# Evidence from Nuclear Masses on Proposed Closed Shells at 20 Nucleons\*

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Combining microwave measurements of sulfur and chlorine masses with mass spectroscopic and nuclear reaction information, the effect of nuclear shell structure on masses in the region of 20 neutrons or protons is investigated. It is found that except for Ca40, nuclei with 20 neutrons or protons do not show any special stability.

HERE are various types of evidence<sup>1-4</sup> that neutron and proton numbers N and Z of 2, 8, 20, 50, 82 and 126 form particularly stable nuclei. These numbers can be associated with the completion of shells of the quantum levels of a single particle in a potential well.<sup>5–7</sup> Such a single particle model has been used with considerable success to correlate nuclear moments, isomerism, and  $\beta$ -decay. However, there has as yet been very little information about nuclear masses at Nor Z of 20, 50, or 82. Hence, the most direct test for closed shells having extra stability has not yet been made.

Combining the recent microwave measurements of masses S<sup>36</sup>, S<sup>34</sup>, Cl<sup>36</sup>, and Cl<sup>37</sup> with mass spectroscopic and nuclear transmutation data, one can determine the masses of a large number of nuclei in the region of N=20 or Z=20. It is, therefore, feasible to test the



FIG. 1. Variation of nuclear mass with fixed Zand varying neutron number.

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- Work supported jointly by the Signal Corps and the ONK.
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  E. Feenberg and K. C. Hammack, Phys. Rev. 75, 1877 (1949).
  L. W. Nordheim, Phys. Rev. 75, 1894 (1949).
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expected variation in proton or neutron binding energy near the supposed closed shell at 20 nucleons. This is most conveniently done by comparison of the experimentally determined masses with those calculated from a semi-empirical formula which allows for most of the known sources of mass variation except shell structure. Such a formula has been given by Bohr and Wheeler<sup>8, 9</sup> as:

 $M_{AZ} = A - 0.00081Z - 0.00611A + 0.014A^{\frac{3}{2}}$  $+0.083(A/2-Z)^2/A+0.000627Z^2A^{-\frac{1}{2}}+\lambda$ 

where A = mass number, Z = nuclear charge,  $\lambda = 0$  for A odd,  $\lambda = -0.036A^{-\frac{3}{4}}$  for A even, Z even,  $\lambda = 0.036A^{-\frac{3}{4}}$ for A odd, Z odd.

This equation thus takes into account the normal odd-even fluctuations of masses. Although in general this expression does not agree with the experimentally measured masses to high accuracy, nevertheless the difference between the experimental and calculated masses plotted as a function of N or Z should give a relatively smooth curve. Any marked deviation from this curve, i.e., a sudden change in slope, might indicate the effects of shell structure.

Figures 1 and 2 show a plot of the difference between



- FIG. 2. Variation of nuclear mass with fixed neutron number and varying Z.
- <sup>8</sup> N. Bohr and J. A. Wheeler, Phys. Rev. 56, 426 (1939). <sup>9</sup> G. B. von Albada, Astrophys. J. 105, 393 (1947).

TABLE I. Mass differences calculated from microwave data and nuclear reactions.

