

### Correlation in the Direction and Polarization of Two Successive Quanta for $Rh^{106}$ , $Co^{60}$ , and $Cs^{134}$

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 April 21, 1950

DEUTSCH and Metzger<sup>1</sup> have reported their observation of the correlation between the polarization of one quantum and the direction of emission of the other for the two successive gamma-rays of  $Rh^{106}$ . We have used a polarimeter of a type similar to that described by Deutsch and have verified their results. In assuming that the two excited states have  $J$ -values 020, the angular correlation function  $W(\theta)$  can be reasonably well duplicated if one uses half the values of the coefficients  $\alpha_2$  and  $\alpha_4$  as given by Hamilton.<sup>2</sup>

By this same device it is possible also for the data to be in agreement with the theoretical curve in the case of the polarization correlation. Spiers,<sup>3</sup> however, has pointed out that by assuming one of the excited levels to consist of two levels which are close together in energy but which have different spins, then it is possible to account for  $W(\theta)$ . The expression  $I_\theta/I_\phi$  for the polarization correlation, which has been developed by Hamilton<sup>4</sup> can also be accounted for by the assumption of the additional level.

The three spin values ( $J_1' = 1, 2, 3$ ) and the two types of transitions (electric dipole and magnetic dipole) suggested by Spiers allow six combinations to be tried for this additional level. Our data, as shown in Fig. 1, indicate that the most likely transitions are  $J_1' = 3$  (magnetic dipole) or  $J_1' = 1$  (magnetic dipole or electric dipole). These choices with the resulting parities are shown in Fig. 2.

The experimental curve for  $Co^{60}$  is shown in Fig. 1. The directional correlation experiments have shown that both transitions are quadrupole. The polarization correlation experiment indicates that both transitions are electric. The parities of the two excited states are then the same as that of the ground state.

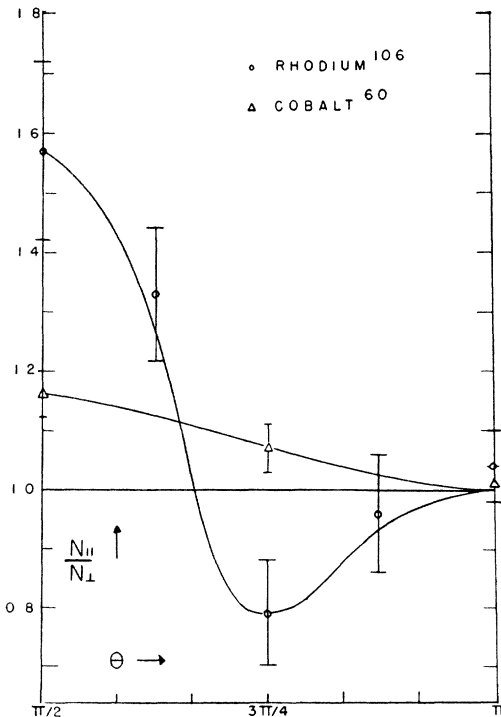


FIG. 1. Polarization of  $Rh^{106}$  and  $Co^{60}$  gamma-rays.

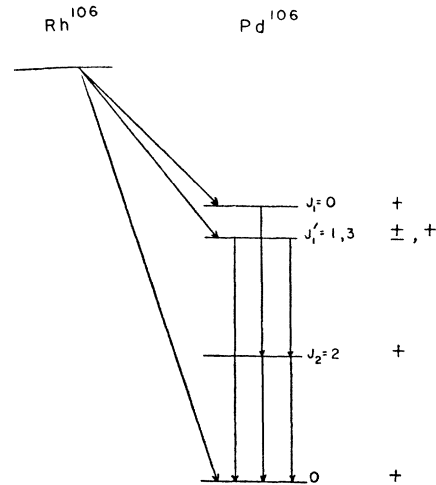


FIG. 2. Energy level scheme for  $Rh^{106}$ . The parity for  $J_1' = 1$  is even or odd, that for  $J_1' = 3$  is even.

