta}} \sum_{(n_x,n_y)\notin C_E^{(3D)}} \left\{ \frac{2^{n_x+n_y-\frac{1}{2}}}{\Gamma\left(\frac{1-n_x}{2}\right)^2\Gamma\left(\frac{1-n_y}{2}\right)^2\Gamma(n_x+1)\Gamma(n_y+1)} e^{\beta(E-\epsilon_{n_x}-\epsilon_{n_y})} \right\},\tag{B4}$$

where  $\eta_{x,y,z}$  is defined in Sec. II. Doing the integration  $\int_{\Lambda}^{\infty} I_{3D}(E,\beta)d\beta$  in both sides of Eq. (B4), we further obtain

$$\int_{\Lambda}^{\infty} I_{3D}(E,\beta) d\beta = B_{3D}^{(2)}(E,\Lambda) + \left(\frac{1}{2\pi}\right)^{3/2} \frac{1}{\sqrt{2\Lambda}},$$
(B5)

where  $B_{3D}^{(2)}(E, \Lambda)$  is defined in Eq. (43).

Substituting Eqs. (B2) and (B5) into Eq. (B1) and then into Eq. (39), and further using Eqs. (36), we can derive Eq. (40).

# APPENDIX C: PROOF OF EQS. (56) and (57)

In this Appendix, we prove Eqs. (56) and (57) in Sec. III, which is related to the behavior of the 2D Green's function  $G_0^{(2D)}(E; \rho)$  in the limit  $|\rho| \rightarrow 0$ . Our approach is similar to the method used in Sec. II A.

As in Sec. II and Appendix A, by choosing  $\rho = \rho \mathbf{e}_z \ (\rho > 0)$  and making direct calculations, we can find that

$$G_0^{(2D)}(E;\rho\mathbf{e}_z) = -\int_0^\infty [K_{2D}(\rho;E,\beta) - Y(\rho;E,\beta)]d\beta + Z(\rho;E),$$
(C1)

which is similar to the expression (31) of the 3D free Green's function. Here the functions  $K_{2D}(\rho; E, \beta)$ ,  $Y(\rho; E, \beta)$ , and  $Z(\rho; E)$  are defined as

$$K_{2D}(\rho; E, \beta) = e^{\beta E} \langle \rho \mathbf{e}_{z} | e^{-\beta \hat{H}_{0}^{(2D)}} | \mathbf{0} \rangle = \left( \prod_{\substack{\alpha = x, z \\ +\infty}} \sqrt{\frac{\eta_{\alpha}}{4\pi \sinh(\eta_{\alpha}\beta)}} \right) \exp\left[ \beta E - \frac{\coth(\beta)}{4} \rho^{2} \right], \tag{C2}$$

$$Y(\rho; E, \beta) = \sum_{n_x \in C_E^{(2D)}} \sum_{n_z=0}^{+\infty} e^{\beta(E-E_{\mathbf{n}})} \langle \rho \mathbf{e}_z | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle,$$
(C3)

$$Z(\rho; E) = \sum_{n_x \in C_E^{(2D)}} \sum_{n_z=0}^{+\infty} \frac{\langle \rho \mathbf{e}_z | \mathbf{n} \rangle \langle \mathbf{n} | \mathbf{0} \rangle}{E - E_{\mathbf{n}}},$$
(C4)

where  $\eta_{x,z}$  and  $\epsilon_{n_{x,z}}$  have the same definition as in Sec. II and  $C_E^{(2D)}$  is defined in Eq. (58). Furthermore, the integration in Eq. (C1) diverges in the limit  $z \to 0$ , and we can remove this divergence using a similar approach as in Sec. III. To this end, we reexpress Eq. (C1) as

$$G_{0}^{(2D)}(E;\rho\mathbf{e}_{z}) = -\int_{0}^{\infty} \frac{1}{4\pi\beta} \exp\left(-\beta\kappa - \frac{1}{4\beta}\rho^{2}\right) d\beta - \int_{0}^{\infty} [\tilde{K}_{2D}(\rho;E,\beta) - Y(\rho;E,\beta)] d\beta + Z(\rho;E),$$
(C5)

with  $\kappa$  being any positive number and

$$\tilde{K}_{2D}(\rho; E, \beta) = K_{2D}(\rho; E, \beta) - \frac{1}{4\pi\beta} \exp\left(-\beta\kappa - \frac{1}{4\beta}\rho^2\right).$$
(C6)

