Electric dipole moments of atoms and molecules produced by enhanced nuclear Schiff moments

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We perform calculations of the *CP*-violating atomic and molecular electric dipole moments (EDMs) induced by the interaction of the nuclear Schiff moments with electrons. EDMs of atoms Eu, Dy, Gd, Ac, Th, Pa, U, Np, and Pu are of special interest since they have isotopes with strongly enhanced nuclear Schiff moments caused by the octupole nuclear deformation or soft octupole vibration mode. These atoms have open 4 *f* or 5 *f* shells making the calculations complicated. We use our special version of the configuration-interaction method combined with the many-body perturbation theory method adopted for an open *f* -shell case. To validate the method we perform similar calculations for simpler atoms (Xe, Hg, TI^+ , Pb, Pb^{2+} , Rn, Ra), where some earlier results are available. In addition, we present the estimates of the *CP*-violating nuclear spin–molecular axis interaction constants for molecules which may be of experimental interest, including AcF, AcN, AcO⁺, EuN, EuO⁺, ThO, PbO, and TlF. We also present updated values of the nuclear Schiff moments and atomic and molecular EDMs expressed in terms of the *CP*-violating π -meson–nucleon interaction constants \bar{g}_0 , \bar{g}_1 , \bar{g}_2 , QCD parameter $\bar{\theta}$, and quark chromo-EDMs. The results may be used to test *CP*-violation theories and search for axion dark matter in atomic, molecular, and solid-state experiments.

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I. INTRODUCTION

A. Nuclear Schiff moments produced by *T, P***-odd nuclear forces and atomic and molecular electric dipole moments produced by electric field of Schiff moments**

Measurements of atomic and molecular time-reversal (*T*) and parity- (*P*) violating electric dipole moments are used to test unification theories predicting *CP* violation. They have already excluded a number of models and significantly reduced parametric space of other popular models including supersymmetry [\[1,2\]](#page-9-0). Another motivation is related to the baryogenesis problem, the matter-antimatter asymmetry in the universe which is produced by an unknown *CP*-violating interaction. The expected magnitude of an electric dipole moment (EDM) is very small; therefore, we are looking for mechanisms that enhance the effects—see, e.g., Refs. [\[3–5\]](#page-9-0).

Schiff demonstrated that the nuclear EDM is completely screened in neutral atoms and molecules and noted that a nonzero atomic EDM still may be produced if the distribution of EDMs and charge in a nucleus are not proportional to each other $[6]$. Further works $[7-11]$ introduced and calculated the so-called Schiff moment, a vector moment presenting electric field inside the nucleus after taking into account nuclear EDM screening by electrons. This Schiff moment electric field polarizes the atom and produces an atomic EDM directed along the nuclear spin. References [\[7,8\]](#page-9-0) calculated the Schiff moment due to the proton EDM. References [\[9–11\]](#page-9-0) calculated (and named) the nuclear Schiff moment produced by the *P*, *T*-odd nuclear forces. It was shown in [\[9\]](#page-9-0) that the contribution of the *P*, *T* -odd forces to the nuclear EDM and Schiff moment is larger than the contribution of a nucleon EDM. In Ref. [\[12\]](#page-9-0) an accurate expression for the Schiff moment electrostatic potential has been derived and the finite nuclear

size corrections to the Schiff moment operator introduced (see also [\[13,14\]](#page-9-0)).

B. Enhancement of Schiff moment due to nuclear octupole deformation and soft octupole vibration mode

A number of nuclei have an opposite parity level with the same spin close to the ground state. This may lead to an enhancement of the nuclear Schiff moment produced by the *P*, *T* -odd nuclear forces which admix this close wave function to the ground state $[9]$.¹ However, the largest enhancement $(\sim 10^2 - 10^3$ times) happens in nuclei with an intrinsic octupole deformation, where both the small energy difference of nuclear levels with opposite parity and the collective effect work together [\[17,18\]](#page-9-0). According to [\[17,18\]](#page-9-0) this happens in some isotopes of Fr, Rn, Ra, and actinide atoms. Atomic and molecular EDMs produced by the Schiff moment increase with the nuclear charge *Z* faster than Z^2 [\[9\]](#page-9-0). This is another reason why EDMs in actinide atoms and their molecules are expected to be significantly larger than in other systems.

The Schiff moment is proportional to the squared octupole deformation parameter $(\beta_3)^2$ [\[18\]](#page-9-0). In nuclei with the octupole deformation $(\beta_3)^2 \sim (0.1)^2$. According to Ref. [\[19\]](#page-9-0), in nuclei with a soft octupole vibration mode the squared dynamical octupole deformation $\langle (\beta_3)^2 \rangle \sim (0.1)^2$, i.e., it is the same as the static octupole deformation. This means that a similar

¹Nuclear EDM and magnetic quadrupole produced by the *T*, *P*-odd nuclear forces are also enhanced due to an opposite parity level with the same spin close to the ground state $[9,15]$. Collective enhancement of the magnetic quadrupole moments in deformed nuclei have been demonstrated in [\[16\]](#page-9-0).

enhancement of the Schiff moment may be due to the dynam-ical octupole effect [\[19–21\]](#page-9-0) in nuclei where $\langle \beta_3 \rangle = 0.2$ This observation significantly increases the list of nuclei where the Schiff moment is enhanced.

In the papers in [\[17,18,22,23\]](#page-9-0) numerical calculations of Schiff moments and estimates of atomic EDM produced by the electrostatic interaction between electrons and these moments have been done for ²²³Ra, ²²⁵Ra, ²²³Rn, ²²¹Fr, ²²³Fr, 225 Ac, and 229 Pa. Unfortunately, these nuclei have a short lifetime. Several experimental groups have considered experiments with ^{225}Ra and ^{223}Rn [\[24](#page-9-0)[–26\]](#page-10-0). The only published EDM measurements $[24,25]$ have been done for 225 Ra which has 15 days half-life. In spite of the Schiff moment enhancement the ²²⁵Ra EDM measurement has not yet reached the sensitivity to the T , P -odd interaction comparable to the Hg EDM experiment [\[27\]](#page-10-0). The experiments continue, however, the instability of 225 Ra and a relatively small number of atoms available may be a problem. In Ref. [\[28\]](#page-10-0) the nuclear Schiff moment of 2^{29} Th nucleus has been estimated since this nucleus has a much longer lifetime (7917 years). In Ref. [\[29\]](#page-10-0) the list of the candidates for the enhanced Schiff moments has been extended to include stable isotopes ¹⁵³Eu, ¹⁶¹Dy, ¹⁶³Dy, ¹⁵⁵Gd, and long lifetime nuclei ²³⁵U, ²³⁷Np, ²³³U, ²²⁹Th, ¹⁵³Sm, ¹⁶⁵Er, ²²⁵Ac, ²²⁷Ac, ²³¹Pa, and ²³⁹Pu. The estimates of the enhanced Schiff moments have been done for the most attractive cases of 153 Eu, 235 U, 237 Np, and 227 Ac.

In this paper we present updated values for many nuclear Schiff moments expressed in terms of the *CP*-violating π meson–nucleon interaction constants \bar{g}_0 , \bar{g}_1 , \bar{g}_2 , QCD parameter θ , and quark chromo-EDMs. We also calculate atomic and molecular EDMs induced by these Schiff moments.

C. Oscillating Schiff moments and atomic and molecular electric dipole moments produced by axion dark matter

The *CP*-violating neutron EDM may be due to the QCD θ term [\[30\]](#page-10-0). It was noted in Ref. [\[31\]](#page-10-0) that the axion dark matter produces oscillating neutron EDMs since the axion field is equivalent to the oscillating $\bar{\theta}$. The QCD θ term also produces *P*, *T* -odd nuclear forces creating nuclear Schiff moments. Correspondingly, the axion field also produces oscillating nuclear Schiff moments [\[32\]](#page-10-0) which are enhanced by the octupole mechanism. To obtain the results for the oscillating Schiff moment it is sufficient to replace the constant $\bar{\theta}$ by $\bar{\theta}(t) = a(t)/f_a$, where f_a is the axion decay constant, $a(t) = a_0 \cos m_a t$, $(a_0)^2 = 2\rho/(m_a)^2$, and ρ is the axion dark matter energy density [\[31,32\]](#page-10-0). Moreover, in the case of a resonance between the frequency of the axion field oscillations and molecular transition frequency there may be an enormous resonance enhancement of the oscillating nuclear Schiff moment effect [\[33\]](#page-10-0). Since an oscillating nuclear Schiff moment and oscillating nuclear EDM may be produced by the axion dark matter, corresponding measurements may be used to search for the dark matter. First results of such a search have been published in Ref. [\[34\]](#page-10-0), where the oscillating neutron EDM and oscillating ¹⁹⁹Hg Schiff moment have been

measured. Search for the effects produced by the oscillating axion-induced Schiff moments in solid-state materials is in progress [\[35\]](#page-10-0).

II. ESTIMATES OF NUCLEAR SCHIFF MOMENTS

In this section we present values of the nuclear Schiff moments for all atoms considered in the present work. We present corrected values if there is a reason to make the corrections, express the results in terms of the π -meson– nucleon interaction constants \bar{g}_0 , \bar{g}_1 , \bar{g}_2 , QCD parameter $\bar{\theta}$, and quark chromo-EDMs (such presentations were often not available), and perform rough estimates for several nuclei where the Schiff moments were unknown.

A. Estimates of enhanced Schiff moments in deformed nuclei

Nuclear spectra of a nucleus with the octupole (pear-shape) deformation are similar to spectra of a diatomic molecule made from different atoms.³ Therefore, looking into the spectra gives us the first indication of the octupole.⁴ Other indications include measured probabilities of the electric octupole and electric dipole transitions. There are also sophisticated nuclear calculations which give us calculated deformation parameters—see, e.g., Refs. [\[38–47\]](#page-10-0).

The Schiff moment is defined by the following expression [\[9\]](#page-9-0):

$$
\mathbf{S} = \frac{e}{10} \left[\langle r^2 \mathbf{r} \rangle - \frac{5}{3Z} \langle r^2 \rangle \langle \mathbf{r} \rangle \right],\tag{1}
$$

where $\langle r^n \rangle \equiv \int \rho(\mathbf{r}) r^n d^3 r$ are the moments of the nuclear charge density ρ . The second term originates from the electron screening and contains nuclear mean-squared charge radius $\langle r^2 \rangle / Z$ and nuclear EDM $d = e \langle \mathbf{r} \rangle$, where *Z* is the nuclear charge.

If a nucleus has an octupole deformation β_3 and a quadrupole deformation β_2 , in the fixed-body (rotating) frame the Schiff moment S_{intr} is proportional to the octupole moment O_{intr} , i.e., it has a collective nature [\[17,18\]](#page-9-0):

$$
S_{\text{intr}} \approx \frac{3}{5\sqrt{35}} O_{\text{intr}} \beta_2 \approx \frac{3}{20\pi\sqrt{35}} eZR^3 \beta_2 \beta_3, \qquad (2)
$$

²Recall an ordinary oscillator where $\langle x \rangle = 0$, while $\langle x^2 \rangle$ is not equal to zero.

³Note that the enhancement of the EDM and Schiff moment in nuclei with the octupole deformation is similar to the enhancement of the *T*, *P*-violating effects in polar molecules with nonzero electron angular momentum which have doublets of the opposite parity levels [\[36\]](#page-10-0).

⁴The doublet splitting in molecules is due to the Coriolis interaction. In nuclei the splitting is dominated by the "tunneling" of the octupole bump to the other side of the nucleus causing change of the valence nucleon spin projection to the nuclear axis. In fact, it is just an octuple vibration mode, so there is no sharp boundary between the static deformation in the minimum of the potential energy and soft octuple vibration when this minimum is very shallow or does not exist. Note that contrary to the Coriolis splitting in diatomic molecules the doublet splitting due to the tunneling does not increase with the rotational angular momentum—see the nuclear spectra in Ref. [\[37\]](#page-10-0).

where R is the nuclear radius. However, in the laboratory frame EDM and Schiff moment are forbidden by the parity and time-reversal invariance. Indeed, EDM and Schiff moment are polar *T* -even vectors which must be directed along the nuclear spin *I* which is a *T* -odd pseudovector.