In the case of  $Cs^{134}$ , no polarization correlation was observed. According to the theory a correlation should be found for two quadrupole transitions, one of which is magnetic and the other electric, only if the experimental arrangement is such that one can distinguish between which of the two quanta goes to the polarimeter or to the individual counter. In our case this discrimination was not possible. If both transitions are electric or both magnetic, this discrimination is not necessary in order to observe a correlation. If one assumes then that both transitions are quadrupole, the conclusion is that one is electric and the other magnetic.

- <sup>1</sup> M. Deutsch and F. Metzger, *Phys. Rev.* **74**, 1542 (1948).  
<sup>2</sup> D. R. Hamilton, *Phys. Rev.* **58**, 122 (1940).  
<sup>3</sup> J. A. Spiers, *Phys. Rev.* **78**, 75 (1950).  
<sup>4</sup> D. R. Hamilton, *Phys. Rev.* **74**, 782 (1948).

### Double Beta-Decay of $Te^{130}$

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 April 27, 1950

IN a previous letter<sup>1</sup> the authors described the preliminary results of a study of the double beta-transition  $Te^{130} \rightarrow Xe^{130}$  by the method of isotopic analysis of xenon extracted from geologically old tellurium ores. The ore on which that work was based was blasted from an outcrop at Mångfallberget, Sweden, and it contained 12 percent by weight of the mineral  $Bi_2Te_3$ . Although the age of the telluride minerals in Mångfallberget and in Boliden, Sweden is known to be  $1500 \pm 500$  million years, there exists an uncertainty in the "xenon age" of the Mångfallberget material owing to the possibility of comparatively recent crystal alteration by percolating surface waters.

As a result of the generous cooperation of Dr. Erland Grip of the Boliden Mining Company, we have recently obtained some excellent samples of 70 percent rich  $Bi_2Te_3$  from the 240-meter level of the Boliden mine. According to geologists, it is improbable that the crystals of this  $Bi_2Te_3$  have been affected by a recent alteration. A 371-g sample of this material, containing 124 g of tellurium, was broken up to pea size and vacuum roasted at temperatures high enough to decompose the mineral and produce vigorous boiling of the molten bismuth and tellurium. The rare gases evolved were purified as before,<sup>1</sup> and were found to consist of  $2.4 \times 10^{-2}$  cc S.T.P. argon plus  $2.6 \times 10^{-7}$  cc S.T.P. xenon. The results of an isotopic analysis of this xenon are presented in Table I. The percentage composition of the excess or radiogenic

TABLE I. Isotopic composition of xenon extracted from 371 g of a  $1.5 \times 10^9$  year old mineral containing 70 percent  $\text{Bi}_2\text{Te}_3$ . The total xenon content was  $2.6 \times 10^{-7}$  cc S.T.P.

Mass	$\text{Bi}_2\text{Te}$ No. 4	Normal	Diff.	Difference normalized
124	<0.004	0.0005	<0.0035	<0.29 —
126	<0.005	0.0005	<0.0045	<0.38 —
128	<0.014	0.011	<0.003	<0.25 —
129	$\approx 1.000$	0.1525	0.8475	$71.66 \pm 0.7$
130	0.0652	0.0236	0.0416	$3.52 \pm 0.17$
131	0.4088	0.1235	0.2853	$24.12 \pm 0.3$
132	0.1566	$\approx 0.1566$	$\approx 0.0000$	$\approx 0.00$ —
134	0.0651	0.0611	0.0040	$0.34 \pm 0.25$
136	0.0562	0.0519	0.0043	$0.36 \pm 0.16$

xenon has been given in the last row of the table. These percentages were calculated by assuming that all of the  $\text{Xe}^{132}$  present was normal (atmospheric), and by subtracting corresponding amounts at the other mass positions. It is evident from the table that, aside from the somewhat questionable excess xenon at mass 134 and 136 (possibly due to uranium fission), the excess xenon is distributed among the isotopes  $\text{Xe}^{129}$ ,  $\text{Xe}^{130}$ , and  $\text{Xe}^{131}$ .

