

The Spectra of Potassium, K IV and K V, and of Calcium Ca V and Ca VI

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From 19 to 36 new lines each have been identified in K IV, K V, Ca V and Ca VI. In K IV and Ca V these include singlet and intercombination lines. The presence in the nebulae of lines due to the forbidden $^3P-^1D$ transitions in S I, Cl II, A III, K IV and Ca V is discussed.

EKEFORS¹ has measured the spectra of potassium and calcium in the range below 1050A with a source capable of exciting all stages of ionization up to the ninth or tenth.

¹ E. Ekefors, *Zeits. f. Physik* **71**, 53 (1931).

In both K IV and Ca V, Ram² classified 30 to 40 lines each as triplet transitions and in K V noted one quartet multiplet of three lines.

In the course of the present investigation

² M. Ram, *Ind. J. Phys.* **8**, 163 (1933).

TABLE I. *Classified lines of K IV, K V, Ca V and Ca VI.*

Int.	λ	ν	Classification	Int.	λ	ν	Classification
<i>Classified lines of K IV</i>				<i>Classified lines of K V</i>			
3	271.820	367891	$s^2p^4\ ^3P_2-5s\ ^3S$	6	294.836	339172	$s^2p^3\ ^4S-4s\ ^4P_{3/2}$
2	273.065	366213	$s^2p^4\ ^3P_1-5s\ ^3S$	6	296.169	337645	$s^2p^3\ ^4S-4s\ ^4P_{1/2}$
1	273.546	365569	$s^2p^4\ ^3P_0-5s\ ^3S$	5	297.064	336628	$s^2p^3\ ^4S-4s\ ^4P_{1/2}$
2	354.139	282375	$s^2p^4\ ^3P_2-4s\ ^1D$	7	300.252	333054	$s^2p^3\ ^3D_{3/2}-4s\ ^2D_{3/2}$
6	354.927	281748	$s^2p^4\ ^1D-4s\ ^1P$	6	300.503	332775	$s^2p^3\ ^3D_{3/2}-4s\ ^2D_{3/2}$
3	356.260	280694	$s^2p^4\ ^3P_1-4s\ ^1D$	6	311.243	321292	$s^2p^3\ ^3D_{3/2}-4s\ ^2P_{1/2}$
2	356.568	277340	$s^2p^4\ ^1D-4s\ ^3P_2$	5	312.770	319724	$s^2p^3\ ^3D_{3/2}-4s\ ^2P_{1/2}$
6	375.955	265989	$s^2p^4\ ^1D-4s\ ^1D$	4	315.181	317278	$s^2p^3\ ^2P_{1/2}-4s\ ^2D_{3/2}$
2	379.279	263658	$s^2p^4\ ^3P_2-3d\ ^3D_1$	3	315.537	316920	$s^2p^3\ ^2P_{1/2}-4s\ ^2D_{3/2}$
5	380.477	262828	$s^2p^4\ ^3P_2-3d\ ^3D_2$	2	327.031	305781	$s^2p^3\ ^3P_{1/2}-4s\ ^2P_{1/2}$
4	381.702	261985	$s^2p^4\ ^3P_1-3d\ ^3D_1$	4	327.376	305459	$s^2p^3\ ^2P_{1/2}-4s\ ^2P_{1/2}$
3	382.487	261447	$s^2p^4\ ^1D-4s\ ^3D_2$	2	328.973	303976	$s^2p^3\ ^3P_{1/2}-4s\ ^2P_{1/2}$
			$s^2p^4\ ^3P_2-3d\ ^1P$	0	329.307	303668	$s^2p^3\ ^2P_{1/2}-4s\ ^2P_{1/2}$
4	382.646	261338	$s^2p^4\ ^3P_0-3d\ ^3D_1$	5	422.178	236867	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
6	382.906	261161	$s^2p^4\ ^3P_1-3d\ ^3D_2$	5	425.159	235206	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
3	384.956	259770	$s^2p^4\ ^3P_1-3d\ ^1P$	7	425.588	234969	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
3	404.412	247273	$s^2p^4\ ^1D-3d\ ^3D_1$	2	452.227	221128	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
2	405.773	246443	$s^2p^4\ ^1D-3d\ ^3D_2$	3	452.900	220799	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
5	408.076	245052	$s^2p^4\ ^1D-3d\ ^1P$	1	455.670	219457	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
3	417.280	239647	$s^2p^4\ ^1D-3d\ ^3P_2$	4	456.328	219141	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
				7	580.319	172319.0	$s^2p^3\ ^2D_{3/2}-sp^4\ ^2P_{1/2}$
				5	585.510	170791.3	$s^2p^3\ ^2D_{3/2}-sp^4\ ^2P_{1/2}$
				8	586.322	170554.8	$s^2p^3\ ^2D_{3/2}-sp^4\ ^2P_{1/2}$
				5	638.668	156575.9	$s^2p^3\ ^3P_{1/2}-sp^4\ ^2P_{1/2}$
				2	639.982	156254.4	$s^2p^3\ ^3P_{1/2}-sp^4\ ^2P_{1/2}$
