Such a ground-to-ground β transition of ⁶⁸Cu to ⁶⁸Zn would imply that the most probable spin of 30-sec ⁶⁸Cu is 1. Bakhru and Mukherjee³ tenuously assigned a spin of 2 to 30-sec ^{68g}Cu since they did not observe a 4.6-MeV ground-to-ground β -rav transition.

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Mean Lifetimes and Branching Ratios of Low-Lying Levels in ³³S⁺

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The mean lifetimes of the levels of ³³S below 3.3-MeV excitation energy have been measured by the Doppler shift attenuation method. The ${}^{30}Si(\alpha, n){}^{33}S$ reaction was used to populate these states. SiO₂ targets (enriched to 95% in 30 Si) evaporated onto Ni backings were bombarded with α -particles ranging in energy 5.5–9.0 MeV. γ -ray spectra were recorded with a 20-cc Ge(Li) detector at 0°, 90°, and 510° to the beam. The following mean lifetimes were found: $\tau(0.842$ -MeV level) = 1.66 ± 0.34 psec, $\tau(1.968) = 182 \pm 22$ fsec, $\tau(2.313) = 183 \pm 25$ fsec, $\tau(2.869) < 15$ fsec, $\tau(2.937) > 4$ psec, $\tau(2.970) = 82 \pm 12$ fsec, $\tau(3.221) < 65$ fsec. The branching ratios of the 2.970-MeV level were determined to be (90 ± 5) % to the ground state and $(10\pm5)\%$ to the second excited state.

I. INTRODUCTION

THE low-lying levels of ³³S have been the subject of a I number of investigations; the available information up to 1967 is summarized by Endt and Van der Leun.¹ The energy levels have been located primarily by the ${}^{32}S(d, p\gamma){}^{33}S$ and ${}^{32}S(n, \gamma){}^{33}S$ reactions.²⁻⁴ However, relatively little is known about the absolute strengths of the γ -ray transitions. Recently, certain collective aspects of ³³S have been studied via the ³¹P(³He, p)³³S reaction,⁵ and Dubois⁶ has used the ${}^{34}S({}^{3}He, \alpha){}^{33}S$ reaction to study the lowest $T = \frac{3}{2}$ states as well as some of the lowlying $T = \frac{1}{2}$ states.

Theoretical calculations have been carried out by Bishop,⁷ Glaudemans et al.,⁸ and Glaudemans, Wilden-

- ⁶ J. Dubois, Nucl. Phys. A117, 533 (1968)
- ⁷G. R. Bishop, Nucl. Phys. **14**, 376 (1959). ⁸P. W. M. Glaudemans, G. Wiechers, and P. J. Brussaard, Nucl. Phys. **56**, 529 (1964); **56**, 548 (1964).

thal, and McGrory.⁹ The calculations for the $(2s_{1/2}, 1d_{3/2})$ shell performed by Glaudemans, Wildenthal, and McGrory⁹ using a surface δ interaction give good agreement with experiment. Wiechers and Brussaard¹⁰ have calculated the M1-transition probability for the $0.842 \rightarrow 0$ transition in ³³S using the wave functions of Glaudemans et al.8

The present paper¹¹ describes a study of the low-lying levels of ³³S up to an excitation energy of 3.3 MeV. The mean lifetimes of the first seven excited states of ³³S were measured using the Doppler-shift attenuation method (DSAM). States of interest were populated by the ${}^{30}\text{Si}(\alpha, n){}^{33}\text{S}$ reaction. The γ -ray decay modes of the 2.97-MeV state were determined, and limits on the branching of other states were obtained.

II. PROCEDURE

A. General

The Triangle Universities Nuclear Laboratory FN tandem Van de Graaff accelerator was used to accelerate

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¹ Present address: University of Colorado, Boulder, Colo.
¹ P. M. Endt and C. Van der Leun, Nucl. Phys. A105, 1 (1967).
² J. A. Becker, L. F. Chase, Jr., D. B. Fossan, and R. E. Mc-Donald, Phys. Rev. 146, 761 (1966).
³ J. M. O'Dell, R. W. Krone, and F. W. Prosser, Jr., Nucl. Phys. 82, 574 (1966).

