

attempts of the electron avalanches to initiate the retrograde streamers which are the fore-runners of negative point breakdown with larger points.¹³ They result from avalanches which would cause breakdown by positive streamers in a larger uniform gap. The streamers cannot advance all the way to the point cathode owing

¹³ L. B. Loeb and J. M. Meek, *Mechanism of the Electrical Spark* (Stanford Press, 1941), pp. 82, 87, 99, 170.

to the positive space charge which accumulates before the whole discharge is choked off. For still larger points the retrograde streamers reach the point and may cause breakdown.

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Zeeman Effect Data and Preliminary Classification of the Spark Spectrum of Praseodymium—Pr II

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The Zeeman effect of praseodymium has been studied at fields up to 95,000 oersteds over the range 2400 to 7100 Å. g and J values have been determined for 74 Pr II levels from resolved Zeeman patterns of 141 lines. With these data, together with new wave-length data from the M.I.T.-W.P.A. Wavelength Project which have been applied to the spectroscopic interval sorter and interval recorder, a quadratic term array has been set up which accounts for 312 lines. This array is consistent with King's temperature classification of the lines and with all previous hyperfine structure observations. It is self-consistent in g values to an average deviation of 0.005 unit, and in wave numbers to an average deviation of 0.08 cm^{-1} . Previously published hyperfine structure measurements were found insufficiently precise to aid the classification, but were of value in its verification. The lowest term of Pr II is found to be $f^3(4I^{\circ}) \cdot s - 5I^{\circ}_4$. Most of the strong lines showing hyperfine structure arise from the f^3s configuration.

THE spectra of praseodymium have long defied classification, both because of the complexity of this rare-earth atom, and because many of its lines show hyperfine structure difficult to measure. White¹ has published extensive and valuable data on the hyperfine structure of Pr II, but these prove to be not quite precise enough to establish a classification.

The *M.I.T. Wavelength Tables*² list 2708 strong lines of praseodymium, and the unpublished W.P.A. card catalog at M.I.T. contains 3454 lines of all intensities assigned to the element. Actually the spectrum is so rich that the number of observable lines is limited only by the difficulty

of distinguishing true lines from bands. The available description of the spectrum between 8000 and 2400 Å is fairly good, though limited in wave-length precision by the difficulties of measurement introduced by partially resolved hyperfine structure patterns (h.f.s.). Unless wave-length values accurate to $\pm 0.003 \text{ Å}$ can be obtained, it is impossible in so complex a spectrum to determine which is the correct quadratic array of the many obtained from the combination

TABLE I. *Fields for various spectrograms.*

SET NO.	FIELD	SET NO.	FIELD
Z 41 <i>H</i>	95,000 oersteds	Z 47	89,900 oersteds
Z 41 <i>L</i>	72,340	Z 48	87,400
Z 42	90,860	Z 63	89,290
Z 44	91,500	Z 73	87,540
Z 45	91,500	Z 75- <i>p</i>	85,510
Z 46	86,400	Z 75- <i>n</i>	87,970

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¹ H. E. White, *Phys. Rev.* **34**, 1397 (1929).

² (John Wiley and Sons, New York, 1939.)

principle with the interval sorter³ and recorder.⁴ In the absence of data of suitable precision regarding wave-length and h.f.s., Zeeman effect measurements appeared to give the most powerful method of attacking the classification of the Pr spectra.

No Zeeman effect measurements on Pr lines appear to have been made hitherto, probably because of the high magnetic and spectrographic resolution needed for lines showing h.f.s. The availability of fields approaching 100,000 oersteds,⁵ combined with spectrographs of high dispersion capable of giving Zeeman exposures in a few minutes, has now made possible the resolution of many patterns. The data thus obtained gave the desired start on the classification, which was then verified and extended by wave-length, h.f.s., and temperature class data.

The three concave gratings and other apparatus used have been described in previous papers dealing with the spectra of other elements.^{6,7} An arc carrying 4 amp. was run between electrodes of silver containing 20 percent of powdered praseodymium chloride. The arc so obtained was the brightest and steadiest we have yet produced in a strong magnetic field with any element, but most of the lines emitted were found to belong to Pr II and Pr III. Eleven sets of spectrograms were obtained, at fields listed in Table I. Each set contained p , n , and no-field exposures on each of from 12 to 24 plates 20 inches long, distributed throughout the spectrum.

All spectrograms were measured in both directions with an automatic comparator,⁸ and were reduced by W.P.A. clerical workers, who have been of great assistance in reading, recording, and averaging the measurements used in calculating the data which follow. The field intensities given in Table I were calculated from measurements on the *raies ultimes* of silver, copper, and calcium, and are believed to be correct to within ± 0.3 percent.

³ G. R. Harrison, *Rev. Sci. Inst.* **4**, 581 (1933).

⁴ G. R. Harrison, *J. Opt. Soc. Am.* **28**, 290 (1938).

⁵ G. R. Harrison and F. Bitter, *Phys. Rev.* **57**, 15 (1940).

⁶ G. R. Harrison and J. Rand McNally, Jr., *Phys. Rev.* **58**, 703 (1940).

⁷ J. P. Molnar and W. J. Hitchcock, *J. Opt. Soc. Am.* **11**, 523 (1940).

⁸ G. R. Harrison, *J. Opt. Soc. Am.* **25**, 169 (1935); *Rev. Sci. Inst.* **9**, 15 (1938); G. R. Harrison and J. P. Molnar, *J. Opt. Soc. Am.* **30**, 343 (1940).

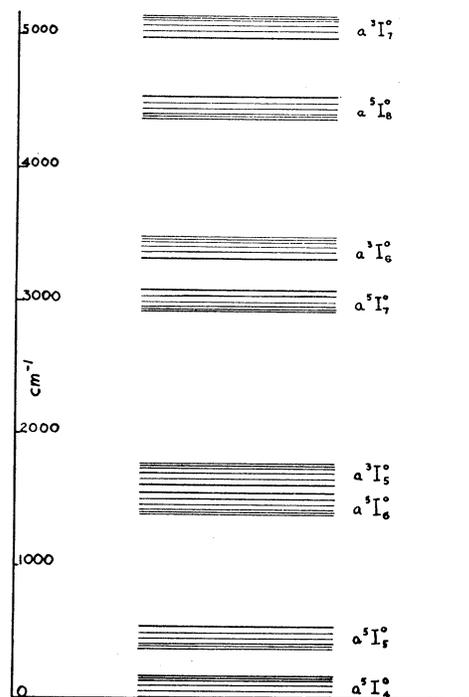


FIG. 1. Lowest terms of Pr II, showing alternate inverted and normal hyperfine-structure multiple terms in pairs.

RESULTS

Table II contains the lines of Pr II which we have classified thus far with reasonable assurance. The wave-lengths in angstroms given in the first column have been taken from the *M.I.T. Wavelength Catalogs*, as have the arc intensities in the second column. The third column gives the temperature class, while the fourth gives the h.f.s. of the line according to White,¹ or in cases marked a , from King.⁹ v and r in this column indicate shaded toward short and long wave-lengths, respectively. The next column contains the observed wave number, while the column headed $o-c$ gives the difference, in 0.01 cm^{-1} units, between the observed wave number and that calculated from the difference between the finally assigned wave numbers of the parent levels. The column headed "Comb." gives the lower and upper terms which appear to produce the line. This is followed by two columns giving the g values for these lower and upper levels as deduced from the observed patterns of the line where these could be interpreted.

⁹ A. S. King, *Astrophys. J.* **68**, 194 (1928).

TABLE II. Pr II Lines.

