Complex-scaled relativistic configuration-interaction study of the *LL* resonances in heliumlike ions: From boron to argon

V. A. Zaytsev D, I. A. Maltsev, I. I. Tupitsyn, and V. M. Shabaev

Department of Physics, St. Petersburg State University, Universitetskaya Naberezhnaya 7/9, 199034 St. Petersburg, Russia

(Received 26 September 2019; published 13 November 2019)

Energies and Auger widths of the *LL* resonances in He-like ions from boron to argon are evaluated by means of a complex-scaled configuration-interaction approach within the framework of the Dirac-Coulomb-Breit Hamiltonian. The nuclear recoil and QED corrections are also taken into account. The results obtained are compared with other calculations based on the complex-scaling method as well as with the related results evaluated by using the stabilization and basis balancing methods.

DOI: 10.1103/PhysRevA.100.052504

I. INTRODUCTION

Autoionizing states of atomic or ionic systems are the excited states which can decay due to the electron-electron interactions via emission of one (or more) electrons. A special place among such states is held by the first autoionizing states of He-like systems; namely, levels of the LL resonance groups. The simplicity of these systems makes them attractive for both theoretical and experimental investigations. The investigations aimed at determining the energies of these levels are of particular interest for plasma diagnostics [1-4] and cosmological [5] and fusion research (see, e.g., the review [6]). A new interest in studying the characteristics of the *LL* resonances was caused by a recent experiment [7]. In this experiment, a new level of accuracy for the energy of the autoionizing states of the He-like carbon ion was reached. Experimental data of such accuracy being complemented by the theoretical predictions of the same precision allow one to set these states as energy-reference standards at synchrotron radiation facilities. The precise theoretical predictions for the energies of the LL resonances are, therefore, highly demanded.

For the accurate evaluation of the energies of autoionizing states, which are strongly affected by electron correlations, the high-precision many-electron methods such as coupled cluster and configuration interaction are required. These methods, although successfully applied for the calculations of boundlevel energies, fail when naively applied for the description of resonances. The energies of such resonances show a strong dependence on the parameters of the basis set, e.g., the convergence of the resonance energy with respect to the number of basis functions, which is one of the basis-set parameters, is very weak or even absent. This is explained by the fact that the autoionizing states are embedded into the positive-energy continuum. As a result, they cannot be described by squareintegrable functions which form the basis set of the coupledcluster and configuration-interaction methods. This problem can be naturally solved with the usage of the complex-scaling approach which is based on the analytical properties of the spectrum of a Hamiltonian being dilated into the complex

plane. The first mathematical analysis of these properties was reported in Refs. [8,9] for the nonrelativistic Hamiltonian and in Refs. [10–12] for the relativistic one. In these works, it was shown that in the spectra of the dilated Hamiltonian the autoionizing states are separated from the continuum. The wave functions of these states, therefore, become square integrable and can be investigated with conventional manyelectron methods. That makes the complex-scaling approach a powerful tool for studying properties of resonances appearing in various systems and processes. As examples, the resonances of nuclei [13–16], few-electron systems [17–23], and molecules [24,25] were investigated by using this method. More applications, as well as the details of the complexscaling approach, can be found in the reviews [26-31]. It is also worth noting that the dilated Hamiltonian is not a Hermitian but a symmetric operator with complex eigenvalues. The real and imaginary parts of the eigenvalues corresponding to the autoionizing states give the energies and Auger widths of the states, respectively.

Apart from the complex-scaling approach, one can apply the stabilization or basis balancing methods. The stabilization method (SM) was pioneered by Holøien and Midtdal [32] and utilized in numerous investigations [33-37]. The basis balancing method (BBM) was worked out by Yerokhin with co-authors just recently [38] and was applied for the calculation of the energies of the autoionizing levels of Li-like ions in a wide range of the nuclear charge number [39]. Both methods are applied to the conventional Hermitian Hamiltonian and, as a result, only the real arithmetic is involved that provides a considerable computational advantage. However, the energy of the autoionizing state obtained within SM or BBM can differ from the exact energy by a shift arising due to the inappropriate treatment of the interaction with the continuum. The advantages of these methods over the complex-scaling approach, thus, can be completely lost in some cases. In view of the considerable progress in experimental accuracy for the energies of the autoionizing states [7], the revision of the applicability of the SM and BBM is required.

In the present paper, we apply the configuration interaction (CI) coupled with the complex-scaling (CS) approach to solve

the Dirac-Coulomb-Breit (DCB) equation for *LL* resonances of He-like ions in the range from boron to argon. The configuration space is spanned by the one-electron Dirac orbitals being constructed from the *B* splines. The DCB energies are supplemented with the quantum electrodynamics (QED), nuclear recoil, and frequency-dependent Breit corrections. We also estimate the difference of the energies obtained within the SM and BBM with ones calculated employing the complexscaling approach. In case of the $2s^2$ level of the He-like carbon ion, it is found that the energy difference between these three methods exceeds the uncertainty reached in the recent experiment [7].

In the paper, for theoretical expressions the Heaviside charge unit $(e^2 = 4\pi\alpha c)$ and $\hbar = 1$ are used meanwhile numerical results are presented in atomic units ($m_e = \hbar = e=1$).

II. BASIC FORMALISM

We start with the formulation of the basic principles of the configuration-interaction with the complex-scaling approach for the solution of the few-electron DCB equation (for a detailed description see, e.g., the review [30]). Here we consider the simplest variant of the CS; namely, the uniform complex rotation. In this case, the radial variable r is transformed as

$$r \to r e^{i\theta},$$
 (1)

with θ being a constant rotation angle. This transformation leads to the following complex rotated DCB Hamiltonian:

$$H_{\text{DCB}}^{(\theta)} = \sum_{j} h_{D}^{(\theta)}(j) + e^{-i\theta} \sum_{j < k} [V_{C}(j,k) + V_{B}(j,k)],$$

$$j,k = 1, \dots, N.$$
(2)

Here N stands for the total number of the electrons and $h_D^{(\theta)}$ is the scaled one-electron Dirac Hamiltonian given by

$$h_D^{(\theta)}(j) = e^{-i\theta} c \boldsymbol{\alpha}_j \cdot \mathbf{p}_j + (\beta_j - 1)m_e c^2 + V_{\text{nuc}}(r_j e^{i\theta}), \quad (3)$$

where α and β are the Dirac matrices, **p** is the momentum operator, and V_{nuc} is the nuclear potential. In the present paper, we use the uniformly charged sphere model for the nuclear charge-density distribution in order to construct the electrostatic nuclear potential [40]. After the dilatation into the complex plane it takes the following form:

$$V_{\rm nuc}(re^{i\theta}) = \begin{cases} -\frac{\alpha Zc}{2R_{\rm nuc}} \left(3 - e^{2i\theta} \frac{r^2}{R_{\rm nuc}^2}\right), & r < R_{\rm nuc} \\ -e^{-i\theta} \frac{\alpha Zc}{r}, & r > R_{\rm nuc}. \end{cases}$$
(4)

In accordance with Eq. (2) the Coulomb and Breit interelectronic-interaction operators are given by

$$V_C(j,k) = \frac{\alpha c}{r_{jk}},\tag{5}$$

$$V_B(j,k) = \alpha c \left\{ \frac{e^{2i\theta}}{2c^2} \left[h_D^{(\theta)}(j), \left[h_D^{(\theta)}(k), r_{jk} \right] \right] - \frac{\boldsymbol{\alpha}_j \cdot \boldsymbol{\alpha}_k}{r_{jk}} \right\}$$
(6)

$$= -\frac{\alpha c}{2r_{jk}} [\boldsymbol{\alpha}_j \cdot \boldsymbol{\alpha}_k + (\boldsymbol{\alpha}_j \cdot \hat{\mathbf{r}}_{jk})(\boldsymbol{\alpha}_k \cdot \hat{\mathbf{r}}_{jk})], \qquad (7)$$

respectively. In Eqs. (5) and (7), $\hat{\mathbf{r}}_{jk} = \mathbf{r}_{jk}/r_{jk}$ with $\mathbf{r}_{jk} = \mathbf{r}_j - \mathbf{r}_k$ and $r_{jk} = |\mathbf{r}_{jk}|$. Having performed the complex rotation of

the DCB Hamiltonian (2) we now proceed to the construction of its eigenfunctions.