Nucleus with an octupole deformation and nonzero nucleon angular momentum has a doublet of close opposite parity rotational states $|I^{\pm}\rangle$ with the same angular momentum $I(|I^{\pm}\rangle = \frac{1}{\sqrt{2}}(|\Omega\rangle \pm |-\Omega\rangle)$, where Ω is the projection of *I* onto the nuclear axis). The states of this doublet are mixed by *P*, *T* -violating interaction *W* . The mixing coefficient is

$$
\alpha_{+-} = \frac{\langle I^- | W | I^+ \rangle}{E_+ - E_-}.
$$
\n(3)

This mixing polarizes nuclear axis **n** along the nuclear spin **I**, $\langle n_z \rangle = 2\alpha_{+-}\frac{I_z}{I_{+}+1}$, and the intrinsic Schiff moment shows up in the laboratory frame [\[17,18\]](#page-9-0):

$$
S = 2\alpha_{+-} \frac{I}{I+1} S_{\text{intr}}.\tag{4}
$$

According to Ref. [\[18\]](#page-9-0) the *T*, *P*-violating matrix element is approximately equal to

$$
\langle I^{-}|W|I^{+}\rangle \approx \frac{\beta_3 \eta}{A^{1/3}} [\text{eV}].
$$
 (5)

Here η is the dimensionless strength constant of the nuclear *T*, *P*-violating potential *W* :

$$
W = \frac{G}{\sqrt{2}} \frac{\eta}{2m} (\sigma \nabla) \rho, \qquad (6)
$$

where *G* is the Fermi constant, *m* is the nucleon mass, and ρ is the nuclear number density. Equations (2)–(5), give an analytical estimate for the Schiff moment [\[18](#page-9-0)[,29\]](#page-10-0):

$$
S \approx 1.0 \times 10^{-4} \frac{I}{I+1} \beta_2 (\beta_3)^2 Z A^{2/3} \frac{[\text{keV}]}{E_- - E_+} e \eta \,\text{[fm}^3]. \tag{7}
$$

This estimate is in agreement with more accurate numerical calculations available for a number of nuclei [\[18\]](#page-9-0). For example, it gives $S = 280 e \eta \text{ fm}^3$ for ²²⁵Ra, which practically coincides with the result of the numerical calculation in Ref. [\[18\]](#page-9-0) $S = 300 e \eta \text{ fm}^3$.

Within the meson exchange theory, the π -meson exchange gives the dominating contribution to the *T*, *P*-violating nuclear forces [\[9\]](#page-9-0). In the standard notations *g* is the strong π -meson–nucleon interaction constant and \bar{g}_0 , \bar{g}_1 , \bar{g}_2 are the π -meson–nucleon *CP*-violating interaction constants in the isotopic channels $T = 0, 1, 2.5$

One can express the results in terms of more fundamental parameters such as the QCD θ -term constant θ using the relation $|g\bar{g}_0| = 0.37|\bar{\theta}|$ from Ref. [\[30\]](#page-10-0) or updated results [\[48–50\]](#page-10-0), $g\bar{g}_0 = 0.21\bar{\theta}$ and $g\bar{g}_1 = -0.046\bar{\theta}$, which give practically the same value of $S(\bar{\theta})$. Alternatively, the results can be expressed via the quark chromo-EDMs \tilde{d}_u and \tilde{d}_d : $g\bar{g}_0 = 0.8 \times 15^{-15}$ $10^{15}(\tilde{d}_u + \tilde{d}_d)$ /cm and $g\bar{g}_1 = 4 \times 10^{15}(\tilde{d}_u - \tilde{d}_d)$ /cm [\[1\]](#page-9-0).

Numerical calculations performed in Ref. [\[22\]](#page-9-0) found the Schiff moment of ²²⁵Ra in terms of $g\bar{g}_{0,1,2}$. We also express the Schiff moment of ²²⁵Ra in terms of $\overrightarrow{\theta}$ and \overrightarrow{d}_u and \overrightarrow{d}_d :

$$
S(^{225}\text{Ra}, g) \approx (-2.6g\bar{g}_0 + 12.9g\bar{g}_1 - 6.9g\bar{g}_2) e \text{ fm}^3,
$$

\n
$$
S(^{225}\text{Ra}, \bar{\theta}) \approx -\bar{\theta} e \text{ fm}^3,
$$

\n
$$
S(^{225}\text{Ra}, \tilde{d}) \approx 10^4 (0.50 \tilde{d}_u - 0.54 \tilde{d}_d) e \text{ fm}^2.
$$
 (8)

The analytical formula for the Schiff moment Eq. (7) gives us dependence of the Schiff moment on the nuclear parameters. Nucleus ²²⁵Ra has the octupole deformation $\beta_3 = 0.099$, quadrupole deformation $\beta_2 = 0.129$, nuclear spin $I = 1/2$, and interval between the opposite parity levels $E(1/2^-)$ – $E(1/2^+) = 55.2$ keV [\[18\]](#page-9-0). Using Eqs. (8) and (7) we present the result for other nuclei with the octupole deformation in the following form [\[28,29\]](#page-10-0):

$$
S(g) \approx K_S(-2.6g\bar{g}_0 + 12.9g\bar{g}_1 - 6.9g\bar{g}_2) e \,\mathrm{fm}^3,\qquad(9)
$$

$$
S(\bar{\theta}) \approx -K_S \,\bar{\theta} \, e \, \text{fm}^3,\tag{10}
$$

$$
S(\tilde{d}) \approx 10^4 K_S (0.50 \tilde{d}_u - 0.54 \tilde{d}_d) e \,\mathrm{fm}^2, \tag{11}
$$

where $K_S = K_I K_\beta K_A K_E$, $K_I = \frac{3I}{I+1}$, $K_\beta = 791 \beta_2 (\beta_3)^2$, $K_A =$ 0.00031*ZA*^{2/3}, and $K_E = \frac{55 \text{KeV}}{E^2 - E^+}$. By definition, numerical factors are chosen such that these coefficients are equal to 1 for 225 Ra and are of the order of unity for other heavy nuclei with octupole deformation. The values of K_S for deformed nuclei with strongly enhanced collective Schiff moments are presented in Table [I.](#page-3-0) Below we present the explanation of how these results were obtained.

The 227 Ac nucleus has a half-life of 21.8 years. It is produced commercially for cancer treatment. The half-life of 237 Np is 2.14 million years. It is produced in macroscopic quantities in nuclear reactors. In Ref. [\[29\]](#page-10-0) we obtained $K_S =$ 10 for ²²⁷Ac and $K_S = 6$ for ²²⁷Np. In Ref. [\[29\]](#page-10-0) we used calculated values $\beta_3 = 0.134$ for ²²⁷Ac and $\beta_3 = 0.12$ for 227 Np. However, in the experimental paper of Ref. [\[51\]](#page-10-0) it was found that β_3 is 0.07 for ²²⁹Ac and 0.1 for ²²⁷Ac and ²²⁵Ac. Updated values of the nuclear Schiff moments for 227Ac $(K_S = 6)$ and ²³⁷Np ($K_S = 4$) are obtained by multiplying the results of Ref. [\[29\]](#page-10-0) by the ratio of $(\beta_3)^2$ from Ref. [\[51\]](#page-10-0) and Ref. [\[29\]](#page-10-0).

 153 Eu is stable with 52% natural abundance. Its nuclear spectra indicate octupole deformation (since they have rotational doublets, with the same values of the moment of inertia for opposite-parity states in the doublets—see details in Ref. [\[29\]](#page-10-0)). According to Ref. [29] for ¹⁵³Eu $K_S = 3.7$.

⁵We also estimated contribution of the exchange by the η -meson which is four times heavier than the π meson and is usually assumed to give a smaller contribution. Indeed, the second power of the meson mass appears in the denominator of the effective interaction constant for the meson-induced nucleon interaction, so the expected suppression is $1/16$. However, the η -meson *CP*-violating exchange constant \bar{g} is an order of magnitude larger than the π -meson constant \bar{g}_0 [\[48\]](#page-10-0). In addition, the *η*-meson *CP*-violating contribution has the same sign for protons and neutrons contrary to the π -meson one. As a result, the suppression of the η -meson contribution to the nuclear Schiff moment is a few times only. Therefore, future

more accurate calculations of the Schiff moment should include the η -meson contribution as well as the finite nuclear size corrections found in Ref. [\[12\]](#page-9-0).

TABLE I. Collective Schiff moments for deformed nuclei containing opposite parity level with the same spin close to the ground state. To obtain the values of the Schiff moments one should use Eqs. (9)–(11) with the scaling factor K_s presented in this table.
For ²²⁵Ra we have $K_s = 1$ by definition—see Eq. (8). For nuclei ¹⁵⁵Gd, ¹⁵³Sm, and ¹⁶⁵Er not presented in this table we may expect $K_S \sim 0.01 - 0.1$.

Ζ	Isotope	K_S
63	$^{153}\mathrm{Eu}$	
66	$\rm ^{161}Dy$	3.7 ≤ 4 ≤ 0.4 3
66	163 Dy	
86	$\mathrm{^{223}Rn}$	
87	$^{221}\mathrm{Fr}$	0.14
87	$^{223}\mathrm{Fr}$	1.6
89	$^{225}\mathrm{Ac}$	3
89	$^{227}\mathrm{Ac}$	
90	$\rm ^{229}Th$	
91	$^{229}\mathrm{Pa}^\mathrm{a}$	6 \leq 2 40 \leq 2 \leq 3 4
92	233 U	
92	235 U	
93	$^{237}\mathrm{Np}$	
94	$\mathrm{^{239}Pu}$	$\lesssim 0.12$

^aEstimate for ²²⁹Pa is presented assuming that the existence of a very close nuclear doublet level will be confirmed.

According to Ref. [\[18\]](#page-9-0) the Schiff moment of 225 Ac is predicted to be three times larger than that of 225Ra, i.e., $K_{S} = 3.6$

Nuclear spectra of 222 Rn nucleus [\[37\]](#page-10-0) indicate octupole deformation. The nucleus 223 Rn has a neutron above 222 Rn nucleus and probably has the octupole deformation too or at least the soft octupole vibration mode. According to Ref. [\[18\]](#page-9-0) the Schiff moment of 223 Rn exceeds the Schiff moment of ²²⁵Ra three times, i.e., $K_S = 3$.

Nuclear spectra of 239 Pu nucleus [\[37\]](#page-10-0) indicate octupole deformation. However, the energy interval between the levels of the doublet in 239 Pu is nine times larger than in 225 Ra: $E(1/2^-) - E(1/2^+) = 470$ KeV. The deformation parameters have been calculated in Ref. [\[38\]](#page-10-0) for the even-even isotope ²⁴⁰Pu: $(\beta_3)^2 = (0.066)^2$ and $\beta_2 = 0.284$.⁷ Using these

parameters and Eq. [\(9\)](#page-2-0) we obtain $K_S \sim 0.12$. Note that ²³⁹Pu may be considered as ²³⁸Pu nucleus plus neutron. The nuclear spectra of ²³⁸Pu are consistent with the octupole deformation. Further study of this problem should bring us more reliable information about the octupole deformation parameter β_3 and more accurate estimates of the ²³⁹Pu Schiff moment.

According to Ref. [\[18\]](#page-9-0) the Schiff moment of 223 Fr is predicted to exceed the Schiff moment of 225Ra 1.6 times. This gives $K_S = 1.6$. Reference [\[18\]](#page-9-0) gives the Schiff moment of ²²¹Fr equal to 0.14 of the ²²⁵Ra Schiff moment, i.e., $K_S =$ 0.14. The nuclear spectra of 232 U (as well as the spectra of 234 U,

 236 U, and 238 U) are consistent with the octupole deformation. The nucleus 233 U may be considered as 232 U nucleus plus neutron. Nuclear spectra of 233 U nucleus [\[37\]](#page-10-0) are also consistent with the octupole deformation. However, the energy interval between the levels of the doublet in 233 U is five times larger than in ²²⁵Ra: $E(5/2^-) - E(5/2^+) = 299$ keV. The deformation parameters have been calculated in Ref. [\[38\]](#page-10-0) for the even-even isotope ²³²U: $(\beta_3)^2 = (0.17)^2$ and $\beta_2 = 0.238$. Using these parameters and Eq. [\(9\)](#page-2-0) we obtain $K_S \sim 2$. This estimate indicates a possibility of a large Schiff moment in 233 U. However, further experimental and theoretical investigation is needed.