λ	<i>I</i>	CLASS	H.F.S.	σ OBS.	<i>o-c</i>	COMB.	g_1	g_2	λ	<i>I</i>	CLASS	H.F.S.	σ OBS.	<i>o-c</i>	COMB.	g_1	g_2
7099.54	2	—	—	14081.54	+ 8	$b^1F_3 - 3s$	—	—	5064.84	2h	—	—	19738.47	+22	$a^3F_6 - 6s$	—	—
7021.54	6	—	—	14237.97	-13	$a^3H_6 - 8s$	—	—	5061.50	1	—	—	19751.50	-14	$b^1F_4 - 26s$	—	—
6850.55	4	—	—	14593.35	-12	$b^1F_4 - 2s$	—	—	5060.40	2	—	—	19755.79	+ 5	$a^3F_3 - z^3I_6$	—	—
6564.632	10	V	—	15228.95	+ 6	$b^1F_4 - z^3K_5$	0.622	0.818	5034.15	30	V	—	19857.76	+ 7	$a^3K_9 - z^3K_9$	1.219	1.219
6475.287	3	—	—	15439.07	+ 2	$b^1F_4 - 5s$	—	—	5015.543	4	—	—	19932.48	-15	$a^3F_3 - 1s$	—	—
6454.865	3	—	—	15487.75	-11	$b^1F_3 - 11s$	—	—	5013.20	2	—	—	19941.79	+ 9	$b^1F_4 - 28s$	—	—
6397.996	12	V	—	15625.58	+13	$b^1F_3 - z^3K_6$	0.926	(0.950)	5004.584	5	—	—	19976.14	- 1	$a^3L_9 - 30s$	—	—
6305.262	9	V	—	15885.39	+ 1	$b^1F_7 - z^3I_7$	—	—	5002.454	8	V	—	19984.63	- 3	$a^3K_7 - 22s$	—	—
6278.675	9	V	—	15922.52	+ 2	$b^1F_6 - z^3K_7$	(1.073)	1.056	4991.774	3	—	—	20027.39	+12	$a^3F_3 - 1s$	—	—
6262.539	9	V	—	15963.56	+10	$b^1F_6 - 17s$	(1.073)	1.120	4991.774	3 ^c	—	—	20027.39	+ 8	$a^3L_8 - 23s$	—	—
6244.344	10	V	—	16010.08	+ 6	$b^1F_6 - z^3I_6$	—	—	4977.578	2h	—	—	20084.50	+10	$a^3L_6 - 11s$	—	—
6200.79	5	—	—	16131.40	- 3	$b^1F_7 - z^3K_8$	1.145	(1.143)	4971.18	4	—	—	20110.35	+18	$a^3K_7 - 24s$	—	—
6182.343	12	V	—	16170.64	- 9	$b^1F_8 - 30s$	1.233	1.073	4968.35	1	—	—	20122.22	+17	$a^3K_8 - 30s$	—	—
6165.945	50	V	—	16213.64	- 3	$b^1F_4 - z^3I_4$	0.620	(0.684)	4956.645	40	V	67a	20169.33	+12	$a^3F_7 - 15s$	(1.143)	1.057
6161.194	50	V	—	16226.15	+ 5	$b^1F_6 - z^3I_6$	0.92	0.89	4943.735	3	—	—	20221.99	- 4	$a^3L_6 - z^3K_6$	—	—
6159.093	3	—	—	16231.67	+18	$a^3H_5 - 17s$	1.04	1.04	4931.28	1	—	—	20273.07	+ 4	$a^3K_6 - 16s$	—	—
6141.508	6	V	57a	16278.15	+10	$a^3H_6 - z^3I_6$	—	—	4925.630	10	V	57a	20296.32	- 1	$a^3F_3 - 2s$	—	—
6025.723	25	V	—	16590.93	- 2	$b^1F_8 - z^3I_8$	—	—	4914.418	3	—	—	20342.63	0	$a^3K_6 - z^3K_7$	—	—
5967.837	15	V	—	16751.86	+ 1	$b^1F_6 - 22s$	1.075	0.991	4912.629	10	V	—	20350.04	- 7	$a^3K_8 - z^3K_7$	1.13	(1.106)
5892.251	10	V	—	16966.81	+ 1	$b^1F_8 - z^3I_7$	1.235	(1.187)	4901.483	5	—	—	20396.31	- 6	$a^3L_9 - z^3I_8$	—	—
5879.253	10	V	—	17004.26	0	$b^1F_7 - z^3K_7$	1.149	(1.106)	4886.045	20	V	—	20460.75	+ 1	$a^3L_7 - z^3K_7$	—	—
5823.723	60	V	2	17166.40	-30	$b^1F_6 - z^3I_6$	—	—	4879.121	30	6r	—	20489.79	0	$a^3F_7 - z^3K_7$	—	—
5813.593	3	—	—	17196.31	-11	$b^1F_7 - z^3I_8$	1.153	(1.154)	4877.81	40	V	6r	20495.30	+17	$a^3F_6 - 9s$	1.046	1.108
5810.622	10	—	—	17205.10	- 2	$b^1F_8 - z^3K_9$	1.227	(1.217)	4876.257	20	V	—	20501.46	-24	$a^3L_7 - 17s$	—	—
5807.55	2	—	—	17214.21	- 2	$b^1F_8 - z^3I_7$	—	—	4865.237	40	V	—	20548.26	+13	$a^3K_7 - z^3H_6$	1.057	(1.073)
5791.382	8	V	—	17262.26	+ 1	$a^3H_5 - 25s$	1.043	1.063	4859.038	40	V	6r	20574.48	-18	$a^3F_6 - 11s$	—	—
5775.290	4h	—	—	17308.47	+ 5	$b^1F_4 - 12s$	—	—	4839.540	25	—	—	20657.37	+ 1	$a^3K_5 - 12s$	—	—
5719.090	10	V	—	17480.46	- 4	$a^3L_8 - 9s$	1.08	(1.107)	4826.649	40	—	6r	20712.54	+25	$a^3F_6 - z^3K_6$	—	—
5717.785	2h	—	—	17484.45	- 1	$a^3K_7 - 9s$	—	—	4826.310	9	—	—	20713.99	-10	$a^3K_7 - z^3K_8$	—	—
5715.37	4	—	—	17491.84	0	$a^3K_8 - 4s$	—	—	4824.655	5	—	—	20721.09	- 2	$a^3K_5 - 13s$	—	—
5711.65	8	—	—	17503.23	0	$b^1F_7 - z^3H_6$	1.147	(1.103)	4814.344	30	V	—	20765.48	0	$a^3K_9 - z^3K_9$	1.20	(1.145)
5695.93	8	—	—	17551.54	- 1	$b^1F_6 - 26s$	1.078	(1.067)	4801.150	40	V	—	20822.54	-14	$a^3L_6 - z^3I_6$	0.73	0.92
5689.185	1	—	—	17572.35	+ 8	$b^1F_7 - z^3I_7$	—	—	4799.94	2	—	—	20827.79	+ 6	$a^3F_3 - 3s$	—	—
5687.19	4h	—	—	17578.51	-20	$a^3K_5 - 1s$	—	—	4794.903	7	—	—	20849.67	+ 2	$a^3K_5 - 14s$	—	—
5685.61	3	—	—	17583.40	+ 5	$a^3H_5 - z^3H_6$	—	—	4783.354	125	V	6v	20900.00	+ 6	$a^3F_7 - 9s$	1.170	1.105
5681.896	7	V	—	17594.89	- 8	$a^3H_5 - z^3H_5$	1.035	(1.101)	4779.196	15	—	—	20918.19	+ 7	$a^3K_8 - z^3I_7$	—	—
5677.04	3	—	—	17609.94	- 1	$b^1F_7 - 4s$	—	—	4778.303	35	—	6v	20922.10	-27	$a^3F_3 - 3s$	—	—
5662.19	2	—	—	17656.12	+ 7	$b^1F_7 - 19s$	—	—	4766.906	35	V	—	20972.13	+ 4	$a^3F_7 - 10s$	—	—
5647.64	2	—	—	17701.61	- 1	$b^1F_7 - z^3K_6$	—	—	4766.322	20	—	6r	20974.69	+ 3	$a^3F_3 - 4s$	—	—
5637.36	2	—	—	17733.64	-14	$b^1F_7 - 27s$	—	—	4765.222	60	V	6r	20979.53	+ 6	$a^3F_3 - 11s$	1.179	(1.156)
5634.93	2	—	—	17741.54	- 7	$b^1F_6 - 28s$	—	—	4762.727	60	V	6r	20990.52	- 6	$a^3F_6 - z^3K_6$	1.038	(0.992)
5621.854	15	V	—	17782.80	-10	$a^3L_6 - 1s$	—	—	4758.910	3	—	—	21007.36	-18	$a^3K_8 - 29s$	—	—
5610.219	6	V	—	17819.68	+10	$a^3H_5 - 2s$	1.043	1.073	4757.937	100	—	6v	21011.66	+16	$a^3K_7 - 29s$	—	—
5582.39	3	—	—	17908.51	+ 2	$b^1F_7 - 22s$	—	—	4754.835	15	—	6v	21026.55	-14	$a^3F_6 - z^3K_5$	—	—
5582.11	2	—	—	17909.41	- 9	$a^3K_5 - z^3K_7$	—	—	4746.635	15	—	6v	21026.25	-14	$a^3F_6 - z^3K_5$	—	—
5580.39	2	—	—	17914.93	+ 3	$a^3K_5 - 6s$	—	—	4746.93	100	V	6r	21060.38	- 1	$a^3K_5 - 21s$	(0.915)	0.886
5571.841	10	V	—	17942.42	+ 1	$a^3K_5 - 2s$	0.684	0.998	4744.925	100	V	6v	21069.28	- 2	$a^3F_3 - 4s$	(1.064)	1.10
5560.21	1	—	—	17979.95	+ 4	$a^3K_7 - z^3K_6$	—	—	4738.622	8d	—	—	21097.30	+10	$a^3F_7 - 20s$	—	—
5551.08	2	—	—	18009.52	-12	$a^3H_5 - 28s$	—	—	4734.177	100	V	6v	21117.09	- 1	$a^3F_7 - z^3K_6$	1.176	0.961
5543.87	2	—	—	18032.95	- 6	$a^3L_7 - 6s$	—	—	4728.630	40	V	6r	21141.88	- 3	$a^3F_3 - 5s$	(0.860)	0.98
5519.39	2	—	—	18112.93	+ 2	$b^1F_3 - z^3K_5$	—	—	4721.910	3	—	—	21171.97	- 1	$a^3K_5 - 22s$	—	—
5509.146	50	V	—	18146.61	+ 1	$a^3L_6 - 2s$	0.723	0.998	4707.939	50	IV	6v	21234.79	+45	$a^3F_3 - 1s$	0.858	0.986
5502.20	2	—	—	18169.51	+13	$a^3K_9 - z^3K_8$	—	—	4707.541	80	V	6v	21236.58	+ 3	$a^3F_6 - 5s$	1.064	(0.994)
5431.135	3	—	—	18407.25	+ 8	$b^1F_3 - z^3H_7$	—	—	4689.341	4	—	—	21319.01	-13	$a^3F_7 - 22s$	—	—
5411.555	25	IV	—	18473.85	+ 4	$a^3K_5 - 3s$	0.660	1.141	4685.447	5	—	—	21336.73	+ 3	$a^3L_7 - 23s$	—	—
5381.262	60	V	—	18577.85	+ 2	$a^3K_5 - z^3K_5$	0.685	0.822	4679.498	2	—	—	21363.94	+ 4	$a^3L_8 - 30s$	—	—
5368.830	5d	—	—	18620.57	+13	$a^3K_5 - 4s$	—	—	4679.112	8	—	6r	21365.61	-14	$a^3F_7 - 23s$	—	—
5352.403	50	V	37a	18678.02	+ 2	$a^3L_6 - 3s$	0.712	1.155	4678.168	12	—	—	21369.93	0	$a^3K_5 - z^3H_4$	—	—
5331.483	6	V	—	18751.31	0	$a^3K_6 - 11s$	0.914	1.157	4672.595	3	—	—	21395.41	+ 2	$a^3F_7 - z^3K_6$	—	—
5322.778	30	V	—	18781.97	- 5	$a^3L_6 - z^3K_5$	0.718	0.820	4672.081	100	IV	6r	21397.76	+ 4	$a^3F_6 - 6s$	0.868	1.070
5321.82	4	—	—	18785.35	-11	$a^3K_8 - 23s$	—	—	4668.454	6	—	—	21414.39	+ 4	$a^3K_6 - 25s$	—	—
5321.086	8	V	—	18787.94	- 5	$a^3K_8 - 5s$	—	—	4668.230	10	—	6r	21415.42	-14	$a^3F_3 - 13s$	—	—
5312.327	5	V	67a	18818.92	- 2	$a^3I_7 - 9s$	—	—	4664.647	100	V	6r	21431.87	- 6	$a^3F_6 - z^3I_6$	1.037	(1.009)
5311.119	10	V	—	18823.20	-10	$a^3K_9 - 30s$	1.217	1.073	4651.517	125	V	6v	21492.35	- 1	$a^3F_6 - 6s$	1.05	1.07
5308.96	3	—	—	18830.86	+ 9	$a^3L_8 - 15s$	—	—	4646.059	50	V	6r	21517.61	- 4	$a^3F_6 - 7s$	0.863	0.993
5300.145	2	—	—	18862.18	+10	$b^1F_6 - z^3H_6$	—	—	4628.751	200	III	6v	21598.06	+ 2	$a^3F_6 - 2s$	0.873	0.894
5300.145	2	—	—	18862.18	+14	$a^3L_7 - 10s$	—	—	4627.05	15	—	—	21606.01	-11	$a^3L_6 - 16s$	—	—
5298.106	25	V	—	18869.44	+ 2	$a^3L_7 - 11s$	0.928	1.160	4612.072	60	IV	67a	21676.19	- 9	$a^3F_4 - 1s$	0.601	0.995
5292.630	50	IV	—	18888.95	+ 1	<											