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$H^1 = 1.008128$	$D^2 = 2.014718, n = 1.0089$	938, α = 4.003880 ma	ss units
Nuclei	Calculated mass difference	Reactions used	Reference
D82 D81	1 00029	dh	
S31 _ D31	$1.00020 \pm 40$ 0.00510 $\pm 17$	ap	a b
D32 _ S32	$0.00319 \pm 17$ 0.00170 $\pm 20$	р 8	C C
C34 D31	$0.00179\pm 20$ 2.00542 $\pm$ 40	μ wh	d
D34 S34	$0.000547 \pm 20$	$\frac{\alpha p}{\beta}$	u o
C33 C32	$0.000547\pm 20$	p db	f
33	$0.99903 \pm 03$	up mic	I G
C33 C34	$0.99977 \pm 30$ 0.00725 $\pm 05$	dh	g f
33	$(0.99723\pm03)$	(mic)	1
C34 C36	$(0.99709 \pm 13)$	(inic.)	h
C135 C32	$2.00034\pm30$	mic.	:
C100-502	$2.99800 \pm 40$	$\alpha p$	1
C100-500	$1.99893 \pm 40$	uα	J
$S^{**} - C^{**}$	$0.000179\pm03$	p	ĸ
$Cl^{\infty} - S^{\infty}$	$0.00549 \pm 07$	$\beta^+$	D
Cl <sup>04</sup> - S <sup>04</sup>	$0.00050 \pm 10$	β	1
$Cl_{39} - Cl_{31}$	$1.99751 \pm 14$	mic.	m
$Cl_{39} - Cl_{36}$	$0.99988 \pm 30$	dp	n
	$1.00017 \pm 40$	mic.	0
$S^{37} - Cl^{37}$	$0.00461\pm20$	$\beta^{-}$	e
$Cl^{38} - Cl^{37}$	$1.00228 \pm 30$	dp	р
$A^{35} - Cl^{35}$	$0.00580 \pm 30$	$\beta^+$	b
$Cl^{36} - A^{36}$	$0.000768 \pm .000006$	$\beta^{-}$	k
$A^{37} - A^{36}$	$0.99961 \pm 05$	dp	q
A <sup>38</sup> -Cl <sup>35</sup>	$2.99443 \pm 40$	αp	r
Cl <sup>38</sup> -A <sup>38</sup>	$0.00558 \pm 20$	$\beta^{-}$	s
K <sup>38</sup> -A <sup>38</sup>	$0.00641 \pm 20$	$\beta^+$	t
		$(\gamma + \beta^+ - ray in)$	
	(	cascade probably)	
$A^{41} - A^{40}$	$1.00249 \pm 05$	dp	u
A41-K41	$0.00274 \pm 20$	$\beta^{-}$	v
K <sup>40</sup> -A <sup>40</sup>	$0.0017 \pm 20$	pn	w
	$\geq 0.00167 \pm 10$	K-capture	x
K <sup>40</sup> -K <sup>39</sup>	$1.00070 \pm 07$	đp	У
Ca <sup>39</sup> -K <sup>39</sup>	$0.00619 \pm 30$	theoretical	-
K40-Ca40	$0.00145 \pm 05$	$\beta^{-}$	aa
	$0.00148 \pm 03$	$\beta^{-}$	bb
Ca41 – Ca40	$0.99997 \pm 05$	d þ	сс
Ca41-K41	$0.00047 \pm 07$	pn	dd
K <sup>39</sup> -Ca <sup>42</sup>	$2.99669 \pm 40$	άΦ	ee
K42-Ca42	$0.00383 \pm 20$	β <sup></sup>	v
K43-Ca43	$0.000868 \pm 03$	$\beta^{-}$	ff
Sc41-Ca41	$0.00630 \pm 20$	B+	b
Sc43-Ca43	$0.00226 \pm 25$	B+	gg
Ca <sup>40</sup> -Sc <sup>43</sup>	$3.00036 \pm 40$	αÞ	ee
K41-K41	$1.00109 \pm 10$	$\alpha p$	ĥĥ
$Ca^{40} - A^{40}$	$0.00032 \pm 08$	mass-spectros	ii
Ca n	0.00002100	mass speece 03.	**

- \* E. C. Pollard, Phys. Rev. 57, 1086 (1940).
  \* D. R. Elliott and L. D. P. King, Phys. Rev. 59, 403 (1941); White, Creutz, Delasso, and Wilson, Phys. Rev. 59, 63 (1941).
  \* K. Siegbahn, Phys. Rev. 70, 127 (1946).
  \* C. Pollard and C. J. Brasefield, Phys. Rev. 51, 8 (1937).
  \* W. Zunti and E. Bleuler, Helv. Phys. Acta 18, 263 (1945).
  \* P. Davison, Phys. Rev. 74, 1233 (1948).
  \* C. H. Townes and S. Geschwind, Phys. Rev. 74, 626 (1948).
  \* W. Low and C. H. Townes, Phys. Rev. 75, 529 (1949).
  \* C. J. Brasefield and E. C. Pollard, Phys. Rev. 75, 529 (1949).
  \* G. H. Towres and S. Geschwind, Phys. Rev. 50, 296 (1936).
  \* See reference 18.
  \* R. T. Overman, Radio Activities Produced in Neutron Irradiation o Chlorine AECD 857 (1947). C. S. Wu and L. Feldman, Phys. Rev. 76, 693 (1949).
  \* Zah Wei Ho, Phys. Rev. 70, 782 (1946).
  \* See references 15 and 16.
  \* See reference 18.
  \* A. Zucker and W. W. Watson, Phys. Rev. 79, 241 (1950).
  \* See reference 18 and 22.
  \* N. Hole, Arkiv. Mat. Astron, Fysik 32A, No, 3 (1945).
  \* Ramsey, Meem, and Mitchell, Phys. Rev. 72, 639 (1947).
  \* E. C. Pollard and P. W. Davison, Phys. Rev. 73, 1241 (1948).
  \* Bleuler, Bollman, and Zunti, Helv. Phys. Rev. 74, 1870 (1946); 20, 96 (1947).
  \* H. T. Richards and R. V. Smith, Phys. Rev. 74, 1870 (1948).