Reference $[15]$ suggested a very interesting case of ²²⁹Pa which probably has a level of opposite parity and the same angular momentum very close to the ground state. The latest measurement [\[53\]](#page-10-0) gave a position of this level at 60 ± 50 eV. According to Ref. [\[18\]](#page-9-0) the Schiff moment of 229 Pa is predicted to exceed the Schiff moment of 225 Ra 40 times. We should note that the estimate $K_S = 40$ is valid if the close level 60 \pm 50 eV really exists and forms the rotational doublet related to the octupole deformation of 229 Pa with the ground state. So far there is no truly convincing evidence for such a case.

The nucleus 161 Dy is stable. The spectrum of 161 Dy is consistent with the octupole deformation; several rotational doublets are seen. The interval between the opposite-parity levels is only 25 keV; it is four times smaller than in 153 Eu. However, we have not found any calculations giving the octupole deformation in ¹⁶¹Dy or even-even nucleus ¹⁶⁰Dy. Therefore, a conservative result for the Schiff moment of 161 Dy (based on the extrapolation from 153 Eu) may be presented as an upper estimate $K_S \lesssim 4$. In another stable isotope 163Dy the interval between the opposite-parity levels is 10 times larger than in 161Dy; therefore, the upper estimate for the Schiff moment is an order of magnitude smaller.

We will try to provide some information for other nuclei of interest. While studying experimental nuclear spectra we noted in Ref. [\[29\]](#page-10-0) a possible trend: adding a proton to an even-even nucleus with octupole deformation usually supports the octupole deformation. Possibly, this is due to the proton increasing the Coulomb repulsion. However, adding a neutron to an even-even nucleus with octupole deformation sometimes blurs the features of the rotational

 6 In Ref. [\[18\]](#page-9-0) we performed calculations of the Schiff moment produced by the contact T , P -violating nuclear potential (6) . Corresponding interaction constants may be expressed in terms of the π -meson exchange constants $g\bar{g}_{0,1,2}$ (see, e.g., Ref. [\[52\]](#page-10-0)), which dominate the *T*, *P*-violating nuclear potential and are widely used now. To avoid lengthy discussion of this problem we often present the result as a ratio of the Schiff moment of a given nucleus to that of ²²⁵Ra, which has been calculated in Ref. [\[22\]](#page-9-0) in terms of $g\bar{g}_{0,1,2}$.

⁷Note that different calculations of the deformation parameters use different nucleon interaction models and may give significantly different results. Moreover, sometimes the authors come to different conclusions about the existence of the octupole deformation—see, e.g., Refs. [\[38–47\]](#page-10-0). This is why we look for the signatures of the octupole deformation in the experimental nuclear excitation spectra. We use deformation parameters calculated in Ref. [\[38\]](#page-10-0), since this paper contains the most comprehensive list of β_3 values for even-

even nuclei. The values of β_3 in different nucleon interaction models in Ref. [\[38\]](#page-10-0) are practically the same (within the accuracy of our estimates of the Schiff moment). Therefore, we present only one number for β_3 .

spectrum for the octupole.⁸ Indeed, nuclear spectra in 2^{29} Th, 2^{35} U, 1^{53} Sm, 1^{65} Er, and 1^{55} Gd do not provide clear evidence for the octupole deformation but have at least one doublet of the opposite-parity levels with the same nuclear spin. The probable outcome for such nuclei is existence of the soft octupole vibrational mode which produces some enhancement of the Schiff moments.

For example, 228Th has signatures of the octupole deformation in its rotational spectrum. However, adding a neutron to this nucleus and forming ²²⁹Th seems to blur the features of the rotational spectrum for the octupole. The estimate for the 229 Th Schiff moment has been done in Ref. [\[28\]](#page-10-0). According to Ref. $[43]$, 229 Th nucleus has the octupole deformation with the parameters $\beta_3 = 0.115$ and $\beta_2 = 0.240$. The nuclear spin $I = 5/2$ and the interval between the opposite parity levels is $E(5/2^-) - E(5/2^+) = 133.3$ keV. Equation [\(9\)](#page-2-0) gives us the value of the Schiff moment which is two times larger than that of 225 Ra. However, the static octupole deformation in 229 Th does not explicitly show up in the nuclear rotational spectra. In fact, we only see one doublet of the opposite parity states with the same spin. In this situation we may only offer an upper limit $K_S \lesssim 2$.

 235 U is practically stable (half life 0.7 billion years), with 0.75% natural abundance. The interval between opposite parity levels which are mixed by the *T*, *P*-odd interaction is $E(\frac{7}{2}) - E(\frac{7}{2}) = 81.7$ keV. This nucleus has a neutron above 234 U nucleus which according to Ref. [\[38\]](#page-10-0) has octupole deformation with $(\beta_3)^2 = (0.17)^2$ and $\beta_2 = 0.25$. However, experimental nuclear excitation spectra of 235 U do not show parity doublets for higher rotational states. Assuming that there is a soft vibrational octupole mode in 235 U we obtained an upper estimate of $K_S \lesssim 3$ [\[29\]](#page-10-0).

Nuclei with a valence neutron 155 Gd, 153 Sm, or 165 Er are close to the area of even-even nuclei with the octupole deformation; they have opposite-parity level with the same spin close to the ground state but their spectra do not indicate octupole deformation. We may expect an order of magnitude enhancement in comparison with the Schiff moments in nuclei which do not have a close level of opposite parity such as 199 Hg and 207 Pb considered below.

B. Schiff moments in Xe, Hg, Tl, and Pb nuclei

Let us now consider Schiff moments of nearly spherical nuclei where the Schiff moments are not enhanced. The Schiff moment of 199 Hg was first calculated in Ref. [\[11\]](#page-9-0). The most complete calculation in terms of π -meson interaction constants $g\bar{g}_{0,1,2}$ has been performed in Ref. [\[54\]](#page-10-0) using five different interaction models. We present an average of these fie values and also express the result in terms of $\bar{\theta}$ and quark chromo-EDMs \tilde{d}_u and \tilde{d}_d . The nucleus ²⁰⁷Pb has the same spin and parity ($I^P = 1/2^-$) and close value of the magnetic moment (0.59 nuclear magnetons) to that of 199 Hg $(0.51$ nuclear magnetons), i.e., the valence nucleons in ^{207}Pb and 199Hg occupy the same orbital. Therefore, within the theoretical accuracy Schiff moments of 199 Hg and 207 Pb are equal:

$$
S(^{207}\text{Pb}, g) \approx S(^{199}\text{Hg}, g)
$$

$$
\approx (0.023g\bar{g}_0 - 0.007g\bar{g}_1 + 0.029g\bar{g}_2) e \text{ fm}^3,
$$
(12)

$$
S(^{207}\text{Pb}, \bar{\theta}) \approx S(^{199}\text{Hg}, \bar{\theta}) \approx 0.005 \bar{\theta} e \,\text{fm}^3,\qquad(13)
$$

$$
S(^{207}\text{Pb}, \tilde{d}) \approx S(^{199}\text{Hg}, \tilde{d}) \approx 5\tilde{d}_d e \,\text{fm}^2. \tag{14}
$$

Note that these results include contributions of the *T*, *P*-odd nuclear forces between the nucleons and the nucleon electric dipole moments (as it was done in Ref. [\[54\]](#page-10-0)); however, the contribution of the proton and neutron EDM to the nuclear Schiff moments is significantly smaller than the contribution of the *T*, *P*-odd nuclear forces (first this was pointed out in Ref. [\[9\]](#page-9-0)). In Refs. $[55,56]$ contributions of the neutron d_n and proton d_p EDM to $S(^{199}Hg)$ have been presented separately: $S(^{207}Pb, d) \approx S(^{199}Hg, d) = (1.9d_n + 0.2d_p)$ fm². However, in Ref. [\[54\]](#page-10-0) the contribution of the neutron EDM (averaged over five interaction models) is three times smaller. Then taking the ratio of the neutron and proton contributions from Ref. [\[55\]](#page-10-0) (where the result is based on the fitting of the neutron and proton contributions to the nuclear magnetic moment) we obtain

$$
S(^{207}\text{Pb}, d) \approx S(^{199}\text{Hg}, d) = (0.6d_n + 0.06d_p) \text{ fm}^2. \quad (15)
$$

Schiff moments of 203 Tl and 205 Tl in the Saxon-Woods nuclear potential have been calculated in Ref. [\[11\]](#page-9-0) in terms of the contact nucleon-nucleon *T*, *P*-violating interaction. Here we express the results in terms of the more fundamental interaction constants:

$$
S(^{203}\text{TI}, g) \approx S(^{205}\text{TI}, g)
$$

 $\approx (0.13g\bar{g}_0 - 0.004g\bar{g}_1 - 0.27g\bar{g}_2) e \text{fm}^3$, (16)

$$
S(^{203}\text{TI}, \bar{\theta}) \approx S(^{205}\text{TI}, \bar{\theta}) \approx 0.027 \bar{\theta} e \text{fm}^3,
$$
 (17)

$$
S(^{203}\mathrm{TI}, \tilde{d}) \approx S(^{205}\mathrm{TI}, \tilde{d}) \approx (12\tilde{d}_d + 9\tilde{d}_u) e \,\mathrm{fm}^2. \tag{18}
$$

In the case of 199 Hg the sophisticated Hartree-Fock-Bogoliubov calculations in Ref. [\[54\]](#page-10-0) gave a smaller value of the nuclear Schiff moment than the calculations in Ref. [\[11\]](#page-9-0) performed in the Saxon-Woods potential. In Hg valence nucleon is neutron; therefore, nuclear EDM and Schiff moment appear due to the nuclear core polarization $[11]$. ²⁰³Tl and 205 Tl nuclei have valence proton; therefore, EDM and Schiff moment appear even without the core polarization. Still the many-body corrections may play an important role [\[11\]](#page-9-0) and provide up to a factor of 3 suppression of $S(203, 205)$ Tl). (See Table II .)

⁸We must admit that a conventional point of view is that the influence of the shell structure dominates the effects of adding an odd particle to a core. Moreover, review [\[41\]](#page-10-0) suggested that both odd proton and odd neutron have a stabilizing effect on octupole deformation. The picture is not completely settled yet. Therefore, we do not automatically assume that the odd nucleon always produces a stabilizing effect on the octupole. Instead, we look for the experimental nuclear spectra with two nearly parallel rational bands (i.e., corresponding to approximately the same moment of inertia) starting from the same value of nuclear spin and having opposite parity (i.e., we look for rotational bands with several opposite-parity doublets).

TABLE II. Schiff moments in Xe, Hg, Tl, and Pb nuclei. $S(g)$ = $(a g \bar{g}_0 + b g \bar{g}_1 c g \bar{g}_2) [e \text{ fm}^3], S(\bar{\theta}) = K_t \bar{\theta} [e \text{ fm}^3], \text{ and } S(\tilde{d}_d) = K_d \tilde{d}_d +$ $K_d \tilde{d}_u[e \text{ fm}^2]$. Many-body corrections are expected to reduce the values of the Schiff moments of Tl nuclei presented in this table.

Isotope	a	S(g) h	c	$S(\theta)$ K,	$S(\tilde{d}_d)$ K_d	K_u
^{129}Xe	-0.03	0.01	-0.04	-0.007	-6	0
203 Tl, 205 Tl	0.13	-0.004	-0.27	0.027	12	9
$^{207}Pb, ^{199}Hg$	0.023	-0.007	0.029	0.005		0

We may also separate the contribution of the valence proton EDM to the nuclear Schiff moment. The result from Refs. [\[7,8\]](#page-9-0), $S = -d_p(R/6)$, depends on the difference $R \equiv$ $r_v^2 - r_q^2$ of the squared values of the spin distribution radius r_v^2 and charge radius r_q^2 . Unfortunately, even the sign of *R* is not firmly established [\[9](#page-9-0)[,52,57\]](#page-10-0). Using the average result of the Hartree Fock calculations with different interactions performed by Brown and presented in Ref. $[57]$, $R = 2.3$ fm², we obtain $S = 0.4d_p$ fm².