TABLE II.—Continued.

λ	<i>I</i>	CLASS	H.F.S.	σ OBS.	<i>o-c</i>	COMB.	<i>g</i> ₁	<i>g</i> ₂	λ	<i>I</i>	CLASS	H.F.S.	σ OBS.	<i>o-c</i>	COMB.	<i>g</i> ₁	<i>g</i> ₂
4493.119	25	—	6v	22250.03	-18	$as^2I^{\circ}_7 - 15r$	—	—	4164.192	200	III	6v	24007.52	-9	$as^2I^{\circ}_6 - 2^5I_6$	—	—
4492.427	25	V	6r	22253.45	-5	$as^2I^{\circ}_6 - 2^5I_6$	—	—	4161.65	8	—	—	24022.18	+1	$as^2I^{\circ}_6 - 2b$	—	—
4487.821	40	V	6r	22276.29	-8	$as^2I^{\circ}_6 - 4c$	—	—	4148.457	15	III	5 γ a	24098.58	+2	$as^2I^{\circ}_6 - 18c$	0.863	1.011
4479.618	35	—	—	22317.09	+16	$as^2I^{\circ}_7 - 29r$	—	—	4143.136	200	III	6v	24129.52	-5	$as^2I^{\circ}_7 - 2^5K_8$	(1.177)	1.141
4472.259	125	IV	6v	22328.85	+8	$as^2I^{\circ}_6 - 11c$	1.054	1.150	4141.257	150	III	6v	24140.47	-16	$as^2I^{\circ}_6 - 2^5I_7$	1.242	1.172
4473.835	30	—	—	22345.93	-5	$as^2I^{\circ}_7 - 29r$	—	—	4132.230	15	IV	6 α a	24193.21	+1	$as^2I^{\circ}_6 - 18c$	—	—
4468.712	125	III	6r	22371.56	-20	$as^2I^{\circ}_6 - 2^5K_6$	0.864	0.963	4118.481	250	III	6v	24273.97	-15	$as^2I^{\circ}_6 - 2^5I_5$	(0.875)	0.920
4458.336	90	V	6v	22423.62	-12	$as^2I^{\circ}_6 - 2^5I_7$	(1.250)	1.125	4111.871	30	IV	5a	24312.99	0	$as^2I^{\circ}_6 - 12c$	0.878	0.914
4454.382	60	IV	6v	22443.52	-10	$as^2I^{\circ}_6 - 5c$	—	—	4100.746	200	III	6v	24378.94	-1	$as^2I^{\circ}_6 - 2^5K_6$	1.256	1.221
4449.867	125	III	6r	22466.30	-10	$as^2I^{\circ}_6 - 2^5K_6$	1.075	0.961	4100.22	15	—	—	24382.08	+1	$as^2I^{\circ}_7 - 27c$	—	—
4429.238	60	III	6r	22570.93	-45	$as^2I^{\circ}_6 - 3c$	0.608	1.165	4096.822	30	III	6 γ 7a	24402.30	-2	$as^2I^{\circ}_6 - 19c$	0.858	1.024
4429.238	100	III	6v	22570.93	+14	$as^2I^{\circ}_7 - 2^5K_7$	(1.177)	1.061	4091.686	50	—	—	24432.93	+7	$as^2I^{\circ}_6 - 20c$	—	—
4421.231	100	IV	6v	22611.80	+5	$as^2I^{\circ}_7 - 17c$	1.175	1.110	4081.018	50	III	6v	24496.80	-16	$as^2I^{\circ}_6 - 19c$	(1.064)	1.035
4413.765	90	III	6 α a	22650.05	0	$as^2I^{\circ}_6 - 2^5K_6$	0.860	0.992	4075.917	15d	—	—	24527.46	-4	$as^2I^{\circ}_6 - 20c$	—	—
4412.155	50	IV	6v	22658.31	0	$as^2I^{\circ}_7 - 2^5I_6$	(1.177)	1.046	4073.305	9	—	—	24543.18	-3	$as^2I^{\circ}_6 - 21c$	—	—
4408.844	125	III	6r	22675.34	-6	$as^2I^{\circ}_6 - 2^5K_6$	0.605	0.823	4062.817	150	III	6r	24606.54	-5	$as^2I^{\circ}_6 - 2^5K_7$	1.035	1.106
4405.849	100	IV	6v	22690.75	-4	$as^2I^{\circ}_6 - 2^5K_6$	(1.250)	1.146	4061.336	4	—	—	24615.51	+4	$as^2I^{\circ}_7 - 2^5K_8$	—	—
4403.605	100	V	6r	22702.31	-3	$as^2I^{\circ}_7 - 30c$	—	—	4056.543	100	III	6r	24644.60	+8	$as^2I^{\circ}_7 - 2^5K_8$	(1.143)	1.149
4395.788	30	—	—	22742.68	-17	$as^2I^{\circ}_6 - 19c$	1.052	(1.025)	4054.845	50	III	6r	24654.92	+12	$as^2I^{\circ}_6 - 22c$	0.862	0.991
4372.385	10	—	—	22864.41	+6	$as^2K^{\circ}_6 - 2^5H_6$	—	—	4044.818	50	III	6 α a	24716.03	-3	$as^2I^{\circ}_6 - 2^5I_5$	0.605	0.910
4368.327	125	III	6r	22885.65	+9	$as^2I^{\circ}_6 - 5c$	0.605	0.994	4039.357	50	III	6v	24749.45	+1	$as^2I^{\circ}_6 - 22c$	1.08	0.98
4365.328	6	—	—	22901.37	+2	$as^2I^{\circ}_7 - 2^5K_7$	—	—	4038.467	15	III	6 γ a	24754.91	-2	$as^2I^{\circ}_6 - 12c$	0.613	0.918
4359.795	100	IV	6r	22930.43	+3	$as^2I^{\circ}_7 - 2^5K_7$	(1.143)	1.102	4034.30	20	—	—	24780.47	+16	$as^2I^{\circ}_6 - 24c$	—	—
4352.752	4	—	—	22967.54	+9	$as^2I^{\circ}_6 - 2^5I_7$	—	—	4033.857	50	III	6v	24783.19	-15	$as^2I^{\circ}_7 - 30c$	(1.177)	1.077
4351.849	80	III	6r	22972.30	-11	$as^2I^{\circ}_6 - 2^5I_5$	(0.860)	0.904	4031.755	50	III	6v	24796.11	+6	$as^2I^{\circ}_6 - 23r$	(1.064)	1.039
4347.490	100	IV	6r	22995.33	0	$as^2I^{\circ}_6 - 22c$	1.047	(0.990)	4015.389	50	III	6 γ a	24897.18	+1	$as^2I^{\circ}_6 - 25c$	0.858	1.073
4338.694	100	IV	6r	23041.98	+4	$as^2I^{\circ}_6 - 23r$	(1.037)	1.075	4008.714	150	III	6r	24938.64	-14	$as^2I^{\circ}_7 - 2^5H_7$	(1.143)	1.209
4333.913	150	IV	6v	23067.37	+32	$as^2I^{\circ}_6 - 2^5I_5$	1.062	0.910	4000.190	50	III	6v	24991.77	-4	$as^2I^{\circ}_6 - 25c$	(1.064)	1.080
4332.487	30	—	—	23074.97	-6	$as^2I^{\circ}_6 - 13c$	—	—	3997.054	100	III	6v	25011.38	-2	$as^2I^{\circ}_7 - 2^5K_7$	(1.177)	1.101
4331.472	5	—	—	23080.37	+21	$as^2I^{\circ}_6 - 2^5H_5$	—	—	3994.834	300	III	6v	25025.28	-28	$as^2I^{\circ}_6 - 2^5H_4$	0.89	0.89