As in the conventional CI method [41,42], the *N*-electron eigenfunction $\Psi(PJM)$ with the parity *P*, total angular momentum *J*, and its projection *M* is expressed as a linear superposition of the configuration-state functions (CSFs) $\Phi(\gamma_r PJM)$:

$$\Psi(PJM) = \sum_{r=1}^{N_{\rm CSF}} c_r \Phi(\gamma_r PJM), \qquad (8)$$

where γ_r stands for all additional quantum numbers which determine uniquely the CSF. The CSFs are eigenstates of the total angular-momentum operators J^2 and J_z , constructed from antisymmetrized products of one-electron Dirac orbitals. Here these orbitals are chosen to be the eigenfunctions of the scaled one-electron Dirac Hamiltonian (3) of the form

$$\psi_{\kappa m}^{(\theta)}(\mathbf{r}) = \frac{e^{-i\theta}}{r} \begin{pmatrix} G_{\kappa}^{(\theta)}(r)\Omega_{\kappa m}(\hat{\mathbf{r}})\\ iF_{\kappa}^{(\theta)}(r)\Omega_{-\kappa m}(\hat{\mathbf{r}}) \end{pmatrix}$$

where $\kappa = (-1)^{l+j+1/2}(j+1/2)$ is the Dirac quantum number determined by the angular momentum *j* and the parity *l*, and $\Omega_{\kappa m}$ is the spinor spherical harmonic [43]. As usual in accordance with the basic principles of the relativistic theory with the DCB approximation, the CSFs are constructed only from positive-energy one-electron Dirac orbitals.

As already mentioned, autoionizing levels after the complex scaling are described by the square-integrable and localized wave functions. To good accuracy these wave functions can be represented by the corresponding solutions of the scaled DCB equation in a spherical cavity of finite radius. For the solution of this equation, the CI method with the CSFs constructed from the one-electron Dirac orbitals is utilized. In the present paper, these orbitals are obtained by solving the one-electron Dirac equation within the dual-kinetic-balance finite basis set approach [44] with the basis functions constructed from *B* splines [45,46]:

$$\begin{pmatrix}
G_{\kappa}^{(\theta)}(r) \\
F_{\kappa}^{(\theta)}(r)
\end{pmatrix} = \sum_{i=1}^{N} C_{\kappa,i}^{(\theta)} \begin{pmatrix}
B_{i}(r) \\
\frac{e^{-i\theta}}{2m_{e}c} \left[\frac{d}{dr} + \frac{\kappa}{r}\right] B_{i}(r) \\
+ \sum_{i=N+1}^{2N} C_{\kappa,i}^{(\theta)} \begin{pmatrix}
\frac{e^{-i\theta}}{2m_{e}c} \left[\frac{d}{dr} - \frac{\kappa}{r}\right] B_{i-N}(r) \\
B_{i-N}(r)
\end{pmatrix}.$$
(9)

Within the framework of this method, the spectrum of the oneelectron Dirac Hamiltonian including positive (and negative) continuum states is represented by a finite number of the quasistates. The interaction of the resonance with the continuum is, thus, incorporated via the inclusion of the interaction with these quasibound and quasicontinuum states. The resulting scaled DCB Hamiltonian is represented by the symmetric matrix with complex eigenvalues $E - i\Gamma_{Aug}/2$ where E and Γ_{Aug} are the energies and Auger widths of the corresponding resonances, respectively.

III. RESULTS AND DISCUSSIONS

A. Comparison of the stabilization and basis balancing methods with the complex-scaling approach

Let us start with a brief description of the principles of the stabilization and basis balancing methods, which are applied to the conventional (Hermitian) Hamiltonian. In the SM [32], the basis-set parameters are chosen in such a way as to provide a minimal value for the rate of change of the energy with respect to a variation of these parameters. Another approach, which consists of manipulating the basis to place the resonance just in the middle between the closest quasicontinuum states in the energy scale, is called the basis balancing method (BBM) [38]. Both these methods utilize the advantages of the finite basis set constructed from the square-integrable functions. As already was mentioned, such basis set functions cannot properly describe the contribution of the continuum to the autoionizing states. That is expressed in the energy shift of the state from the exact value. The size of this shift is, however, strongly resonance dependent and may be negligible in some cases. Here we estimate the difference between the results of the complex-scaling approach with results from the stabilization and basis balancing methods considering the state which is known to be significantly coupled with the continuum; namely, the $2s^2$ autoionizing state of the He-like carbon ion (Z = 6). For this purpose, we choose the radial grid, which uniquely defines the basis functions constructed from the *B* splines, as in Ref. [38]:

$$t_i = t_0 e^{A(i/N)^{\gamma}},\tag{10}$$

where $A = \ln (t_{\text{max}}/t_0)$, t_{max} is the radial size of the spherical cavity, t_0 is the radius of the nucleus, and γ is the basis set parameter. The energies of the autoionizing and quasicontinuum states depend strongly on the parameter γ and form γ -parametric trajectories, which are analyzed in accordance with the SM and BBM. For the sake of simplicity, we include only the CSFs being constructed from one-electron s and p Dirac orbitals. Figure 1 presents the γ -parametric energy trajectories for the $2s^2$ state of the He-like carbon (Z = 6) ion obtained in the basis of 30 B splines. This figure also presents the energies obtained with the use of the SM and BBM for each γ -parametric trajectory. From Fig. 1, it is seen that, at γ less than 0.5, the results of the SM and BBM are very close to each other. We note that the most accurate results are expected at small γ for which the density of the quasicontinuum states in the vicinity of resonance is sufficiently high. On the other hand, the states have to be well separated from one another to allow clear identification of the autoionizing state. In our case, both of these requirements are fulfilled for γ from 0.3 to 0.5. Before we proceed to the investigation of the convergence with respect to the number of B splines, let us explore how the results obtained within the CS approach depend on the γ parameter.

As was discussed in the preceding section, in the uniform complex rotation approach, the Hamiltonian depends on the θ parameter. Energies of the bound and quasibound states in this method are, however, θ independent for $\theta_c < \theta < \pi/2$, where θ_c is the critical angle, which can be approximately evaluated



FIG. 1. Energy of the $2s^2$ state of the He-like carbon (Z = 6) ion as a function of the parameter γ [see Eq. (10)]. The CSFs are constructed from one-electron *s* and *p* Dirac orbitals obtained in the basis of 30 *B* splines. The size of the spherical box was chosen to be 15 a.u. Blue circles and red squares correspond to the γ parameters chosen in accordance with the stabilization and basis balancing methods, respectively.

with the use of the nonrelativistic expression [8,47]

$$\theta_c = \frac{1}{2} \arctan\left\{ \Gamma_{\text{Aug}} / [2(E - E_t)] \right\}.$$
(11)

Here Γ_{Aug} and *E* are the Auger width and the energy of the level of interest, respectively, and E_t is the autoionization threshold energy, which for the $2s^2$ state is provided by the ground state of the corresponding H-like ion. Note that the energies do not depend on θ only if the complete or large basis set is utilized. In practice, however, one has to deal with an incomplete basis set that requires a search of an optimal angle for the uniform complex rotation. Although this angle should be larger than θ_c , we start the search from zero degrees. The optimal angle corresponds to the stationary point of the θ -parametric energy curve in the complex plane. In our case,



FIG. 2. Dependence of the *s* function (in a.u.) given by Eq. (12) on the θ and γ parameters for the $2s^2$ state of the He-like carbon (Z = 6) ion. The CSFs are constructed from one-electron *s* and *p* Dirac orbitals obtained in the basis of 30 *B* splines. The size of the spherical box was chosen to be 15 a.u.