Finally, we present the Schiff moment of 129 Xe which has a valence neutron. First, it was calculated in the Saxon-Woods potential in Ref. [\[11\]](#page-9-0). Then Ref. [\[58\]](#page-10-0) claimed a very strong suppression of the Schiff moments by the RPA corrections. However, Ref. $[54]$ demonstrated that in the case of 199 Hg a more complete Hartree-Fock-Bogoliubov calculation does not show such strong suppression as in the RPA method. In Ref. [\[11\]](#page-9-0) the ratio $S^{(129}\text{Xe})/S^{(199)}_{(199-199)} = -1.25$. We use this result and Eq. [\(12\)](#page-4-0) for the $S(^{199}Hg)$ to estimate the ¹²⁹Xe Schiff moment:

$$
S(^{129}\text{Xe}, g) \approx (-0.03g\bar{g}_0 + 0.01g\bar{g}_1 - 0.04g\bar{g}_2) e \,\text{fm}^3, \tag{19}
$$

$$
S(129Xe, \bar{\theta}) \approx -0.007 \bar{\theta} e \, \text{fm}^3,\tag{20}
$$

$$
S({}^{129}\text{Xe}, \tilde{d}) \approx -6\tilde{d}_d e \,\text{fm}^2. \tag{21}
$$

We also present contributions of the neutron and proton EDMs to the 129 Xe Schiff moment from Ref. [\[59\]](#page-10-0):

$$
S(^{129}\text{Xe}, d) \approx (0.63d_n + 0.13d_p) \text{ fm}^2. \tag{22}
$$

III. ATOMIC EDM CALCULATIONS

We have done calculations of the atomic electric dipole moments for few atoms in our earlier works [\[55,60–63\]](#page-10-0). The EDM of an atom induced by *CP*-odd Hamiltonian H_{CP} is given by

$$
\mathbf{d}_{\text{at}} = 2 \sum_{n} \frac{\langle 0 | H_{CP} | n \rangle \langle n | \mathbf{D} | 0 \rangle}{E_0 - E_n},
$$
 (23)

where $\mathbf{D} = \sum_i \mathbf{d}_i = -|e| \sum_i \mathbf{r_i}$ is the electric dipole operator with summation over all atomic electrons, $|0\rangle$ is the atomic ground state, summation in (23) goes over a complete set of states $|n\rangle$, and E_n are the energies of atomic states.

In this work we focus on atoms with enhanced nuclear Schiff moments. Therefore, we consider only one type of the *CP*-odd Hamiltonian, the Hamiltonian of the interaction of atomic electrons with nuclear Schiff moment ($H_{CP} \equiv H_{SM}$).

Taking into account the finite nuclear size effect, this Hamiltonian has the following form [\[12\]](#page-9-0):

$$
H_{\rm SM} = \sum_{i} h_i^{\rm SM} = -\sum_{i} \frac{3\mathbf{S} \cdot \mathbf{r}_i}{B} \rho(r),\tag{24}
$$

where $B = \int \rho(r)r^4 dr$, ρ is nuclear density, and **S** is the vector of the nuclear Schiff moment. If H_{SM} in (23) is replaced by the electric dipole operator **D**, the expression gives dipole static polarizability of the atom. This gives us a good test of the accuracy of our calculations.

The ways to reduce (23) to expressions containing only single-electron integrals depends on electron structure of the atom. Below we consider particular cases which cover almost all atoms in the Periodic Table. Note that the accuracy of the nuclear calculations of the Schiff moment is never better than a factor of 2 (see, e.g., Ref. [\[2\]](#page-9-0)). Therefore, 30% accuracy of the atomic calculations is sufficient.

A. Closed-shell atoms

For closed-shell atoms, such as Xe, Rn, Fr^+ , Ac³⁺, and $Th⁴⁺$, it is sufficient to use the so-called random-phase approximation (RPA). The RPA equations present a linear response of the Hartree-Fock atomic states to a perturbation by an external field. They can be written in the form

$$
(H_0 - \epsilon_c)\delta\psi_c = -(F + \delta V^F)\psi_c.
$$
 (25)

Here H_0 is the relativistic Hartree-Fock (HF) Hamiltonian, ψ_c is the HF electron state in the core, $\delta \psi_c$ is the correction to the HF state in the core induced by the external field F , and δV^F is the correction to the self-consistent HF potential due to the corrections to all core states. Index *c* numerates states in the core. The operator of the external field in our case is either H_{SM} or **D** operators ($F = h^{SM}$ or $F = d$). RPA equations (25) are solved self-consistently for all states in the core. The EDM of the closed-shell atom is given then by either of expressions

$$
\mathbf{d}_{\rm at} = \frac{2}{3} \sum_{c} \langle \psi_c | \mathbf{d} | \delta \psi_c^{\rm SM} \rangle \tag{26}
$$

$$
\mathbf{d}_{\rm at} = \frac{2}{3} \sum_{c} \langle \psi_c | h^{\rm SM} | \delta \psi_c^{\rm D} \rangle, \tag{27}
$$

where summation goes over all core states, $\delta \psi_c^{\text{SM}}$ is the solution of the RPA equations with the Schiff moment operator h^{SM} , and $\delta \psi_c^{\text{D}}$ is the solution of the RPA equations with the electric dipole operator *d*. Expressions (26) and (27) give identical answers. Note that replacing $\delta \psi_c^{\text{SM}}$ in (26) by $\delta \psi_c^{\text{D}}$ or h^{SM} in (27) by *d* gives the dipole static polarizability of the closed-shell atom [compare with Eq. (23)].

Expressions (26) and (27) do not include interelectron correlations beyond core polarization, which is treated by the RPA equations. They give accurate results for noble-gas atoms such as Xe and Rn, where such correlations are small. Formally, they could be used to calculate the EDM of such atoms as Yb, Hg, and Ra, which have upper closed s^2 shell. However, the results would be less accurate. This is because in these atoms correlations between external electrons (6*s* or 7*s*) and between external electrons and electrons in the core

or

TABLE III. Static scalar polarizabilities (in a.u.) and EMDs due to Schiff moment (in units of 10−17[*e* cm]*S*/[*e* fm³]) for selected atoms and ions.

Configuration			Polarizability			EDM					
					This work		Other ^a		This work		Other
Z	Atom	Core	Valence	Core	Valence	Total		Core	Valence	Total	
54	Xe	$[Pd]$ 5s ² 5p ⁶	None	27	None	27	27.32(20)	0.38	None	0.38	0.378 [74]
63	$\mathop{\mathrm{Eu}}\nolimits$	[Ba] $4f^7$	$6s^2$	$10\,$	178	188	184(20)	0.22	-1.85	-1.63	
63	Eu^{3+}	[Ba] $4f6$	None	7	None	τ		0.33	None	0.33	
64	Gd	[Ba] $4f^7$	$6s^25d$	9	150	159	158(20)	0.24	-2.36	-2.22	
66	Dy	[Ba] $4f^{10}$	$6s^2$	8	156	164	163(15)	0.27	-2.50	-2.23	
70	Yb	$[Ba]4f^{14}$	$6s^2$	6	141	147	139(6)	0.37	-2.26	-1.88	-1.903 [74]; -1.9 [61]; -2.12 [62]
80	Hg	[Yb] $5d^{10}$	$6s^2$	$\overline{7}$	27	34	33.91(34)	0.37	-2.87	-2.50	-2.63 [62]; -5.07 [75]
											-2.914 [76]; -2.3 [77]
81	$T1$ ⁺	[Yb] $5d^{10}$	$6s^2$	5	14	19		0.40	-3.19	-2.79	
82	Pb	[Hg]	$6s^26p^2$	\mathfrak{Z}	43	46	47(3)	0.39	0.15	0.54	0.766 [76]
82	Pb^{2+}	[Hg]	$6s^2$	3	10	13		0.39	-3.38	-2.99	-3.08 [76]
86	Rn	[Hg] $6p^6$	None	35	None	35	35(2)	3.3	None	3.3	3.3 [62]
87	Fr^+	[Rn]	7s	20	None	20		2.87	None	2.87	
88	Ra	[Rn]	$7s^2$	12	238	250	246(4)	1.75	-10.0	-8.25	-8.093 [74]; -8.5 [55]; -8.8 [62]
89	$\mathop{\mathrm{Ac}}$	[Rn]	$7s^26d$	9	186	195	203(12)	1.5	-11.6	-10.1	
	Ac^+	[Rn]	$7s^2$	9	112	121		1.5	-11.3	-9.8	
	Ac^{3+}	[Rn]	None	10	None	$10\,$		2.5	None	2.5	
90	${\rm Th}$	[Rn]	$7s^26d^2$	6.7			217(54)	1.4	-10.6	-9.2	
	$Th2+$	[Rn]	$7s^2$	6.7	45.7	52.4		1.4	-8.33	-6.93	
	Th^{4+}	[Rn]	None	7.8	None	7.8		2.4	None	2.4	
91	Pa	[Rn] $5f^2$	$7s^26d$	20	150	170	154(20)	0.5	-11.9	-11.4	
92	$\mathbf U$	[Rn] $5f^3$	$7s^26d$	17	148	165	129(17)	0.7	-12.8	-12.1	
93	Np	[Rn]5f ⁴	$7s^26d$	14	146	160	151(20)	0.9	-8.4	-7.5	
94	Pu	[Rn] $5f^6$	$7s^2$	21	123	144	132(20)	2.3	-11.5	-9.2	-10.9 [55]

^aRecommended values from Ref. [\[78\]](#page-10-0).

are not small. Inclusion of these correlations is considered in the next section.

B. Atoms with few valence electrons

In this section we consider atoms like Hg, Yb, Ra, and Pb. Although some of these atoms can be considered as closedshell atoms, the correlations affecting external electrons are important and should be taken into account for accurate results. This can be illustrated by calculation of atomic polarizability. The RPA value for the static polarizability of Hg is 44 a.u. [\[64\]](#page-10-0). Inclusion of correlations for two external electrons reduces the value to 34 a.u., bringing it to excellent agreement with experiment (see Table III). The trend is confirmed by other calculations (see, e.g., Refs. [\[55,65\]](#page-10-0)). The standard approach is to divide all electrons into two groups: most electrons go into the closed-shell core, while the few remaining electrons are treated as valence ones. The usual method of the calculations, $CI + MBPT$, is the configuration-interaction (CI) technique enhanced by the many-body perturbation theory (MBPT) approach for the more accurate calculation of the CI matrix elements.

The EDM is calculated as a sum of the core and valence contributions. Core contribution is calculated as in the previous section with one important amendment which is often called the core-valence contribution. Summation over excited states must not include states occupied by valence electrons due to the Pauli exclusion principle. Since we do not perform

the summation directly but rather solve the RPA equations, we take into account the Pauli principle by imposing the orthogonality condition on the RPA corrections to the core wave functions [see Eq. (25)]

$$
\delta \psi'_c = \delta \psi_c - \langle \delta \psi_c | v \rangle \langle v |, \tag{28}
$$

where v is the state occupied by an external electron.

For valence electrons we use the $CI + MBPT$ method [\[60\]](#page-10-0). The effective Hamiltonian has one- and two-electron terms:

$$
H^{CI} = \sum_{i} [H_0(r_i) + \Sigma_1(r_i)] + \sum_{i < j} \left(\frac{e^2}{r_{ij}} + \Sigma_2(r_i, r_j)\right). \tag{29}
$$

Here H_0 is the relativistic HF Hamiltonian; Σ_1 and Σ_2 are the operators responsible for the core-valence correlations [\[60\]](#page-10-0). Summation in (29) goes over valence electrons. We use the *B*-spline technique [\[66\]](#page-10-0) to construct single-electron orbitals in valence space. The wave functions for the valence electrons are found as eigenstates of the CI Hamiltonian (29). They are used to calculate valence contribution to the EDM using formula [\(23\)](#page-5-0).

