

Evidence from Nuclear Masses on Proposed Closed Shells at 20 Nucleons*

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Combining microwave measurements of sulfur and chlorine masses with mass spectroscopic and nuclear reaction information, the effect of nuclear shell structure on masses in the region of 20 neutrons or protons is investigated. It is found that except for Ca^{40} , nuclei with 20 neutrons or protons do not show any special stability.

THERE are various types of evidence¹⁻⁴ that neutron and proton numbers N and Z of 2, 8, 20, 50, 82 and 126 form particularly stable nuclei. These numbers can be associated with the completion of shells of the quantum levels of a single particle in a potential well.⁵⁻⁷ Such a single particle model has been used with considerable success to correlate nuclear moments, isomerism, and β -decay. However, there has as yet been very little information about nuclear masses at N or Z of 20, 50, or 82. Hence, the most direct test for closed shells having extra stability has not yet been made.

Combining the recent microwave measurements of masses S^{36} , S^{34} , Cl^{36} , and Cl^{37} with mass spectroscopic and nuclear transmutation data, one can determine the masses of a large number of nuclei in the region of $N=20$ or $Z=20$. It is, therefore, feasible to test the

expected variation in proton or neutron binding energy near the supposed closed shell at 20 nucleons. This is most conveniently done by comparison of the experimentally determined masses with those calculated from a semi-empirical formula which allows for most of the known sources of mass variation except shell structure. Such a formula has been given by Bohr and Wheeler^{8,9} as:

$$M_{AZ} = A - 0.00081Z - 0.00611A + 0.014A^{\frac{1}{2}} + 0.083(A/2 - Z)^2/A + 0.000627Z^2A^{-\frac{1}{2}} + \lambda,$$

where A = mass number, Z = nuclear charge, $\lambda = 0$ for A odd, $\lambda = -0.036A^{-\frac{1}{2}}$ for A even, Z even, $\lambda = 0.036A^{-\frac{1}{2}}$ for A odd, Z odd.

This equation thus takes into account the normal odd-even fluctuations of masses. Although in general this expression does not agree with the experimentally measured masses to high accuracy, nevertheless the difference between the experimental and calculated masses plotted as a function of N or Z should give a relatively smooth curve. Any marked deviation from this curve, i.e., a sudden change in slope, might indicate the effects of shell structure.

Figures 1 and 2 show a plot of the difference between

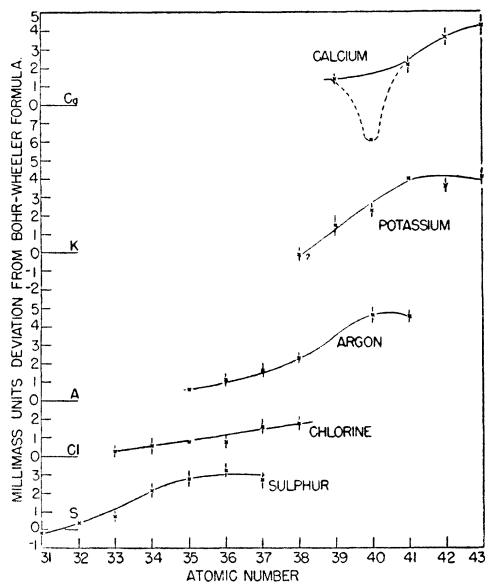


FIG. 1. Variation of nuclear mass with fixed Z and varying neutron number.

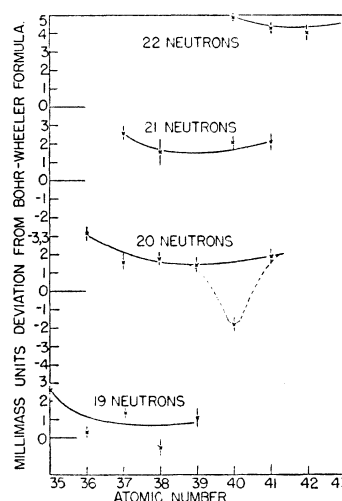


FIG. 2. Variation of nuclear mass with fixed neutron number and varying Z .

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TABLE I. Mass differences calculated from microwave data and nuclear reactions.

