tion as determined from our angular distribution assuming set I for the T=1 phase shifts.

All solutions listed in Table I have been obtained for the moving-neutron angular distribution in the formfactor impulse approximation. We have verified the existence of similar solutions for the moving- and stationary-neutron angular distributions with the closure and form-factor impulse approximations. In all solutions for the moving-neutron angular distributions, a better χ^2 is obtained with the form-factor method. To allow for all the uncertainties caused by the deuteron, the errors estimated from χ^2 minimization have been doubled.

We conclude: Both the standard and CD methods of analysis give equivalent results and establish a Yangtype phase-shift solution for the T=0 channel. The data are consistent with S- and P-wave K-nucleon scattering in both the T=0 and 1 channels. When

fitting the angular distribution alone, the CD method usually gives fewer solutions for a given number of parameters; in particular, the CD method gives no acceptable T=0 Fermi solution. The T=0 Yang solution is also consistent with the ratio $K_S^0 p/(\Delta \pi^+ + 2\Sigma^0 \pi^+)$ observed in $K_L^0 p$ scattering at comparable energies.¹⁵

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¹⁵ J. A. Kadyk, Y. Oren, G. Goldhaber, S. Goldhaber, and G. H. Trilling, Phys. Rev. Letters **17**, 599 (1966); J. K. Kim, *ibid*. **19**, 1074 (1967).

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Dilatation Covariance and Exact Solutions in Local Relativistic Field Theories*

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Consideration is given to the invariance of field equations under space-time dilatations with (induced) transformation of constant parameters therein. Analysis of such "dilatation covariance" facilitates the determination of rigorous closed-form singularity-free solutions to essentially nonlinear classical field equations. More over, for essentially nonlinear quantized (boson) field theories, dilatation-covariance considerations facilitate the determination of rigorous closed-form solutions to the Schrödinger stationary-state equation.

 $\mathbf{B}^{\mathrm{ECAUSE}}$ it has been customary to regard mass constants and coupling constants that appear in Lagrangians as being absolutely fixed, space-time dilatation (scale) transformations have played a minor role in the analysis of local relativistic field theories. Only special classical field theories feature dilatationtransformation invariance with absolutely constant physical parameters in their Lagrangians, and such theories do not ordinarily retain dilatation invariance when subject to quantization.¹ On the other hand, dilatation covariance on both the classical and the quantum level is a property of all local relativistic field theories with the appropriate (induced) dilatation transformation of parameters in the field equations. Our purpose in this article is to show that dilatationcovariance considerations not only facilitate the determination of rigorous closed-form solutions to classical field equations, but also facilitate the determination of rigorous closed-form solutions to the Schrödinger

stationary-state equation in local relativistic quantum field theories.

We illustrate the general notion and utility of dilatation covariance by discussing a specific model theory, the self-interacting complex scalar field theory based on the Lagrangian density

$$\mathcal{L} = -\left(\frac{\partial \psi^*}{\partial x^{\mu}}\right)\left(\frac{\partial \psi}{\partial x_{\mu}}\right) - m^2 \psi^* \psi + g \psi^* \psi \ln \psi^* \psi, \quad (1)$$

where m^2 and g are real physical constants and a system of convenient physical units is employed. The derived classical field equation

$$\frac{\partial^2 \psi}{\partial x^{\mu} \partial x_{\mu}} - m^2 \psi + g(1 + \ln \psi^* \psi) \psi = 0 \qquad (2)$$

admits solutions $\psi = \psi(x; m^2, g)$ that depend parametrically on m^2 and g with $x = (x^0, x^1, x^2, x^3)$ a point in space-time. For distinct pairs of values of m^2 and g, the manifolds of solutions² to Eq. (2) are homologous, the

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¹ For the implications of dilatation invariance in a special nonlinear field model, see H. Mitter, Nuovo Cimento **32**, 1789 (1964).

² Suitable regularity conditions associated with the existence of global solutions must supplement Eq. (1) for definition of the manifold of classical (c-number) solutions; the existence of global solutions to the generic equation $\partial^2 \psi / \partial x^{\mu} \partial x_{\mu} - U'(\psi^* \psi) \psi = 0$ has been the subject of numerous recent investigations: K. Jörgens, Math. Z. 77, 295 (1961); I. Segal, Proc. Symp. Appl. Math. 17, 210 (1965); C. S. Morawetz, Proc. Roy. Soc. (London) **306A**, 291 (1968).

correspondence of solutions being given by

$$\psi(x; m^2, g) = \lambda \psi(\xi x; \xi^{-2}(m^2 - 2g \ln \lambda), \xi^{-2}g), \quad (3)$$

with λ and ξ arbitrary positive constant parameters. Alternatively, the content of Eq. (3) is expressed by the covariance of Eq. (2) under the space-time dilatation transformations

$$\begin{split} \psi(x) &\to \lambda \psi(\xi x) ,\\ m^2 &\to \xi^{-2}(m^2 - 2g \ln \lambda) ,\\ g &\to \xi^{-2}g . \end{split} \tag{4}$$