C. Atoms with open *f* **shell**

These atoms are the main subject of the present work. This is because these atoms have isotopes in which the Schiff moment is strongly enhanced due to nuclear octupole deformation. These include rare-earth and actinide atoms 153 Eu, 233 U, 237 Np, etc. [\[29\]](#page-10-0). Accurate calculations for such atoms are difficult due to the large number of electrons in open shells. One possible approach is to include all these electrons into valence space (see, e.g., Refs. [\[67,68\]](#page-10-0)). This works well for energy levels and transition amplitudes [\[69–73\]](#page-10-0). However, calculation of atomic characteristics, such as polarizabilities, EDM, etc., for which summation over a complete set of intermediate states is needed, is problematic. This is because the methods are constructed for low-energy states and not appropriate for high-energy states needed for completeness.⁹ Therefore, we use an alternative approach developed in Ref. [\[63\]](#page-10-0) for polarizabilities of open-shell atoms. It was shown in [\[63\]](#page-10-0) that calculations can be performed as for two or three valence electron systems by attributing *f* electrons to the core. Calculations for the core containing open *f* shell are done as for a closed-shell system but with the use of the fractional occupation numbers for the *f* shell. Calculations for the valence electrons are done with the use of the $CI +$ MBPT method as in the previous section. Note that atomic polarizability is given by the square of the electric dipole operator matrix element. In fact, the approached developed in [\[63\]](#page-10-0) is valid for any pair of vector operators, such as, e.g., Schiff moment operator and electric dipole operator as in the present work.

D. Results of atomic calculations

The results of atomic calculations are presented in Ta-ble [III.](#page-6-0) In addition to atomic EDMs calculated here we included atoms for which previous calculations are available. This is done to illustrate the accuracy of the calculations. Therefore, the list of previous calculations is far from being complete. In most cases we have included most recent or most accurate calculations. The table also includes the data on the calculated and experimental polarizabilities of neutral atoms. Studying these data provides another way of judging the accuracy of the calculations. Note, however, that atomic EDMs are sensitive to the wave functions in the vicinity of the nucleus while polarizabilities are not sensitive to it. Therefore, having good accuracy for polarizabilities is a necessary but not sufficient condition of having good accuracy for EDMs. For this reason we have studied the sensitivity of the calculated EDMs to the variation of the single-electron basis. We found that more accurate results can be obtained if most important valence states (6*s*, 6*p* for atoms from Eu to Pb and 7*s*, 7*p* for atoms from Ra to Pu) are calculated by the relativistic Hartree-Fock computer code without using *B*-splines. For higher virtual states we use linear combinations of *B*-splines which are eigenstates of the relativistic HF Hamiltonian. The same results can be achieved if the number of *B*-splines is increased significantly and *B*-splined states are used everywhere, including the 6*s*, 6*p*, 7*s*, and 7*p* states. This modification of the basis is the main reason for some difference (∼10%) with

TABLE IV. Schiff moments (S) and EDMs (d_A) of some atoms in terms of the OCD θ -term constant $\bar{\theta}$. We remind the reader that the current experimental limit is $|\bar{\theta}| < 10^{-10}$.

		S	$d_A[e \text{ cm}]$	
Ζ	Atom	$[e \, \text{fm}^3 \bar{\theta}]$	$10^{-17} S[e fm^3]$	10^{-17} $\bar{\theta}$
63	$^{153}\mathrm{Eu}$	-3.7	-1.63	6
63	153 Eu ³⁺	-3.7	0.33	-1.2
66	161 Dy	≤ 4	-2.23	≤ 9
80	199 Hg	0.005	-2.50	-0.013
81	$205,203$ Tl ⁺	0.02	-2.79	-0.06
82	$^{207}Ph^{2+}$	0.005	-2.99	-0.015
86	223 _{Rn}	-3	3.3	-10
87	223 Fr ⁺	-1.6	2.87	-4.6
88	^{225}Ra	-1	-8.25	8
89	227 Ac	-6	-10.1	60
89	$^{227}Ac^{+}$	-6	-9.8	60
90	$^{229}Th^{2+}$	\leq 2	-6.93	\lesssim 14
91	$^{229}Pa^a$	-40	-11.4	460
92	233 U	\leq 2	-12.1	\lesssim 20
93	237 Np	-4	-7.5	30
94	^{239}Pu	$\lesssim 0.1$	-9.2	\lesssim 1

^aEstimates for ²²⁹Pa are presented assuming that the existence of a very close nuclear doublet level will be confirmed.

our earlier results of Refs. [\[55,61,62\]](#page-10-0). Taking this difference as an estimation of the accuracy of the calculations we see that the uncertainty is less than 30%. This is a satisfactory accuracy since the accuracy of the interpretation of the EDM measurements in terms of fundamental parameters of the *T P*-odd interactions is limited by nuclear physics, where the accuracy is significantly lower.

Comparing our results with the results of other groups is another way to estimate uncertainties. The most striking difference is about two times disagreement of our result for Hg with the result of Latha *et al.* [\[75\]](#page-10-0). Note, however, that the later result of Latha with different co-workers is about 1.7 times smaller [\[76\]](#page-10-0). This value as well as the results of other Hg EDM calculations agree with our results. Results of Ref. [\[76\]](#page-10-0) for Pb and Pb^{2+} are also in agreement with our results. For other atoms the differences with results of other groups are within declared uncertainty.

Table [III](#page-6-0) also includes some ions. Corresponding atoms can be found in molecules or crystals used in an experimental search for *CP*-odd nuclear forces. The effect of the *CP*-violating forces in such systems can be reduced to the EDM of the ion of the heavy element. For example, the Ac^+ ion is a part of the AcF molecule suggested for the EDM measurements [\[29\]](#page-10-0), the Pb^{2+} ion is a part of the PbO molecule and PbTiO₃ solid used in the CASPEr experiment [\[35\]](#page-10-0), and the Th^{2+} ion is a part of the ThO molecule used in the experiment in Ref. [\[79\]](#page-10-0). Note that the ground state of the isolated Th^{2+} ion is [Rn] 5*f* 6*d* ${}^{3}\text{H}_{4}^{o}$. However, in the ThO molecule it is $[{\rm Rn}]7s^2$ ¹S₀ [\[79\]](#page-10-0).

To help in estimation of the sensitivity of different experiments, in Table IV we presented values of the Schiff moments and atomic EDM for a number of most interesting atoms and ions in terms of the QCD θ -term constant θ . Dependence on other *CP*-violating parameters for all atoms and nuclear

⁹For atoms with one or few valence electrons it may be more convenient to solve HF equations in an external electric field and avoid summation over intermediate states.

isotopes considered in the present paper may be found by the multiplication of the results of the atomic calculations from Table [III](#page-6-0) by the values of the Schiff moments presented in Sec. [II.](#page-1-0)

IV. MOLECULAR *T, P***-VIOLATING SPIN-AXIS INTERACTION CONSTANTS**

Molecules have very close opposite-parity rotational levels or rotational doublets. Polar molecules also have large intrinsic electric dipole moments and may be nearly completely polarized by an external electric field. In this case an internal molecular field (which for a polarized molecule is directed along the external electric field) exceeds an external field by many orders of magnitude and dramatically increases sensitivity to the *T*, *P*-violating interactions [\[7,](#page-9-0)[36\]](#page-10-0). For example, experiment with ThO molecule [\[80\]](#page-10-0) gives a limit on the electron EDM which is two orders of magnitude better than the limit from the atomic EDM measurement. Sandars [\[7\]](#page-9-0) suggested using this molecular property to measure *T*, *P*violating interaction of the nuclear spin with molecular axis. Similar to the atomic EDM, this interaction is induced by the electrostatic interaction of the nuclear Schiff moment with molecular electrons which are able to reach heavy nucleus. The effect increases faster than Z^2 [\[9\]](#page-9-0) so actinide molecules which have enhanced Schiff moment may have an advantage.

The interaction constant W_S for the effective T, P violating interaction in molecules is defined by the following expression:

$$
W_{T,P} = W_S \frac{S}{J} \mathbf{J} \cdot \mathbf{n},\tag{30}
$$

where **J** is the nuclear spin and **n** is the unit vector along the molecular axis in linear molecules.¹⁰

Molecule AcF has a valence electron structure similar to TlF. Indeed assuming that the upper valence electron goes to F atom, Ac^+ and Tl^+ ions have valence $7s^2$ and $6s^2$ correspondingly. The ratio of the atomic EDM for Ac^+ and TI^+ is 3.5 (see Table [III\)](#page-6-0), so we have

$$
W_S(\text{AcF}) = 3.5 W_S(\text{TIF}) \sim 100000 \tag{31}
$$

in atomic units (here a.u. $= e/a_B^4$). Here we used $W_S(TIF)$ = 45810 from Ref. [\[81\]](#page-10-0), which includes many-body effects reducing the result by 22%. Previous calculations for TlF have been published in Refs. [\[8,](#page-9-0)[82,](#page-10-0)[83\]](#page-11-0). Note, however, that our previous experience with the extrapolation from TlF to RaO (see Refs. $[84,85]$) indicates that $W_S(AcF)$ in Eq. (31) may be strongly overestimated since Tl and Ac atoms' chemical properties are very different.

Substitution of the Schiff moment from Table [I](#page-3-0) to the energy shift $W_{T,P} = W_S \frac{S}{J} \mathbf{J} \cdot \mathbf{n}$ gives for the fully polarized molecule the energy difference between the $J_z = J$ and $J_z =$ −*J* states:

$$
2W_S(\text{AcF})S(^{227}\text{Ac}) \sim 10^8 \bar{\theta} \, h \, \text{Hz},\tag{32}
$$

where *h* is the Plank constant. This is 1000 times larger than the energy shift for TlF:

$$
2W_S(\text{TIF})S(^{205}\text{TI}) = 1.0 \times 10^5 \bar{\theta} \, h \,\text{Hz.}
$$
 (33)

Using the Hg atom EDM measurements Ref. [\[27\]](#page-10-0) and Eqs. [\(12\)](#page-4-0) and [\(15\)](#page-4-0) we can extract the limits $|\bar{\theta}| < 5 \times 10^{-11}$, $|d_n|$ < 10^{−26}*e* cm, and $|d_p|$ < 5 × 10^{−25}*e* cm. The neutron EDM measurements give limit $|\bar{\theta}| < 10^{-10}$. With this limit the maximal shift in AcF is $\sim 10^{-3}$ Hz. The measured shift in the 1991 TlF experiment [\[86\]](#page-11-0) was $(-1.3 \pm 2.2) \times 10^{-4}$ Hz, i.e., such accuracy is already sufficient. It is expected that a new generation of molecular experiments will improve this accuracy by several orders of magnitude [\[87\]](#page-11-0). Therefore, we may expect a very significant improvement of the current limit $|\bar{\theta}| < 10^{-10}$ and also improvement of the limits on other fundamental parameters of the *CP*-violation theories such as the πNN interaction constants \bar{g} and the quark chromo-EDMs \tilde{d} .

Other interesting examples include molecules AcN and $AcO⁺$, which have electronic structure similar to RaO (naively, they may be considered as ion molecules $Ac^{3+}N^{3-}$, $Ac^{3+}O^{2-}$, and $Ra^{2+}O^{2-}$, where corresponding atomic ions have the same electron states). The ratio of Ac^+ EDM to Ra EDM is $9.8/8.25 = 1.18$ (see Table [III\)](#page-6-0), so we may use the results of the RaO calculation W_S (RaO) = 45192 from Ref. $[85]$ to obtain reliable estimates for AcN and AcO⁺:

$$
W_S(\text{AcN}) \approx W_S(\text{AcO}^+) \approx 53000. \tag{34}
$$

If we assume that in the PbO molecule two electrons go from Pb to O, it has electronic structure similar to TfF. The ratio of atomic EDM for Pb^{2+} and Tl⁺ is 1.06 (see Table [III\)](#page-6-0), so we have

$$
W_S(\text{PbO}) = 1.06W_S(\text{TIF}) = 49000. \tag{35}
$$

A more accurate direct calculation in Ref. [\[88\]](#page-11-0) gives 47250. Reference $[88]$ also gives values of W_s in the solid PbTiO₃ used in the CASPEr experiment [\[35\]](#page-10-0):

$$
W_S(\text{PbTiO}_3) = 30270. \tag{36}
$$

The Schiff moment of $207Pb$ is slightly smaller than that of 205Tl since 207Pb has valence neutron.