				10	644.963	155047.7	$s^2p^3\ ^3P_{1/2}-sp^4\ ^2P_{1/2}$
<i>Classified lines of Ca V</i>				<i>Classified lines of Ca VI</i>			
3	184.280	542652	$s^2p^4\ ^3P_2-5s\ ^3P_2$	7	228.628	437392	$s^2p^3\ ^4S-4s\ ^4P_{3/2}$
1	184.415	542255	$s^2p^4\ ^3P_2-5s\ ^3P_1$	7	229.734	435286	$s^2p^3\ ^4S-4s\ ^4P_{1/2}$
2	185.102	540243	$s^2p^4\ ^3P_1-5s\ ^3P_2$	5	230.495	433849	$s^2p^3\ ^4S-4s\ ^4P_{1/2}$
1	185.288	539700	$s^2p^4\ ^3P_1-5s\ ^3P_1$	6	232.282	430511	$s^2p^3\ ^3D_{3/2}-4s\ ^2D_{3/2}$
2	185.540	538967	$s^2p^4\ ^3P_0-5s\ ^3P_1$	5	232.531	430050	$s^2p^3\ ^3D_{3/2}-4s\ ^2D_{3/2}$
4	190.363	525312	$s^2p^4\ ^1D-5s\ ^1P$	0	239.296	417892	$s^2p^3\ ^3D_{3/2}-4s\ ^2P_{1/2}$
5	190.457	525053	$s^2p^4\ ^3P_2-5s\ ^3D_3$	7	239.535	417476	$s^2p^3\ ^3D_{3/2}-4s\ ^2P_{1/2}$
3	190.558	524775	$s^2p^4\ ^3P_2-5s\ ^3D_2$	6	240.721	415419	$s^2p^3\ ^3D_{3/2}-4s\ ^2P_{1/2}$
3	191.439	522360	$s^2p^4\ ^3P_1-5s\ ^3D_2$	3	242.265	412771	$s^2p^3\ ^3P_{1/2}-4s\ ^2D_{3/2}$
2	191.480	522248	$s^2p^4\ ^3P_1-5s\ ^3D_1$	5	242.592	412215	$s^2p^3\ ^3P_{1/2}-4s\ ^2D_{3/2}$
2	191.801	521374	$s^2p^4\ ^3P_0-5s\ ^3D_1$	3	242.631	412148	$s^2p^3\ ^3P_{1/2}-4s\ ^2D_{3/2}$
5	196.970	507692	$s^2p^4\ ^1D-5s\ ^1D$	4	249.914	400138	$s^2p^3\ ^3P_{1/2}-4s\ ^2P_{1/2}$
2	197.531	506250	$s^2p^4\ ^1D-5s\ ^3D_3$	3	250.265	399576	$s^2p^3\ ^3P_{1/2}-4s\ ^2P_{1/2}$
2	197.648	505950	$s^2p^4\ ^1D-5s\ ^3D_2$	1	251.465	397670	$s^2p^3\ ^3P_{1/2}-4s\ ^2P_{1/2}$
6	199.553	501120	$s^2p^4\ ^3P_2-5s\ ^3S$	4	251.816	397115	$s^2p^3\ ^3P_{1/2}-4s\ ^2P_{1/2}$
3	199.890	500275	$s^2p^4\ ^1S-5s\ ^1P$	6	339.463	294583	$s^2p^3\ ^3D_{3/2}-3d\ ^2D_{3/2}$
5	200.512	498723	$s^2p^4\ ^3P_2-5s\ ^3S$	4	339.940	294170	$s^2p^3\ ^3D_{3/2}-3d\ ^2D_{3/2}$
3	200.860	497859	$s^2p^4\ ^3P_1-5s\ ^3S$	8	340.037	294086	$s^2p^3\ ^3D_{3/2}-3d\ ^2D_{3/2}$
3	200.860	497859	$s^2p^4\ ^3P_0-5s\ ^3S$	4	340.528	293662	$s^2p^3\ ^3D_{3/2}-3d\ ^2D_{3/2}$
3	266.863	374724	$s^2p^4\ ^3P_2-4s\ ^1D$	2	361.234	276829	$s^2p^3\ ^2P_{1/2}-3d\ ^2D_{3/2}$
8	267.772	373452	$s^2p^4\ ^1D-4s\ ^1P$	4	362.612	275777	$s^2p^3\ ^2P_{1/2}-3d\ ^2D_{3/2}$
2	268.583	372328	$s^2p^4\ ^3P_1-4s\ ^1D$	5	370.022	270254	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
2	270.570	369590	$s^2p^4\ ^3P_2-4s\ ^3D_1$	7	373.418	267796	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
4	271.141	368812	$s^2p^4\ ^1D-4s\ ^3P_2$	2	373.997	267382	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
1	271.440	368406	$s^2p^4\ ^1D-4s\ ^3P_1$	7	396.055	252490	$s^2p^3\ ^2D_{3/2}-3d\ ^2P_{1/2}$
8	280.992	355882	$s^2p^4\ ^1D-4s\ ^1D$	2	396.917	251942	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
2	284.794	351131	$s^2p^4\ ^1D-4s\ ^3D_3$	0	399.925	250047	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
9dB	286.965	348475	$s^2p^4\ ^1S-4s\ ^1P$	2	400.824	249486	$s^2p^3\ ^2P_{1/2}-3d\ ^2P_{1/2}$
0	301.139	332073	$s^2p^4\ ^1D-4s\ ^3S$	2	629.594	158833	$s^2p^3\ ^4S-sp^4\ ^4P_{1/2}$
6	321.609	310937	$s^2p^4\ ^3P_2-3d\ ^3D_1$	2	633.815	157775	$s^2p^3\ ^4S-sp^4\ ^4P_{3/2}$
6	330.937	302172	$s^2p^4\ ^3P_2-3d\ ^1P$	2	641.883	155792	$s^2p^3\ ^4S-sp^4\ ^4P_{3/2}$
4	333.570	299787	$s^2p^4\ ^3P_1-3d\ ^1P$				
6	334.545	298913	$s^2p^4\ ^3P_0-3d\ ^1P$				
4	343.640	291002	$s^2p^4\ ^1D-3d\ ^3D_2$				
9	352.915	283354	$s^2p^4\ ^1D-3d\ ^1P$				
5	356.246	280705	$s^2p^4\ ^1D-3d\ ^3P_1$				
5	387.077	258347	$s^2p^4\ ^1S-3d\ ^1P$				

