

conjunction with a low noise input circuit described by Good<sup>1</sup> for the 87 kc/sec amplifier and a low pass filter gives a sensitivity of  $1 \times 10^{-7}$  nepers/cm as judged by the ability to detect all of the weak  $N^2H_3$  lines tabulated by Kisliuk and Townes.<sup>2</sup>

The observed spectrum consists of some 55 lines extending from 19,940 Mc/sec to 25,656 Mc/sec with only a few scattered lines between 23,790 Mc/sec and 24,980 Mc/sec suggesting two groupings of lines. The region from 25,656 Mc/sec to 25,920 Mc/sec which is the present upper limit of our equipment is void of lines. The complete tabulation of lines is shown in Table I, where the

TABLE I. Methyl mercaptan microwave spectra. Frequencies are known to an estimated 5 Mc/sec.

Frequency (Mc/sec)	Frequency (Mc/sec)	Frequency (Mc/sec)
19,909	22,414	24,485
19,996	22,558.0	24,995
20,052	22,561.2 <sup>a</sup>	25,015
20,136	22,663	25,125
20,163	22,830	25,145.0
20,380	23,071.5	25,147.0 <sup>a</sup>
20,645	23,075.0 <sup>a</sup>	25,150.5
20,712.5	23,233 <sup>b</sup>	25,152.0 <sup>a</sup>
20,714.0 <sup>a</sup>	23,260	25,210
21,058	23,340	25,225
21,522	23,500	25,290.0
21,735	23,525.0	25,290.4 <sup>a</sup>
21,770	23,532.3 <sup>a</sup>	25,565
21,866	23,565	25,660
21,878	23,622	
21,903	23,780	
21,945	23,805	
21,976	23,995	
22,270	24,072	
22,333	24,420	
	24,455	

<sup>a</sup> For each pair of lines so marked the separation of the two lines of the pair is known to an estimated 20 percent and the mean frequency of the pair to 5 Mc/sec.

<sup>b</sup> The accuracy of this frequency assignment is considered to be good to 0.5 Mc/sec. This is known by reference to an ammonia line at 23,232.20 Mc/sec.

frequencies are given to an estimated 5 Mc/sec as measured by a wavemeter calibrated against the known frequencies of the  $NH_3$  lines.<sup>2</sup>

The interpretation of this spectrum has not yet been achieved. The  $CH_3SH$  molecule is a slightly asymmetric rotator with moments of inertia for the  $S^{32}$  isotope<sup>3</sup> consistent with a pure rotational transition  $J_0=0 \rightarrow J_0=1$  at approximately 24,600 Mc/sec. Because of the small Boltzmann factor for all vibrational levels other than  $v_i=0$  this transition would be expected to yield a single strong line accompanied by a large number of extremely weak lines due to the excited vibrational levels. The 4.18 percent naturally occurring  $S^{34}$  would give a similar spectrum of intensity ratio 1 to 24 with respect to that of  $S^{32}$ . The observed spectrum does not fit this scheme. There exists, however, the splitting of the torsional vibration  $v=0$  level into three levels as a result of hindered internal rotation. This splitting depends upon the quantum number  $K$  in a complex manner.<sup>4</sup> Transitions between these levels could easily yield the large number of lines observed.

\* Physics Department.

† College of Engineering.

<sup>1</sup> W. E. Good, Westinghouse Research Paper 1538.

<sup>2</sup> P. Kisliuk and C. H. Townes, J. Research Natl. Bur. Standards **44**, 611 (1950).

<sup>3</sup> Russel, Osborne, and Yost, J. Am. Chem. Soc. **64**, 165 (1942).

<sup>4</sup> Koehler and Dennison, Phys. Rev. **57**, 1006 (1940).

### $\beta$ -Recoil Experiments with $Kr^{89}$

O. KOFOED-HANSEN AND P. KRISTENSEN  
Institute for Theoretical Physics, University of Copenhagen,  
Copenhagen, Denmark

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THE maximum energy and the average value of the energy divided by the charge of the recoil from the  $\beta$ -decay of  $Kr^{89}$  have been measured. The experimental method is similar to

that used previously<sup>1</sup> in the investigation of recoils from  $Kr^{88}$ . However, the technique has been improved in order to permit a study of short-lived activities since the half-life of  $Kr^{89}$  (3.18 min) is much shorter than that of  $Kr^{88}$  (2.77 hr).