The ground state of Eu^{3+} ion has zero electron angular momentum. Possibly, some Eu molecules may have ground or metastable state with the zero angular momentum too. Unfortunately, we have not found a specific example yet. Possibly such a state exists in the molecule EuN, where three electrons from the Eu atom can make a close shell on the N atom. Another example is molecular ion $EuO⁺$, which has electronic structure similar to EuN. The estimate for the energy shift is

$$
2W_S(\text{EuN})S(^{153}\text{Eu}) \approx 2W_S(\text{EuN})S(^{153}\text{Eu}) \sim 10^7 \bar{\theta} \, h \, \text{Hz.}
$$
\n(37)

This is 100 times larger than the energy shift for TlF.

The Th atom has two extra $6d^2$ electrons in comparison with the Ra atom. $6d^2$ electrons give a very small contribution to atomic EDM; therefore, the atomic calculation of the Th EDM gives the result close to that for Th^{2+} and Ra EDM. $Th²⁺$ and Ra have similar electronic structure with filled $7s²$ subshell and close values of atomic EDM in units of the

¹⁰Some older molecular calculations presented constant $X = W_S/6$.

Schiff moment (see Table [III\)](#page-6-0). Therefore, W_S for the ThO molecule is comparable to that of the RaO molecule calculated in Ref. [\[85\]](#page-11-0):

$$
W_S
$$
(ThO) ~ W_S (RaO) = 45192. (38)

Another possibility may be to use the doublet in ${}^{3}\Delta_1$ metastable state of ²²⁹ThO (used to improve the limit on the electron EDM in Ref. [\[79\]](#page-10-0)) and the ground-state doublet ${}^{3}\Delta_1$ in ThF^+ .

DeMille and Fleig suggested recently to make cold molecule AgRa for EDM measurements from cold atoms cooled by laser [\[89\]](#page-11-0). Some atoms which we considered may be cooled, including Fr, Ac, and Ra. There is a number of atoms which may be cooled and combined with Fr, Ac, and Ra atoms to form a cold molecule. Corresponding *T*, *P*-odd interaction constant W_S is determined mainly by the heavy atoms and may be estimated using the molecular calculations for TlF and RaO using the results for Fr^+ , Ac⁺, Ra²⁺, and Tl⁺ EDM.

Finally, in a recent paper [\[90\]](#page-11-0) it was suggested that linear molecules MOH, molecular ions MOH^+ (M is a heavy atom, e.g., Ra in the molecule $RaOH⁺$). and symmetric top molecules (such as MCH_3 or $MOCH_3$) may be better systems than molecules MO, since such polyatomic molecules have a doublet of close opposite-parity energy levels in the bending mode and may be polarized by a weak electric field. The reduction of the strength of the necessary electric field simplifies the experiment and dramatically reduces systematic effects. These molecules may be cooled by a laser.

V. CONCLUSION

In this paper we presented estimates for the nuclear Schiff moments, calculations of the atomic EDMs produced by these Schiff moments, and estimates of the *T*, *P*-odd nuclear spin–molecular axis interaction constants for molecules containing these atoms. For the comparison and estimate of the accuracy we included atoms where the measurements as well as the atomic and nuclear calculations have been done previously: 199 Hg, 129 Xe, and 225 Ra.

Then we presented results for atoms and ions where the nuclear Schiff moments and atomic EDMs expressed in terms of the CP -violating π -meson–nucleon interaction constants \bar{g}_0 , \bar{g}_1 , \bar{g}_2 , QCD parameter $\bar{\theta}$, and quark chromo-EDM were not available including Pb, TI^+ , and Pb²⁺. These ions may be considered as parts of the TlF and PbO molecules and $PbTiO₃$ solid used in the CASPEr experiment [\[35\]](#page-10-0) searching for the axion dark matter which induces Pb^{2+} oscillating EDM.

The main part of the work is the calculations for stable or very long lifetime rare-earth and actinide atoms and ions where the nuclear Schiff moments are enhanced up to three orders of magnitude: Eu, Eu^{3+} , Gd, Dy, Ac, Ac⁺, Ac³⁺, Th, Th^{2+} , Th⁴⁺, Pa, U, Np, and Pu. In the molecules containing corresponding ions, e.g., in ²²⁷AcF or ²²⁷AcN, the *T*, *P*violating effects are up to three orders of magnitude larger than in TlF, where the experiments have been performed earlier [\[86\]](#page-11-0) and are carried out now [\[87\]](#page-11-0).

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- [1] M. Pospelov and A. Ritz, [Ann. Phys. \(Amsterdam\)](https://doi.org/10.1016/j.aop.2005.04.002) **[318](https://doi.org/10.1016/j.aop.2005.04.002)**, [119](https://doi.org/10.1016/j.aop.2005.04.002) [\(2005\)](https://doi.org/10.1016/j.aop.2005.04.002).
- [2] [J. Engel, M. J. Ramsey-Musolf, and U. van Kolck,](https://doi.org/10.1016/j.ppnp.2013.03.003) Prog. Part. Nucl. Phys. **[71](https://doi.org/10.1016/j.ppnp.2013.03.003)**, [21](https://doi.org/10.1016/j.ppnp.2013.03.003) [\(2013\)](https://doi.org/10.1016/j.ppnp.2013.03.003).
- [3] I. B. Khriplovich, *Parity Nonconservation in Atomic Phenomena* (Gordon and Breach, Amsterdam, 1991).
- [4] I. B. Khriplovich and S. K. Lamoreaux, *CP Violation without Strangeness* (Springer-Verlag, Berlin, 1997).
- [5] J. S. M. Ginges and V. V. Flambaum, [Phys. Rep.](https://doi.org/10.1016/j.physrep.2004.03.005) **[397](https://doi.org/10.1016/j.physrep.2004.03.005)**, [63](https://doi.org/10.1016/j.physrep.2004.03.005) [\(2004\)](https://doi.org/10.1016/j.physrep.2004.03.005).
- [6] L. I. Schiff, [Phys. Rev.](https://doi.org/10.1103/PhysRev.132.2194) **[132](https://doi.org/10.1103/PhysRev.132.2194)**, [2194](https://doi.org/10.1103/PhysRev.132.2194) [\(1963\)](https://doi.org/10.1103/PhysRev.132.2194).
- [7] P. G. H. Sandars, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.19.1396) **[19](https://doi.org/10.1103/PhysRevLett.19.1396)**, [1396](https://doi.org/10.1103/PhysRevLett.19.1396) [\(1967\)](https://doi.org/10.1103/PhysRevLett.19.1396).
- [8] E. A. Hinds and P. G. H. Sandars, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.21.471) **[21](https://doi.org/10.1103/PhysRevA.21.471)**, [471](https://doi.org/10.1103/PhysRevA.21.471) [\(1980\)](https://doi.org/10.1103/PhysRevA.21.471).
- [9] O. P. Sushkov, V. V. Flambaum, and I. B. Khriplovich, Zh. Eksp. Teor. Fiz. **87**, 1521 (1984) [Sov. JETP **60**, 873 (1984)].
- [10] [V. V. Flambaum, I. B. Khriplovich, and O. P. Sushkov,](https://doi.org/10.1016/0370-2693(85)90908-6) *Phys.* Lett. B **[162](https://doi.org/10.1016/0370-2693(85)90908-6)**, [213](https://doi.org/10.1016/0370-2693(85)90908-6) [\(1985\)](https://doi.org/10.1016/0370-2693(85)90908-6).
- [11] [V. V. Flambaum, I. B. Khriplovich, and O. P. Sushkov,](https://doi.org/10.1016/0375-9474(86)90331-3) Nucl. Phys. A **[449](https://doi.org/10.1016/0375-9474(86)90331-3)**, [750](https://doi.org/10.1016/0375-9474(86)90331-3) [\(1986\)](https://doi.org/10.1016/0375-9474(86)90331-3).
- [12] V. V. Flambaum and J. S. M. Ginges, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.65.032113) **[65](https://doi.org/10.1103/PhysRevA.65.032113)**, [032113](https://doi.org/10.1103/PhysRevA.65.032113) [\(2002\)](https://doi.org/10.1103/PhysRevA.65.032113).
- [13] V. V. Flambaum and A. Kozlov, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.85.068502) **[85](https://doi.org/10.1103/PhysRevC.85.068502)**, [068502](https://doi.org/10.1103/PhysRevC.85.068502) [\(2012\)](https://doi.org/10.1103/PhysRevC.85.068502).
- [14] V. V. Flambaum and A. Kozlov, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.85.022505) **[85](https://doi.org/10.1103/PhysRevA.85.022505)**, [022505](https://doi.org/10.1103/PhysRevA.85.022505) [\(2012\)](https://doi.org/10.1103/PhysRevA.85.022505).
- [15] W. C. Haxton and E. M. Henley, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.51.1937) **[51](https://doi.org/10.1103/PhysRevLett.51.1937)**, [1937](https://doi.org/10.1103/PhysRevLett.51.1937) [\(1983\)](https://doi.org/10.1103/PhysRevLett.51.1937).
- [16] V. V. Flambaum, [Phys. Lett. B](https://doi.org/10.1016/0370-2693(94)90646-7) **[320](https://doi.org/10.1016/0370-2693(94)90646-7)**, [211](https://doi.org/10.1016/0370-2693(94)90646-7) [\(1994\)](https://doi.org/10.1016/0370-2693(94)90646-7).
- [17] N. Auerbach, V. V. Flambaum, and V. Spevak, *[Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.76.4316)* **[76](https://doi.org/10.1103/PhysRevLett.76.4316)**, [4316](https://doi.org/10.1103/PhysRevLett.76.4316) [\(1996\)](https://doi.org/10.1103/PhysRevLett.76.4316).
- [18] V. Spevak, N. Auerbach, and V. V. Flambaum, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.56.1357) **[56](https://doi.org/10.1103/PhysRevC.56.1357)**, [1357](https://doi.org/10.1103/PhysRevC.56.1357) [\(1997\)](https://doi.org/10.1103/PhysRevC.56.1357).
- [19] J. Engel, J. L. Friar, and A. C. Hayes, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.61.035502) **[61](https://doi.org/10.1103/PhysRevC.61.035502)**, [035502](https://doi.org/10.1103/PhysRevC.61.035502) [\(2000\)](https://doi.org/10.1103/PhysRevC.61.035502).
- [20] V. V. Flambaum and V. G. Zelevinsky, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.68.035502) **[68](https://doi.org/10.1103/PhysRevC.68.035502)**, [035502](https://doi.org/10.1103/PhysRevC.68.035502) [\(2003\)](https://doi.org/10.1103/PhysRevC.68.035502).
- [21] N. Auerbach, V. F. Dmitriev, V. V. Flambaum, A. Lisetskiy, R. A. Sen'kov, and V. G. Zelevinsky, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.74.025502) **[74](https://doi.org/10.1103/PhysRevC.74.025502)**, [025502](https://doi.org/10.1103/PhysRevC.74.025502) [\(2006\)](https://doi.org/10.1103/PhysRevC.74.025502).
- [22] J. Engel, M. Bender, J. Dobaczewski, J. H. de Jesus, and P. Olbratowski, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.68.025501) **[68](https://doi.org/10.1103/PhysRevC.68.025501)**, [025501](https://doi.org/10.1103/PhysRevC.68.025501) [\(2003\)](https://doi.org/10.1103/PhysRevC.68.025501).
- [23] [J. Dobaczewski, J. Engel, M. Kortelainen, and P. Becker,](https://doi.org/10.1103/PhysRevLett.121.232501) *Phys.* Rev. Lett. **[121](https://doi.org/10.1103/PhysRevLett.121.232501)**, [232501](https://doi.org/10.1103/PhysRevLett.121.232501) [\(2018\)](https://doi.org/10.1103/PhysRevLett.121.232501).
- [24] R. H. Parker, M. R. Dietrich, M. R. Kalita, N. D. Lemke, K. G. Bailey, M. Bishof, J. P. Greene, R. J. Holt, W. Korsch, Z.-T. Lu, P. Mueller, T. P. O'Connor, and J. T. Singh, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.114.233002) **[114](https://doi.org/10.1103/PhysRevLett.114.233002)**, [233002](https://doi.org/10.1103/PhysRevLett.114.233002) [\(2015\)](https://doi.org/10.1103/PhysRevLett.114.233002).
- [25] M. Bishof, R. H. Parker, K. G. Bailey, J. P. Greene, R. J. Holt, M. R. Kalita, W. Korsch, N. D. Lemke, Z.-T. Lu, P. Mueller,

T. P. O'Connor, J. T. Singh, and M. R. Dietrich, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.94.025501) **[94](https://doi.org/10.1103/PhysRevC.94.025501)**, [025501](https://doi.org/10.1103/PhysRevC.94.025501) [\(2016\)](https://doi.org/10.1103/PhysRevC.94.025501).

