

Lifetimes of the μ^- -Meson Stopped in Li, Be, and C

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THE time distributions of delays between the stopping of cosmic-ray μ -mesons in several light elements and the emission of decay electrons have been determined with the apparatus previously used¹ for a measurement of the μ^+ -meson lifetime. The measurements reported here were obtained by substituting, in turn, absorbers of pure lithium, beryllium, and carbon to stop μ -mesons of the sea-level mixture, in place of the iron used previously. Details of the absorbers are given in Table I.

For each element about 30,000 meson decays, with delays in the range 1 to 11 μ sec, were counted. After making the appropriate corrections¹ the decay curve for each of the three absorbers, when plotted, appears to be a single exponential. The contribution to each curve from μ^+ -decays was estimated by computing the relative number of μ^+ -mesons stopped in the absorber, brass counter walls, and iron supports, (assuming a sea-level positive excess of 20 percent for momenta of 500–1000 Mev/c) and attributing to the μ^+ -meson a mean life, $\tau^+ = 2.22 \pm 0.02 \mu$ sec.¹ Figure 1 shows the measured composite ($\mu^+ + \mu^-$) decay curve for carbon, together with the μ^- -decay curve obtained by subtracting the μ^+ -component.

In Table I are given values, for each absorber, of the fraction f^- of the total initial electron emission rate estimated as being due to μ^- -mesons, the apparent mean life τ' from the composite

TABLE I. Details of the absorbers and results.*

Absorber	Physical form	Mass (kg)	f^-	τ' (μ sec)	τ^- (μ sec)	A (10^4 sec^{-1})
Lithium	Metal cast in thin-walled steel box	6.5	0.27	2.20 ± 0.02	2.15 ± 0.09	<3.5
Beryllium	Metal bricks	12.4	0.42	2.15 ± 0.02	2.05 ± 0.06	3.7 ± 1.5
Carbon	Graphite block	25.4	0.42	2.12 ± 0.02	1.98 ± 0.06^b	5.5 ± 1.5

* Preliminary results with a sodium fluoride absorber give $\tau_{\text{NaF}}^- = 1.4 \pm 0.2 \mu$ sec, in agreement with Ticho's value $1.28 \pm 0.12 \mu$ sec (see reference 3).

^b H. A. Morewitz and M. H. Shamos [Phys. Rev. **87**, 241 (1952)] have recently reported a measurement of τ^- for carbon which is slightly lower than that given here.

curve, the mean life τ^- for the μ^- -meson, and the inferred nuclear capture rate A for the μ^- -meson.

The difference $(\tau^+ - \tau_{\text{Li}}^-) = +0.07 \pm 0.09 \mu$ sec indicates that the μ^+ and μ^- decay lifetimes are equal within the experimental error. The effect of nuclear capture by the Li nucleus is expected to be small, the correction to be added to τ_{Li}^- to get τ_{decay}^- being $A_{\text{Li}}(\tau_{\text{Li}}^-)^2$ if $A_{\text{Li}} \ll 1/\tau_{\text{decay}}^-$. By using the relation $A \propto Z^4$ proposed by Wheeler² to extrapolate from results^{3,4} obtained at higher atomic number (Z), where capture predominates over decay, we estimate that $A_{\text{Li}} \sim 0.3 \times 10^4 \text{ sec}^{-1}$, and, therefore, $A_{\text{Li}}(\tau_{\text{Li}}^-)^2 \sim 0.01 \mu$ sec. On this basis the difference $(\tau^+ - \tau_{\text{decay}}^-)$ is equal to $+0.06 \pm 0.09 \mu$ sec; it could only differ significantly from zero (becoming negative) should the above estimate of A_{Li} be too low by a factor $\gg 10$.