Furthermore, using the result

$$\int_{0}^{\infty} \frac{1}{4\pi\beta} \exp\left(-\beta\kappa - \frac{1}{4\beta}\rho^{2}\right) d\beta = -\frac{1}{2\pi} \ln\rho - \frac{1}{2\pi}\gamma - \frac{1}{4\pi} \ln\frac{\kappa}{4} + O(\rho) \quad \text{(for } \rho > 0\text{)}, \tag{C7}$$

with  $\gamma = 0.5772...$  being the Euler's constant, we derive a result with the same form of Eq. (55):

$$\lim_{\rho \to 0} G_0^{(2D)}(E; \boldsymbol{\rho}) = \frac{1}{2\pi} \ln \rho - 2 \bigg[ W_{2D}(E) + \int_0^{+\infty} I_{2D}(E, \beta) d\beta \bigg].$$
(C8)

In this step, the functions  $W_{2D}(E)$  and  $I_{2D}(E, \beta)$  are given by

$$W_{2D}(E) = -\frac{1}{4\pi}\gamma - \frac{1}{8\pi}\ln\frac{\kappa}{4} - \frac{1}{2}Z(0, E),$$
(C9)

$$I_{2D}(E,\beta) = \frac{1}{2} [\tilde{K}_{2D}(0;E,\beta) - Y(0;E,\beta)].$$
(C10)

Moreover, with the method in Appendix A, we can derive the alternative expressions of  $W_{2D}(E)$  and  $I_{2D}(E, \beta)$ , i.e., Eqs. (56) and (57).

# APPENDIX D: TECHNIQUES FOR FAST CALCULATION OF $J_{2D}(E)$

In this Appendix, we generalize the techniques shown in Sec. II B and Appendix B to the 2D case. We first generalize Eqs. (40)–(43) to the 2D cases and prove Eq. (62). This can be done via direct calculations with the method shown in Appendix B. We separate the integration in Eq. (60) into two parts, i.e.,

$$\int_0^\infty I_{2\mathrm{D}}(E,\beta)d\beta = \int_0^\Lambda I_{2\mathrm{D}}(E,\beta)d\beta + \int_\Lambda^\infty I_{2\mathrm{D}}(E,\beta)d\beta, \tag{D1}$$

with  $\Lambda$  being an arbitrary positive number. In addition, using the definition (57) of  $I_{2D}(E, \beta)$ , we immediately obtain

$$\int_0^{\Lambda} I_{2\mathrm{D}}(E,\beta) d\beta$$
$$= \int_0^{\Lambda} A_{2\mathrm{D}}(E,\beta) d\beta - \sum_{n_x \in \mathcal{C}_E^{(2\mathrm{D})}} \int_0^{\Lambda} \frac{2^{n_x - \frac{5}{2}} \sqrt{\eta_x} e^{\beta(E - \epsilon_{n_x})}}{\Gamma\left(\frac{1 - n_x}{2}\right)^2 \Gamma(n_x + 1) \sqrt{\sinh\beta}} d\beta$$

$$= \int_{0}^{\Lambda} A_{2D}(E,\beta) d\beta + \sum_{(n_{x})\in C_{E}^{(2D)}} \frac{2^{n_{x}-3}\sqrt{\eta_{x}}}{\Gamma\left(\frac{1-n_{x}}{2}\right)^{2}\Gamma(n_{x}+1)} \left\{ -\frac{\sqrt{\pi}\Gamma\left(\frac{1}{4}-\frac{E-\epsilon_{n_{x}}}{2}\right)}{\Gamma\left(\frac{3}{4}-\frac{E-\epsilon_{n_{x}}}{2}\right)} + \frac{\Gamma\left(\frac{1}{4}-\frac{E-\epsilon_{n_{x}}}{2}\right)}{\Gamma\left(\frac{5}{4}-\frac{E-\epsilon_{n_{x}}}{2}\right)} e^{(E-\epsilon_{n_{x}}-\frac{3}{2})\Lambda}\sqrt{e^{2\Lambda}-1} \times_{2}F^{1}\left[1,\frac{3}{4}-\frac{E-\epsilon_{n_{x}}}{2},\frac{5}{4}-\frac{E-\epsilon_{n_{x}}}{2},e^{-2\Lambda}\right]\right\},$$
(D2)

with  $A_{2D}(E, \beta)$  defined in Eq. (63).

