

FIG. 1. Delayed coincidences as a function of delay time.

Using a high speed synchroscope developed by G. G. Kelley² photographs of the individual signals from the Type 5819 multi-

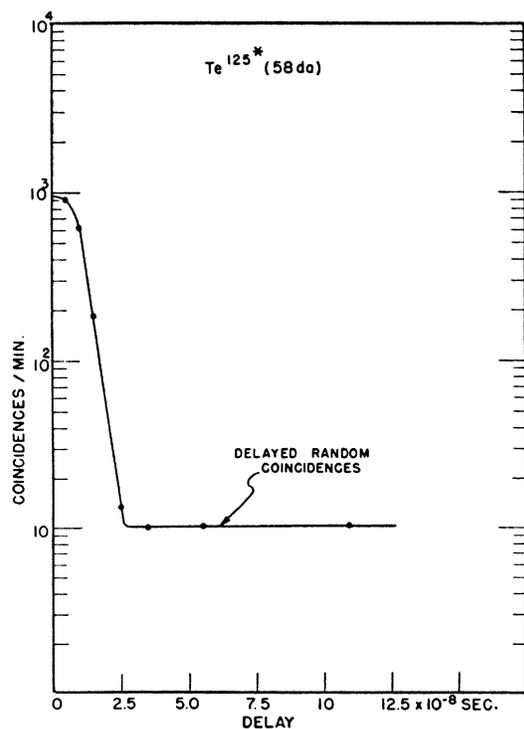


FIG. 2. Coincidences as a function of delay time obtained with a source of Te^{125} .

plier showed the presence of delayed signals due to the decay of the isomeric state.

Hg^{197} as listed in the Table of Isotopes³ is known to decay by K -capture with periods 23 hr. and 64 hr. All points on the decay curve of Au^{197*} appeared to decay according to the 23-hr. period. Huber *et al.*⁴ have observed that 4 percent of the disintegrations of Hg^{197} (23 hr.) lead to a 7.5 sec. isomeric state in Au^{197*} and that 96 percent of the disintegrations are followed by γ -ray transitions corresponding to 165 and 135 kev which are strongly converted.

A delayed coincidence scintillation spectrometer¹ is being used to investigate both the radiation spectrum announcing the formation of the isomeric state and the spectrum resulting from its decay. These results and a discussion of the disintegration scheme will be reported later.

* This document is based on work performed under Contract No. W-7405, eng. 26 for the Atomic Energy Project at Oak Ridge National Laboratory.

¹ F. K. McGowan, Phys. Rev. 76, 1730 (1949).

² G. G. Kelley (to be published).

³ G. T. Seaborg and I. Perlman, Rev. Mod. Phys. 20, 630 (1948).

⁴ Huber, Steffen and Humbel, Helv. Phys. Acta 21, 192 (1948).

Lifetimes of Excited States of Au^{197} , Fe^{57} , Cd^{111} , Hg^{198} , and Tl^{203} *

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November 10, 1949

WE have applied the method of delayed coincidences to the measurement of nuclear lifetimes between 5×10^{-9} sec. and 1×10^{-6} sec. in several radioactive decays. Scintillation counters using stilbene or anthracene with RCA 5819 photomultiplier tubes were used with wide band amplifiers of a design due to Elmore and a fast electronic coincidence circuit. Beta-beta-, beta-gamma-, and gamma-gamma-coincidences were studied. In most cases the lifetimes of the excited states were too short to be measured by our instrument and only upper limits can be set. Table I summarizes some of our results. The 8×10^{-8} sec. state in Cd^{111} had been previously reported by Deutsch and Stevenson¹ in the decay of In^{111} . We have now also found it in the decay of the $43'$ isomeric state of Cd^{111} . The new 8×10^{-9} sec. states in