B, blend.

TABLE II. Term values of K IV, K V, Ca V and Ca VI.

Term values of K IV			Term values of Ca V			
$\Sigma l = 3$	4	5	$\Sigma l = 3$	4	5	
	$s^2p^4\ ^3P_2$	0	$s^2p^4\ ^3P_2$	350914	$s^2p^5\ ^3P_2$	154664
$4s\ ^3D_1$	$s^2p^4\ ^3P_1$	1673	$4s\ ^3D_1$	369590	$s^2p^5\ ^3P_1$	156756
$4s\ ^3D_2$	$s^2p^4\ ^3P_0$	2324	$4s\ ^3D_2$	369696	$s^2p^5\ ^3P_0$	157897
$4s\ ^3D_3$	$s^2p^4\ ^1D$	16386	$4s\ ^3D_3$	369959	$s^2p^4\ ^1D$	18831
$4s\ ^1D$			$4s\ ^1D$	374728	$s^2p^4\ ^1S$	43847
$4s\ ^3P_0$			$4s\ ^3P_0$	387039		
$4s\ ^3P_1$			$4s\ ^3P_1$	387226	$3d\ ^3P_2$	298204
$4s\ ^3P_2$			$4s\ ^3P_2$	387652	$3d\ ^3P_1$	299535
$4s\ ^1P$			$4s\ ^1P$	392283	$3d\ ^1P$	302184
					$3d\ ^3D_2$	309834
					$3d\ ^3D_1$	310945
$5s\ ^3S$			$5s\ ^3S$	501127		
			$5s\ ^3D_1$	524651		
			$5s\ ^3D_2$	524770		
			$5s\ ^3D_3$	525053		
			$5s\ ^1D$	526523		
			$5s\ ^3P_1$	542249		
			$5s\ ^3P_2$	542650		
			$5s\ ^1P$	544143		