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O. Hizel and H. Waffler, Helv. Phys. Acta 19, 216 (1946).
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See pow. Kopjova, and Vorobjov, Phys. Rev. 69, 538 (1946).
D. E. Alburger, Phys. Rev. 79, 236 (1950).
See reference 21.

the experimental and computed masses against A for various Z and N. Tables I and II summarize data employed for determination of the masses, the mass values, estimates of their probable errors, and the references. Figure 3 indicates the experiments which relate the various nuclear masses.

TABLE II. Atomic masses based on Table I.\*

Standard: S <sup>22</sup> =31.98089, A <sup>40</sup> =39.97516						
	Mass.		computed from Bohr-	Δ differ- ence		
Nucleus	based on Table I	Comment	Wheeler formula	in m.m.u.		
P31	$30.98239 \pm 40$	Average of $P^{31}\alpha pS^{34}$ and $P^{31}dpP^{32}$	30.98218	0.21		
T) 20	24.00260.00	reactions				
Paz	$31.98268 \pm 20$		31.98315	-0.47		
P34	$33.98326\pm25$		33.98219	0.07		
S <sup>31</sup>	$30.98758\pm50$		30.98756	0.02		
S <sup>32</sup>	31.98089	Standard all nuclei from P <sup>31</sup> to A <sup>40</sup> calculated with respect to S <sup>32</sup>	31.98051	0.38		
\$33	$32.98058 \pm 10$	respect to b	32 08011	0.47		
Š34	$33.97780\pm20$	Average of microwave and S <sup>32</sup> -dp-S <sup>33</sup> reaction	33.97566	2.14		
S35	34.97907 + 45		34,97649	2 58		
S36	35.97834 + 40		35,97510	3 24		
S37	$3698111 \pm 50$		36 97866	2 4 5		
ငိုးအ	$32.98607 \pm 10$		32 08575	0 32		
C134	$3308436\pm 25$		33 08301	0.52		
C135	$33.90430\pm 23$ $34.07880\pm 40$	Average of	34 07817	0.43		
CI	J4.97889±40	S <sup>32</sup> - $\alpha p$ -Cl <sup>35</sup> and S <sup>33</sup> - $d\alpha$ -Cl <sup>35</sup>	54.97817	0.72		
Cl <sup>36</sup>	$35.97892 \pm 50$	Average of $Cl^{35}dpCl^{36}$ and microwave data	35.97855	+0.37		
$Cl^{37}$	$36.97640 \pm 50$		36.97503	1.37		
Cl38	$37.97868 \pm 55$		37.97728	1.40		
A 35	$34.98468 \pm 50$		34 98406	0.63		
A 36	35.97822 + 55		35 97717	1 05		
A <sup>38</sup>	$37.97320\pm60$	Average of $Cl^{35}\alpha p A^{38}$ and $Cl^{38}\beta^-A^{38}$ reactions	37.97138	1.82		
<b>A</b> 40	$30.07516 \pm 26$	Touctions	30 07020	4 06		
	07.710101101	Standard, reevaluated using references 10, 11 (Mattauch)	59.97020	4.90		
A41	$40.97765 \pm 05$		40.97293	4.72		
K <sup>38</sup>	37.97905±70	Assuming $\beta^+$ and $\gamma$ are in cascade	37.98035	-1.30		
$K^{39}$	$38.97613 \pm 30$		38.97460	+1.53		
$K^{40}$	$39.97683 \pm 10$		39.97446	2.37		
K41	$40.97491 \pm 25$		40.97075	4.16		
K42	$41.97600 \pm 30$		41.97235	3.65		
$K^{43}$	$42.97433 \pm 65$		42.97038	3.95		
A37	$36.9783 \pm 60$		35,97634	1 49		
Ca <sup>39</sup>	$38.98232\pm40$	Theoretical value using $\beta^+ = 4.8$ Mev.	38.98100	+1.32		
Ca <sup>40</sup>	$39.97546 \pm 08$		39.97701	-1.55		
Ca41	$40.97543 \pm 12$		40.97299	2.44		
Ca42	$41.97217 \pm 35$		41.96829	3.88		
Ca43	42.97346 + 55		42.96888	4 58		
Sc41	$40.98173 \pm 35$		40 97955	2.18		
Sc <sup>43</sup>	$42.97572\pm50$		42.97141	4.29		

\* Note that masses and errors given are not absolute values, but are relative to the masses chosen for  $S^{22}$  and  $A^{40}$ .