4329.415	50	IV	6 γ 7a	23091.34	-6	$as^2I^{\circ}_6 - 2^5I_5$	(0.860)	1.009	3989.718	200	III	6v	25057.29	-27	$as^2I^{\circ}_6 - 16c$	0.873	0.982
4327.698	3	—	—	23100.50	-8	$as^2K^{\circ}_6 - 26c$	—	—	3982.063	125	III	6r	25105.53	-3	$as^2I^{\circ}_6 - 2^5H_6$	(1.037)	1.106
4323.922	35	—	6r	23120.67	-17	$as^2I^{\circ}_6 - 24c$	—	—	3972.164	125	III	6v	25168.11	-1	$as^2I^{\circ}_6 - 17c$	0.877	1.118
4323.551	100	IV	6r	23122.65	+9	$as^2I^{\circ}_7 - 2^5I_5$	(1.143)	1.167	3971.164	100	III	6r	25174.44	-16	$as^2I^{\circ}_6 - 2^5I_7$	(1.037)	1.187
4314.74	2	—	—	23169.87	+20	$as^2I^{\circ}_6 - 13c$	—	—	3966.573	100	III	6v	25203.57	+1	$as^2I^{\circ}_7 - 2^5I_8$	(1.177)	1.154
4313.843	4	—	—	23174.69	-4	$as^2I^{\circ}_6 - 8c$	—	—	3965.263	100	III	6v	25211.90	+8	$as^2I^{\circ}_6 - 2^5I_7$	(1.064)	1.128
4308.457	20	—	—	23203.66	+9	$as^2I^{\circ}_6 - 14c$	—	—	3964.825	125	III	6v	25214.68	0	$as^2I^{\circ}_6 - 2^5I_6$	(0.875)	1.039
4305.763	150	III	6 α a	23218.18	-6	$as^2I^{\circ}_6 - 2^5I_4$	0.877	0.683	3964.261	60	III	6r	25218.27	0	$as^2I^{\circ}_6 - 2^5H_5$	0.858	1.071
4302.100	60	—	—	23237.95	+25	$as^2I^{\circ}_6 - 25c$	(1.037)	1.065	3962.445	60	III	6 γ a	25229.83	-6	$as^2I^{\circ}_6 - 2^5H_5$	0.869	1.108
4297.764	50	III	6r	23261.40	+10	$as^2I^{\circ}_6 - 7c$	0.601	0.895	3953.516	150	III	6v	25286.81	+7	$as^2I^{\circ}_6 - 2^5K_8$	(1.250)	1.145
4292.351	4	—	—	23290.72	+8	$as^2K^{\circ}_6 - 28c$	—	—	3949.438	150	III	6v	25312.92	+1	$as^2I^{\circ}_6 - 2^5H_6$	1.08	1.08
4290.99	5	—	—	23298.11	-10	$as^2I^{\circ}_6 - 14c$	—	—	3947.653	125	III	6v	25324.49	-4	$as^2I^{\circ}_6 - 2^5H_5$	(1.064)	1.125
4288.830	8	—	—	23300.85	-19	$as^2K^{\circ}_6 - 2^5K_8$	—	—	3935.823	125	III	6 γ a	25400.48	+21	$as^2I^{\circ}_6 - 18c$	(0.875)	0.990
4282.440	75	IV	6v	23344.63	+7	$as^2I^{\circ}_6 - 30c$	1.245	1.071	3927.454	80	III	6 γ a	25454.60	+10	$as^2I^{\circ}_6 - 26c$	0.859	1.066
4272.271	50	III	6v	23400.18	+4	$as^2I^{\circ}_6 - 22c$	1.177	0.991	3925.456	125	III	3a	25467.56	+6	$as^2I^{\circ}_6 - 2^5H_4$	0.603	0.903
4263.805	50	IV	6v	23446.65	-10	$as^2I^{\circ}_6 - 23r$	—	—	3920.524	30d	III	2a	25499.60	+10	$as^2I^{\circ}_6 - 16c$	(0.605)	0.985
4261.796	15	IV	6 α a	23457.71	0	$as^2I^{\circ}_6 - 2^5I_7$	—	—	3918.856	100	III	4a	25510.45	+8	$as^2I^{\circ}_7 - 2^5H_6$	(1.177)	1.094
4259.64	5d	—	—	23469.58	+22	$as^2L^{\circ}_7 - 2^5I_7$	—	—	3912.898	150	III	4a	25549.28	+15	$as^2I^{\circ}_6 - 26c$	(1.064)	1.097
4254.240	35	IV	6r	23498.38	-3	$as^2I^{\circ}_7 - 2^5I_7$	(1.143)	1.191	3906.093	100	IV	5 γ a	25581.10	+10	$as^2I^{\circ}_6 - 2^5H_7$	(1.250)	1.220
4249.484	20	V	6 α a	23525.67	+2	$as^2I^{\circ}_7 - 24c$	(1.177)	0.98	3899.555	10	—	—	25636.71	-2	$as^2I^{\circ}_6 - 27c$	—	—
4247.662	60	III	6v	23535.75	-9	$as^2I^{\circ}_6 - 11c$	0.878	1.156	3889.330	150	IV	5a	25704.11	+13	$as^2I^{\circ}_6 - 19c$	(0.875)	1.025
4243.528	20	III	6r	23558.69	-11	$as^2I^{\circ}_6 - 2^5H_5$	1.022	(1.073)	3885.190	100	IV	4a	25731.51	+14	$as^2I^{\circ}_6 - 27c$	1.058	1.068
4241.019	50	IV	6v	23572.62	0	$as^2I^{\circ}_6 - 2^5K_7$	(1.250)	1.108	3884.741	4	—	—	25734.48	-9	$as^2I^{\circ}_6 - 20c$	—	—
4236.210	20	III	6v	23599.38	-13	$as^2I^{\circ}_6 - 15r$	(1.064)	1.085	3884.039	5	—	—	25739.13	-7	$as^2I^{\circ}_6 - 28c$	—	—
4225.327	50	III	6r	23660.17	-1	$as^2I^{\circ}_6 - 2^5I_4$	0.604	0.685	3878.307	—	—	—	25777.17	-11	$as^2I^{\circ}_6 - 29c$	—	—
4222.98	125	III	6v	23673.31	-16	$as^2I^{\circ}_6 - 2^5K_6$	0.877	0.961	3868.578	15	—	—	25841.99	-22	$as^2I^{\circ}_6 - 18c$	—	—
4213.96	12	—	—	23723.99	+14	$as^2I^{\circ}_6 - 2^5H_4$	—	—	3868.125	4	—	—	25845.02	+10	$as^2I^{\circ}_6 - 21c$	—	—
4208.305	18	III	6 α a	23755.87	+2	$as^2I^{\circ}_6 - 16c$	(0.860)	0.980	3826.708	—	—	—	26124.74	+3	$as^2L^{\circ}_6 - 2^5H_7$	—	—
4206.739	50	III	6v	23767.71	-7	$as^2I^{\circ}_6 - 2^5I_8$	1.256	1.160	3823.571	10	—	—	26146.17	+20	$as^2I^{\circ}_6 - 19c$	—	—
4191.615	40	III	6v	23850.45	-4	$as^2I^{\circ}_6 - 16c$	(1.064)	0.985	3803.110	50d	—	—	26286.83	-3	$as^2I^{\circ}_6 - 21c$	—	—
4189.518	100	III	6v	23862.39	-13	$as^2I^{\circ}_7 - 2^5I_7$	1.175	1.121	3792.495	4	—	—	26360.82	-12	$as^2I^{\circ}_6 - 2^5K_7$	—	—
4180.68	8	—	—	23812.84	-13	$as^2I^{\circ}_6 - 2^5I_6$	—	—	3769.695	20d	V	2a	26519.84	+14	$as^2I^{\circ}_6 - 2^5H_6$	—	