G. Van Middelkoop and H. Gruppelaar, Nucl. Phys. 80, 321

^{(1966).}

⁵ R. S. Cox, R. W. West, and R. J. Ascuitto, Phys. Rev. 175, 1419 (1968).

⁹ P. W. M. Glaudemans, B. H. Wildenthal, and J. B. McGrory, Phys. Letters 21, 427 (1966). ¹⁰ G. Wiechers and P. J. Brussaard, Nucl. Phys. **73**, 604 (1965).

¹¹ A preliminary version of this work was presented at a recent American Physical Society meeting: N. R. Roberson, C. E. Ragan, III, C. E. Moss, R. V. Poore, G. P. Lamaze, G. E. Mitchell, and D. R. Tilley, Bull. Am. Phys. Soc. **14**, 629 (1969).





FIG. 1. Schematic drawing of the scattering chamber showing the beam path to the target, the collimating system, and the lead used to shield the 20-cc Ge(Li) detector.

 α particles to energies ranging from 5.5–9.0 MeV. In Fig. 1a schematic drawing of the experimental setup is shown. A series of three Ta collimators served to define the beam. Each collimator was followed by a coaxial lead tube for shielding purposes. The stainless-steel beam pipe immediately before the collimators was lined with Ta sheet.

The target chamber (a glass cylinder approximately 13.5 cm high \times 3.8 cm diam) was connected to the beam pipe by a Ta-lined glass tube. The chamber was also lined with Ta which served as a Faraday cup. A vacuum of 5×10^{-7} Torr was maintained in the target chamber by a Vac-Ion pump connected to the top of the chamber.

The γ rays were detected with a 20-cc Ge(Li) detector placed ~8 cm from the target. The resolution of the detector was ~4.5 keV [full width half-maximum (FWHM)] for 1.33-MeV γ rays. Spectra were recorded with a 2048-channel analog-to-digital converter in conjuction with a DDP-224 on-line computer. After a spectrum was recorded at each angle it was stored on magnetic tape for later analysis.

B. Experimental Procedure

Targets were prepared by evaporating SiO₂ (enriched to 95% in ³⁰Si) from a Ta boat onto 1.5- μ Ni foils. The thickness of the SiO₂ targets ranged 15–150 μ g/cm². The target thickness was estimated by noting the thin film interference pattern formed on a glass slide during evaporation. Another estimate of the target thickness was obtained by measuring the energy lost by α particles in the SiO₂ target. This value was converted into a target thickness by use of the energy-loss data of Whaling.¹² (The energy loss was determined by measuring the energy difference between α particles which had been elastically scattered at 135° from the Ni foil and those which had been scattered from the Ni foil after passing through the target.) The results of the two methods of measurement agree to within 10%.

The particular form of the DSAM used follows closely that described by Warburton, Olness, and Poletti.¹³ The heavy-ion recoils were produced by using the endothermic reaction ${}^{30}\text{Si}(\alpha, n){}^{33}\text{S}$ with a Q value of -3.504 MeV. These recoils were confined to a narrow cone about the beam axis by using an α -particle energy slightly above the threshold energy. Beam energies 0.5-1.0 MeV above threshold produced half-angles of the heavy-ion recoils of the order of 10°. The velocities of the ${}^{33}\text{S}$ recoils, v/v_0 , ranged from 0.88-1.11, where $v_0 = c/137$.

For most of the transitions studied, data were acquired at several different bombarding energies and with several different target thicknesses. The γ rays were detected at 0°, 90°, and 150° to the beam axis in order to obtain several values of the experimental Doppler shift ΔE_{γ} and thus several values of $F(\tau)$ for each experimental condition. (Data were accumulated at each angle for several hours at a time, and then the runs at each angle were combined.) The energy calibration of the detector was based on the γ rays from ⁵⁶Co.