The excess  $\text{Xe}^{129}$  and  $\text{Xe}^{131}$  is probably caused by  $(n, \gamma)$  reactions on  $\text{Te}^{128}$  and  $\text{Te}^{130}$ . To account for the neutron "flux" required to produce this much xenon in  $1.5 \times 10^9$  years, it is necessary to assume that there was considerable uranium in the immediate neighborhood of the tellurium mineral. Dr. Grip informs us that unusually large amounts of the uranium mineral thucholite have been found in the stope from which the  $\text{Bi}_2\text{Te}_3$  was taken. Thus we ascribe the excess  $\text{Xe}^{129}$  and  $\text{Xe}^{131}$  to the proximity of such a deposit.

An interesting discrepancy appears in the ratio of  $\text{Xe}^{129}$  to  $\text{Xe}^{131}$  found in the sample. The present measurements show this to be 3.0, whereas one would expect a ratio of 0.6 from Seren's values<sup>2</sup> for the tellurium cross sections. One possible explanation is that  $\text{Xe}^{129}$  was also produced by the decay of small amounts of  $\text{I}^{129}$  ( $\sim 3 \times 10^7$  yr.) present in the mineral. This nuclide, as yet undetected in nature, may have been formed originally in amounts comparable to that of  $\text{I}^{127}$ .

The excess  $\text{Xe}^{130}$  is attributed to double beta-decay of  $\text{Te}^{130}$ . There appears to be no other explanation for its formation. Assuming an age of  $1.5 \times 10^9$  years for the  $\text{Bi}_2\text{Te}_3$ , the excess  $\text{Xe}^{130}$  present corresponds to double beta-decay of  $\text{Te}^{130}$  with a half-life of  $1.4 \times 10^{21}$  years. This result is to be compared with theoretical half-lives of  $6 \times 10^{14}$  years and  $10^{24}$  years, the former computed from the Majorana theory of the neutrino, and the latter from the Dirac theory. Both calculations are for allowed transitions with 1.6 Mev of available energy.

<sup>1</sup> M. G. Inghram and J. H. Reynolds, Phys. Rev. **76**, 1265 (1949).

<sup>2</sup> Seren, Friedlander, and Turkel, Phys. Rev. **72**, 888 (1947).

## The Effect of the Compressibility of the Earth on Its Magnetic Field

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April 25, 1950

IN his recent review of geomagnetism associated with the earth's interior Elsasser<sup>1</sup> has noted that the equation for the magnetic induction  $\mathbf{B}$  in a fluid moving with velocity  $\mathbf{v}$  reduces to

$$D\mathbf{B}/Dt = \mathbf{B} \cdot \text{grad} \mathbf{v} - \mathbf{B} \text{ div} \mathbf{v}, \quad (1)$$

provided displacement currents and decay may be neglected. He remarks that "we may safely drop the last term, considering the fluid as incompressible," so that (1) reduces to a form "analogous to the well-known Helmholtz equations for the conservation of vorticity . . .," whence results analogous to the Helmholtz theorems follow.

While the neglect of volume changes is doubtless justified for motions at uniform depth, the great pressure gradients conjectured to exist in the earth's interior make it rather unlikely that Elsasser's approximation is justified if there be any considerable vertical motion. In some of the models suggested by Bullard<sup>2</sup> a part of the flow is nearly vertical. The neglect of compressibility effects is quite unnecessary, however, to obtain the conservation theorems. Equation (1) as it stands is a youthful discovery of Lagrange,<sup>3</sup> and Nanson<sup>4</sup> observed that by using Euler's continuity equation

$$\text{div} \mathbf{v} = -D \log \rho / Dt, \quad (2)$$

where  $\rho$  is the density, one can reduce it to the form

$$D(\mathbf{B}/\rho) / Dt = (\mathbf{B}/\rho) \cdot \text{grad} \mathbf{v}. \quad (3)$$

Hence the analogs of the Helmholtz theorems for the present instance may be stated in the following form: (a) the lines of induction are material lines, (b) the flux of induction,<sup>5</sup>  $\int_S \mathbf{B} \cdot d\mathbf{S}$ , is constant in time for a material surface  $S$ .