TABLE III. Pr II terms.

TERM	WAVE NUMBER	h.f.s.	No. COMB.	g LS (THEOR.)	g (MEAS.)	No. PATTERNS RESOLVED	TERM	WAVE NUMBER	h.f.s.	No. COMB.	g LS (THEOR.)	g (MEAS.)	No. PATTERNS RESOLVED
$a^3I^o_4$	0.00	6i	15	0.600	0.605	12	z^3K_6	24115.41	—	10	0.905	0.959	5
$a^3I^o_5$	441.94	6n	23	0.900	0.875	12	z^3K_6	24393.70	—	6	0.857	0.992	1
$a^3I^o_6$	1649.01	6n	32	1.071	1.064	3	z^3I_5	24716.06	—	8	0.900	0.911	2
$a^3I^o_5$	1743.65	6i	30	0.833	0.860	16	12i	24754.93	—	4	—	0.911	2
$a^3I^o_7$	2998.31	6n	24	1.179	1.177	12	13i	24818.68	—	5	—	—	0
$a^3I^o_6$	3403.12	6i	26	1.024	1.037	3	z^3I_5	24835.05	—	4	1.024	1.009	2
$a^3L^o_6$	3893.38	—	14	0.714	0.721	5	14i	24947.22	—	4	—	—	0
$a^3K^o_5$	4097.57	—	16	0.667	0.683	3	15i	25248.52	—	5	—	1.07	0
$a^3I^o_8$	4437.09	6n	10	1.250	1.250	7	z^3H_4	25467.50	—	4	0.900	0.905	1
$a^3I^o_7$	5079.31	6i	20	1.143	1.143	1	16i	25499.50	—	7	—	0.984	3
$a^3L^o_7$	5108.36	—	16	0.911	0.926	1	z^3K_7	25569.10	—	10	1.054	1.062	1
$a^3K^o_6$	5226.47	—	12	0.905	0.915	1	17i	25610.06	—	9	—	1.116	2
$a^3K^o_7$	6413.79	—	14	1.054	1.042	0	z^3I_5	25656.62	—	9	1.071	1.042	2
$a^3L^o_8$	6417.75	—	6	1.042	1.06	0	18i	25842.21	—	6	—	0.999	2
$b^3I^o_5$	7446.51	—	9	0.600	0.622	1	19i	26145.97	—	4	—	1.025	2
$a^3K^o_8$	7659.60	—	7	1.153	1.161	0	20i	26176.51	—	4	—	—	0
$a^3L^o_9$	7805.50	—	4	1.133	1.115	0	21i	26286.86	—	9	—	0.89	0
$b^3I^o_9$	8489.94	—	7	0.900	0.93	0	22i	26398.45	—	9	—	0.990	3
$a^3K^o_9$	8958.35	—	5	1.222	1.219	1	23i	26445.06	—	8	—	1.08	0
$a^3L^o_{10}$	9254.94	—	1	1.200	1.19	0	24i	26523.96	—	4	—	1.02	1
$a^3H^o_{10}$	9378.57	—	8	1.033	1.044	0	25i	26640.82	—	6	—	1.075	1
$b^3I^o_6$	9646.60	—	11	1.071	1.073	1	z^3I_7	26860.83	—	10	1.179	1.123	3
$b^3I^o_5$	11005.45	—	7	1.179	1.153	1	z^3H_5	26961.92	—	9	1.167	1.073	2
$b^3I^o_8$	11610.92	—	6	1.250	1.231	2	z^3H_5	26973.54	—	5	1.100	1.101	1
1s	21676.28	—	6	—	1.001	2	z^3K_8	27127.88	—	9	1.153	1.143	1
2s	22039.98	—	7	—	0.997	4	26i	27198.15	—	7	—	1.067	1
3s	22571.38	—	7	—	1.163	4	27i	27380.38	—	6	—	1.06	0
z^3K_3	22675.40	—	6	0.667	0.821	4	28i	27388.21	—	8	—	0.886	1
4s	22718.31	—	7	—	1.10	0	29i	27425.29	—	6	—	—	0
5s	22885.56	—	7	—	0.994	2	30i	27781.65	—	8	—	1.075	4
6s	23141.37	—	5	—	1.062	1	z^3K_7	28009.71	—	8	1.018	1.106	3
7s	23261.30	—	4	—	0.896	3	z^3I_8	28201.87	—	8	1.250	1.154	1
8s	23616.67	—	5	—	1.068	1	z^3H_6	28508.68	—	6	1.214	1.103	1
z^3I_4	23660.18	—	5	0.600	0.684	2	z^3I_7	28577.72	—	7	1.143	1.187	3
9i	23898.25	—	6	—	1.107	2	z^3K_9	28816.04	—	4	1.222	1.217	0
10s	23970.40	—	4	—	1.14	0	z^3K_8	29723.83	—	7	—	1.145	1
11i	23977.78	—	10	—	1.156	5	z^3H_7	30018.09	—	7	1.286	1.215	0