Term values of K V			Term values of Ca VI			
$\Sigma l = 2$	3	4	$\Sigma l = 2$	3	4	
	$s^2p^3\ ^4S$	0	$4s\ ^4P_{3/2}$	433849	$s^2p^3\ ^4S$	0
$4s\ ^4P_{3/2}$	$s^2p^3\ ^2D_{3/2}$	24000	$4s\ ^4P_{1/2}$	435286	$s^2p^3\ ^2D_{1/2}$	27000
$4s\ ^4P_{1/2}$	$s^2p^3\ ^2D_{5/2}$	24237	$4s\ ^4P_{3/2}$	437392	$s^2p^3\ ^2D_{3/2}$	27417
$4s\ ^4P_{3/2}$	$s^2p^3\ ^2P_{3/2}$	39745	$4s\ ^4P_{5/2}$	442423	$s^2p^3\ ^2P_{3/2}$	44754
$4s\ ^2P_{3/2}$	$s^2p^3\ ^2P_{1/2}$	40064	$4s\ ^2P_{3/2}$	442423	$s^2p^3\ ^2P_{1/2}$	45310
$4s\ ^2P_{1/2}$			$4s\ ^2P_{1/2}$	444890		
$4s\ ^2D_{3/2}$			$4s\ ^2D_{3/2}$	457458	$3d\ ^2P_{1/2}$	294798
$4s\ ^2D_{1/2}$			$4s\ ^2D_{1/2}$	457525	$3d\ ^2P_{3/2}$	297250
					$3d\ ^2D_{3/2}$	321084
					$3d\ ^2D_{1/2}$	321584

plates of these spectra were taken in the region below 600A on a grazing incidence spectrograph of two meter focus. With the additional lines found on these long exposure plates to supplement the observations of Ekefors, it has been possible to make the further identifications of lines of K IV, K V, Ca V and Ca VI listed in Table I. All terms of these ions that have been fixed by either Ram's or the present identifications are given in Table II.

Since a large part of the K IV and Ca V lines, listed in the present paper, are singlet and intercombination lines it is now possible to fix accurately the relative positions of the 1D and 3P terms of the ground configuration of these ions. Ruedy³ has likewise determined the relative positions of these terms in S I. Furthermore by interpolation between S I and K IV and Ca V the relative positions of these terms in Cl II and A III can now be predicted with an uncertainty of at most 200 cm^{-1} . The predicted wave-lengths of the forbidden transitions from the metastable $s^2p^4\ ^1D$ state to the $s^2p^4\ ^3P$ states corresponding to these relative positions are given in Table III. The only one of these wave-lengths that corre-

TABLE III. Forbidden transitions from metastable states.

	$^3P_2 - ^1D$ λ	$^3P_1 - ^1D$ λ	$^3P_2 - ^3P_1$ $\Delta\nu$ (cm^{-1})
S I	10820.	11306. \pm 10	397.4 \pm 5
Cl II	8589.	9132. \pm 150	692. \pm 5
A III	7141.	7761. \pm 100	1113. \pm 5
K IV	6101.	6795. \pm 5	1673. \pm 8
Ca V	5309.	6086. \pm 5	2404. \pm 10

sponds to an observed nebular line is $^3P_2 - ^1D$ of A III which falls near the strong unidentified line at 7135.6A. The decision as to the reality of this correlation will depend largely on whether the companion $^3P_1 - ^1D$ line, which should appear at 7755A, can be found. *Nova Pictoris*⁴ shows unidentified lines at the positions of both the Ca V forbidden lines. However the ratio between the intensities of the two lines is neither constant nor at any time equal to the predicted ratio. Furthermore the observed lines persist with increasing intensity after other forbidden lines have disappeared. Since the $^3P_2 - ^1D$ and $^3P_1 - ^1D$ are, respectively, the strongest and second strongest forbidden lines that these ions emit, the failure of these lines to be observed points definitely to the low abundance of the ions involved.

⁴ H. Spencer Jones, M. N. R. A. S. 91, 777 (1931), 92 728 (1932).

³ J. E. Ruedy, Phys. Rev. 44, 757 (1933).