- [26] E. R. Tardiff, J. A. Behr, T. E. Chupp, K. Gulyuz, R. S. Lefferts, W. Lorenzon, S. R. Nuss-Warren, M. R. Pearson, N. Pietralla, G. Rainovski, J. F. Sell, and G. D. Sprouse, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.77.052501) **[77](https://doi.org/10.1103/PhysRevC.77.052501)**, [052501\(R\)](https://doi.org/10.1103/PhysRevC.77.052501) [\(2008\)](https://doi.org/10.1103/PhysRevC.77.052501).
- [27] [B. Graner, Y. Chen, E. G. Lindahl, and B. R. Heckel,](https://doi.org/10.1103/PhysRevLett.116.161601) *Phys. Rev.* Lett. **[116](https://doi.org/10.1103/PhysRevLett.116.161601)**, [161601](https://doi.org/10.1103/PhysRevLett.116.161601) [\(2016\)](https://doi.org/10.1103/PhysRevLett.116.161601).
- [28] V. V. Flambaum, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.99.035501) **[99](https://doi.org/10.1103/PhysRevC.99.035501)**, [035501](https://doi.org/10.1103/PhysRevC.99.035501) [\(2019\)](https://doi.org/10.1103/PhysRevC.99.035501).
- [29] V. V. Flambaum and H. Feldmeier, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.101.015502) **[101](https://doi.org/10.1103/PhysRevC.101.015502)**, [015502](https://doi.org/10.1103/PhysRevC.101.015502) [\(2020\)](https://doi.org/10.1103/PhysRevC.101.015502).
- [30] R. J. Crewther, P. di Vecchia, G. Veneziano, and E. Witten, [Phys. Lett. B](https://doi.org/10.1016/0370-2693(80)91025-4) **[91](https://doi.org/10.1016/0370-2693(80)91025-4)**, [487](https://doi.org/10.1016/0370-2693(80)91025-4) [\(1980\)](https://doi.org/10.1016/0370-2693(80)91025-4).
- [31] P. W. Graham and S. Rajendran, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.84.055013) **[84](https://doi.org/10.1103/PhysRevD.84.055013)**, [055013](https://doi.org/10.1103/PhysRevD.84.055013) [\(2011\)](https://doi.org/10.1103/PhysRevD.84.055013).
- [32] Y. V. Stadnik and V. V. Flambaum, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.89.043522) **[89](https://doi.org/10.1103/PhysRevD.89.043522)**, [043522](https://doi.org/10.1103/PhysRevD.89.043522) [\(2014\)](https://doi.org/10.1103/PhysRevD.89.043522).
- [33] V. V. Flambaum and H. B. T. Tan, [Phys. Rev. D](https://doi.org/10.1103/PhysRevD.100.111301) **[100](https://doi.org/10.1103/PhysRevD.100.111301)**, [111301\(R\)](https://doi.org/10.1103/PhysRevD.100.111301) [\(2019\)](https://doi.org/10.1103/PhysRevD.100.111301).
- [34] C. Abel, N. J. Ayres, G. Ban, G. Bison, K. Bodek, V. Bondar, M. Daum, M. Fairbairn, V. V. Flambaum, P. Geltenbort *et al.*, [Phys. Rev. X](https://doi.org/10.1103/PhysRevX.7.041034) **[7](https://doi.org/10.1103/PhysRevX.7.041034)**, [041034](https://doi.org/10.1103/PhysRevX.7.041034) [\(2017\)](https://doi.org/10.1103/PhysRevX.7.041034).
- [35] D. Budker, P. W. Graham, M. Ledbetter, S. Rajendran, and A. O. Sushkov, [Phys. Rev. X](https://doi.org/10.1103/PhysRevX.4.021030) **[4](https://doi.org/10.1103/PhysRevX.4.021030)**, [021030](https://doi.org/10.1103/PhysRevX.4.021030) [\(2014\)](https://doi.org/10.1103/PhysRevX.4.021030).
- [36] O. P. Sushkov and V. V. Flambaum, Zh. Eksp. Teor. Fiz. **75**, 1208 (1978) [Sov. JETP **48**, 608 (1978)].
- [37] [https://www.nndc.bnl.gov/nudat2.](https://www.nndc.bnl.gov/nudat2)
- [38] S. E. Agbemava, A. V. Afanasjev, and P. Ring, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.93.044304) **[93](https://doi.org/10.1103/PhysRevC.93.044304)**, [044304](https://doi.org/10.1103/PhysRevC.93.044304) [\(2016\)](https://doi.org/10.1103/PhysRevC.93.044304).
- [39] L. M. Robledo and G. F. Bertsch, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.84.054302) **[84](https://doi.org/10.1103/PhysRevC.84.054302)**, [054302](https://doi.org/10.1103/PhysRevC.84.054302) [\(2011\)](https://doi.org/10.1103/PhysRevC.84.054302).
- [40] L. M. Robledo and P. A. Butler, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.88.051302) **[88](https://doi.org/10.1103/PhysRevC.88.051302)**, [051302\(R\)](https://doi.org/10.1103/PhysRevC.88.051302) [\(2013\)](https://doi.org/10.1103/PhysRevC.88.051302).
- [41] P. A. Butler and W. Nazarewicz, [Rev. Mod. Phys.](https://doi.org/10.1103/RevModPhys.68.349) **[68](https://doi.org/10.1103/RevModPhys.68.349)**, [349](https://doi.org/10.1103/RevModPhys.68.349) [\(1996\)](https://doi.org/10.1103/RevModPhys.68.349).
- [42] W. Zhang, Z. P. Li, S. Q. Zhang, and J. Meng, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.81.034302) **[81](https://doi.org/10.1103/PhysRevC.81.034302)**, [034302](https://doi.org/10.1103/PhysRevC.81.034302) [\(2010\)](https://doi.org/10.1103/PhysRevC.81.034302).
- [43] N. Minkov and A. Palffy, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.118.212501) **[118](https://doi.org/10.1103/PhysRevLett.118.212501)**, [212501](https://doi.org/10.1103/PhysRevLett.118.212501) [\(2017\)](https://doi.org/10.1103/PhysRevLett.118.212501).
- [44] [K. Nomura, R. Rodriguez-Guzman, and L. M. Robledo,](https://doi.org/10.1103/PhysRevC.92.014312) *Phys.* Rev. C **[92](https://doi.org/10.1103/PhysRevC.92.014312)**, [014312](https://doi.org/10.1103/PhysRevC.92.014312) [\(2015\)](https://doi.org/10.1103/PhysRevC.92.014312).
- [45] [R. N. Bernard, L. M. Robledo, and T. R. Rodríguez,](https://doi.org/10.1103/PhysRevC.93.061302) *Phys. Rev.* C **[93](https://doi.org/10.1103/PhysRevC.93.061302)**, [061302\(R\)](https://doi.org/10.1103/PhysRevC.93.061302) [\(2016\)](https://doi.org/10.1103/PhysRevC.93.061302).
- [46] Y. Fu, H. Wang, L.-J. Wang, and J. M. Yao, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.97.024338) **[97](https://doi.org/10.1103/PhysRevC.97.024338)**, [024338](https://doi.org/10.1103/PhysRevC.97.024338) [\(2018\)](https://doi.org/10.1103/PhysRevC.97.024338).
- [47] B. Bucher, S. Zhu, C. Y. Wu, R. V. F. Janssens, R. N. Bernard, L. M. Robledo, T. R. Rodríguez, D. Cline, A. B. Hayes, A. D. Ayangeakaa *et al.*, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.118.152504) **[118](https://doi.org/10.1103/PhysRevLett.118.152504)**, [152504](https://doi.org/10.1103/PhysRevLett.118.152504) [\(2017\)](https://doi.org/10.1103/PhysRevLett.118.152504).
- [48] J. de Vries, E. Mereghetti, and A. Walker-Loud, *[Phys. Rev. C](https://doi.org/10.1103/PhysRevC.92.045201)* **[92](https://doi.org/10.1103/PhysRevC.92.045201)**, [045201](https://doi.org/10.1103/PhysRevC.92.045201) [\(2015\)](https://doi.org/10.1103/PhysRevC.92.045201).
- [49] J. Bsaisou, J. de Vries, C. Hanhart, S. Liebig, U.-G. Meissner, D. Minossi, A. Nogga, and A. Wirzba, [J. High Energy Phys. 03](https://doi.org/10.1007/JHEP03(2015)104) [\(2015\)](https://doi.org/10.1007/JHEP03(2015)104) [104.](https://doi.org/10.1007/JHEP03(2015)104)
- [50] N. Yamanaka, B. K. Sahoo, N. Yoshinaga, T. Sato, K. Asahi, and B. P. Das, [Eur. Phys. J. A](https://doi.org/10.1140/epja/i2017-12237-2) **[53](https://doi.org/10.1140/epja/i2017-12237-2)**, [54](https://doi.org/10.1140/epja/i2017-12237-2) [\(2017\)](https://doi.org/10.1140/epja/i2017-12237-2).
- [51] E. Verstraelen, A. Teigelhofer, W. Ryssens, F. Ames, A. Barzakh, M. Bender, R. Ferrer, S. Goriely, P.-H. Heenen, M. Huyse, P. Kunz, J. Lassen, V. Manea, S. Raeder, and P. Van Duppen, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.100.044321) **[100](https://doi.org/10.1103/PhysRevC.100.044321)**, [044321](https://doi.org/10.1103/PhysRevC.100.044321) [\(2019\)](https://doi.org/10.1103/PhysRevC.100.044321).
- [52] [V. V. Flambaum, D. DeMille, and M. G. Kozlov,](https://doi.org/10.1103/PhysRevLett.113.103003) *Phys. Rev.* Lett. **[113](https://doi.org/10.1103/PhysRevLett.113.103003)**, [103003](https://doi.org/10.1103/PhysRevLett.113.103003) [\(2014\)](https://doi.org/10.1103/PhysRevLett.113.103003).
- [53] I. Ahmad, R. R. Chasman, J. P. Greene, F. G. Kondev, and S. Zhu, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.92.024313) **[92](https://doi.org/10.1103/PhysRevC.92.024313)**, [024313](https://doi.org/10.1103/PhysRevC.92.024313) [\(2015\)](https://doi.org/10.1103/PhysRevC.92.024313).
- [54] S. Ban, J. Dobaczewski, J. Engel, and A. Shukla, *[Phys. Rev. C](https://doi.org/10.1103/PhysRevC.82.015501)* **[82](https://doi.org/10.1103/PhysRevC.82.015501)**, [015501](https://doi.org/10.1103/PhysRevC.82.015501) [\(2010\)](https://doi.org/10.1103/PhysRevC.82.015501).
- [55] V. A. Dzuba, V. V. Flambaum, J. S. M. Ginges, and M. G. Kozlov, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.66.012111) **[66](https://doi.org/10.1103/PhysRevA.66.012111)**, [012111](https://doi.org/10.1103/PhysRevA.66.012111) [\(2002\)](https://doi.org/10.1103/PhysRevA.66.012111).
- [56] V. F. Dmitriev and R. A. Senkov, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.91.212303) **[91](https://doi.org/10.1103/PhysRevLett.91.212303)**, [212303](https://doi.org/10.1103/PhysRevLett.91.212303) [\(2003\)](https://doi.org/10.1103/PhysRevLett.91.212303).
- [57] P. V. Coveney and P. G. H. Sandars, [J. Phys. B](https://doi.org/10.1088/0022-3700/16/20/009) **[16](https://doi.org/10.1088/0022-3700/16/20/009)**, [3727](https://doi.org/10.1088/0022-3700/16/20/009) [\(1983\)](https://doi.org/10.1088/0022-3700/16/20/009).