<sup>dd</sup> H. T. Richards and R. V. Smith, Phys. Rev. 74, 1257 (1948).
<sup>ee</sup> E. C. Pollard and E. J. Brasefield, Phys. Rev. 51, 8 (1937).
<sup>ef</sup> G. T. Seaborg and I. Perlman, Rev. Mod. Phys. 20, 589 (1948).
<sup>gg</sup> Hibdon, Pool, and Kurbatov, Phys. Rev. 67, 289 (1945).
<sup>hb</sup> See reference 21a.
<sup>ii</sup> See reference 20a.



FIG. 3. Known connections between masses of certain nuclei.

### **II. DISCUSSION OF MASS DETERMINATIONS**

All masses from P<sup>31</sup> to A<sup>40</sup> have been calculated with respect to the mass of  $S^{32} = 31.98089$ . The masses from argon to scandium have been computed with respect to A<sup>40</sup>. The mass of A<sup>40</sup> has been taken as 39.97516  $\pm 0.00028$  determined from the doublet<sup>10</sup> (Ne-A/2)  $=111.42 \times 10^{-4} \pm 0.38$  and  $using^{11}$  Ne<sup>20</sup> = 19.99872  $\pm 0.00013$ . All reactions used for the determination of masses have been recalculated using Bainbridge's values<sup>11</sup> for the masses of H, n, D, and  $\alpha$ . The mass values of a few nuclei require further explanations as given below.

S<sup>31</sup>.—S<sup>31</sup> is positron active and has been determined with respect to S32 through the mass of P31, as is indicated in Fig. 3. The P31 mass obtained is  $30.98238 \pm 45$  or  $30.98240 \pm 40$  for the reaction paths involving S34 or P32 respectively. The S31 mass is obtained immediately from P<sup>31</sup>.

Cl<sup>35</sup>.—The mass of Cl<sup>35</sup> can be determined from two different reactions as shown in Fig. 3. From the  $\alpha p$  reaction Cl<sup>35</sup>=34.97889 ±40 and from the  $d\alpha$ -reaction Cl<sup>35</sup>=34.97888±40. For the  $\alpha p$ reactions Bethe's revised Q values have been used throughout.12

Cl<sup>37</sup>.—The Cl<sup>37</sup> mass is determined from the mass ratio of Cl<sup>35</sup>/Cl<sup>37</sup>. Unfortunately there are different values for this ratio from mass spectroscopy and microwave spectroscopy. A recent mass spectroscopic determination<sup>13</sup> of this ratio gives the value as 0.9459441±65 (Aston<sup>14</sup> had found earlier the ratio as 0.9459806, but his experimental error is stated as  $\pm 300$ ). This value disagrees with the values obtained from microwave measurements<sup>15, 16</sup> in ICl

<sup>11</sup> K. Bainbridge, Isotopic Weights of the Fundamental Isotopes, Nuclear Science Series (Nat. Res. Council, 1949).
 <sup>12</sup> M. S. Livingston and H. A. Bethe, Rev. Mod. Phys. 9, 300

(1937).

<sup>13</sup> Okuda, Ogata, Aoki, and Sugawara, Phys. Rev. 58, 578 (1940).
 <sup>14</sup> F. W. Aston, Proc. Roy. Soc. 162A, 191 (1937).
 <sup>15</sup> Townes, Merritt, and Wright, Phys. Rev. 73, 1334 (1948).

and FCl which give  $0.9459801\pm50$  and  $0.9459775\pm40$  respectively. Isotopic frequencies for Cl<sup>35</sup>, Cl<sup>36</sup>, and Cl<sup>37</sup> in the molecule ClCN have also been measured.<sup>16a</sup> With these and a knowledge of the mass difference  $Cl^{35} - Cl^{36}$  (obtained from  $Cl^{35} dp Cl^{36}$ reaction) one finds this ratio Cl<sup>35</sup>/Cl<sup>37</sup>=0.9459906±120 in agreement with other microwave results. Moreover, an evaluation of this ratio from transmutation data<sup>17, 18</sup> leads to a value in essential agreement with microwave results. Hence the most recent mass spectroscopic value must be in error. Table III summarizes these results.