The nuclear capture rates A were calculated from the relation

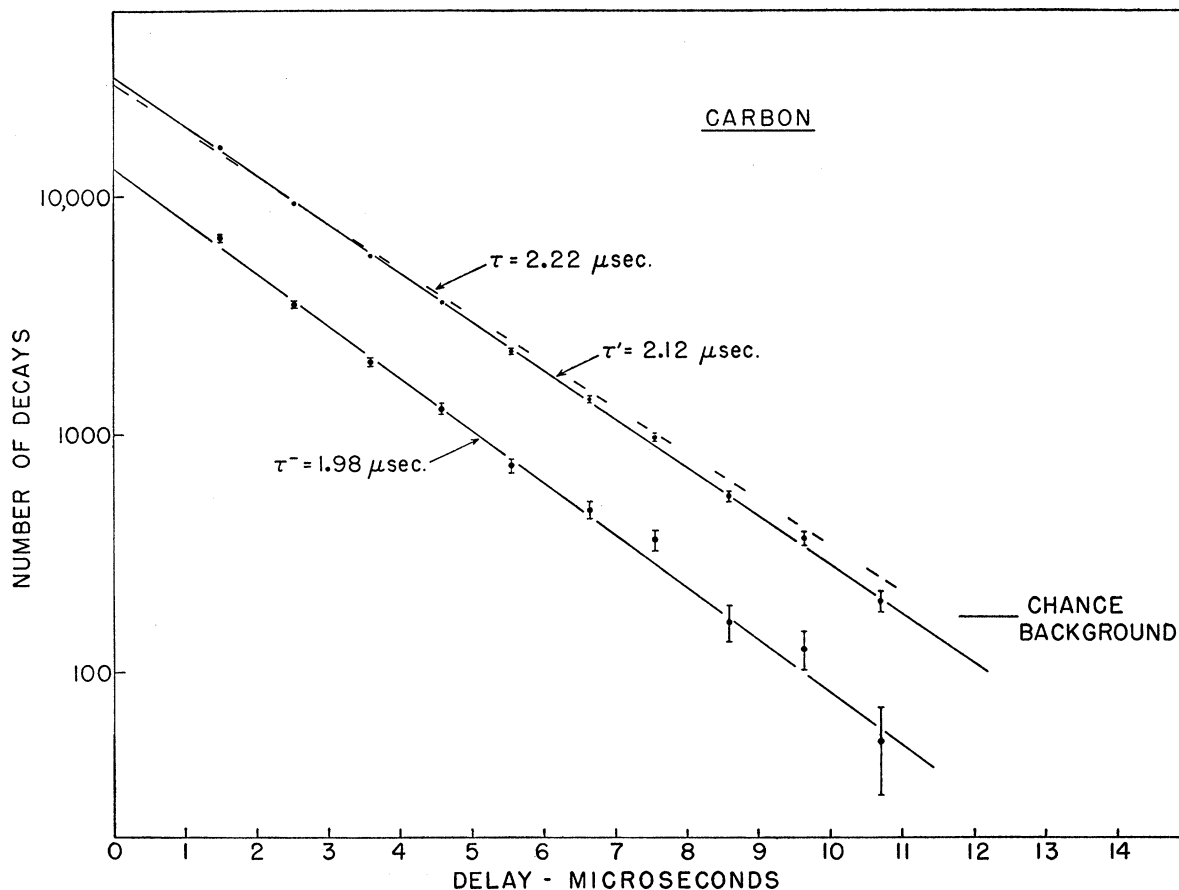


FIG. 1. Differential decay curves obtained with a carbon absorber. The upper curve is for the ($\mu^+ + \mu^-$) mixture, and the lower one is the corrected curve for μ^- -mesons only. The errors shown are standard deviations, and the best straight lines were determined by least squares fit. The dashed curve with $\tau = 2.22 \mu$ sec is shown for comparison.

$A = 1/\tau^- - 1/\tau^-_{\text{decay}}$, assuming that τ^-_{decay} is, in fact, equal to $2.22 \pm 0.02 \mu\text{sec}$. They do not depart significantly from the values obtained by extrapolating with a Z^4 dependence the results of previous measurements^{3,4} for $Z > 7$. Our experimental value of A for carbon may be compared with the values calculated by Preston and Duret⁶ assuming a charge-exchange reaction with a coupling constant equal to that determined from the neutron β -decay. The comparison indicates that the coupling constants for μ -capture and β -decay are equal to within about 30 percent, strengthening the argument for a universal interaction between pairs of fermions.

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¹ W. E. Bell and E. P. Hincks, *Phys. Rev.* **84**, 1243 (1951).