$H^1=1.008128, D^2=2.014718, n=1.008938, \alpha=4.003880$ mass units			
Nuclei	Calculated mass difference	Reactions used	Reference
$P^{32}-P^{31}$	1.00028 ± 40	$d\beta$	a
$S^{31}-P^{31}$	0.00519 ± 17	β^+	b
$P^{32}-S^{32}$	0.00179 ± 20	β^-	c
$S^{34}-P^{31}$	2.99542 ± 40	$\alpha\beta$	d
$P^{34}-S^{34}$	0.000547 ± 20	β^-	e
$S^{33}-S^{32}$	0.99965 ± 05	$d\beta$	f
	0.99977 ± 30	mic.	g
$S^{33}-S^{34}$	0.99725 ± 05	$d\beta$	i
	(0.99709 ± 15)	(mic.)	
$S^{34}-S^{36}$	2.00054 ± 30	mic.	h
$Cl^{35}-S^{32}$	2.99800 ± 40	$\alpha\beta$	i
$Cl^{35}-S^{33}$	1.99893 ± 40	$d\alpha$	j
$S^{35}-Cl^{35}$	0.000179 ± 03	β^-	k
$Cl^{33}-S^{33}$	0.00549 ± 07	β^+	b
$Cl^{34}-S^{34}$	0.00656 ± 10	β^+	l
$Cl^{35}-Cl^{37}$	1.99751 ± 14	mic.	m
$Cl^{36}-Cl^{36}$	0.99988 ± 30	$d\beta$	n
	1.00017 ± 40	mic.	o
$S^{37}-Cl^{37}$	0.00461 ± 20	β^-	e
$Cl^{38}-Cl^{37}$	1.00228 ± 30	$d\beta$	p
$A^{35}-Cl^{35}$	0.00580 ± 30	β^+	b
$Cl^{36}-A^{36}$	$0.000768 \pm .000006$	β^-	k
$A^{37}-A^{38}$	0.99961 ± 05	$d\beta$	q
$A^{38}-Cl^{35}$	2.99443 ± 40	$\alpha\beta$	r
$Cl^{38}-A^{38}$	0.00558 ± 20	β^-	s
$K^{38}-A^{38}$	0.00641 ± 20	β^+	t
		($\gamma+\beta^+$ -ray in cascade probably)	
$A^{41}-A^{40}$	1.00249 ± 05	$d\beta$	u
$A^{41}-K^{41}$	0.00274 ± 20	β^-	v
$K^{40}-A^{40}$	0.0017 ± 20	pn	w
	$\geq 0.00167 \pm 10$	K-capture	x
$K^{40}-K^{39}$	1.00070 ± 07	$d\beta$	y
$Ca^{39}-K^{39}$	0.00619 ± 30	theoretical	
$K^{40}-Ca^{40}$	0.00145 ± 05	β^-	aa
	0.00148 ± 03	β^-	bb
$Ca^{41}-Ca^{40}$	0.99997 ± 05	$d\beta$	cc
$Ca^{41}-K^{41}$	0.00047 ± 07	pn	dd
$K^{39}-Ca^{42}$	2.99669 ± 40	$\alpha\beta$	ee
$K^{42}-Ca^{42}$	0.00383 ± 20	β^-	v
$K^{43}-Ca^{43}$	0.000868 ± 03	β^-	ff
$Sc^{41}-Ca^{41}$	0.00630 ± 20	β^+	b
$Sc^{43}-Ca^{43}$	0.00226 ± 25	β^+	gg
$Ca^{40}-Sc^{43}$	3.00036 ± 40	$\alpha\beta$	ee
$K^{41}-K^{41}$	1.00109 ± 10	$\alpha\beta$	hh
$Ca^{40}-A^{40}$	0.00032 ± 08	mass-spectros.	ii

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the experimental and computed masses against A for various Z and N . Tables I and II summarize data employed for determination of the masses, the mass values, estimates of their probable errors, and the references. Figure 3 indicates the experiments which relate the various nuclear masses.

TABLE II. Atomic masses based on Table I.*

Standard: $S^{32}=31.98089, A^{40}=39.97516$				
Nucleus	Mass, based on Table I	Comment	Masses computed from Bohr-Wheeler formula	Δ difference in m.m.u.
P^{31}	30.98239 ± 40	Average of $P^{31}\alpha\beta S^{34}$ and $P^{31}d\beta P^{32}$ reactions	30.98218	0.21
P^{32}	31.98268 ± 20		31.98315	-0.47
P^{34}	33.98326 ± 25		33.98219	0.07
S^{31}	30.98758 ± 50		30.98756	0.02
S^{32}	31.98089	Standard all nuclei from P^{31} to A^{40} calculated with respect to S^{32}	31.98051	0.38
S^{33}	32.98058 ± 10		32.98011	0.47
S^{34}	33.97780 ± 20	Average of microwave and $S^{32}-d\beta-S^{33}$ reaction	33.97566	2.14
S^{35}	34.97907 ± 45		34.97649	2.58
S^{36}	35.97834 ± 40		35.97510	3.24
S^{37}	36.98111 ± 50		36.97866	2.45
Cl^{33}	32.98607 ± 10		32.98575	0.32
Cl^{34}	33.98436 ± 25		33.98391	0.45
Cl^{35}	34.97889 ± 40	Average of $S^{32}-\alpha\beta-Cl^{35}$ and $S^{33}-d\alpha-Cl^{35}$	34.97817	0.72
Cl^{36}	35.97892 ± 50	Average of $Cl^{35}d\beta Cl^{36}$ and microwave data	35.97855	+0.37
Cl^{37}	36.97640 ± 50		36.97503	1.37
Cl^{38}	37.97868 ± 55		37.97728	1.40
A^{35}	34.98468 ± 50		34.98406	0.63
A^{36}	35.97822 ± 55		35.97717	1.05
A^{38}	37.97320 ± 60	Average of $Cl^{35}\alpha\beta A^{38}$ and $Cl^{38}\beta-A^{38}$ reactions	37.97138	1.82
A^{40}	39.97516 ± 26	Standard, reevaluated using references 10, 11 (Mattauch)	39.97020	4.96
A^{41}	40.97765 ± 05		40.97293	4.72
K^{38}	37.97905 ± 70	Assuming β^+ and γ are in cascade	37.98035	-1.30
K^{39}	38.97613 ± 30		38.97460	+1.53
K^{40}	39.97683 ± 10		39.97446	2.37
K^{41}	40.97491 ± 25		40.97075	4.16
K^{42}	41.97600 ± 30		41.97235	3.65
K^{43}	42.97433 ± 65		42.97038	3.95
A^{37}	36.9783 ± 60		35.97634	1.49
Ca^{39}	38.98232 ± 40	Theoretical value using $\beta^+=4.8$ Mev.	38.98100	+1.32
Ca^{40}	39.97546 ± 08		39.97701	-1.55
Ca^{41}	40.97543 ± 12		40.97299	2.44
Ca^{42}	41.97217 ± 35		41.96829	3.88
Ca^{43}	42.97346 ± 55		42.96888	4.58
Sc^{41}	40.98173 ± 35		40.97955	2.18
Sc^{43}	42.97572 ± 50		42.97141	4.29

