

Ripple Theory of Surface Tension in Superfluid Helium[†]

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The ripple theory of surface tension in superfluid helium, originated by Atkins and developed by Brouwer and Pathria, implies a surface instability.

The ripple theory of surface tension in superfluid helium was first introduced by Atkins¹ and has recently² been considerably refined. The basis of the theory is that a liquid of mass density ρ and surface tension γ undergoes ripple oscillations³ with the spectrum

$$\omega_k^2 = \gamma k^3 / \rho . \quad (1)$$

If these oscillations are quantized, then the ripple excitations obey Bose statistics:

$$\bar{n}_k = (e^{\beta\omega_k} - 1)^{-1}, \quad \beta = \hbar/k_B T , \quad (2)$$

and the surface free energy is easily computed.^{1,2}

The purpose of this paper is to point out that the above theory implies an instability in the superfluid surface. Let $\zeta(x, y)$ represent the height of the surface above the equilibrium position in the (x, y) plane. This displacement can be expanded into normal modes

$$\zeta(\vec{r}) = A^{-1/2} \sum_{\vec{k}} \zeta_{\vec{k}} e^{i\vec{k} \cdot \vec{r}} , \quad (3)$$

where A is the area of the surface. Since $\zeta_{\vec{k}}$ is the complex oscillator coordinate for a single mode, the thermal fluctuation in $\zeta_{\vec{k}}$ is easily

computed

$$\gamma k^2 \langle \zeta_{\vec{k}}^\dagger \zeta_{\vec{k}} \rangle = (\bar{n}_k + \frac{1}{2}) \hbar \omega_k , \quad (4a)$$

$$\gamma k^2 \langle \zeta_{\vec{k}}^\dagger \zeta_{\vec{k}} \rangle = \frac{1}{2} \hbar \omega_k \coth(\frac{1}{2} \beta \omega_k) . \quad (4b)$$

In the long-wavelength limit

$$\langle \zeta_{\vec{k}}^\dagger \zeta_{\vec{k}} \rangle \propto 1/k^2, \quad k \rightarrow 0 . \quad (5)$$

Long-wavelength ($1/k^2$) divergences plague much of two-dimensional physics.⁴ In the present case, the surface thickness δ is easily shown to be infinite:

$$\delta^2 = \langle |\zeta(\vec{r})|^2 \rangle, \quad (6a)$$

$$\delta^2 = A^{-1} \sum_{\vec{k}} \langle \zeta_{\vec{k}}^\dagger \zeta_{\vec{k}} \rangle, \quad (6b)$$

$$\delta^2 = (2\pi)^{-2} \int d^2k \langle \zeta_{\vec{k}}^\dagger \zeta_{\vec{k}} \rangle, \quad (6c)$$

$$\delta^2 = \infty . \quad (6d)$$

Equation (6d) follows from Eqs. (5) and (6c). The divergence of δ at finite temperatures implies a surface instability.

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