That so many values of $o-c$ are less than 10 (i.e., less than 0.1 cm^{-1}) is gratifying, as it is difficult to make exact wave-length measurements on lines with partially resolved h.f.s. 132 lines are found to agree with the wave number calculated from the combination principle to within $\pm 0.05 \text{ cm}^{-1}$, and 78 more to within ± 0.10 , leaving only 108 with deviations greater than 0.10 cm^{-1} .

In the first column of Table III are listed the term assignments, followed by the wave number of each and the hyperfine structure deduced from all lines considered as arising from the term, where i indicates wider separations at lower wave numbers, and n at higher. The fourth column gives the number of combinations with the term in the present array. The fifth column contains the theoretical g value of a term of the designation given, calculated from Landé's formula for LS coupling. The next column contains the measured g value of the term, averaged from measurements on the number of resolved patterns given, or on unresolved patterns reduced from several plates. The observed g values are in most cases believed to be correct to within ± 0.005 unit.

LOW CONFIGURATIONS OF Pr II

In accordance with the Bohr-Stoner theory the normal praseodymium atom ($Z=59$) can be expected to have three electrons outside a closed shell. Pr II, with two such electrons, would have a simpler spectrum than Pr I, but in both cases two $6s$ electrons are present which are very loosely bound, and it is necessary in Pr II to take four electrons into account when predicting terms. Among the odd configurations to be expected are fds^2 , fd^2s , f^3s and f^3d . Even configurations which should be important are f^2ds , f^2s^2 , f^3p , and f^4 .

Many of the strongest lines of Pr II, which presumably arise from low levels, show wide h.f.s. patterns. These usually arise when a single s electron occurs in a configuration, making

TABLE IV. Expected important terms of Pr II.

$4f^36s$	$5,3(IGFDS)^o$, $3,1(LKIH_2G_2F_2D_2P)^o$
$4f^35d$	$5,3(LKIHG, IHGFD, HGFDP, GFDPS, D)^o$ $3,1(NML_2K_4I_6H_7G_9F_9D_7P_6S_3)^o$
$4f^36p$	$5,3(KIH, HGF, GFD, FDP, P)$ $3,1(ML_2K_3I_4H_5G_6F_6D_5P_3S)$

TABLE V. Principal supermultiplet of Pr II.*

	$a^5I^{\circ}_8$	$a^5I^{\circ}_7$	$a^5I^{\circ}_6$	$a^5I^{\circ}_5$	$a^5I^{\circ}_4$	$a^3I^{\circ}_7$	$a^3I^{\circ}_6$	$a^3I^{\circ}_5$
z^5K_9	200 III 6v	X	X	X	X	X	X	X
8	100 IV 6v	200 III 6v	X	X	X	150 IV 6r	X	X
7	—	100 III 6v	200 III 6v	X	X	30 — 6r	200 III 6r	X
6	X	100 V 6v	125 III 6v	125 III 6v	X	10 V 5 ² a	40 — 6r	125 III 6r
5	X	X	15 — 6v	200 I III 6v	125 III 6r	X	—	—
z^3K_8	150 III 6v	—	X	X	X	100 III 6r	X	X
7	50 IV 6v	100 III 6v	4 — —	X	X	100 IV 6r	150 III 6r	X
6	X	3 — —	—	—	X	60 V —	60 V 6r	90 III 6ra
z^5I_8	50 III 6v	100 III 6v	X	X	X	100 IV 6r	X	X
7	90 V 6v	100 III 6v	100 III 6v	X	X	5d — —	15 IV 6ra	X
6	X	50 IV 6v	200 III 6v	125 III 6v	X	—	25 V 6r	8 — —
5	X	X	150 IV 6v	250 III 6v	50 III 6ra	X	X	80 III 6r
4	X	X	X	150 III 6va	50 III 6r	X	X	6 — —
z^3I_7	150 III 6v	—	—	X	X	35 IV 6r	100 III 6r	X
6	X	40 V 6v	—	—	X	2 — —	100 V 6r	50 IV 6r?a
5	X	X	—	—	—	X	—	—
z^5H_7	100 IV 5 ² a	12 — —	—	X	X	150 III 6r	—	X
6	X	100 III 4a	—	—	X	—	125 III 6r	60 III 6?a
5	X	X	125 IV 6v	—	—	X	—	12 — —
4	X	X	X	300 III 6v	125 III 3a	X	X	X
3	X	X	X	X	—	X	X	—
z^3H_6	X	75 III 6v	150 III 6v	20d V 2a	X	30 — —	20 III 6r	60 III 6r
5	X	X	X	—	—	X	X	—
4	X	X	X	—	—	X	X	—

* X—indicates line forbidden by selection principle for ΔJ . — indicates line not found.

plausible the assumption that the lowest terms of Pr II arise from such a configuration. Three likely configurations are then fd^2s , f^2ds and f^3s . Resulting terms of the latter are given in Table IV. When Hund's theory and Meggers and Laporte's rule are used the lowest terms to be expected from these three configurations are $^5I^{\circ}_4$, 5K_5 , and $^5I^{\circ}_4$, respectively.

Our analysis indicates that the level $f^3(4I^{\circ}) \cdot s^5I^{\circ}_4$ is the lowest. As will be seen from Fig. 1, all eight levels of the low $^5I^{\circ}$ and $^3I^{\circ}$ terms arising from the addition of an s electron to the parent $4I^{\circ}$ term of Pr III have been found. Combinations of these terms with the even triads 5KIH and 3KIH which arise from the f^3p configuration give rise to the strongest lines of the spectrum. The *raies ultimes* of praseodymium are usually given¹¹ as 4179.422 and 4062.817A. These lines we have classified as $a^5I^{\circ}_6 - z^5K_7$ and $a^3I^{\circ}_6 - z^3K_7$. From the usual multiplet intensity rules one would expect the leading line of the supermultiplet to be strongest, with a prediction of $a^5I^{\circ}_8 - z^5K_9$ as the *raie ultime*. This line, at 4100.746, is of intensity 200, but other lines are stronger. The estimated intensities of the lines of the basic supermultiplet are given in Table V, where an irregularity in intensities is to be noted. A similar irregularity is found in the homologous super-

multiplet of Nd II.¹² The intensities need more careful study before the cause of this anomaly can be explained.

HYPERFINE STRUCTURE

The eight levels of the terms $f^3(4I^{\circ}) \cdot s^5I^{\circ}$ and $^3I^{\circ}$ have marked hyperfine structure which appears to be due mainly to the interaction of the $6s$ electron with the moment of the nucleus. It is of interest to apply to this case the equations given by Goudsmit and Bacher.¹³ The change in atomic energy due to the interactions between electronic and nuclear moments (h.f.s.) is given by

$$\Gamma_F = \frac{1}{2}A[F(F+1) - I(I+1) - J(J+1)], \quad (1)$$

where J represents the mechanical moment of the electrons, I that of the nucleus ($= 5/2$ for Pr), F the total moment, and A is a number independent of F .