A<sup>38</sup>.—A<sup>38</sup> may be determined from Cl<sup>35</sup> with an  $\alpha$ , *p* reaction and from the beta-decay of Cl38. The values are 37.97332 and  $33.97322 \pm 60$  respectively.

K<sup>38</sup>.—K<sup>38</sup> gives off a positron of 2.53 Mev and a  $\gamma$ -ray of 2.15 Mev. We assume that these are in cascade.

Ca<sup>39</sup>.—Ca<sup>39</sup> belongs to a class of nuclei with Z-N=1 which are positron active. These "mirror" nuclei have been studied extensively; their half-lives and positron energies form a very regular and apparently well-understood series. The half-life of Ca<sup>39</sup> has been measured recently<sup>19</sup> and is 1.06 sec. in good agreement with expectations from this series. Hence the predicted<sup>20</sup> but still unmeasured<sup>†</sup> value of 4.8 Mev for the positron energy seems to us very reliable for obtaining the mass difference K<sup>39</sup>-Ca<sup>39</sup>.

Ca<sup>40</sup>.—The mass difference Ca<sup>40</sup>-A<sup>40</sup> has recently been obtained as  $(3.2\pm0.8)\times10^{-4}$  from mass spectroscopy<sup>20a</sup> and as  $(2.7\pm2.1)$  $\times 10^{-4}$  from nuclear reactions.<sup>21</sup>

K43, Ca42, 43, Sc43, Ca42.-The Ca42 mass has been determined from beta-decay of K42. From the measurement by Sailor<sup>21a</sup> of the mass difference K42-K41 one obtains for the mass of Ca42 =41.97217 $\pm$ 35. If the K<sup>39</sup>  $\alpha p$  Ca<sup>42</sup> reaction reported by Pollard and Brasefield<sup>22</sup> is used, a value  $41.97293 \pm 50$  results. This gives support to Bethe's suggestion<sup>12</sup> that the  $\alpha p$  reaction measured does not represent a transition to the ground state. Similarly, it is not quite certain whether the observed  $Ca^{40} \alpha p Sc^{43}$  reaction represents a transition to the ground state, and therefore, the

TABLE III. Value of Cl<sup>35</sup>/Cl<sup>37</sup> ratio.

Name	Method	Value
Aston*	Mass spectroscopy	$0.9459806 \pm 300$
Okuda et al.b	Mass spectroscopy	$0.9459452 \pm 65$
Townes et al.º	Microwave (ICl)	$0.9459801 \pm 50$
Roberts et al.d	Microwave (FCl)	$0.9459775 \pm 40$
Townes et al.º	Microwave (ClCN)	$0.9459906 \pm 120$
Pollard	Nuclear reaction	$0.9459893 \pm 110$

See reference 14.
See reference 13.
See reference 15.
See reference 16.
Townes and Shulman, private communication.
f See references 17 and 18.

<sup>16</sup> Gilbert, Roberts, and Griswold, Phys. Rev. **76**, 1721 (1949). <sup>16a</sup> We are grateful for the help of Dr. R. G. Shulman in meas-uring the difference in  $B_0$  values for Cl<sup>35</sup>CN and Cl<sup>36</sup>CN as 63.51±0.02 Mc. The similar difference for Cl<sup>35</sup>CN and Cl<sup>37</sup>CN is 123.56 Mc. [Townes, Holden, and Merritt, Phys. Rev. **74**, 1123 (1948).7

<sup>17</sup> E. C. Pollard, Phys. Rev. 57, 1086 (1940).

J. H. Shrader and E. C. Pollard, Phys. Rev. 59, 277 (1941).
 <sup>19</sup> Huber, Leinhardt, Scherrer, and Waffler, Helv. Phys. Acta

 16, 33 (1943).
 <sup>20</sup> E. Wigner, private communication to Walke, Phys. Rev. 57, 177 (1940).

† L. Motz has pointed out to us the determination of the  $(\gamma, n)$  threshold for Ca<sup>40</sup> as 15.9±0.4 Mev by McElhinney, Hanson, Becker, Duffield, and Diven, Phys. Rev. 75, 542 (1949). This would give Ca<sup>39</sup> the surprisingly large mass of  $38.9362\pm40$ , but

<sup>20a</sup> T. R. Roberts and A. O. C. Nier, Phys. Rev. 79, 198 (1950).
 <sup>21</sup> V. L. Sailor, Phys. Rev. 75, 1836 (1949).
 <sup>21a</sup> V. L. Sailor, Phys. Rev. 77, 749 (1950).