TABLE I. Energy (in a.u.) of the $2s^2$ state of the He-like C (Z = 6) ion obtained within the stabilization method (SM), the basis balancing method (BBM), and the complex-scaling (CS) approach. The CSFs are constructed from one-electron *s* and *p* Dirac orbitals obtained in the basis of *N* functions. The size of the spherical box was chosen to be 15 a.u. Parameter γ is varied in the range between 0.3 and 0.5. The calculations within the CS approach are performed for θ varying from 20° to 25°.

			CS				
Ν	SM	BBM	$\operatorname{Re}(E)$	$\operatorname{Im}(E) \times 10^3$			
30 40 50 60	$\begin{array}{r} -8.291\ 30(3)\\ -8.291\ 34(4)\\ -8.291\ 37(3)\\ -8.291\ 40(2)\end{array}$	$\begin{array}{r} -8.2924(2) \\ -8.2921(1) \\ -8.29197(3) \\ -8.2919(1) \end{array}$	$\begin{array}{r} -8.291450(4) \\ -8.2914499(2) \\ -8.29144998(9) \\ -8.29144998(9) \end{array}$	$\begin{array}{r} -3.529(4) \\ -3.5290(3) \\ -3.52906(15) \\ -3.52907(13) \end{array}$			

one needs to find the stationary point of the (γ, θ) -parametric energy surface in the complex plane. That is equivalent to the search for the minimum of the function

$$s(\gamma,\theta) \equiv \sqrt{\left|\frac{dE}{d\theta}\right|^2 + \left|\frac{dE}{d\gamma}\right|^2}.$$
 (12)

Figure 2 presents the *s* function (12) for the $2s^2$ state of the He-like carbon (Z = 6) ion obtained in the basis of 30 *B* splines. From this figure it is seen that the $s(\gamma, \theta)$ function takes minimal values at γ from 0.3 to 0.5 and θ from 20° to 30°. For γ and θ changing within this area, the energy of the $2s^2$ state exhibits very stable behavior.

We now turn to the investigation of the convergence of the results obtained within the SM, BBM, and CS methods with respect to the number of basis functions. Table I presents the energy of the $2s^2$ state of the He-like C (Z = 6) ion for different numbers of *B* splines.

The calculations within the stabilization and basis balancing methods are performed for the γ parameter varying from 0.3 to 0.5. In Table I, we present the average values of the energies originating from different energetic curves corresponding to this γ interval (see Fig. 1). The uncertainty reflects the dependence of the results on the choice of the curve. From Table I, it is seen that the BBM results depend more strongly on the energetic curve than do the SM results. This can be due to the fact that, in the BBM, the resonance position is balanced with respect to the closest quasicontinuum states whereas in the SM the whole spectra is effectively taken into account. For both methods, the dependence on the energetic curve strongly masks the convergence with respect to the number of basis functions and gives the main source of the uncertainty. The calculations within the CS approach are performed for γ varying from 0.3 to 0.5 and θ varying in the range between 20° and 25° . The dependence of the energy on the γ and θ parameters forms the uncertainty indicated in Table I. It is seen that the energy obtained within the CS approach exhibits extremely fast convergence with respect to the number of basis functions. It is also seen that the energies obtained within the SM and BBM differ from the energy calculated by using the complex-scaling approach by more than 1 and 10 meV, respectively, the values which actually define accuracy limits of the SM and BBM. We note also that, working with SM and BBM, one needs to reselect the basis set parameters each time when the number of basis functions is enlarged. The necessity of this procedure drastically increases the required computation time and, thus, strongly reduces the advantage of the real arithmetic.

B. Energies and Auger widths of the LL resonances

We now apply the configuration-interaction complexscaling method for the calculation of the energies and Auger width of *LL* resonances of the He-like ions from boron (Z =5) to argon (Z = 18). The simplicity of the system studied allows performing the full CI calculations, i.e., the configuration space is formed from all possible combinations of the one-electron Dirac orbitals appearing for a given number of *B* splines. In the present paper, *B* splines of order 11 are utilized. Such a high order of *B* splines is chosen to guarantee the correct behavior of the one-electron Dirac orbitals with orbital angular momenta up to L = 8 at the origin. The one-electron orbitals with proper behavior at the origin appear to be less

TABLE II. Energy *E* and Auger width Γ_{Aug} of the $2s^2$ state of the He-like carbon (*Z* = 6) ion obtained within the configuration-interaction complex-scaling method. The CSFs are constructed from one-electron Dirac orbitals with orbital angular momenta up to L_{max} being obtained in the basis of *N B* splines. The size of the spherical box was chosen to be 15 a.u., $\gamma = 0.3$, and $\theta = 20^{\circ}$. The values listed after the second row are the increments obtained by successively adding configurations while increasing L_{max} .

		<i>E</i> [a.u.]		$\Gamma_{\rm Aug} imes 10^3$ [a.u.]			
L _{max}	N = 30	N = 40	N = 50	N = 30	N = 40	N = 50	
1	-8.291 450 6	-8.291 450 0	-8.291 449 9	7.056 59	7.058 01	7.058 00	
2	-0.0007961	-0.0007960	-0.0007960	-0.08175	-0.08155	-0.08152	
3	-0.0001180	-0.0001183	-0.0001183	-0.01888	-0.01872	-0.01870	
4	-0.0000375	-0.0000378	-0.0000378	-0.00672	-0.00658	-0.00656	
5	-0.0000159	-0.0000161	-0.0000162	-0.00302	-0.00291	-0.00289	
6	-0.0000078	-0.0000081	-0.0000081	-0.00157	-0.00149	-0.00148	
7	-0.0000043	-0.0000045	-0.0000045	-0.00091	-0.00085	-0.00083	
8	-0.0000025	-0.0000027	-0.0000027	-0.00057	-0.00052	-0.00051	
9–∞	-0.0000057	-0.0000068	-0.0000071	-0.00176	-0.00149	-0.00142	
Total	-8.2924383	-8.2924402	-8.2924407	6.941 41	6.943 91	6.944 07	

TABLE III. Energies E_{tot} and Auger widths Γ_{Aug} of the *LL* resonances of the He-like ions from boron (*Z* = 5) to argon (*Z* = 18), in a.u. The CS DCB energy, the QED correction, and the nuclear recoil correction are explicitly shown. Energies E_{tot} are supplemented with the total uncertainties from all calculated and uncalculated contributions. The nuclear charge radii are taken from Ref. [59].