If we assume that the hyperfine structure is due to the s electron and that there is JJ coupling between the latter and the ion, then, as shown by Goudsmit and Bacher,

$$A = a \frac{J(J+1) + s(s+1) - J'(J'+1)}{2J(J+1)}, \quad (2)$$

¹² McNally, Harrison, and Albertson, to be published shortly.

¹³ S. Goudsmit and R. F. Bacher, Phys. Rev. **34**, 1501 (1929).

¹¹ M. I. T. Wavelength Tables.

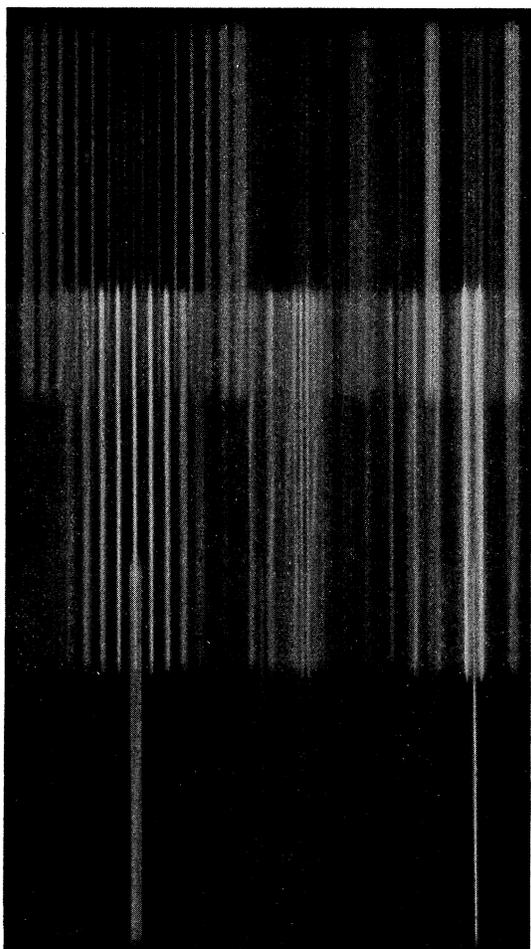


FIG. 2. Zeeman patterns of Pr II line showing hyperfine structure. Below, no-field pattern; middle, n components; above, p components. The strong line at the left is $\lambda 4368.327A$.

where s is the spin of the electron ($=\frac{1}{2}$), J' the moment of the ion, and a a constant which measures the strength of the interaction between the s electron and the nuclear moment.

There are two cases to be considered. For the levels ${}^5I_{5,6,7,8}^{\circ}$

$$J' = J - \frac{1}{2}$$

and (2) gives

$$A = a/2J. \quad (3)$$

For the levels ${}^5I_4^{\circ}$ and ${}^3I_{5,6,7}^{\circ}$

$$J' = J + \frac{1}{2}$$

and

$$A = -a/[2(J+1)]. \quad (4)$$

The difference in signs of (3) and (4) is in agree-

ment with the observed difference in the shading of the h.f.s. for lines involving the two groups of levels.

For $J > 1$, the total width of the hyperfine structure $\Delta\sigma$ is found from (1) to be given by

$$\Delta\sigma = \Gamma_{J+I} - \Gamma_{J-I} = AI(2J+1). \quad (5)$$

In attempting to make a quantitative comparison of theory and observation, one is confronted with the difficulty that there is a great deal of variation in the measured h.f.s. data¹ for lines going to a common level that is responsible for the observable h.f.s. One can take an average over these lines to get information about the level in question, but it is to be expected that some uncertainty will remain.

Table VI gives the $\Delta\sigma$ values for the eight levels as determined from White's h.f.s. data for those lines the classification of which we have confirmed with Zeeman measurements. A minus sign indicates that the h.f.s. is inverted. From the $\Delta\sigma$ in each case the corresponding value of a was calculated by means of (5) and (3) or (4). The various calculated values differ somewhat,

TABLE VI. Values of $\Delta\sigma$ from White's h.f.s. data.

LEVEL	$\Delta\sigma(\text{OBS.})$	$a(\text{CALC.})$	$\Delta\sigma(\text{CALC.})$
${}^5I_4^{\circ}$	-0.80	+0.36	-0.90
${}^3I_5^{\circ}$	-0.74	0.32	-0.92
${}^3I_6^{\circ}$	-0.84	0.36	-0.93
${}^3I_7^{\circ}$	-0.91	0.39	-0.94
${}^5I_5^{\circ}$	+1.13	0.41	+1.10
${}^5I_6^{\circ}$	+1.10	0.41	+1.08
${}^5I_7^{\circ}$	+1.05	0.39	+1.07
${}^5I_8^{\circ}$	+1.11	0.42	+1.06

TABLE VII. g sums for $f^3({}^4I^{\circ}) \cdot s$ levels of Pr II.

TERM	$g(\text{MEAS.})$	$g(\text{LS})$	$g(J_i)$	J_i	SUM $g(\text{MEAS.})$	SUM $g(\text{THEOR.})$
${}^6I_4^{\circ}$	0.605	0.600	0.600	$(4\frac{1}{2}, \frac{1}{2})$	0.605	0.600
${}^6I_5^{\circ}$	0.875	0.900	0.855	$(4\frac{1}{2}, \frac{1}{2})$	1.735	1.733
${}^6I_6^{\circ}$	0.860	0.833	0.878	$(5\frac{1}{2}, \frac{1}{2})$		
${}^6I_6^{\circ}$	1.064	1.071	1.051	$(5\frac{1}{2}, \frac{1}{2})$	2.101	2.095
${}^6I_6^{\circ}$	1.037	1.024	1.044	$(6\frac{1}{2}, \frac{1}{2})$		
${}^6I_7^{\circ}$	1.177	1.179	1.171	$(6\frac{1}{2}, \frac{1}{2})$	2.320	2.322
${}^6I_7^{\circ}$	1.143	1.143	1.151	$(7\frac{1}{2}, \frac{1}{2})$		
${}^6I_8^{\circ}$	1.250	1.250	1.250	$(7\frac{1}{2}, \frac{1}{2})$	1.250	1.250

TABLE VIII. *g* sums for even levels of Pr II.

	9		8		7		6		5		4	
	<i>g</i> _M	<i>g</i> _T										
2^5K_5									0.821	0.667		
2^5I_4											0.684	0.600
2^5K_6							0.959	0.905				
2^3K_6							0.992	0.857				
2^5I_5									0.911	0.900		
2^3I_6							1.009	1.024				
2^5H_4											0.905	0.900
2^5K_7					1.062	1.054						
2^5I_6							1.042	1.071				
2^5I_7					1.123	1.179						
2^3H_6							1.073	1.167				
2^5H_5									1.101	1.100		
2^5K_8			1.143	1.153								
2^3K_7					1.106	1.018						
2^5I_8			1.154	1.250								
2^5H_6							1.103	1.214				
2^3I_7					1.187	1.143						
2^5K_9	1.217	1.222										
2^5H_7					1.215	1.286						
Sums	1.217	1.222	2.297	2.403*	5.693	5.680	6.178	6.238	2.833	2.667*	1.589	1.500*

* Indicates that all terms of that particular *j* value derivable from parent term $^4I + p$ have not been found and *g* sum rule is not applicable. This assumes that 4I is perfect *LS* coupling.

partly because of errors in the values of $\Delta\sigma$ and partly, probably, because of incomplete *JJ* coupling, etc. By taking

$$a = 0.40 \text{ cm}^{-1},$$

the value of $\Delta\sigma$ was calculated in each case on the basis of the above equations. It is evident that the agreement is quite satisfactory.