* Note that masses and errors given are not absolute values, but are relative to the masses chosen for S^{32} and A^{40} .

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masses of Ca⁴³, K⁴³, and Sc⁴³ given in Table I may be somewhat too high.

III. DISCUSSION OF RESULTS

Figures 1 and 2 show that although there is considerable deviation of the measured masses from the Bohr-Wheeler formula, all deviations with the exception of the value for Ca⁴⁰ lie on relatively smooth curves. The deviations from the curves are not generally larger than 0.5 Mev. There seems to be no abrupt change of mass or of slope of the curves of mass *versus* neutron number near S³⁶, Cl³⁷, A³⁸, or K³⁹, all of which are atoms with 20 neutrons. Similarly, there is no evidence of shell structure at 20 protons from the curves of mass *versus* proton number, except in the case of Ca⁴⁰ which shows a striking deviation from the smooth curves of about

3.5 Mev. This deviation is far greater than the probable error of mass determinations or than the deviation of any other nucleus plotted.

The absence of any change in slope at 20 nucleons makes it rather questionable whether 20 nucleons should be regarded as the closing point of a major shell. The well-known stability of Ca⁴⁰ seems not to be simply connected with the completion of a shell at 20 nucleons. This case indicates that some large deviations from the stability curve may be encountered for other nuclei which are not attributable to neutron or proton shells alone but depend on the combination of neutron and proton numbers. Perhaps the large mass spread of stable Ca isotopes is due to this exceptional stability of Ca⁴⁰ and to a shell at 28 neutrons which makes Ca⁴⁸ stable rather than to a general stability of 20 protons.

On a Difference Equation Method in Cosmic-Ray Shower Theory*

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The recent results of Snyder and Bhabha-Chakrabarty for the cascade theory of cosmic-ray showers are shown to be derivable from a general approach involving the use of the Laplace and Mellin transforms, and a general and powerful method, due to Snyder, for solving the resulting difference equations. Boundary conditions are introduced in a natural and automatic way, and the accuracy of the solution is limited by the possible ways of evaluating the resulting triple complex integral.

I. INTRODUCTION

SNYDER¹ has recently obtained numerical results for the cascade theory of electron-photon showers which appear to be considerably more accurate than those of Bhabha and Chakrabarty.² It is the object of this paper to present a general method of solving the shower equations which yield both of the above-mentioned solutions, and which should be applicable to a number of other problems.

II. THEORY

Using Snyder's¹ notation, we write the diffusion equation for $P(E, t)$, the mean energy spectrum of electrons at depth t , and $\gamma(E, t)$, the mean energy spectrum of photons:

$$\frac{\partial P(E, t)}{\partial t} = \lim_{\delta \rightarrow 0} \left[\int_{E+\delta}^{\infty} P(E', t) R(E', E' - E) \frac{E' - E}{E'^2} dE' - P(E, t) \int_{\delta}^E R(E, E') \frac{E' dE'}{E^2} \right] + \beta \frac{\partial P(E, t)}{\partial E} + 2 \int_E^{\infty} \gamma(E', t) R(E, E') \frac{dE'}{E'}, \quad (1)$$

$$\frac{\partial \gamma(E, t)}{\partial t} = \int_E^{\infty} P(E', t) R(E', E) \frac{E dE'}{E'^2} - \gamma(E, t) \int_0^E R(E', E) \frac{dE'}{E}. \quad (2)$$

In these equations $R(E', E)$ is a function which yields the elementary probabilities per unit path length of the pair-production and bremsstrahlung processes. In the case of high energies, the asymptotic form of R is that of a function homogeneous in E/E' ; this is the only case dealt with in the present treatment.

* This work was started in 1940 (Ph.D. Thesis, University of Michigan, 1941). It was completed at the Brookhaven National Laboratory, under the auspices of the AEC.

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