Ion	Resonance	J	DCB	Recoil	QED	$E_{ m tot}$	$\Gamma_{ m Aug}$
¹¹ B ³⁺	$2s_{1/2}^2$	0	-5.662 877 1	0.000 282 1	0.000 085 6	-5.662509(24)	$6.674(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-5.4702350	0.000 280 5	-0.0000007	-5.4699552(88)	<10 ⁻⁶
	$2n^2_{2}$	0	-5,145,761,9	0.0002760	0.000 020 3	-5.145466(42)	$3.0(2) \times 10^{-4}$
	-P 3/2	2	-5.4694360	0.000 280 4	0.000.001.3	-5.4691543(48)	$< 2 \times 10^{-6}$
	$2s_{1/2}2n_{1/2}$	0	-56151794	0.000.281.1	0.000.054.1	-5.614.844(13)	$3314(7) \times 10^{-4}$
	251/22p1/2	1	-56149173	0.000.281.1	0.000.054.8	-5.614582(13)	$3.277(7) \times 10^{-4}$
	$2s_{1/2}2n_{2/2}$	1	-5.3817344	0.0002710	0.000 046 5	-5.381417(54)	$3.09(4) \times 10^{-3}$
	-51/2-P 5/2	2	-5.6143278	0.000 281 0	0.000 056 2	-5.613991(13)	$3.241(5) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-5.4699408	0.000 280 5	0.0	-5.4696604(36)	<10 ⁻⁶
	$-r^{1/2}-r^{5/2}$	2	-5.4044733	0.000 271 1	0.000 001 8	-5.404200(30)	$5.52(1) \times 10^{-3}$
${}^{12}C^{4+}$	$2s_{1/2}^2$	0	-8.2924407	0.000 378 6	0.000 167 9	-8.291894(41)	$6.944(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-8.057 969 3	0.000 377 0	-0.0000015	-8.057594(13)	$< 2 \times 10^{-7}$
	$2p_{2/2}^2$	0	-7.6535338	0.0003724	0.000 041 3	-7.653120(54)	$3.1(1) \times 10^{-4}$
	-r 3/2	2	-8.056242.6	0.000 377 0	0.000.002.8	-8.0558628(66)	$6.9(9) \times 10^{-7}$
	$2s_{1/2}2n_{1/2}$	0	-8.2349686	0.000 377 7	0.000 105 8	-8.234485(22)	$3.377(6) \times 10^{-4}$
	-51/2-P1/2	1	-8.2343998	0.000 377 6	0.000 107 2	-8.233915(22)	$3.32(1) \times 10^{-4}$
	$2s_{1/2}2p_{2/2}$	1	-7.9435675	0.000 366 1	0.000 094 0	-7.943107(71)	$3.35(4) \times 10^{-3}$
	-~1/2-r 5/2	2	-8.2331403	0.0003776	0.000 110 2	-8.232653(22)	$3.27(1) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-8.0573405	0.000 377 0	0.0	-8.0569635(38)	<10 ⁻⁶
	F 1/2 F 5/2	2	-7.9712525	0.000 366 7	0.000 003 1	-7.970883(40)	$6.017(9) \times 10^{-3}$
$^{14}N^{5+}$	$2s_{1/2}^2$	0	-11.4234480	0.000 447 0	0.000 295 3	-11.422706(64)	$7.146(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-11.1468905	0.0004456	-0.0000027	-11.146448(19)	$1.1(6) \times 10^{-7}$
	$2p_{2/2}^2$	0	-10.6620276	0.000 441 2	0.000 074 5	-10.661512(64)	$3.3(2) \times 10^{-4}$
	- r 3/2	2	-11.1435964	0.000 445 5	0.000 005 4	-11.1431455(93)	$1.8(6) \times 10^{-6}$
	$2s_{1/2}2p_{1/2}$	0	-11.3564130	0.000 446 2	0.000 185 6	-11.355781(34)	$3.444(2) \times 10^{-4}$
	-~1/2-r 1/2	1	-11.3553256	0.000 446 2	0.000 188 3	-11.354691(34)	$3.363(6) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-11.0062113	0.000 434 3	0.000 168 9	-11.005608(88)	$3.54(4) \times 10^{-3}$
	-~1/2-r 5/2	2	-11.3529409	0.000 446 1	0.000 193 9	-11.352301(34)	$3.289(5) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-11.145 698 1	0.000 445 6	0.0	-11.1452525(40)	<10 ⁻⁶
	1 1/2 1 5/2	2	-11.038 563 4	0.000 435 2	0.000 005 1	-11.038 123(40)	$6.39(2) \times 10^{-3}$
$^{16}O^{6+}$	$2s_{1/2}^2$	0	-15.0564866	0.000 515 5	0.000 480 2	-15.055 491(96)	$7.304(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-14.737 464 3	0.000 514 2	-0.0000045	-14.736955(27)	$2.3(7) \times 10^{-7}$
	$2p_{3/2}^2$	0	-14.171 660 3	0.000 509 9	0.000 123 3	-14.171027(76)	$3.4(2) \times 10^{-4}$
	1 5/2	2	-14.731 723 8	0.000 514 0	0.000 009 4	-14.731200(14)	$3.9(3) \times 10^{-6}$
	$2s_{1/2}2p_{1/2}$	0	-14.980 155 6	0.000 514 8	0.000 300 9	-14.979340(50)	$3.516(5) \times 10^{-4}$
	-//-	1	-14.9782594	0.000 514 8	0.000 305 6	-14.977439(50)	$3.409(6) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-14.5702588	0.000 502 7	0.000 279 0	-14.56948(11)	$3.69(5) \times 10^{-3}$
	/ 1 /	2	-14.9741209	0.000 514 6	0.000 315 2	-14.973 291(50)	$3.304(5) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-14.7353908	0.000 514 1	0.000 000 3	-14.7348764(44)	$< 10^{-6}$
		2	-14.6067695	0.0005037	0.0000080	-14.606258(44)	$6.67(3) \times 10^{-3}$
$^{19}F^{7+}$	$2s_{1/2}^2$	0	-19.1922303	0.000 553 3	0.0007355	-19.19094(14)	$7.434(3) \times 10^{-3}$
	$2p_{1/2}^2$	0	-18.8302321	0.0005522	-0.0000069	-18.829687(37)	$4.8(6) \times 10^{-7}$
	$2p_{3/2}^{2'}$	0	-18.1828693	0.0005482	0.000 191 0	-18.182 130(88)	$3.5(3) \times 10^{-4}$
	-/-	2	-18.8208917	0.000 551 9	0.000 015 3	-18.820324(19)	$8(1) \times 10^{-6}$
	$2s_{1/2}2p_{1/2}$	0	-19.1069359	0.000 552 8	0.0004597	-19.105923(70)	$3.600(7) \times 10^{-4}$
	, ,	1	-19.1038519	0.0005527	0.0004672	-19.102832(70)	$3.463(7) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-18.6363161	0.000 541 1	0.000 432 5	-18.63534(13)	$3.81(4) \times 10^{-3}$
		2	-19.0971277	0.0005525	0.0004828	-19.096092(70)	$3.326(8) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-18.8268543	0.000 552 1	0.0000007	-18.826 301 5(50)	$<\!2 \times 10^{-6}$
		2	-18.6762343	0.0005422	0.000 012 3	-18.675680(50)	$6.90(4) \times 10^{-3}$
²⁰ Ne ⁸⁺	$2s_{1/2}^2$	0	-23.8314470	0.0006527	0.0010752	-23.82972(19)	$7.542(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-23.4258147	0.0006516	-0.0000100	-23.425 173(50)	$9.5(4) \times 10^{-7}$
	$2p_{3/2}^{2'}$	0	-22.6961226	0.0006474	0.000 281 1	-22.695 19(10)	$3.6(2) \times 10^{-4}$
	- , -	2	-23.4114149	0.0006512	0.000 023 5	-23.410740(28)	$1.47(5) \times 10^{-5}$
	$2s_{1/2}2p_{1/2}$	0	-23.7375939	0.0006522	0.0006701	-23.736272(96)	$3.690(4) \times 10^{-4}$
		1	-23.7328442	0.000 652 1	0.000 681 5	-23.731 511(96)	$3.518(3) \times 10^{-4}$