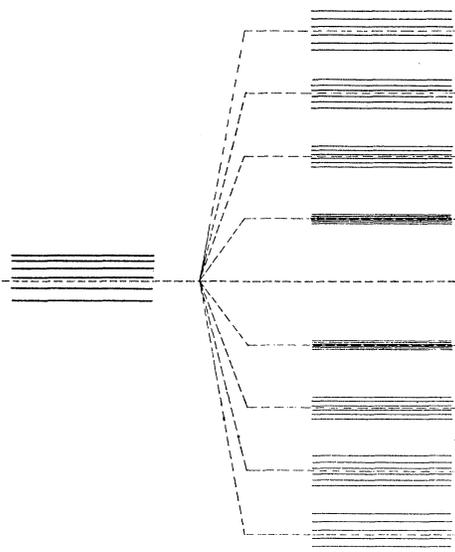


FIG. 3. Effect of a strong magnetic field in breaking a normal hyperfine-multiple level, $a^5I_4^o$, containing six levels, up into 54 levels.

In the presence of the magnetic field which is used to produce the Zeeman effects, the character of the h.f.s. changes completely. The magnetic field used in practice is a strong field from the standpoint of the h.f.s. and hence it destroys the coupling between *I* and *J*, so that the component of each in the direction of the field is quantized nearly independently (Back-Goudsmit effect).¹⁴

In this case the change in atomic energy due to the external magnetic field and the magnetic moment of the nucleus is given by

$$\Delta E = (-g_I M_I + gM) \frac{He}{4\pi mc^2} + A M_I M, \quad (6)$$

where *M* and *M_I* are the components along the field of *J* and *I*, respectively, *A* is the same constant as in (1), *g* is the usual *g* factor for the extra-nuclear electrons, and *g_I* is the nuclear *g* factor. Since *g_I* is usually small compared to *g*, one can take

$$\Delta E = gM \frac{He}{4\pi mc^2} + A M_I M. \quad (7)$$

The first term in the right-hand member corresponds to the ordinary Zeeman effect, the

¹⁴ E. Back and S. Goudsmit, *Zeits. f. Physik* **47**, 174 (1928).

second gives the h.f.s. In this case, the width of the h.f.s. is given by

$$\Delta\sigma = (\Delta E)_{M_I = +I} - (\Delta E)_{M_I = -I}. \quad (8)$$

Hence in a Zeeman pattern, the sharpest lines should be found near the center of the p and of each s branch, for which cases $|M|$ is a minimum. An example of this is to be seen in Fig. 2, which shows the no-field line 4368.327A and its Zeeman pattern. A diagram of the h.f.s. for the lower parent level of this line, a^5I_4 without and with a magnetic field, according to Eqs. (1) and (7), is given in Fig. 3, where the h.f.s. spacing has been exaggerated in comparison with the separations of the Zeeman components.

CALCULATION OF g SUMS

In Table VII are given the g sums for Pr II levels arising from the configuration $f^3(^4I^\circ) \cdot s$. It will be noted that the agreement between the measured and theoretical sums is usually within a few tenths of a percent, though perturbations much greater than this cause deviations of the individual terms. In Table VIII, which gives the g sums for even levels, the agreement is not quite so satisfactory.

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On the ${}^2\Pi_u \rightarrow {}^2\Pi_g$ Bands of CO_2^+ . Part I

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The band spectrum appearing in emission in the region $\lambda 2900$ – 4300 and believed to belong to the CO_2^+ or possibly CO_2 molecule has been studied and the excitation conditions found to be in substantial agreement with the former results of Duffendack and collaborators and of Smyth. The bands have been obtained with great intensity and photographs of most of them have been made in the second order of the 30-foot grating (actually obtained resolving power of 350,000). The rotational structure and the excitation conditions show that most of the bands belong to an extensive ${}^2\Pi \rightarrow {}^2\Pi$ system of bands of the molecule CO_2^+ . The molecule is linear in both states; the lower ${}^2\Pi$ appears to be the ground state ${}^2\Pi_g$ and the upper ${}^2\Pi$ is the first excited state ${}^2\Pi_u$ of this molecule predicted by Mulliken. The complete rotational and vibrational analysis of this band system is still in progress; in this

paper the analysis of 5 double bands of the $v''_1 = v''_2 = v''_3 = 0$ progression of the symmetrical vibration (v'_1 varying, $v'_2 = v'_3 = 0$) is presented. The results of the analysis are: $\nu_0^{(0,0)} = 28,532.60$ (${}^2\Pi_{3/2} \rightarrow {}^2\Pi_{3/2}$) and $\nu_0^{(0,0)} = 28,468.48$ (${}^2\Pi_1 \rightarrow {}^2\Pi_1$). The vibrational intervals (i.e., the distances between the origins of successive bands in the progression) are: $\Delta G'_1 = 1126.71$; $\Delta G'_2 = 1122.66$; $\Delta G'_3 = 1120.22$; $\Delta G'_4 = 1120.04$ (all ${}^2\Pi_{3/2u}$) and $\Delta G'_1 = 1125.97$; $\Delta G'_2 = 1120.79$; $\Delta G'_3 = 1116.09$; $\Delta G'_4 = 1111.76$ (all ${}^2\Pi_{1u}$). Further for ${}^2\Pi_{3/2g}B''_0 = 0.3796$; for ${}^2\Pi_{3/2g}B''_0 = 0.3812$; for ${}^2\Pi_{3/2u}B'_0 = 0.3485$; $B'_1 = 0.3475$; $B'_2 = 0.3465$; $B'_3 = 0.3457$; $B'_4 = 0.3453$; for ${}^2\Pi_{1u}B'_0 = 0.3501$; $B'_1 = 0.3492$; $B'_2 = 0.3483$; $B'_3 = 0.3475$; $B'_4 = 0.3466$. The Λ -doubling is observable only in the ${}^2\Pi_1 \rightarrow {}^2\Pi_1$ sub-bands in this progression; in ${}^2\Pi_u$ it is bigger than in ${}^2\Pi_g$ ($p''_0 = 0.004$ for $v''_1 = 0$) and increases fast with the vibrational energy.

INTRODUCTION

WHEN a gas through which a discharge is passing contains carbon dioxide (or CO), in addition to the bands belonging to CO and CO^+ a great number of bands appear in the spectrum in the region $\lambda 2800$ – 4500 : the most

prominent of these is the double band at $\lambda 2883$ – 2896 , which appears whenever the smallest traces of CO_2 (or CO) are present. All these bands have been observed by many investigators (some of them as far back as 1802), but there was considerable uncertainty about their emitter,

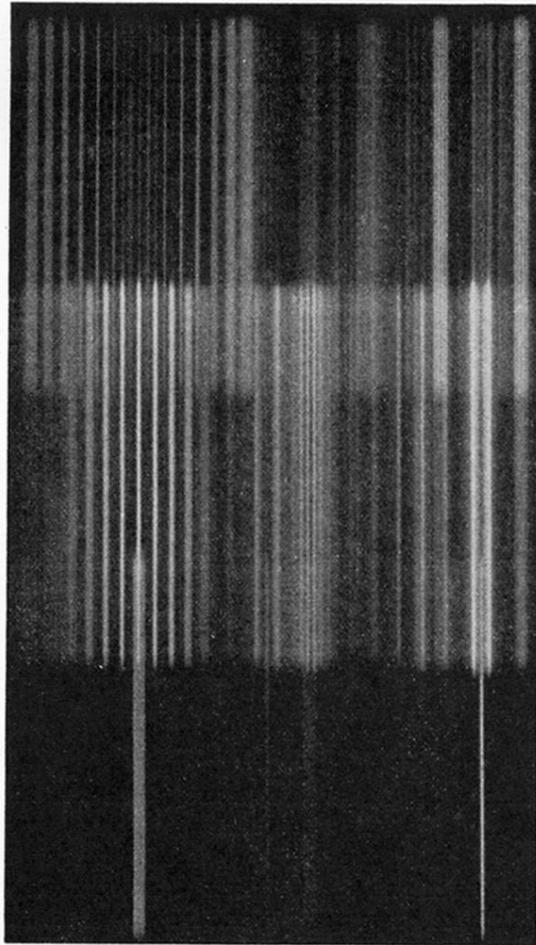


FIG. 2. Zeeman patterns of Pr II line showing hyperfine structure. Below, no-field pattern; middle, n components; above, p components. The strong line at the left is $\lambda 4368.327\text{\AA}$.