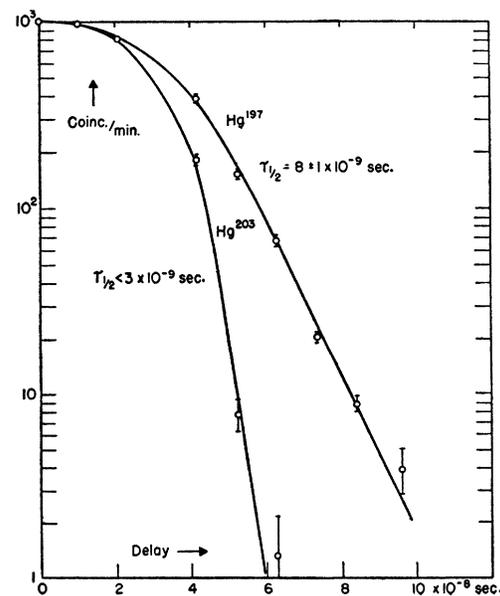


FIG. 3. Typical decay curves.

TABLE I. Life-times of excited states.

	Parent	$\tau_{\frac{1}{2}} \times 10^8$ sec.	E_{γ} Mev
Au ¹⁹⁷	Hg ¹⁹⁷	0.8 ± 0.1	0.13
Fe ⁵⁷	Co ⁵⁷	11 ± 1	0.014
Cd ¹¹¹	In ¹¹¹ Cd ^{111*}	8 ± 1	0.247
Hg ¹⁹⁸	Au ¹⁹⁸	<0.4	0.411
Tl ²⁰⁸	Hg ²⁰⁸	<0.3	0.28

An¹⁹⁷ appears in the 23 hr. Hg¹⁹⁷ decay but not in the 18 hr. Pt¹⁹⁷. The probable energy of 135 kev is assigned on the basis of the decay scheme of Frauenfelder *et al.*,² and the fact that the delayed conversion electrons are more readily absorbed than those preceding the delay. The 1.1×10^{-7} sec. state in Fe⁵⁷ appears in the decay of 270 d Co⁵⁷. The 14-kev gamma-ray has been established by magnetic spectrometer and proportional counter measurements of Hedgran and Deutsch (to be published). Assignment of the observed life time to this transition is based on absorption of the delayed and preceding radiations. Most of the counts in both components are due to iron K x-rays which are quite well detected with anthracene crystals. The upper limit of 4×10^{-9} sec. on the lifetime of the 411-kev state in Hg¹⁹⁸ is established beyond doubt. It contradicts the result of MacIntyre³ on the same transition. Besides the activities listed in Table I we studied and failed to find measurable periods in the decay of Ir¹⁹², Te^{121, 123, 125} (Sb and *d*), Au¹⁹⁹, Cs¹³⁴, Co⁶⁰, Rh¹⁰⁶. Bittencourt and Goldhaber⁴ had reported a 3×10^{-8} sec. state in the decay of Te¹²¹ which we are not able to confirm. Figure 1 shows typical decay curves obtained.

Some of the radionuclides were obtained from the AEC, the others were prepared in the MIT cyclotron.

* This work was supported in part by the Joint Program of the AEC and the ONR.

¹ Deutsch and Stevenson, Phys. Rev. **76**, 184 (1949).

² Frauenfelder *et al.*, Helv. Phys. Acta **20**, 238 (1947).

³ W. J. MacIntyre, Phys. Rev. **76**, 312 (1949).

⁴ Bittencourt and Goldhaber, Phys. Rev. **70**, 780 (1946).

Applications of Total Reflection of Neutrons*

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November 17, 1949

TOTAL reflection of neutrons provides a powerful technique for the measurement of coherent scattering cross sections and phases. Some examples of its possible application are:

1. *Low cross section.*—A typical application would be to vanadium. Shull and Wollan¹ using crystal diffraction have been able only to set an upper limit of 0.1 barn for σ_{coh} and are unable to determine the sign of the amplitude. The critical angle for total reflection is given by

$$\theta_c = \lambda(NC/\pi)^{\frac{1}{2}},$$

where λ is the neutron wave-length, N is the number of nuclei per cc, and C is the coherent amplitude (or the average of the coherent amplitudes of the constituent nuclei in the case of compounds). Assuming $\sigma_{\text{coh}} = 0.01$ barn, for $\lambda = 4.4 \text{ \AA}$ we find $\theta_c \sim 7$ minutes. The presence or absence of total reflection determines the sign of the amplitude without any precise measurements. Even for $\sigma_{\text{coh}} = 0.01$ barn, θ_c will be ~ 4 minutes. Any impurities in the vanadium of the mirror could cause large errors. Another possible mirror would be vanadium carbide; for a vanadium cross section of 0.1 barn, the two possible signs of the amplitude give critical angles differing by ~ 20 percent.

2. *Highly absorbing materials.*—It is difficult to apply crystal diffraction methods to measurement of σ_{coh} for strong absorbers.

But the real part of the index of refraction, and the critical angle depend very little on the absorption cross section.² With $\sigma_{\text{coh}} = 1$ barn and $\sigma_{\text{abs}}/\lambda = 2 \times 10^8$ barns/Å, the effect of the absorption on the critical angle is only \sim three percent. By mirror experiments it should be possible to measure σ_{coh} for Cd, B, and other strong absorbers. We should note that the shape of the reflection curve depends on the absorption.³ In the presence of absorption we no longer obtain "total" reflection; the curve of reflectivity smooths out as the absorption increases.

3. *Coherent amplitude of hydrogen.*—Measurement of critical angle for a given wave-length can be applied to a determination of the coherent amplitude of hydrogen by using hydrogen compounds as mirrors. This method has one great advantage over those previously used (ortho-parahydrogen scattering and crystal diffraction) in that no complicated calculations and no uncertain corrections are necessary. In total reflection we are dealing with the forward-scattered beam (zero-order Bragg reflection) so that form factors due to molecular structure or thermal agitation are always unity.

4. *Neutron-electron interaction.*—Here again the form factor for the electron scattering is unity so that the neutron-electron interaction contributes fully. For a tungsten mirror a neutron-electron interaction of 5000 volts will contribute 1.4 percent to θ_c . The nuclear amplitude would have to be determined by a separate experiment in which the electron form factor is made to approach zero (e.g., by use of short wave-lengths).

* Condensed from Argonne National Laboratory Report-4298 (May 12, 1949).

¹ C. G. Shull and E. O. Wollan (unpublished).

² M. L. Goldberger and F. Seitz, Phys. Rev. **71**, 294 (1947).

³ A. H. Compton and S. K. Allison, *X-Rays in Theory and Experiment* (D. Van Nostrand Company, Inc., New York, 1935), p. 305 ff.

Note on Rotational Universe

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November 10, 1949

THE model of rotational universe was proposed by Gamow,¹ and studied in detail by Gödel.² The angular velocity calculated by Gödel is $2(\pi\kappa\rho)^{\frac{1}{2}}$ where κ is Newton's gravitational constant and ρ is the density of the universe. This result is rather strange because the angular momentum is not conserved during expanding. Assume that the universe is an ellipsoid of major axes "a" and minor axis "b", the rotation is along the minor axis, then the angular momentum in classical sense is

$$m = \frac{2}{3}Ma^22(\pi\kappa\rho)^{\frac{1}{2}},$$

where M is the mass of the universe. During expanding, ρ is only decreasing inversely as a^2 or a^3 and so m will be increasing as a or $a^{\frac{1}{2}}$. This is impossible except introducing new concept or new mechanism.

¹ G. Gamow, Nature **158**, 549 (1946).

² Kurt Gödel, Rev. Mod. Phys. **21**, 447 (1949).

On the Hyperfine Structure of the ²P_{1/2} State of Tl²⁰⁵ and Tl²⁰³

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November 21, 1949

IT has been found that an atomic beam of thallium may be evaporated from an iron oven of the type conventionally used in molecular beam research and that the beam may be detected by the formation of positive ions on a hot oxidized tungsten filament. It has therefore been possible to observe the hyperfine