TABLE III. (Continued.)							
Ion	Resonance	J	DCB	Recoil	QED	$E_{ m tot}$	Γ_{Aug}
	$2s_{1/2}2p_{3/2}$	1	-23.2050294	0.000 639 7	0.000 638 1	-23.20375(16)	$3.91(4) \times 10^{-3}$
		2	-23.7224673	0.0006518	0.0007056	-23.721 110(96)	$3.346(7) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-23.4205846	0.0006514	0.000 001 6	-23.4199316(61)	$<\!2 \times 10^{-6}$
		2	-23.2473363	0.0006409	0.0000186	-23.246677(59)	$7.08(4) \times 10^{-3}$
13Na9+	$2s_{1/2}^2$	0	-28.975002	0.000 690	0.001 514	-28.97280(25)	$7.635(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-28.524916	0.000 689	-0.000013	-28.524240(66)	$1.72(8) \times 10^{-6}$
	$2p_{3/2}^2$	0	-27.711922	0.000685	0.000 397	-27.71084(12)	$3.7(2) \times 10^{-4}$
	1 3/2	2	-28.503662	0.000688	0.000 035	-28.502939(39)	$2.7(1) \times 10^{-5}$
	$2s_{1/2}2p_{1/2}$	0	-28.873071	0.000 690	0.000 941	-28.87144(13)	$3.790(5) \times 10^{-4}$
	-~ 1/2-F 1/2	1	-28.866072	0.000 690	0.000 957	-28.86443(13)	$3.581(2) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-28.277097	0.000 677	0.000 905	-28.27551(20)	$3.99(4) \times 10^{-3}$
	-51/2-P 5/2	2	-28.850706	0.000689	0.000 993	-28.84902(13)	$3.367(3) \times 10^{-4}$
	$2n_{1/2}2n_{2/2}$	1	-28517138	0.000.689	0.000.003	-285164466(78)	$< 2 \times 10^{-6}$
	<i>2p</i> _{1/2} <i>2p</i> _{3/2}	2	-28.320470	0.000.679	0.000.003	-28.310763(70)	$7.23(4) \times 10^{-3}$
$^{24}M\sigma^{10+}$	$2s^2$	0	-34 623 863	0.000.079	0.000.028	-34.621.01(33)	$7.25(4) \times 10^{-3}$ $7.716(1) \times 10^{-3}$
IVIE	$\frac{23_{1/2}}{2n^2}$	0	_34 128 326	0.000790	-0.002000	-34.127.554(85)	$2.91(7) \times 10^{-6}$
	$\frac{2p_{1/2}}{2r^2}$	0	-34.128.320	0.000 785	-0.000 017	-34.127334(63)	$2.91(7) \times 10^{-4}$
	$2p_{3/2}$	0	-35.250 798	0.000 783	0.000.342	-55.22947(14)	$5.9(2) \times 10^{-5}$
	2 2	2	-34.098.066	0.000 /88	0.000 049	-34.097 229(54)	$4.49(6) \times 10^{-3}$
	$2s_{1/2}2p_{1/2}$	0	-34.514412	0.000 /90	0.001 281	-34.51234(17)	$3.901(4) \times 10^{-4}$
		l	-34.504 470	0.000790	0.001 304	-34.50238(17)	$3.652(5) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-33.853 273	0.000777	0.001 243	-33.851 25(24)	$4.06(4) \times 10^{-3}$
		2	-34.482472	0.000789	0.001 355	-34.48033(17)	$3.38/(7) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-34.117 134	0.000789	0.000 005	-34.116340(10)	$<2 \times 10^{-6}$
27 11 .		2	-33.896 040	0.000778	0.000 040	-33.895 222(86)	$7.34(4) \times 10^{-3}$
$^{27}\mathrm{Al}^{11+}$	$2s_{1/2}^2$	0	-40.779 109	0.000 827	0.002752	-40.77553(42)	$7.788(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-40.236929	0.000 826	-0.000020	-40.23612(11)	$4.66(6) \times 10^{-6}$
	$2p_{3/2}^2$	0	-39.253305	0.000 822	0.000718	-39.25177(17)	$4.0(2) \times 10^{-4}$
		2	-40.195135	0.000 825	0.000067	-40.194 243(72)	$7.4(1) \times 10^{-5}$
	$2s_{1/2}2p_{1/2}$	0	-40.662765	0.000 827	0.001 699	-40.66024(21)	$4.025(8) \times 10^{-4}$
		1	-40.649071	0.000 827	0.001730	-40.64651(21)	$3.738(6) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-39.934371	0.000814	0.001 661	-39.931 90(29)	$4.12(5) \times 10^{-3}$
		2	-40.618454	0.000 826	0.001 801	-40.61583(21)	$3.412(7) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-40.221253	0.000 826	0.000 009	-40.220418(14)	$<\!2 imes 10^{-6}$
		2	-39.974458	0.000 816	0.000 058	-39.97358(10)	$7.42(4) \times 10^{-3}$
28Si12+	$2s_{1/2}^2$	0	-47.441930	0.000 928	0.003 583	-47.43742(52)	$7.853(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-46.851709	0.000 927	-0.000021	-46.85080(13)	$7.1(1) \times 10^{-6}$
	$2p_{2/2}^2$	0	-45.780017	0.000 923	0.000 928	-45.778 17(19)	$4.1(2) \times 10^{-4}$
	r 3/2	2	-46.795465	0.000 926	0.000 089	-46.794451(95)	$1.16(1) \times 10^{-4}$
	$2s_{1/2}2p_{1/2}$	0	-47.319385	0.000 928	0.002 206	-47.31625(26)	$4.157(2) \times 10^{-4}$
	1/2 1/2	1	-47.301016	0.000 928	0.002.247	-47.297.84(26)	$3.84(1) \times 10^{-4}$
	$2s_{1/2}2n_{2/2}$	1	-46.521263	0.000914	0.002.171	-46.51818(34)	$4.17(5) \times 10^{-3}$
	-51/2-P3/2	2	-47 259 405	0.000.927	0.002.343	-4725614(26)	$3437(9) \times 10^{-4}$
	$2n_{1/2}2n_{2/2}$	- 1	-46 830 238	0.000.926	0.000.014	-46829298(20)	$6(3) \times 10^{-7}$
	-P1/2-P3/2	2	-46556124	0.000.916	0.000.082	-4655513(13)	$748(4) \times 10^{-3}$
$^{31}\mathbf{p}^{13+}$	$2s^2$	0	-54 613 632	0.000910	0.000.002	-5460809(64)	$7.10(1) \times 10^{-3}$ $7.911(1) \times 10^{-3}$
1	$\frac{23_{1/2}}{2n^2}$	0	53 073 757	0.000 965	0.004.017	53 072 81(16)	$1.05(2) \times 10^{-5}$
	$\frac{2p_{1/2}}{2n^2}$	0	-53.975757 52.811.508	0.000 904	-0.000 017	-53.972.01(10) 52.800.37(22)	$1.03(2) \times 10^{-4}$
	$2p_{3/2}$	2	-52.811.508	0.000 939	0.001175	-52.80957(22)	$4.3(2) \times 10^{-4}$
	2	2	-35.899739	0.000 962	0.000113	-33.89808(12)	$1.77(2) \times 10$ $4.205(1) \times 10^{-4}$
	$2s_{1/2}2p_{1/2}$	0	-54.485 055	0.000 965	0.002.810	-54.48180(32)	$4.305(1) \times 10^{-4}$
	2 2	1	-54.461 556	0.000 964	0.002 862	-54.45773(32)	$3.946(7) \times 10^{-3}$
	$2s_{1/2}2p_{3/2}$	1	-53.614 881	0.000 952	0.002783	-53.61115(40)	$4.21(5) \times 10^{-3}$
	2 2	2	-54.406 141	0.000963	0.002991	-54.402 19(32)	$3.4/(2) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-53.944 897	0.000 963	0.000 021	-53.943 913(28)	$8(4) \times 10^{-7}$
22 - 14	- 2	2	-53.641418	0.000 953	0.000114	-53.64035(17)	$7.50(5) \times 10^{-3}$
$^{52}S^{14+}$	$2s_{1/2}^2$	0	-62.295 647	0.001 066	0.005757	-62.288 82(78)	$7.963(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-61.604283	0.001 065	-0.000006	-61.603 22(20)	$1.48(2) \times 10^{-5}$
	$2p_{3/2}^2$	0	-60.348353	0.001 060	0.001 458	-60.345 83(26)	$4.5(2) \times 10^{-4}$
		2	-61.230680	0.001 063	0.000 144	-61.22947(21)	$7.48(5) \times 10^{-3}$