In most of the spectra small baseline or gain shifts were noted. These shifts were monitored by recording γ rays of known energy from radioactive sources at the same time as γ rays from the reaction. In most cases, the energy of the "source" γ ray was chosen to be very close to the energy of the "reaction" γ rays; thus, any observed shifts could be corrected by treating them as baseline shifts. The source γ rays were chosen to be lower in energy than the reaction γ rays in order to minimize background contributions to the γ rays of interest. The primary sources used were ¹³⁷Cs, ⁶⁰Co, and ⁸⁸Y. The 2.937-MeV γ ray from the fifth excited state of ³³S, which was determined to be a long-lived state in many preliminary runs, was also used to monitor the stability of the system.

¹² W. Whaling, in *Handbuch der Physik*, edited by S. Flügge (Springer-Verlag, Berlin, 1958), Vol. 34, p. 193.

¹³ E. K. Warburton, J. W. Olness, and A. R. Poletti, Phys. Rev. **160**, 938 (1967).



FIG. 2. Theoretical curves of $F(\tau)$ versus τ in seconds for two different target thicknesses. The upper curve is for a 150- $\mu g/cm^2 \operatorname{SiO}_2$ target evaporated onto a Ni backing, and the lower curve is for a 15- $\mu g/cm^2 \operatorname{SiO}_2$ target on a Ni backing. The two experimental values of $F(\tau)$ for the 2.97-MeV state are shown, and the lifetimes are seen to agree rather well. α_B and α_T are the characteristic slowing down times for ³³S ions in Ni and SiO₂, respectively.

In order to determine the branching ratios, spectra were recorded with the detector at 60° using a bombarding energy of 8.75 MeV and a target thickness of 150 μ g/cm². The efficiency of the Ge(Li) detector was determined by using the known energies and intensities¹⁴ of the γ rays from a ⁵⁶Co source placed ~8 cm from the detector.

C. DSAM

The DSAM has become a widely used method of measuring lifetimes of nuclear levels. The details of the method outlined here follow closely those given by Warburton, Olness, and Poletti.¹³ The experimental Doppler-shift attenuation factor $F_{exp}(\tau)$ is given by

$$F_{\exp}(\tau) = \Delta E_{\gamma} / E_{\gamma 0} \beta(0) \left(\cos\theta_1 - \cos\theta_2 \right), \qquad (1)$$

where ΔE_{γ} is the experimentally observed shift between the two detector angles θ_1 and θ_2 , $E_{\gamma 0}$ is the unshifted γ -ray energy, and $\beta(0) = v(0)/c$, where $v(0) = v_z(0)$ is the initial recoil velocity of the excited nucleus. The initial recoil velocity of ³³S along the beam direction $v_z(0)$ can be calculated from the center-of-mass velocity $v_{\text{e.m.}}$. Since the outgoing neutrons in the ³⁰Si(α , n)³³S reaction have some ditribution in the c.m. system, the value of $v_z(0)$ has some upper and lower limit. These limits can be computed from

$$\langle v_{\mathbf{z}}(0) \rangle = v_{\mathbf{c.m.}} [1 + \gamma^{-1} \langle \cos \theta_{\mathbf{c.m.}} \rangle],$$
 (2)

where $\theta_{0.m.}$ is the c.m. angle of the outgoing ³³S nucleus, and γ^{-1} is the ratio of the speed of the outgoing ³³S nucleus in the c.m. to the speed of the c.m. in the laboratory system. Following Warburton *et al.*,¹³ $\langle cos \theta_{o.m.} \rangle$ was set equal to 0±0.33. Thus, Eq. (2) becomes

$$\langle v_z(0) \rangle \simeq v_{c.m.} (1 \pm 0.33 \gamma^{-1}).$$
 (3)

For an ensemble of N nuclei with lifetime τ recoiling into a backing there are a certain number dN which decay in time dt, i.e.,

$$dN/dt = -N/t.$$
 (4)