Among the finest proofs of these results are those of Kirchhoff,<sup>6</sup> who derived them directly from Cauchy's formula<sup>7</sup>

$$\mathbf{B}/\rho = \mathbf{B}_0/\rho_0 \cdot \text{GRAD} \mathbf{r}, \quad (4)$$

where  $\mathbf{B}_0$  and  $\rho_0$  are the values of  $\mathbf{B}$  and  $\rho$  at some arbitrary initial instant  $t=0$ , and  $\text{GRAD} \mathbf{r}$  is the gradient of the present position field  $\mathbf{r}$  with respect to the initial position field  $\mathbf{R}$ . It is easy to see that (4) is the general solution of (3).

In fact, however, Zorawski<sup>8</sup> showed directly that a condition of the form (a) is both necessary and sufficient for the validity of theorems of the Helmholtzian type for the field  $\mathbf{B}$ , irrespective of its physical significance.

Equivalent to (b) is the conservation of the circulation of the magnetic vector potential around a material circuit.

For motions in the earth's interior the case when the flow is rotationally symmetric and the field  $\mathbf{B}$  is perpendicular to the axial planes might be relevant for part of the motion. Equation (4) then reduces to a result analogous to the vorticity convection theorem of Svanberg:<sup>9</sup>

$$B/r\rho = \text{const.} \quad (5)$$

for each particle,  $r$  being the distance from the axis of symmetry. In motions of this type the effect of density changes appears in a particularly lucid way, tending to counteract the increase in  $B$  incident upon approaching the axis of symmetry.

<sup>1</sup> W. M. Elsasser, Rev. Mod. Phys. **22**, 1 (1950). See p. 30.

<sup>2</sup> E. C. Bullard, Proc. Roy. Soc. **A197**, 433 (1949). See Section 8.

<sup>3</sup> J. L. Lagrange, Misc. Taur. **2**<sup>o</sup> (1760-1761), 196-298 (1762) = Oeuvres **1**, 365-468. See Chapter XLII.

<sup>4</sup> E. J. Nanson, Mess. Math. **3**, 120-121 (1874).

<sup>5</sup> This result was discovered by Cowling. See W. M. Elsasser, Phys. Rev. **72**, 821 (1947), see p. 827.

<sup>6</sup> G. Kirchhoff, *Vorlesungen über mathematische Physik: Mechanik* (Leipzig, 1876); second edition (1877); third edition (1883). See Vorlesung 15, §3.

<sup>7</sup> A.-L. Cauchy, *Théorie de la propagation des ondes à la surface d'un fluide pesant d'une profondeur indéfinie* (1815), Mem. Divers Savants (2) **1** (1816), 3-312 = Oeuvres (1) **1**, 5-318. See part I, Section 1, §4.

<sup>8</sup> K. Zorawski, Bull. Acad. Sci. Cracovie, Comptes Rendus 335-342 (1900). See R. Prim and C. Truesdell, Proc. Am. Math. Soc. **1**, 32-34 (1950).

<sup>9</sup> A. F. Svanberg, *Om fluides rörelse*, K. Vetenskaps-Acad. Handlingar (1839), pp. 139-154 (1841)—trans. "Sur le mouvement des fluides," J. Reine Angew. Math. **24**, 153-163 (1842).

## The Production of $\pi^+$ -Mesons by Protons on Protons in the Direction of the Beam\*

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April 26, 1950

A SIMPLE method<sup>1</sup> has recently been developed for measuring the differential production cross sections of positive and negative  $\pi$ -mesons when various nuclei are struck by high energy charged particles from the 184-in. synchrocyclotron. In this