<sup>22</sup> E. C. Pollard and C. J. Brasefield, Phys. Rev. 51, 8, (1937).

<sup>&</sup>lt;sup>10</sup> J. Mattauch, Nuclear Physics Tables (Interscience Publishers,

masses of Ca<sup>43</sup>,  $K^{43}$ , and Sc<sup>43</sup> given in Table I may be somewhat too high.

#### **III. DISCUSSION OF RESULTS**

Figures 1 and 2 show that although there is considerable deviation of the measured masses from the Bohr-Wheeler formula, all deviations with the exception of the value for  $Ca^{40}$  lie on relatively smooth curves. The deviations from the curves are not generally larger than 0.5 Mev. There seems to be no abrupt change of mass or of slope of the curves of mass *versus* neutron number near S<sup>36</sup>, Cl<sup>37</sup>, A<sup>38</sup>, or K<sup>39</sup>, all of which are atoms with 20 neutrons. Similarly, there is no evidence of shell structure at 20 protons from the curves of mass *versus* proton number, except in the case of Ca<sup>40</sup> which shows a striking deviation from the smooth curves of about 3.5 Mev. This deviation is far greater than the probable error of mass determinations or than the deviation of any other nucleus plotted.

The absence of any change in slope at 20 nucleons makes it rather questionable whether 20 nucleons should be regarded as the closing point of a major shell. The well-known stability of  $Ca^{40}$  seems not to be simply connected with the completion of a shell at 20 nucleons. This case indicates that some large deviations from the stability curve may be encountered for other nuclei which are not attributable to neutron or proton shells alone but depend on the combination of neutron and proton numbers. Perhaps the large mass spread of stable Ca isotopes is due to this exceptional stability of  $Ca^{40}$ and to a shell at 28 neutrons which makes  $Ca^{48}$  stable rather than to a general stability of 20 protons.

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## On a Difference Equation Method in Cosmic-Ray Shower Theory\*

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The recent results of Snyder and Bhabha-Chakrabarty for the cascade theory of cosmic-ray showers are shown to be derivable from a general approach involving the use of the Laplace and Mellin transforms, and a general and powerful method, due to Snyder, for solving the resulting difference equations. Boundary conditions are introduced in a natural and automatic way, and the accuracy of the solution is limited by the possible ways of evaluating the resulting triple complex integral.

### I. INTRODUCTION

**S**NYDER<sup>1</sup> has recently obtained numerical results for the cascade theory of electron-photon showers which appear to be considerably more accurate than those of Bhabha and Chakrabarty.<sup>2</sup> It is the object of this paper to present a general method of solving the shower equations which yield both of the abovementioned solutions, and which should be applicable to a number of other problems.

#### II. THEORY

Using Snyder's<sup>1</sup> notation, we write the diffusion equation for P(E, t), the mean energy spectrum of electrons at depth t, and  $\gamma(E, t)$ , the mean energy spectrum of photons:

$$\frac{\partial P(E,t)}{\partial t} = \lim_{\delta \to 0} \left[ \int_{E+\delta}^{\infty} P(E',t) R(E',E'-E) \frac{E'-E}{E'^2} dE' - P(E,t) \int_{\delta}^{E} R(E,E') \frac{E'dE'}{E^2} \right] + \beta \frac{\partial P(E,t)}{\partial E} + 2 \int_{E}^{\infty} \gamma(E',t) R(E,E') \frac{dE'}{E'}, \quad (1)$$

$$\frac{\partial \gamma(E, t)}{\partial t} = \int_{E}^{\infty} P(E', t) R(E', E) \frac{EdE'}{E'^{2}} - \gamma(E, t) \int_{0}^{E} R(E', E) \frac{dE'}{E}.$$
 (2)

In these equations R(E', E) is a function which yields the elementary probabilities per unit path length of the pair-production and bremsstrahlung processes. In the case of high energies, the asymptotic form of R is that of a function homogeneous in E/E'; this is the only case dealt with in the present treatment.

<sup>\*</sup> This work was started in 1940 (Ph.D. Thesis, University of Michigan, 1941). It was completed at the Brookhaven National Laboratory, under the auspices of the AEC. <sup>1</sup> H. S. Snyder, Phys. Rev. **76**, 1563 (1949).

<sup>&</sup>lt;sup>2</sup> H. J. Bhabha and S. K. Chakrabarty, Phys. Rev. 74, 1352 (1948).