If the velocity as a function of time is known, then this expression can be converted to an expression for dN(V)/dV, where $V \equiv v_z(t)/v(0)$. The general theory for the stopping of heavy ions in different backings has been given by Lindhard, Scharff, and Schiøtt (LSS).¹⁵ Their expression for the energy loss can be approximated¹³ by

$$-M_1(dv_z/dt) = K_e(v_z/v_0) + K_n(v_z/v_0)^{-1}, \quad (5)$$

where M_1 is the mass of the recoiling nucleus and K_o and K_n are constants determined from the LSS value and experimental data. Using Eq. (5) the expression for dN(V)/dV becomes

$$dN(V)/dV = \left[N_0/(\gamma_i^{-2} + 1)^{x/2} \times \left[xV(\gamma_i^{-2} + V^2)^{x/2-1} + \gamma_i^{-x}\delta(V) \right], \quad (6)$$

where $x = \alpha/\tau$, $\alpha = M_1 v_0/K_e \rho$, $\gamma_i^2 = (K_e/K_n) [v(0)/v_0]^2$, and ρ is the density of the backing.

TABLE I. γ -ray branching ratios for the low-lying levels of ³³S.

	Ei	Er	Branching ratios	
Level	(MeV)	(MeV)	Ref. 1	Present
2	1.968	0	100	100
		0.842	<0.5	<1.5
3	2.313	0	35	34 ± 5
		0.842	65	66 ± 5
		1.968	<3	<6
4	2.869	0	100	100
		0.842	<3	<3
		1.968	<3	<3
		2.313	•••	<4
5	2.937	0	50	•••
		0.842	<2	<2
		1.968	50	•••
		2.313	•••	<2
6	2.970	0	•••	90 ± 5
		0.842	•••	<3
		1.968	•••	10 ± 5
		2.313	•••	<2
7	3.221	0	40	38 ± 5
		0.842	60	62 ± 5
		1.968	•••	<2
		2.313	•••	<5
		2.869	•••	<7
		2.937	•••	<2
		2.970	•••	<2

¹⁵ J. Lindhard, M. Scharff, and H. E. Schiøtt, Kgl. Danske Videnskab. Selskab, Mat.-Fys. Medd. 33, No. 14 (1963).

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¹⁴ J. B. Marion, Nucl. Data A4, 301 (1968).

Level number	E_{x} (MeV)	E_{lpha} (MeV)	Target thickness (µg/cm ²)	$F(au)^{ a}$	Mean lifetimeª (fsec)	$\langle { m Mean} \ { m lifetime} angle^{ m b} \ ({ m fsec})$
1	0.842	5.60	15	0.095 ± 0.013	1.60 _{-0.21} ^{+0.29} psec	1.66 ± 0.34 psec
			150	0.111 ± 0.031	2.07 _{-0.47} ^{+0.88} psec	•
2	1.968	7.20	150	0.671 ± 0.012	182±8	182 ± 22
3	2.313	7.20	150	0.651 ± 0.028	196 ± 18	
		8.08	15	0.566 ± 0.046	180 ± 28	183 ± 25
		8.08	150	0.700 ± 0.037	164 ± 22	
4	2.869	8.08	15	1.019 ± 0.019	<10	
		8.08	150	0.999 ± 0.013	<10	<15
		8.75	150	1.004 ± 0.022	<14	
5	2.937	8.08	150	0.00 ± 0.038	>4.7 psec	>4 psec
6	2.970	8.08	15	$0.778 {\pm} 0.018$	80 ± 7	-
		8.08	150	$0.858 {\pm} 0.007$	79 ± 4	82 ± 12
		8.75	150	0.831 ± 0.023	91 ± 7	
7	3.221	8.75	150	0.944 ± 0.050	<60	<65

TABLE II. Mean lifetimes of the levels of ³³S.

^a Statistical errors.

^b Total errors calculated as explained in the text.