The following results were obtained. The maximum recoil energy is equal to  $115 \pm 5$  ev corresponding to a maximum  $\beta$ -energy of  $3.9 \pm 0.1$  Mev. This agrees with the value 4.0 Mev found from absorption measurements.<sup>2</sup> The average value of the energy divided by the charge of the recoils amounts to  $58 \pm 2$  ev. The charge is always  $\geq 1$  and consequently 58 ev is a lower limit for the average recoil energy. Because of the uncertainty of the charge of the recoil atoms and the incomplete knowledge of the decay scheme the results do not permit a detailed comparison with the various possible angular correlations in  $\beta$ -decay.<sup>3</sup> However, certain possibilities may be excluded; in particular the data seem difficult to reconcile with the assumption of a backward neutrino emission with respect to the direction of emission of the  $\beta$ -particle. A forward neutrino emission would also in general be expected if the  $\beta$ -decay is forbidden, as seems to be the case for  $Kr^{89}$  judging from the ft value.

The half-life of  $Kr^{89}$  was measured, and the result, 3.14 min, agrees with the result found in the mass spectroscopic investigation.<sup>2</sup> Also the relative fission yield of mass numbers 88 and 89 was estimated. The result is  $y_{88}/y_{89} = 1.5 \pm 0.2$ .

A more detailed account of these experiments will be published elsewhere.<sup>4</sup>

We wish to express our gratitude to Professor N. Bohr and to Professor J. C. Jacobsen for their interest in our work.

<sup>1</sup> J. C. Jacobsen and O. Kofoed-Hansen, Phys. Rev. **73**, 675 (1948).

<sup>2</sup> O. Kofoed-Hansen and K. O. Nielsen, Kgl. Danske Videnskab. Selskab Mat.-fys. Medd. See also O. Kofoed-Hansen and K. O. Nielsen, Phys. Rev., following letter.

<sup>3</sup> D. R. Hamilton, Phys. Rev. **71**, 457 (1947).

<sup>4</sup> O. Kofoed-Hansen and P. Kristensen, Kgl. Danske Videnskab. Selskab Mat.-fys. Medd. (to be published).

### Short-Lived Krypton Isotopes and Their Daughter Substances

O. KOFOED-HANSEN AND K. O. NIELSEN  
Institute for Theoretical Physics, University of Copenhagen,  
Copenhagen, Denmark

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THE isotopes  $Kr^{89}$ ,  $Kr^{90}$ ,  $Kr^{91}$ , and their daughter substances have been investigated. Krypton formed in fission of uranium was pumped through a 10-m long tube directly from the cyclotron into the ion source of the isotope separator. The cyclotron and the isotope separator were operated simultaneously, and the counting could begin immediately after the interruption of the separation. The rubidium and strontium daughter substances were separated chemically; strontium was precipitated as carbonate. Half-lives were measured and an absorption analysis of the radiations was carried out. The results are given in Table I.

TABLE I. Observed radiations.

Isotope	Half-life	Radiation	$E_{\beta}^{max}$	Spectrum
$Kr^{89}$	3.18 min (2.6)	$\beta^-$ , $\gamma$	4.0 Mev	Complex
$Kr^{90}$	33 sec (33)	$\beta^-$ , $\gamma$	3.2 Mev	Complex
$Rb^{90}$	2.74 min	$\beta^-$ , $\gamma$	5.7 Mev	Complex
$Kr^{91}$	10 sec (9.8)	$\beta^-$ , $\gamma$ probable	$\sim 3.6$ Mev	Complex
$Rb^{91}$	100 sec	$\beta^-$ , $\gamma$	4.6 Mev	Complex
$Rb^{91}$	14 min	$\beta^-$ , $\gamma$	3.0 Mev	Complex

Previous data (see N.B.S. Circular 499: Nuclear Data) are given in parentheses.

It was found that at least 35 percent of the decays of  $Kr^{89}$  lead to an excited state of  $Rb^{89}$  which lies  $\sim 2$  Mev above the ground state. This result is of importance for the interpretation of the  $\beta$ -recoil experiments with this krypton isotope.<sup>1</sup>