Furthermore, using the method in Appendix B, we find that the function  $I_{2D}(E,\beta)$  defined in Eq. (57) has an alternative expression,

$$I_{2D}(E,\beta) = -\frac{1}{8\pi\beta}e^{-\kappa\beta} + \sqrt{\frac{\eta_x}{4\sinh\beta}} \sum_{n_x \notin C_E^{(2D)}} \left\{ \frac{2^{n_x - \frac{3}{2}}}{\Gamma\left(\frac{1 - n_x}{2}\right)^2 \Gamma(n_x + 1)} e^{\beta(E - \epsilon_{n_x})} \right\},\tag{D3}$$

which is similar to Eq. (B4). Thus, doing the integration  $\int_{\Lambda}^{\infty} I_{2D}(E,\beta)d\beta$  in both sides of Eq. (D3), we further obtain

$$\int_{\Lambda}^{\infty} I_{2\mathrm{D}}(E,\beta) d\beta = B_{2\mathrm{D}}^{(2)}(E,\Lambda) - \frac{\Gamma(0,\kappa\Lambda_2)}{8\pi},\tag{D4}$$

where  $\Gamma[a, z]$  is the incomplete Gamma function and  $B_{2D}^{(2)}(E, \Lambda)$  is defined in Eq. (65). As in Appendix B, substituting Eqs. (D2) and (D4) into Eq. (D1) and then into Eq. (60), and further using Eqs. (56), we can derive Eq. (62).

In addition, the second technique shown in Sec. II B is the one based on Eqs. (44) and (45). It is clear that this technique can be directly generalized to the 2D case.

- [1] T. Busch, B.-G. Englert, K. Rzażewski, and M. Wilkens, Found. Phys. 28, 549 (1998).
- [2] Z. Idziaszek and T. Calarco, Phys. Rev. A 71, 050701(R) (2005).
- [3] Z. Idziaszek and T. Calarco, Phys. Rev. A 74, 022712 (2006).
- [4] J.-J. Liang and C. Zhang, Phys. Scr. 77, 025302 (2008).
- [5] D. Blume, Rep. Prog. Phys. 75, 046401 (2012).
- [6] S. Grishkevich and A. Saenz, Phys. Rev. A 80, 013403 (2009).
- [7] S. Grishkevich, S. Sala, and A. Saenz, Phys. Rev. A 84, 062710 (2011).
- [8] S. Sala, G. Zürn, T. Lompe, A. N. Wenz, S. Murmann, F. Serwane, S. Jochim, and A. Saenz, Phys. Rev. Lett. 110, 203202 (2013).
- [9] S. Sala and A. Saenz, Phys. Rev. A 94, 022713 (2016).
- [10] B. Sun, W. X. Zhang, S. Yi, M. S. Chapman, and L. You, Phys. Rev. Lett. 97, 123201 (2006).
- [11] L. M. A. Kehrberger, V. J. Bolsinger, and P. Schmelcher, Phys. Rev. A 97, 013606 (2018).
- [12] G. Bougas, S. I. Mistakidis, and P. Schmelcher, Phys. Rev. A 100, 053602 (2019).
- [13] L. Budewig, S. I. Mistakidis, and P. Schmelcher, Mol. Phys. 117, 2043 (2019).
- [14] A. N. Wenz, G. Zürn, S. Murmann, I. Brouzos, T. Lompe, and S. Jochim, Science 342, 457 (2013).
- [15] D. Blume and C. H. Greene, Phys. Rev. A 66, 013601 (2002).
- [16] X.-J. Liu, H. Hu, and P. D. Drummond, Phys. Rev. Lett. 102, 160401 (2009).
- [17] J. von Stecher, C. H. Greene, and D. Blume, Phys. Rev. A 77, 043619 (2008).
- [18] K. M. Daily and D. Blume, Phys. Rev. A 81, 053615 (2010).
- [19] S. E. Gharashi, K. M. Daily, and D. Blume, Phys. Rev. A 86, 042702 (2012).