The quantity $F(\tau)$, in terms of dN(V)/dV, is given by

$$F(\tau) = \int_0^1 V \frac{dN(V)}{dV} dV \bigg/ \int_0^1 \frac{dN(V)}{dV} dV.$$
(7)

Using Eqs. (6) and (7) gives

$$F(\tau) = \frac{x\gamma_i^x}{(1+\gamma_i^2)^{x/2}} \int_0^1 V^2 (\gamma_i^{-2} + V^2)^{x/2 - 1} dV.$$
(8)

If the recoiling ions are stopping in two different media and v_c is the velocity at which the ions cross from one media to the other, then

$$F(\tau) = \frac{1}{A_1^{x_1/2}} \left[\frac{x_2 C_1^{x_1/2}}{C_2^{x_2/2}} \int_0^{V_\sigma} V^2 B_2^{x_2/2-1} dV + x_2 \int_{V_\sigma}^1 V^2 B_1^{x_1/2-1} dV \right], \quad (9)$$

where $V_c = v_c/v_c(0)$, $A_j = (1 + \gamma_{ij}^{-2})$, $B_j = (V^2 + \gamma_{ij}^{-2})$, and $C_j = (V_c^2 + \gamma_{ij}^{-2})$. The subscript 1 refers to the target and 2 refers to the backing.

Since an ensemble of nuclei recoiling into a backing have a certain velocity distribution, the emitted γ rays have a certain energy distribution, and this gives rise to a particular line shape for the detected γ rays. The theoretical line shape can be determined by folding the response function of the γ -ray detector into the distribution given by Eq. (6). If the response function of the detector is Gaussian, then the theoretical distribution is given by

$$(dN/dV)_{\rm th} = C_n \int_0^1 \left[dN(U)/dU \right] \\ \times \exp[-(U-V)^2/\sigma^2] dU, \quad (10)$$

where the normalization C_n and σ are determined from

the experimental distribution. The theoretical distribution given in Eq. (10) is fitted to the experimental distribution by normalizing the two distributions at the unshifted energy $E_{\gamma 0}$ and then varying x and γ_i^2 in discrete steps. The best fit is determined from the minimum value of χ^2 .

D. Analysis

The centroids of the peaks of interest in each spectrum were determined by a computer program which calculated the first moment and its statistical uncertainty after subtracting background. The background was approximated by least-squares fitting an exponential to the portions of the spectrum near the peaks. The Doppler shifts were then computed and corrected for any baseline shifts by observing the shifts of the "source" γ rays. The experimental Doppler shift attenuation factor $F_{exp}(\tau)$ was then computed from Eq. (1).

In order to extract the lifetime from the experimental value of $F_{exp}(\tau)$, Eq. (8) was used to evaluate $F(\tau)$ as a function of the lifetime. The value of K_e , the electronic stopping parameter, was obtained for ³³S ions recoiling in Ni by increasing the LSS value of $(-dE/dx)_{electronic}$ at $v_z = v_0$ by 15%. Evidence presented by several investigators¹⁶⁻¹⁹ indicates that for Z=16 ions in any backing the value of K_e is ~15% larger than the LSS value. The value of K_n , the nuclear-stopping parameter, was obtained directly from the LSS theory by using the value of $(-dE/dx)_{\text{nuclear}}$ at $v_z = v_0$.

¹⁶ J. H. Omrod and H. E. Duckworth, Can. J. Phys. 41, 1424 ¹⁶ J. H. Omrod and H. E. Deckard, (1963).
¹⁷ J. H. Omrod, J. R. MacDonald, and H. E. Duckworth, Can. J. Phys. 43, 275 (1965).
¹⁸ B. Fastrup, P. Hvelplund, and C. A. Sautter, Kgl. Danske Videnskab. Selskab, Mat.-Fys. Medd. 35, No. 10 (1966).
¹⁹ P. Hvelplund and B. Fastrup, Phys. Rev. 165, 408 (1968).



FIG. 3. 0.842-MeV γ ray from the ³³S 0.842 \rightarrow 0 transition observed at 0°. The state was populated by bombarding a 15- μ g/cm² SiO₂ target evaporated onto a Ni backing with 5.60-MeV α particles. The spectrum has a dispersion of 0.435 keV/channel. Background has been subtracted. The solid curve is a theoretical fit to the line shape and was obtained as explained in the text. The dashed curve is the line shape observed with the detector at 90°. The parameters of the theoretical fit are given in the figure.