² J. A. Wheeler, *Revs. Modern Phys.* **21**, 133 (1949).

³ H. K. Ticho, *Phys. Rev.* **74**, 1337 (1948).

⁴ An account of existing data, with references, may be found in B. Rossi, *High Energy Particles* (Prentice-Hall, Inc., New York, 1952), Sec. 4.9.

⁵ M. A. Preston and M. F. Duret, following Letter [*Phys. Rev.* **88**, 1425 (1952)].

The Coupling Constant of the Nucleon- μ -Meson Interaction

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ON the hypothesis of a "universal" interaction for particles of spin $\frac{1}{2}\hbar$, μ -capture ($P + \mu \rightarrow N + \nu$) and β -decay will be described by interaction Hamiltonians containing the same five coupling constants g_K , $K = S, V, T, A, P$. We write $g_S^2 + g_V^2 = (1-x)g^2$, $g_T^2 + g_A^2 = xg^2$, $G^2 = (1+2x)g^2$. Recent β -decay evidence¹ suggests $\frac{1}{2} \lesssim x \lesssim \frac{3}{4}$. The decay of the neutron gives $G\beta = (3.12 \pm 0.4) \times 10^{-49} \text{ erg cm}^3$. From μ -capture data, Tiomno and Wheeler² have estimated $G_\mu \sim 10^{-49} \text{ erg cm}^3$.

More recently, Kennedy has used a shell model to study the experimental results with heavy elements³ and finds $G_\mu \sim 3 \times 10^{-49} \text{ erg cm}^3$. We have attempted to obtain a more reliable value of G_μ based on the μ -capture rate in C^{12} , since this has been measured and considerable information about the momentum distribution in this nucleus is now available.

We have expressed A , the transition probability for μ -meson capture, in terms of the nucleon momentum distribution. With a "gas" model approximation, we have

$$A = M\hbar^{-4}G_\mu^2\Psi^2 \int_0^{\alpha_0} dq q \int_0^{\infty} \rho W(\mathbf{p}_P)[1 - W(\mathbf{p}_N)] y dy.$$

Here \mathbf{p}_P , \mathbf{p}_N , and \mathbf{q} are proton, neutron and neutrino momenta, $\alpha_0 = \mu c^2 - \Delta$, $y = |\mathbf{p}_P \times \mathbf{q}/q|$, and $\mu c^2 = cq + \Delta + (\mathbf{p}_N^2 - \mathbf{p}_P^2)/2M$, where $\Delta = \text{mass of } \text{B}^{12} \text{ atom} - \text{mass of } \text{C}^{12} \text{ atom} + \text{meson binding energy} + mc^2$. The momentum distribution is $\rho W(\mathbf{p})$, where ρ is a density of order (nuclear dimensions/ \hbar)³ determined explicitly by the normalization of the momentum distribution and $W(\mathbf{p})$ is the probability that a nucleon with momentum \mathbf{p} is in one of the states which is occupied in the initial nucleus. The average over the nucleus of the K -shell meson probability density is Ψ^2 .

High energy (n, d),⁴ (p, p),⁵ and (p, π)⁶ experiments on C^{12} agree with a momentum distribution represented by (a) a Gaussian of average energy 14–19 Mev or (b) the Chew-Goldberger ($\alpha^2 + \beta^2$)⁻² distribution cutoff at 72 Mev. Taking $G_\mu = 3.12 \times 10^{-49} \text{ erg cm}^3$, we find for A the values in the second column of Table I.

TABLE I. Calculated transition probabilities and the values of G_μ .

Model	$A (10^4 \text{ sec}^{-1})$ with $G_\mu = 3.12$ $\times 10^{-49}$	$G_\mu (10^{-49} \text{ erg cm}^3)$
Gaussian: $\bar{E} = 18 \text{ Mev}$	8.65	2.5 ± 0.3
Chew-Goldberger-Temmer	8.86	2.5 ± 0.3
Shell: $x = 0.8$	4.67	3.4 ± 0.4
Shell: $x = 0.5$	4.48	3.5 ± 0.4
Shell: $x = 0.8$	2.11*	$5.0 \pm 0.7^*$
Shell: $x = 0.5$	1.92*	$5.3 \pm 0.8^*$