- [20] X.-J. Liu, Phys. Rep. 524, 37 (2013).
- [21] X. Y. Yin, D. Blume, P. R. Johnson, and E. Tiesinga, Phys. Rev. A **90**, 043631 (2014).
- [22] S.-G. Peng, S.-H. Zhao, and K. Jiang, Phys. Rev. A 89, 013603 (2014).
- [23] X. Y. Yin and D. Blume, Phys. Rev. A 92, 013608 (2015).
- [24] D. Blume, M. W. C. Sze, and J. L. Bohn, Phys. Rev. A 97, 033621 (2018)
- [25] X. Y. Yin, H. Hu, and X.-J. Liu, Phys. Rev. Lett. 124, 013401 (2020).
- [26] F. Scazza, C. Hofrichter, M. Höfer, P. C. De Groot, I. Bloch, and S. Fölling, Nat. Phys. 10, 779 (2014).
- [27] G. Cappellini, M. Mancini, G. Pagano, P. Lombardi, L. Livi, M. Siciliani de Cumis, P. Cancio, M. Pizzocaro, D. Calonico, F. Levi, C. Sias, J. Catani, M. Inguscio, and L. Fallani, Phys. Rev. Lett. 113, 120402 (2014).
- [28] M. A. Norcia, A. W. Young, and A. M. Kaufman, Phys. Rev. X 8,041054 (2018).
- [29] A. Cooper, J. P. Covey, I. S. Madjarov, S. G. Porsev, M. S. Safronova, and M. Endres, Phys. Rev. X 8, 041055 (2018).
- [30] G. Cappellini, L. F. Livi, L. Franchi, D. Tusi, D. Benedicto Orenes, M. Inguscio, J. Catani, and L. Fallani, Phys. Rev. X 9,011028 (2019).
- [31] Q. Guan, V. Klinkhamer, R. Klemt, J. H. Becher, A. Bergschneider, P. M. Preiss, S. Jochim, and D. Blume, Phys. Rev. Lett. 122, 083401 (2019).
- [32] L. R. Liu, J. D. Hood, Y. Yu, J. T. Zhang, N. R. Hutzler, T. Rosenband, and K.-K. Ni, Science **360**, 900 (2018).
- [33] L. Anderegg, L. W. Cheuk, Y. Bao, S. Burchesky, W. Ketterle, K.-K. Ni, and J. M. Doyle, Science 365, 1156 (2019).
- [34] J. D. Hood, Y. Yu, Y.-W. Lin, J. T. Zhang, K. Wang, L. R. Liu, B. Gao, and K.-K. Ni, Phys. Rev. Research 2, 023108 (2020).

- [35] K. Wang, X. He, R. Guo, P. Xu, C. Sheng, J. Zhuang, Z. Xiong, M. Liu, J. Wang, and M. Zhan, Phys. Rev. A 100, 063429 (2019).
- [36] D. E. Chang, J. S. Douglas, A. González-Tudela, C.-L. Hung, and H. J. Kimble, Rev. Mod. Phys. 90, 031002 (2018).
- [37] Y. Meng, A. Dareau, P. Schneeweiss, and A. Rauschenbeutel, Phys. Rev. X 8, 031054 (2018).
- [38] A. Dareau, Y. Meng, P. Schneeweiss, and A. Rauschenbeutel, Phys. Rev. Lett. **121**, 253603 (2018).
- [39] C. Ospelkaus, S. Ospelkaus, L. Humbert, P. Ernst, K. Sengstock, and K. Bongs, Phys. Rev. Lett. 97, 120402 (2006).
- [40] T. Stöferle, H. Moritz, K. Günter, M. Köhl, and T. Esslinger, Phys. Rev. Lett. 96, 030401 (2006).
- [41] M. J. Mark, E. Haller, K. Lauber, J. G. Danzl, A. J. Daley, and H.-C. Nägerl, Phys. Rev. Lett. **107**, 175301 (2011).
- [42] L. Riegger, N. Darkwah Oppong, M. Höfer, D. R. Fernandes, I. Bloch, and S. Fölling, Phys. Rev. Lett. **120**, 143601 (2018).

annurin V. Vie, M. I. Van de Graaff, I. S. Donouvski, I. D.

PHYSICAL REVIEW A 101, 053624 (2020)

- [43] R. Chapurin, X. Xie, M. J. Van de Graaff, J. S. Popowski, J. P. D'Incao, P. S. Julienne, J. Ye, and E. A. Cornell, Phys. Rev. Lett. 123, 233402 (2019).
- [44] P. Massignan and Y. Castin, Phys. Rev. A **74**, 013616 (2006).
- [45] R. Zhang and P. Zhang, Phys. Rev. A 98, 043627 (2018).
- [46] R. Zhang and P. Zhang, Phys. Rev. A 100, 063607 (2019).
- [47] D. Xiao, R. Zhang, and P. Zhang, Few-Body Syst. 60, 63 (2019).
- [48] R. Zhang and P. Zhang, Phys. Rev. A 101, 013636 (2020).
- [49] B. J. Verhaar, J. P. H. W. van den Eijnde, M. A. J. Voermans, and M. M. J. Schaffrath, J. Phys. A: Math. Gen. 17, 595 (1984).
- [50] Notice that the definitions of the 2D scattering length in Ref. [1] and in this work are different.
- [51] Y. Nishida and S. Tan, Phys. Rev. Lett. 101, 170401 (2008).
- [52] G. Bougas, S. I. Mistakidis, G. M. Alshalan, and P. Schmelcher, arXiv:2001.10722.