The target thickness was taken into account explicitly in calculating the theoretical value of $F(\tau)$. The target was divided into 10 layers, and $F_i(\tau)$ was calculated for each layer i by using Eqs. (8) and (9). The velocity of the ion as it crossed into the backing V_c was calculated for each layer of the target using an expression for the distance along the beam direction as a function of V. The cross section for the reaction was assumed to be uniform throughout the target, and the average value of $F(\tau)$ was calculated from the 10 values of $F_i(\tau)$ for values of τ ranging 10⁻¹⁵-10⁻¹¹ sec. Figure 2 shows $F(\tau)$ plotted against τ for two different target thicknesses. The experimental values of $F(\tau)$ for the 2.970-MeV state are shown for two target thicknesses, and the lifetimes are seen to agree rather well. The characteristic slowing down times for ³³S ions in Ni and SiO₂ (α_B and α_T) are also indicated.

Most of the γ rays studied in this work exhibited line shapes as well as shifts. Information on the stopping parameters of the backing as well as the lifetime can be extracted from a line-shape analysis. However, for all but one of the γ rays the shoulder caused by the Doppler shift is on the low-energy side of the peak. This introduces difficulties in folding in the detector response function since the Ge(Li) detector exhibited a pronounced low-energy tail. The response function is not Gaussian, and the line shape is degraded by the tailing. The 0.842-MeV γ ray had a shoulder on the high-energy side, and a line-shape analysis was carried out. In calculating the statistical error in $F(\tau)$ and τ the statistical uncertainty in the position of each centroid was doubled, and this uncertainty was then used in determining a statistical error in $F(\tau)$ and τ . The total error was then calculated by assuming the following errors in addition to the statistical errors: a 15% error in the electronic stopping parameter (K_e) of the target and the backing, a 20% error in the nuclear stopping parameter (K_n) of the target thickness, and a variation in the initial recoil velocity of ~5% determined from Eq. (3). These assumptions provide error estimates that are rather conservative.

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III. RESULTS

A. Branching Ratios

Branching ratios were determined by using the spectra recorded with the detector at 60° since this is the nearest angle measured to 55°, the zero of $P_2(\cos\theta)$. A limit on the branching of a given state was obtained by summing the region of the spectrum at which the γ ray was expected. This sum plus two standard deviations was used to set a limit on the branch. The branching ratios of the first seven excited states of ³³S are given in Table I. The 0.969-MeV γ ray resulting from the transition 2.937 \rightarrow 1.968 was broadened by a contaminant γ ray, and the branching of the 2.937-MeV state could not be determined. Assuming that the branching is 50% to the ground state and 50% to the second excited state,1 a limit of < 2% is obtained for the 2.937 \rightarrow 2.313 transition. The information obtained on the branching of the 2.970-MeV state is new. The branching ratios for the 3.221- and 2.313-MeV states are in good agreement with the previous work, and limits on all possible branches of the 3.221-MeV state have been obtained.



FIG. 4. Full-energy loss peaks of the 1.968- and the 2.313-MeV γ rays corresponding to the 1.968- \rightarrow 0 and 2.313- \rightarrow 0 transitions, respectively. The levels were populated by bombarding a 150- μ g/cm² SiO₂ target evaporated onto a Ni backing with 7.20-MeV α particles. Spectra were recorded with the detector at 0°, 90°, and 150°. The experimental shifts in keV are shown. These spectra have a dispersion of 0.768 keV/channel.