Ion	Resonance	J	DCB	Recoil	QED	$E_{ m tot}$	$\Gamma_{ m Aug}$
	$2s_{1/2}2p_{1/2}$	0	-62.162977	0.001 066	0.003 522	-62.158 39(39)	$4.468(1) \times 10^{-4}$
	-//-	1	-62.132053	0.001 066	0.003 588	-62.12740(39)	$4.078(6) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-61.216210	0.001 052	0.003 508	-61.211 65(48)	$4.24(5) \times 10^{-3}$
	, _ ,	2	-62.059542	0.001 064	0.003 756	-62.05472(39)	$3.50(2) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-61.566 101	0.001 064	0.000 031	-61.565007(38)	$1.0(3) \times 10^{-6}$
	, ,	2	-61.508840	0.001 053	0.000 157	-61.50763(17)	$2.63(3) \times 10^{-4}$
35Cl15+	$2s_{1/2}^2$	0	-70.489528	0.001 103	0.007 135	-70.48129(94)	$8.012(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-69.744626	0.001 101	0.000017	-69.74351(24)	$2.01(2) \times 10^{-5}$
	$2p_{3/2}^2$	0	-68.391 105	0.001 097	0.001778	-68.38823(30)	$4.7(2) \times 10^{-4}$
	- 5/2	2	-69.324 198	0.001 099	0.000178	-69.32292(25)	$7.43(4) \times 10^{-3}$
	$2s_{1/2}2p_{1/2}$	0	-70.352997	0.001 103	0.004 351	-70.34754(47)	$4.641(1) \times 10^{-4}$
	, _ ,	1	-70.313 991	0.001 102	0.004 433	-70.30846(47)	$4.225(8) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-69.326288	0.001 089	0.004 357	-69.320 84(56)	$4.27(6) \times 10^{-3}$
	, ,	2	-70.220551	0.001 101	0.004 650	-70.21480(47)	$3.53(2) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-69.694787	0.001 101	0.000 044	-69.693642(52)	$1.3(8) \times 10^{-6}$
		2	-69.623669	0.001 090	0.000 214	-69.62236(23)	$3.79(3) \times 10^{-4}$
$^{40}Ar^{16+}$	$2s_{1/2}^2$	0	-79.196961	0.001 084	0.008 730	-79.187 1(11)	$8.056(1) \times 10^{-3}$
	$2p_{1/2}^2$	0	-78.396270	0.001 083	0.000 060	-78.395 13(28)	$2.66(3) \times 10^{-5}$
	$2p_{3/2}^{2}$	0	-76.940290	0.001 078	0.002 134	-76.93708(34)	$4.9(3) \times 10^{-4}$
	-/-	2	-77.922194	0.001 080	0.000 215	-77.92090(29)	$7.33(4) \times 10^{-3}$
	$2s_{1/2}2p_{1/2}$	0	-79.057381	0.001 084	0.005 310	-79.05099(56)	$4.833(1) \times 10^{-4}$
		1	-79.008977	0.001 084	0.005 408	-79.00249(56)	$4.400(6) \times 10^{-4}$
	$2s_{1/2}2p_{3/2}$	1	-77.946209	0.001 071	0.005 344	-77.93979(64)	$4.30(6) \times 10^{-3}$
		2	-78.890181	0.001 082	0.005 685	-78.88341(56)	$3.57(2) \times 10^{-4}$
	$2p_{1/2}2p_{3/2}$	1	-78.331956	0.001 082	0.000 061	-78.330813(70)	$1.7(7) \times 10^{-6}$
		2	-78.245356	0.001 072	0.000288	-78.244 00(31)	$5.31(4) \times 10^{-4}$

TABLE III. (Continued.)

dependent on the choice of the complex rotation angle θ and, thus, provide more accurate results. The accuracy of the DCB eigenvalues, apart from the choice of the θ and γ parameters, depends on the number of B splines and the number of the orbital angular momenta L included. To estimate the uncertainty arising from the number of orbital angular momenta we carry out the CI calculations for $L \leq 8$ and estimate the tail contributions via polynomial least square fitting of the increments in powers of 1/L, as in Refs. [38,48,49]. An example of such an uncertainty analysis is presented in Table II for the $2s^2$ state of the carbon (Z = 6) ion. In this table, the results obtained solely for $L_{max} = 0$ are not presented since the inclusion of the p orbitals leads to a drastic change of the energy and Auger width. From Table II, it is seen that, for the basis of more than 40 B splines, the dominant contribution to the uncertainty of the DCB eigenvalues is provided by the configuration states with orbital angular momenta $L \ge 9$, whose contributions are taken into account by extrapolation. Therefore, in what follows we solve the complex rotated DCB equation in the configuration space formed from all possible combinations of the one-electron Dirac orbitals constructed out of 40 or 50 B splines.

To obtain the energies of the *LL* resonances with an accuracy at a few meV level, we supplement the solutions of the complex rotated DCB equation with the nuclear recoil and QED corrections. Both corrections are obtained with the use of the conventional (Hermitian) DCB Hamiltonian. The nuclear recoil effect arising due to the finite nuclear mass M admits fully relativistic treatment only within the framework of QED [50,51]. Here we account for this effect in the

lowest-order relativistic approximation and to first order in m_e/M via the inclusion of the mass shift operator [50,52]

$$H_{\rm MS} = \frac{1}{2M} \sum_{i,j} \left\{ \mathbf{p}_i \cdot \mathbf{p}_j - \frac{\alpha Z}{r_i} \left[\boldsymbol{\alpha}_i + \frac{(\boldsymbol{\alpha}_i \cdot \mathbf{r}_i)\mathbf{r}_i}{r_i^2} \right] \cdot \mathbf{p}_j \right\}$$
(13)

into the DCB Hamiltonian. The nuclear recoil correction to the energy of the particular *LL* resonance is given by the first-order perturbation theory with respect to this additional term [42]. As already mentioned, in addition to the nuclear recoil corrections we supplement the complex rotated DCB energies with the QED corrections. The ab initio evaluation of these corrections still remains a challenging task even for He-like systems for which the methods of the QED calculations are currently well established (see, e.g., Refs. [53-55] and references therein). We thus compute the two-electron QED effects on the energies of the autoionizing states. In the present paper, we evaluate the QED corrections utilizing the model QED operator [56], constructed with the usage of the QEDMOD package [57]. We evaluate the QED correction as the difference between the CI results obtained with and without the model QED operator included into the DCB Hamiltonian. This approach has shown its efficiency in numerous investigations [38,39,49,58]. However, in the QED model operator method, the screened QED corrections are taken into account only approximately. These corrections as well as the QED part of the two-photon-exchange contributions give rise to another source of uncertainty. We also note that the frequency-dependent Breit correction was found to be of