We have also made calculations using the shell model. Here we would expect the results to be reliable for transitions to low-lying bound states only. We use spherical potential wells to represent the nuclei. Their depths are found from the known separation energies of the $1p$ protons of C^{12} and $1p$ neutrons of B^{12} . On this model there are then no higher bound single particle states in B^{12} , and the matrix elements required are those from the various $s_{\frac{1}{2}}$ and $p_{\frac{1}{2}}$ states of C^{12} to the $p_{\frac{1}{2}}$ states of B^{12} . Using the above value for G_μ , we obtain for the transition probability to the bound states:

$$(1+2x)^{-1}[3.62(1-x) + 4.76(3x)] \times 10^4 \text{ sec}^{-1}.$$

Some of these states, which are bound on the shell model, have excitations of $\sim 11 \text{ Mev}$. Since the shell model is not reliable in these cases, we have also shown the transition probabilities excluding these states. (These values are labelled with an asterisk in Table I.)

Comparing the calculated values in the second column of Table I with the value $A = 5.5 \pm 1.5 \times 10^4 \text{ sec}^{-1}$ determined by Bell and Hincks,⁷ we find the values of G_μ given in column three. We may safely conclude that G_μ lies in the range (2.5 to 5.0) $\times 10^{-49} \text{ erg cm}^3$, and for reasons discussed in a fuller account (to be submitted to the Canadian Journal of Physics), we feel that the most likely value of G_μ is about $3.2 \times 10^{-49} \text{ erg cm}^3$.

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¹ O. Kofoed-Hanson and A. Winther, *Phys. Rev.* **86**, 428 (1952); R. Nataf and R. Bouchez, *Phys. Rev.* **87**, 155 (1952); J. M. Blatt, unpublished.

² J. Tiomno and J. A. Wheeler, *Revs. Modern Phys.* **21**, 153 (1949).

³ Kenuffel, Harrison, Godfrey, and Reynolds, *Phys. Rev.* **87**, 942 (1952); J. M. Kennedy, *Phys. Rev.* **87**, 953 (1952).

⁴ G. F. Chew and M. L. Goldberger, *Phys. Rev.* **77**, 470 (1950).

⁵ G. M. Temmer, *Phys. Rev.* **83**, 1067 (1951); Cladis, Hess, and Moyer, *Phys. Rev.* **87**, 425 (1952); P. A. Wolff, *Phys. Rev.* **87**, 434 (1952).

⁶ E. M. Henley and R. H. Huddleston, *Phys. Rev.* **82**, 754 (1951).

⁷ W. E. Bell and E. P. Hincks, preceding Letter [*Phys. Rev.* **88**, 1424 (1952)].

The g Factor of Ferrites

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THE microwave resonance in Ni ferrite and Mn ferrite was observed in polycrystalline specimens at several frequencies from 9450 Mc/sec to 47,000 Mc/sec, and it was found that the frequency dependence of the apparent g factors could be explained in relation to the anisotropic internal field in the material.

Many spherical specimens whose diameter varied from about 3 mm to 0.3 mm were polished from a sintered block, and after studying the size effect¹ on the g factor at each frequency, the resonance fields H_z were determined by extrapolating to zero diameter. These are shown in Table I. The g factor at each frequency, designated g^* in the table, was calculated by Kittel's formula,²

$$\nu = (\gamma^*/2\pi)H_z, \quad (1)$$

where $\gamma^* = g^*e/2mc$. This g factor was found to depend on the frequency, approaching the value of $g^* = 2$ at shorter wavelengths.

As Landé's factor is a materially constant factor, it should have the same value at any frequency. We, therefore, substituted

TABLE I. g^* calculated by Kittel's formula $\nu = (\gamma^*/2\pi)H_z$.

Material	Frequency (Mc/sec)	Resonance field H_z (oersteds)	g^*
NiOFe ₂ O ₄	9450	2780	2.43
	18,400	5930	2.22
	23,500	7660	2.19
	47,000	15,870	2.12
MnOFe ₂ O ₄	9450	3120	2.16
	18,400	6320	2.08
	23,500	8190	2.05
	47,000	16,620	2.02