E_{x} (MeV)	Ey (MeV)	$J_{i^{\pi}}$	J_f^{\star}	δ Reference 2	δ Reference 3	δ	Multipole	M ² (Weisskopf units)
0 842	0.842	<u>+</u>	<u>3</u> ↓			+0 18ª	<i>M</i> 1	3 1×10 ⁻²
0.012	0.012	2 1	2 1			10.10	E2	6.0
1.968	1.968	<u></u> 5+	$\frac{3}{2}$ +	$0.38 < \delta < 1.04$	0.79 ± 0.26	0.79	$\overline{M1}$	1.4×10^{-2}
		2.	-				E2	9.3
2.313	2.313	$\frac{3}{2}+$	$\frac{3}{2}+$			0 ^b	<i>M</i> 1	4.9×10 ⁻³
		-	-			ω ^b	E2	3.7
2.313	1.471	$\frac{3}{2}+$	$\frac{1}{2}+$	0.13<δ<1.33		0.47°	<i>M</i> 1	2.9×10 ⁻²
				$-8.1 < \delta < -0.73$			E2	11.9
2.937	2.937	$\frac{7}{2}$ —	$\frac{3}{2}+$	0.09 ± 0.27	$0.48 {\pm} 0.09$	0ь	M2	$10^{-3} M ^2 < (2.5)^d$
				$2.1 < \delta < \infty$	1.64 ± 0.45	ω p	E3	$0.7 < M ^2 < (2 \times 10^3)^d$
2.937	0.969	$\frac{7}{2}$ —	$\frac{5}{2}+$	0.0 ± 0.01	0.08 ± 0.09	0ь	E1	$5 \times 10^{-8} < M ^2 < (10^{-4})^d$
3.221	3.221	$\frac{3}{2}$ —	$\frac{3}{2}+$			0ь	E1	$>1.7 \times 10^{-4}$
3.221	2.379	$\frac{3}{2}$ —	$\frac{1}{2}$ +			0ь	E1	>6.5×10-4

TABLE III. Electromagnetic transition strengths in ³³S.

^a Calculated from τ and $B(E2\uparrow)$.

^b Assum**ed.**

Assumed near midpoint of the range of $\arctan\delta$.

 $^{\rm tl}$ Upper limit on τ of 10^{-8} sec from Ref. 2.

B. Lifetimes

Table II is a summary of the measured values $F(\tau)$ and τ . Columns 1-4 list the level number, the excitation energy, the bombarding energy, and the target thickness. Columns 5 and 6 list the weighted averages of $F(\tau)$ and τ based on shifts between different combinations of angles for each experimental condition. Column 7 gives the weighted averages of the lifetimes given in column 6. In columns 5 and 6 the errors are statistical, but in column 7 the total error is given. In setting limits on lifetimes the values of $F(\tau)$ used were obtained by adding or subtracting the total error on $F(\tau)$ to the value of $F(\tau)$ shown in Table II.

0.842-MeV Level

As mentioned previously the 0.842-MeV γ ray shown in Fig. 3 was the only one on which a line-shape analysis was carried out. The solid curve is the best fit to the 0° data minus the background using the method described in Sec. II C. Because of the low-energy tail, the theoretical distribution was fitted only to the points from $E_{\gamma 0}-0.5$ (FWHM) to the last data point shown. The dashed line shown in Fig. 3 is the line shape observed with the detector at 90°. The maximum shift in this case was only 5.44 keV, and the shoulder is so small that this shape was used only to check the previous values of the stopping parameters. The minimum value of χ^2 was obtained with the stopping parameters of Sec. II D; the lifetime of 1.60 psec is consistent with the value of 1.66±0.34 psec obtained from the shift measurement.

1.968- and 2.313-MeV Levels

Sample spectra obtained in the investigation of these two levels are shown in Fig. 4. These spectra were obtained with the detector at 0° , 90° , and 150° with a

beam energy of 7.20 MeV and a target thickness of 150 μ g/cm². The γ rays are the ground-state branches of these two levels. Neither of the γ rays are shifted by the maximum amount, as can be seen by comparing the observed shift shown in the figure to the maximum shift [which is just the denominator of Eq. (1)]. Both of the states have a value of $F(\tau) \approx 0.66$, which implies nearly equal lifetimes. The lifetime of the 1.968-MeV level is 182 ± 22 fsec, while the lifetime of the 2.313-MeV level is 183 ± 25 fsec.

Triplet at 2.9 MeV

Sample spectra of the ground state branches of the triplet at approximately 2.9 MeV are shown in Fig. 5. These spectra were recorded under the same conditions as those of Fig. 4, except for an increase of the beam energy to 8.08 MeV. The lifetimes of the 2.869- and the 2.970-MeV levels were determined to be less than 15 and 82 ± 12 fsec, respectively. As mentioned previously the 2.937-MeV state is long lived and the ground-state branch of this level was used to correct the spectra in Fig. 5 for baseline shifts. In setting a lower limit of 4 psec on the lifetime of the 2.937-MeV level an error of 1.0 channel was assumed for the error in the experimental Doppler shift. This corresponds to a F(r) of 0.0 ± 0.04 .