			This	work	Other theory	
Ζ	Resonance	J	E	Γ_{Aug}	E	$\Gamma_{ m Aug}$
5	$2s_{1/2}^2$	0	-5.662 502(24)	$6.674(1) \times 10^{-3}$	-5.66 088ª	6.650×10^{-3a}
	$2p_{3/2}^2$	0	-5.145465(42)	$3.0(2) \times 10^{-4}$	-5.14461^{a}	3.010×10^{-4a}
	$2s_{1/2}2p_{1/2}$	0	-5.614844(13)	$3.314(7) \times 10^{-4}$	-5.61299^{a}	3.208×10^{-4a}
	-//-	1	-5.614 581(13)	$3.277(7) \times 10^{-4}$		
	$2s_{1/2}2p_{3/2}$	2	-5.613991(13)	$3.241(5) \times 10^{-4}$		
6	$2s_{1/2}^2$	0	-8.291 878(40)	$6.944(1) \times 10^{-3}$	-8.28820^{a}	6.910×10^{-3a}
	$2p_{3/2}^2$	0	-7.653 119(54)	$3.1(1) \times 10^{-4}$	-7.65106^{a}	3.210×10^{-4a}
	$2s_{1/2}2p_{1/2}$	0	-8.234 485(22)	$3.377(6) \times 10^{-4}$	-8.23029^{a}	3.220×10^{-4a}
	1/2 1 1/2				-8.234 485 ^b	3.392×10^{-4b}
		1	-8.233915(22)	$3.32(1) \times 10^{-4}$	-8.233 914 ^b	3.327×10^{-4b}
	$2s_{1/2}2p_{3/2}$	2	-8.232652(22)	$3.27(1) \times 10^{-4}$	-8.232654^{b}	3.269×10^{-4b}
7	$2s_{1/2}^2$	0	-11.422672(64)	$7.146(1) \times 10^{-3}$	-11.41546^{a}	7.100×10^{-3a}
	$2p_{3/2}^2$	0	-10.661 511(64)	$3.3(2) \times 10^{-4}$	-10.65732^{a}	3.340×10^{-4a}
	$2s_{1/2}2p_{1/2}$	0	-11.355781(34)	$3.444(2) \times 10^{-4}$	-11.34755^{a}	3.230×10^{-4a}
	1/2 1 1/2	1	-11.354 691(34)	$3.363(6) \times 10^{-4}$		
	$2s_{1/2}2p_{3/2}$	2	-11.352301(34)	$3.289(5) \times 10^{-4}$		
8	$2s_{1/2}^2$	0	-15.055 424(96)	$7.304(1) \times 10^{-3}$	-15.04266^{a}	7.250×10^{-3a}
	$2p_{3/2}^{2}$	0	-14.171026(76)	$3.4(2) \times 10^{-4}$	-14.16345^{a}	3.440×10^{-4a}
	$2s_{1/2}2p_{1/2}$	0	-14.979 340(50)	$3.516(5) \times 10^{-4}$	-14.96481^{a}	3.235×10^{-4a}
	1/2 1 1/2	1	-14.977439(49)	$3.409(6) \times 10^{-4}$		
	$2s_{1/2}2p_{3/2}$	2	-14.973291(50)	$3.304(5) \times 10^{-4}$		
9	$2s_{1/2}^2$	0	-19.19081(14)	$7.434(3) \times 10^{-3}$	-19.16983^{a}	7.365×10^{-3a}
	$2p_{3/2}^2$	0	-18.182 130(88)	$3.5(3) \times 10^{-4}$	-18.16951^{a}	3.520×10^{-4a}
	$2s_{1/2}2p_{1/2}$	0	-19.105923(70)	$3.600(7) \times 10^{-4}$	-19.08204^{a}	3.240×10^{-4a}
	1/2 1 1/2	1	-19.102832(70)	$3.463(7) \times 10^{-4}$		
	$2s_{1/2}2p_{3/2}$	2	-19.096092(70)	$3.326(8) \times 10^{-4}$		
10	$2s_{1/2}^2$	0	-23.82944(19)	$7.542(1) \times 10^{-3}$	-23.79699^{a}	7.460×10^{-3a}
	$2p_{3/2}^2$	0	-22.695 19(10)	$3.6(2) \times 10^{-4}$	-22.67551^{a}	3.585×10^{-4a}
	$2s_{1/2}2p_{1/2}$	0	-23.736271(96)	$3.690(4) \times 10^{-4}$	-23.69927^{a}	3.243×10^{-4a}
	1/2 1 1/2	1	-23.731 510(96)	$3.518(3) \times 10^{-4}$		
	$2s_{1/2}2p_{3/2}$	2	-23.721 110(96)	$3.346(7) \times 10^{-4}$		

TABLE IV. The comparison of the calculated energies E and Auger widths Γ_{Aug} of the *LL* resonances of the He-like ions with other nonrelativistic [60] and relativistic results [7]. All data are given in atomic units.

^aHo [60].

^bMüller et al. [7].

minor importance for the systems under investigation and, therefore, its contribution can be omitted.

Table III presents the energies and Auger widths of the LL resonances of the He-like ions from boron (Z = 5) to argon (Z = 18). In this table, the complex rotated DCB energy, the QED correction, and the nuclear recoil correction are shown explicitly. The presented Auger widths Γ_{Aug} were calculated only by means of the CS DCB Hamiltonian. The smallness of the Auger widths of the $2p_{1/2}^2$ (*J* = 0), $2p_{3/2}^2$ (*J* = 2), and $2p_{1/2}2p_{3/2}$ (J = 1) resonances is explained by the fact that the Auger decay of the ${}^{3}P_{0}$, ${}^{3}P_{2}$, and ${}^{3}P_{1}$ states corresponding to these resonances in the LS-coupling scheme, respectively, is strictly forbidden in the nonrelativistic limit. Indeed, due to parity and total angular-momentum conservation, the Auger decay is allowed only to the 1ses and 1sed configurations. In the nonrelativistic case, the transition to these configurations is forbidden by the conservation of the orbital angular momentum. Energies E_{tot} are supplemented with the total uncertainties from all calculated contributions as well as from uncalculated high-order QED corrections. The uncertainty due to the uncalculated QED corrections was estimated by

analysis of the related contributions for the ground and singleexcited states in He-like ions [53]. In most cases, the accuracy of the present calculations is limited by the uncertainties from the QED contributions. Using the presented results with the available high-precision data for the energies of the ground and lowest excited states (see Refs. [53–55]), one can easily find the corresponding transition energies.

In Table IV, we compare some of our results with other nonrelativistic [60] and relativistic calculations [7]. In Ref. [60], the calculations were performed by using the complex-scaling technique in combination with Hylleraas-type functions without taking into account the QED corrections. Since the nonrelativistic method cannot resolve the fine structure of the $2s_{2p}$ resonance, for our three values for the $2s_{1/2}2p_{1/2}$ (J = 0, 1) and $2s_{1/2}2p_{3/2}$ (J = 2) states there is only one corresponding value from Ref. [60]. As one can see from the table, our results are in reasonable agreement with the nonrelativistic results. We also compared the values obtained for the carbon ion (Z = 6) with the recent relativistic calculations of Ref. [7]. These calculations were performed by employing the many-body perturbation theory in an all-order

formulation with the complex-scaling technique (see Ref. [30] and references therein). The QED corrections were taken into account by using the Welton method, which is different from the QED model operator approach. However, the results of Ref. [7] are in excellent agreement with our values.

IV. CONCLUSION

The energies and Auger widths of the *LL* resonances of the He-like ions from boron (Z = 5) to argon (Z = 18) have been evaluated by means of the complex-scaled configuration-interaction method. The systematic analysis of the uncertainty arising from the limited size of the configuration space was performed. The obtained energies have been compared with

those calculated by using the stabilization and basic balancing methods. It was found that the energies obtained with these methods differ from the complex-scaling results by a shift that varies from about 1 to 10 meV.

The nuclear recoil and QED corrections were evaluated separately and added to the complex rotated Dirac-Coulomb-Breit energies. As the result, the most accurate theoretical predictions for the energies of the *LL* resonances are obtained. In most cases, the accuracy of the total results is limited by the uncertainties from the higher-order QED corrections.

ACKNOWLEDGMENT

This work was supported by the Russian Science Foundation (Grant No. 17-12-01097).