- [58] V. F. Dmitriev, R. A. Senkov, and N. Auerbach, [Phys. Rev. C](https://doi.org/10.1103/PhysRevC.71.035501) **[71](https://doi.org/10.1103/PhysRevC.71.035501)**, [035501](https://doi.org/10.1103/PhysRevC.71.035501) [\(2005\)](https://doi.org/10.1103/PhysRevC.71.035501).
- [59] [V. A. Dzuba, V. V. Flambaum, and P. G. Silvestrov,](https://doi.org/10.1016/0370-2693(85)90564-7) *Phys. Lett.* B **[154](https://doi.org/10.1016/0370-2693(85)90564-7)**, [93](https://doi.org/10.1016/0370-2693(85)90564-7) [\(1985\)](https://doi.org/10.1016/0370-2693(85)90564-7).
- [60] V. A. Dzuba, V. V. Flambaum, and M. G. Kozlov, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.54.3948) **[54](https://doi.org/10.1103/PhysRevA.54.3948)**, [3948](https://doi.org/10.1103/PhysRevA.54.3948) [\(1996\)](https://doi.org/10.1103/PhysRevA.54.3948).
- [61] [V. A. Dzuba, V. V. Flambaum, and J. S. M. Ginges,](https://doi.org/10.1103/PhysRevA.76.034501) *Phys. Rev.* A **[76](https://doi.org/10.1103/PhysRevA.76.034501)**, [034501](https://doi.org/10.1103/PhysRevA.76.034501) [\(2007\)](https://doi.org/10.1103/PhysRevA.76.034501).
- [62] V. A. Dzuba, V. V. Flambaum, and S. G. Porsev, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.80.032120) **[80](https://doi.org/10.1103/PhysRevA.80.032120)**, [032120](https://doi.org/10.1103/PhysRevA.80.032120) [\(2009\)](https://doi.org/10.1103/PhysRevA.80.032120).
- [63] V. A. Dzuba, A. Kozlov, and V. V. Flambaum, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.89.042507) **[89](https://doi.org/10.1103/PhysRevA.89.042507)**, [042507](https://doi.org/10.1103/PhysRevA.89.042507) [\(2014\)](https://doi.org/10.1103/PhysRevA.89.042507).
- [64] V. A. Dzuba, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.93.032519) **[93](https://doi.org/10.1103/PhysRevA.93.032519)**, [032519](https://doi.org/10.1103/PhysRevA.93.032519) [\(2016\)](https://doi.org/10.1103/PhysRevA.93.032519).
- [65] T. Fleig, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.99.012515) **[99](https://doi.org/10.1103/PhysRevA.99.012515)**, [012515](https://doi.org/10.1103/PhysRevA.99.012515) [\(2019\)](https://doi.org/10.1103/PhysRevA.99.012515).
- [66] W. R. Johnson and J. Sapirstein, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.57.1126) **[57](https://doi.org/10.1103/PhysRevLett.57.1126)**, [1126](https://doi.org/10.1103/PhysRevLett.57.1126) [\(1986\)](https://doi.org/10.1103/PhysRevLett.57.1126).
- [67] [T. Fleig, H. Jørgen Aa. Jensen, J. Olsen, and L. Visscher,](https://doi.org/10.1063/1.2176609) J. Chem. Phys. **[124](https://doi.org/10.1063/1.2176609)**, [104106](https://doi.org/10.1063/1.2176609) [\(2006\)](https://doi.org/10.1063/1.2176609).
- [68] V. A. Dzuba, J. C. Berengut, C. Harabati, and V. V. Flambaum, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.95.012503) **[95](https://doi.org/10.1103/PhysRevA.95.012503)**, [012503](https://doi.org/10.1103/PhysRevA.95.012503) [\(2017\)](https://doi.org/10.1103/PhysRevA.95.012503).
- [69] V. A. Dzuba, V. V. Flambaum, and S. Schiller, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.98.022501) **[98](https://doi.org/10.1103/PhysRevA.98.022501)**, [022501](https://doi.org/10.1103/PhysRevA.98.022501) [\(2018\)](https://doi.org/10.1103/PhysRevA.98.022501).
- [70] [B. G. C. Lackenby, V. A. Dzuba, and V. V. Flambaum,](https://doi.org/10.1103/PhysRevA.98.022518) *Phys.* Rev. A **[98](https://doi.org/10.1103/PhysRevA.98.022518)**, [022518](https://doi.org/10.1103/PhysRevA.98.022518) [\(2018\)](https://doi.org/10.1103/PhysRevA.98.022518).
- [71] [B. G. C. Lackenby, V. A. Dzuba, and V. V. Flambaum,](https://doi.org/10.1103/PhysRevA.98.042512) *Phys.* Rev. A **[98](https://doi.org/10.1103/PhysRevA.98.042512)**, [042512](https://doi.org/10.1103/PhysRevA.98.042512) [\(2018\)](https://doi.org/10.1103/PhysRevA.98.042512).
- [72] [B. G. C. Lackenby, V. A. Dzuba, and V. V. Flambaum,](https://doi.org/10.1103/PhysRevA.99.042509) *Phys.* Rev. A **[99](https://doi.org/10.1103/PhysRevA.99.042509)**, [042509](https://doi.org/10.1103/PhysRevA.99.042509) [\(2019\)](https://doi.org/10.1103/PhysRevA.99.042509).
- [73] [B. G. C. Lackenby, V. A. Dzuba, and V. V. Flambaum,](https://doi.org/10.1103/PhysRevA.101.012514) *Phys.* Rev. A **[101](https://doi.org/10.1103/PhysRevA.101.012514)**, [012514](https://doi.org/10.1103/PhysRevA.101.012514) [\(2020\)](https://doi.org/10.1103/PhysRevA.101.012514).
- [74] S. M. Ramachandran and K. V. P. Latha, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.90.042503) **[90](https://doi.org/10.1103/PhysRevA.90.042503)**, [042503](https://doi.org/10.1103/PhysRevA.90.042503) [\(2014\)](https://doi.org/10.1103/PhysRevA.90.042503).
- [75] [K. V. P. Latha, D. Angom, B. P. Das, and D. Mukherjee,](https://doi.org/10.1103/PhysRevLett.103.083001) *Phys.* Rev. Lett. **[103](https://doi.org/10.1103/PhysRevLett.103.083001)**, [083001](https://doi.org/10.1103/PhysRevLett.103.083001) [\(2009\)](https://doi.org/10.1103/PhysRevLett.103.083001).
- [76] S. M. Ramachandran, K. V. P. Latha, and N. Meenakshisundaram, [J. Phys. B](https://doi.org/10.1088/1361-6455/aa77a4) **[50](https://doi.org/10.1088/1361-6455/aa77a4)**, [145003](https://doi.org/10.1088/1361-6455/aa77a4) [\(2017\)](https://doi.org/10.1088/1361-6455/aa77a4).
- [77] L. V. Skripnikov (private communication).
- [78] P. Schwerdtfeger and J. K. Nagle, [Mol. Phys.](https://doi.org/10.1080/00268976.2018.1535143) **[117](https://doi.org/10.1080/00268976.2018.1535143)**, [1200](https://doi.org/10.1080/00268976.2018.1535143) [\(2019\)](https://doi.org/10.1080/00268976.2018.1535143).
- [79] C. D. Panda, B. R. O'Leary, A. D. West, J. Baron, P. W. Hess, C. Hoffman, E. Kirilov, C. B. Overstreet, E. P. West, D. DeMille, J. M. Doyle, and G. Gabrielse, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.93.052110) **[93](https://doi.org/10.1103/PhysRevA.93.052110)**, [052110](https://doi.org/10.1103/PhysRevA.93.052110) [\(2016\)](https://doi.org/10.1103/PhysRevA.93.052110).
- [80] J. Baron *et al.* (ACME collaboration), [Science](https://doi.org/10.1126/science.1248213) **[343](https://doi.org/10.1126/science.1248213)**, [269](https://doi.org/10.1126/science.1248213) [\(2014\)](https://doi.org/10.1126/science.1248213).
- [81] A. N. Petrov, N. S. Mosyagin, T. A. Isaev, A. V. Titov, V. F. Ezhov, E. Eliav, and U. Kaldor, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.88.073001) **[88](https://doi.org/10.1103/PhysRevLett.88.073001)**, [073001](https://doi.org/10.1103/PhysRevLett.88.073001) [\(2002\)](https://doi.org/10.1103/PhysRevLett.88.073001).
- [82] F. A. Parpia, [J. Phys. B](https://doi.org/10.1088/0953-4075/30/17/022) **[30](https://doi.org/10.1088/0953-4075/30/17/022)**, [3983](https://doi.org/10.1088/0953-4075/30/17/022) [\(1997\)](https://doi.org/10.1088/0953-4075/30/17/022).
- [83] [H. M. Quiney, J. K. Laerdahl, K. Faegri, Jr., and T. Saue,](https://doi.org/10.1103/PhysRevA.57.920) Phys. Rev. A **[57](https://doi.org/10.1103/PhysRevA.57.920)**, [920](https://doi.org/10.1103/PhysRevA.57.920) [\(1998\)](https://doi.org/10.1103/PhysRevA.57.920).
- [84] V. V. Flambaum, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.77.024501) **[77](https://doi.org/10.1103/PhysRevA.77.024501)**, [024501](https://doi.org/10.1103/PhysRevA.77.024501) [\(2008\)](https://doi.org/10.1103/PhysRevA.77.024501).
- [85] A. D. Kudashov, A. N. Petrov, L. V. Skripnikov, N. S. Mosyagin, A. V. Titov, and V. V. Flambaum, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.87.020102) **[87](https://doi.org/10.1103/PhysRevA.87.020102)**, [020102\(R\)](https://doi.org/10.1103/PhysRevA.87.020102) [\(2013\)](https://doi.org/10.1103/PhysRevA.87.020102).
- [86] D. Cho, K. Sangster, and E. A. Hinds, [Phys. Rev. A](https://doi.org/10.1103/PhysRevA.44.2783) **[44](https://doi.org/10.1103/PhysRevA.44.2783)**, [2783](https://doi.org/10.1103/PhysRevA.44.2783) [\(1991\)](https://doi.org/10.1103/PhysRevA.44.2783).
- [87] E. B. Norrgard, D. J. McCarron, M. H. Steinecker, M. R. Tarbutt, and D. DeMille, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.116.063004) **[116](https://doi.org/10.1103/PhysRevLett.116.063004)**, [063004](https://doi.org/10.1103/PhysRevLett.116.063004) [\(2016\)](https://doi.org/10.1103/PhysRevLett.116.063004).
- [88] L. V. Skripnikov and A. V. Titov, [J. Chem. Phys.](https://doi.org/10.1063/1.4959973) **[145](https://doi.org/10.1063/1.4959973)**, [054115](https://doi.org/10.1063/1.4959973) [\(2016\)](https://doi.org/10.1063/1.4959973).
- [89] T. Fleig, talk at the workshop "Electric Dipole Moments," Canberra, 2019 (unpublished).
- [90] I. Kozyryev and N. R. Hutzler, [Phys. Rev. Lett.](https://doi.org/10.1103/PhysRevLett.119.133002) **[119](https://doi.org/10.1103/PhysRevLett.119.133002)**, [133002](https://doi.org/10.1103/PhysRevLett.119.133002) [\(2017\)](https://doi.org/10.1103/PhysRevLett.119.133002).