3.221-MeV Level

This level was populated by bombarding the target with 8.75-MeV α particles. Only the thick target (150 μ g/cm²) yield was high enough for a lifetime analysis. The value of $F(\tau)$ is sufficiently close to one that only an upper limit of 65 fsec can be set on the lifetime. The value of $F(\tau)$ quoted in Table II (0.944) corresponds to a lifetime of 34 fsec.



FIG. 5. Full-energy loss peaks of the ground-state branches of the triplet at approximately 2.9 MeV. The levels were populated by bombarding a 150-µg/cm² SiO₂ barget evaporated onto a Ni backing with 8.08-MeV α particles. Spectra were recorded with the detector at 0°, 90°, and 150°. The experimental shifts in keV are indicated. These spectra have a dispersion of 0.854 keV/channel.

IV. DISCUSSION

The calculated values of the electromagnetic transition strengths in Weisskopf units²⁰ (W.u.) are given in Table III for the levels of ³³S below 3.3 MeV with known spins. The values of the measured mixing ratios^{2,3} are given in columns 5 and 6, while the values used to calculated the transition strengths are shown in column 7. The branching ratios used to calculate the transition strengths are those given in Ref. 1. For the decays from the 2.937-MeV level, only pure transitions were assumed since only upper and lower limits of the lifetime are known. The value of the mixing ratio $|\delta| = 0.18$ for the first excited state at 0.842 MeV was obtained from the value of $B(E2\uparrow)$ from Coulomb excitation measurements²¹ and from the lifetime determined by this experiment. This yields an M1 strength of 3.1×10^{-2} W.u. and an E2 strength of 6.0 W.u.

Wiechers and Brussaard¹⁰ have calculated M1transition probabilities in the $(2s_{1/2}, 1d_{3/2})$ shell using the wave functions of Glaudemans et al.8 For the 0.842-MeV transition in ³³S these authors give values of $\Gamma(M1)$ of 0.023 and 0.013 meV, where the two values are calculated using single-particle and effective nucleon g factors, respectively. The transition strengths are then 1.9×10^{-3} , and 1.0×10^{-3} W.u. and are smaller than the experimental value of 3.1×10^{-2} by approximately a factor of 10. This difference is of the same order of magnitude as that observed for the 1.28-MeV level in ²⁹Si and the 1.26-MeV level in ³¹P.

The two levels of ³³S at 2.869 and 2.970 MeV do not have spin assignments. On the basis of the present lifetime measurements, the spins and parities $\frac{1}{2}\pm$, $\frac{3}{2}\pm$, $\frac{5}{2}\pm$, and $\frac{7}{2}^+$ are allowed.

As was pointed out by Becker et al.,² the decay of the 2.937-MeV level $(J^{\pi} = \frac{7}{2})$ via E3/M2 radiation to the ground state and E1 radiation to the 1.968 level is unusual. The lower limit of 4 psec set on the lifetime by this experiment indicates that either the recoil-distance²² or electronic timing method would be suitable for an accurate lifetime measurement. Knowledge of the transition strengths for this level should provide a sensitive test for theoretical calculations as well as allow one to determine whether a large inhibition of the E1or an enhancement of the E3/M2 is resonsible for this unusual competition.

Following a preliminary report of this work,¹¹ we became aware of a similar study²³ of the levels of ³³S.

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²⁰ D. H. Wilkinson, in Nuclear Spectroscopy, edited by F. Ajenberg-Selove (Academic Press Inc., New York, 1960), Part

B, p. 862 ff. ²¹ I. Kh. Lemberg, in *Reactions Between Complex Nuclei*, edited by A. Zucker, F. T. Howard, and E. C. Halbert (Wiley-Interscience, Inc., New York, 1960), p. 112.

²² K. W. Jones, A. Z. Schwarzschild, E. K. Warburton, and D. B. Fossan, Phys. Rev. 178, 1773 (1969).
 ²³ J. E. Cummings and D. J. Donahue (to be published).