- P. Beiersdorfer, T. Phillips, V. L. Jacobs, K. W. Hill, M. Bitter, S. von Goeler, and S. M. Kahn, Astrophys. J. 409, 846 (1993).
- [2] K. Widmann, P. Beiersdorfer, V. Decaux, S. R. Elliott, D. Knapp, A. Osterheld, M. Bitter, and A. Smith, Rev. Sci. Instrum. 66, 761 (1995).
- [3] H.-J. Kunze, *Introduction to Plasma Spectroscopy*, Springer series on atomic, optical, and plasma physics (Springer-Verlag, Berlin, 2009).
- [4] G. J. Tallents, An Introduction to the Atomic and Radiation Physics of Plasmas (Cambridge University Press, Cambridge, 2018).
- [5] D. Porquet, J. Dubau, and N. Grosso, Space Sci. Rev. 157, 103 (2010).
- [6] A. C. Fabian and I. M. George, in *Iron Line Diagnostics in X-Ray Sources*, edited by A. Treves, G. C. Perola, and L. Stella (Springer-Verlag, Berlin, 1991), p. 169.
- [7] A. Müller, E. Lindroth, S. Bari, A. Borovik, Jr., P.-M. Hillenbrand, K. Holste, P. Indelicato, A. L. D. Kilcoyne, S. Klumpp, M. Martins, J. Viefhaus, P. Wilhelm, and S. Schippers, Phys. Rev. A 98, 033416 (2018).
- [8] E. Balslev and J. M. Combes, Commun. Math. Phys. 22, 280 (1971).
- [9] J. Aguilar and J. M. Combes, Commun. Math. Phys. 22, 269 (1971).
- [10] R. A. Weder, J. Math. Phys. 15, 20 (1974).
- [11] R. A. Weder, J. Funct. Anal. 20, 319 (1975).
- [12] P. Seba, Lett. Math. Phys. 16, 51 (1988).
- [13] A. T. Kruppa, P.-H. Heenen, H. Flocard, and R. J. Liotta, Phys. Rev. Lett. 79, 2217 (1997).
- [14] S. Aoyama, T. Myo, K. Kato, and K. Ikeda, Prog. Theor. Phys. 116, 1 (2006).
- [15] K. Arai, Phys. Rev. C 74, 064311 (2006).
- [16] J.-Y. Guo, M. Yu, J. Wang, B.-M. Yao, and P. Jiao, Comput. Phys. Commun. 181, 550 (2010).
- [17] E. Lindroth, Phys. Rev. A 49, 4473 (1994).
- [18] E. Lindroth, Phys. Rev. A 52, 2737 (1995).
- [19] S. I. Themelis and C. A. Nicolaides, J. Phys. B: At., Mol. Opt. Phys. 28, L379 (1995).
- [20] N. Brandefelt and E. Lindroth, Phys. Rev. A 59, 2691 (1999);
 65, 032503 (2002).

- [21] G. Pestka, M. Bylicki, and J. Karwowski, J. Phys. B 39, 2979 (2006); 40, 2249 (2007).
- [22] M. Bylicki, G. Pestka, and J. Karwowski, Phys. Rev. A 77, 044501 (2008).
- [23] S. B. Zhang and D. L. Yeager, Phys. Rev. A 85, 032515 (2012).
- [24] M. N. Medikeri and M. K. Mishra, Chem. Phys. Lett. 246, 26 (1995); J. Chem. Phys. 103, 676 (1995).
- [25] S. Mahalakshmi, A. Venkatnathan, and M. K. Mishra, J. Chem. Phys. 115, 4549 (2001).
- [26] W. P. Reinhardt, Annu. Rev. Phys. Chem. 33, 223 (1982).
- [27] B. R. Junker, Adv. At. Mol. Phys. 18, 207 (1982).
- [28] Y. K. Ho, Phys. Rep. 99, 1 (1983).
- [29] N. Moiseyev, Phys. Rep. 302, 211 (1998).
- [30] E. Lindroth and L. Argenti, Adv. At. Mol. Opt. Phys. 63, 247 (2012).
- [31] B. Simon, Int. J. Quantum Chem. 14, 529 (1978).
- [32] E. Holøien and J. Midtdal, J. Chem. Phys. 45, 2209 (1966); 45, 3897 (1966).
- [33] A. D. Isaacson and D. G. Truhlar, Chem. Phys. Lett. 110, 130 (1984).
- [34] R. Moccia and P. Spizzo, J. Phys. B: At. Mol. Phys. 20, 1423 (1987).
- [35] R. Lefebvre, J. Phys. B: At., Mol. Opt. Phys. 21, L709 (1988).
- [36] P. Froelich, M. Höghede, and S. A. Alexander, J. Chem. Phys. 91, 1702 (1989).
- [37] V. A. Mandelshtam, T. R. Ravuri, and H. S. Taylor, Phys. Rev. Lett. 70, 1932 (1993).
- [38] V. A. Yerokhin, A. Surzhykov, and A. Müller, Phys. Rev. A 96, 042505 (2017); 96, 069901(E) (2017).
- [39] V. A. Yerokhin and A. Surzhykov, J. Phys. Chem. Ref. Data 47, 023105 (2018).
- [40] D. Andrae, Phys. Rep. **336**, 413 (2000).
- [41] M. H. Chen, K. T. Cheng, W. R. Johnson, and J. Sapirstein, Phys. Rev. A 52, 266 (1995).
- [42] I. I. Tupitsyn, V. M. Shabaev, J. R. Crespo López-Urrutia, I. Draganić, R. S. Orts, and J. Ullrich, Phys. Rev. A 68, 022511 (2003).
- [43] D. A. Varshalovich, A. N. Moskalev, and V. K. Khersonskii, *Quantum Theory of Angular Momentum* (World Scientific, Singapore, 1988).

- [44] V. M. Shabaev, I. I. Tupitsyn, V. A. Yerokhin, G. Plunien, and G. Soff, Phys. Rev. Lett. 93, 130405 (2004).
- [45] J. Sapirstein and W. R. Johnson, J. Phys. B: At., Mol. Opt. Phys. 29, 5213 (1996).
- [46] H. Bachau, E. Cormier, P. Decleva, J. E. Hansen, and F. Martín, Rep. Prog. Phys. 64, 1815 (2001).
- [47] B. Simon, Commun. Math. Phys. 27, 1 (1972); Ann. Math. 97, 247 (1973).
- [48] V. A. Yerokhin and A. Surzhykov, Phys. Rev. A 86, 042507 (2012).
- [49] M. Y. Kaygorodov, Y. S. Kozhedub, I. I. Tupitsyn, A. V. Malyshev, D. A. Glazov, G. Plunien, and V. M. Shabaev, Phys. Rev. A 99, 032505 (2019).
- [50] V. M. Shabaev, Teor. Mat. Fiz. 63, 394 (1985) [Theor. Math. Phys. 63, 588 (1985)]; Yad. Fiz. 47, 107 (1988) [Sov. J. Nucl. Phys. 47, 69 (1988)]; Phys. Rev. A 57, 59 (1998).
- [51] G. S. Adkins, S. Morrison, and J. Sapirstein, Phys. Rev. A 76, 042508 (2007).

- [52] C. W. P. Palmer, J. Phys. B: At. Mol. Phys. 20, 5987 (1987).
- [53] A. N. Artemyev, V. M. Shabaev, V. A. Yerokhin, G. Plunien, and G. Soff, Phys. Rev. A 71, 062104 (2005).
- [54] V. A. Yerokhin and K. Pachucki, Phys. Rev. A 81, 022507 (2010).
- [55] A. V. Malyshev, Y. S. Kozhedub, D. A. Glazov, I. I. Tupitsyn, and V. M. Shabaev, Phys. Rev. A 99, 010501(R) (2019).
- [56] V. M. Shabaev, I. I. Tupitsyn, and V. A. Yerokhin, Phys. Rev. A 88, 012513 (2013).
- [57] V. M. Shabaev, I. I. Tupitsyn, and V. A. Yerokhin, Comput. Phys. Commun. 189, 175 (2015); 223, 69 (2018).
- [58] I. I. Tupitsyn, M. G. Kozlov, M. S. Safronova, V. M. Shabaev, and V. A. Dzuba, Phys. Rev. Lett. **117**, 253001 (2016).
- [59] I. Angeli and K. P. Marinova, At. Data Nucl. Data Tables 99, 69 (2013).
- [60] Y. K. Ho, Phys. Rev. A 23, 2137 (1981).