Notwithstanding the fact that Eq. (3) involves transformed values of m^2 and g on the right-hand side, the general group-theoretic method³ for obtaining selfsimilar solutions to Eq. (2) is applicable; it leads, for example, to rigorous, spatially localized, singularity-free solutions of the form

$$\psi(x; m^2, g) = \exp\{ik_{\mu}x^{\mu} - \frac{1}{2}g[x_{\mu}x^{\mu} - (k_{\mu}x^{\mu})^2/(k_{\nu}k^{\nu})] + \frac{1}{2}g^{-1}(m^2 + k_{\mu}k^{\mu}) + 1\}, \quad (5)$$

where k_{μ} is a real timelike four-vector constant of integration related to the solution's canonical energymomentum four-vector⁴

$$p_{\mu} = (\pi/g)^{3/2} [2(-k_{\nu}k^{\nu})^{1/2} + g(-k_{\nu}k^{\nu})^{-1/2}] \\ \times \{ \exp[g^{-1}(m^{2} + k_{\nu}k^{\nu}) + 2] \} k_{\mu}.$$
(6)

Quantization of a scalar (boson) field that satisfies an essentially nonlinear dynamical equation like (1) is achieved by evoking the field-diagonal representation.⁵ Then we have the equal-time canonical commutation relations

$$[\psi(\mathbf{x}),\pi(\mathbf{y})] = [\psi^*(\mathbf{x}),\pi^*(\mathbf{y})] = i\delta(\mathbf{x}-\mathbf{y})$$
(7)

satisfied, with $\psi(\mathbf{x})$ and $\psi^*(\mathbf{x})$ acting effectively as c-number fields and the associated momentum densities prescribed as the (commuting) functional differential operators

$$\pi(\mathbf{x}) = -i\delta/\delta\psi(\mathbf{x}), \quad \pi^*(\mathbf{x}) = -i\delta/\delta\psi^*(\mathbf{x}). \tag{8}$$

We obtain the operator form of Eq. (1) by combining the Heisenberg equations of motion for $\psi(\mathbf{x})$ and $\pi(\mathbf{x})$,

⁴ It should be noted that the rigorous solution (5) is formally invariant under the dilatation transformations (4) if $k_{\mu} \rightarrow \xi^{-1}k_{\mu}$, or, equivalently, if $p_{\mu} \rightarrow \xi \lambda^{-2} p_{\mu}$; in general, the constants of integration in a solution acquire specific dilatation transformation character if one requires invariance of the solution.

⁶ Recent applications of the field-diagonal representation have been presented: G. Rosen, Phys. Rev. **173**, 1680 (1968), and works cited therein. Although a representation of the field in terms of creation and annihilation operators diagonalizes the energy in the case of a linear field theory and provides a practical basis for perturbation calculations in the case of a quasilinear field theory, it is expedient to work with the field-diagonal representation for an essentially nonlinear (boson) field theory, just as it is expedient to work with the coordinate-diagonal representation (and the Schrödinger equation) for a particle system involving a potential energy that is not simply quadratic in the particle coordinates.

in which the canonical Hamiltonian operator appears as

$$H = \int [\pi^*(\mathbf{x})\pi(\mathbf{x}) + \nabla \psi^*(\mathbf{x}) \cdot \nabla \psi(\mathbf{x}) + m^2 \psi^*(\mathbf{x})\psi(\mathbf{x}) -g\psi^*(\mathbf{x})\psi(\mathbf{x}) \ln\psi^*(\mathbf{x})\psi(\mathbf{x}) - u]d^3x, \quad (9)$$

with the disposable constant u inserted to compensate the zero-point quantum energy density. Stationary states are complex-valued functionals of $\psi(\mathbf{x})$ and $\psi^*(\mathbf{x})$ in the field-diagonal representation

$$\Psi = \Psi [\psi(\mathbf{x}), \psi^*(\mathbf{x}); m^2, g, u, E], \qquad (10)$$

satisfying the Schrödinger functional differential equation $H\Psi = E\Psi$. The Hamiltonian operator (9) is homologous with respect to the one-parameter subgroup of dilatation transformations (4) for which $\lambda = \xi$:

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$$\begin{aligned}
\psi(\mathbf{x}) &\to \xi \psi(\xi \mathbf{x}), \\
\pi(\mathbf{x}) &\to \xi^2 \pi(\xi \mathbf{x}), \\
m^2 &\to \xi^{-2}(m^2 - 2g \ln \xi), \\
g &\to \xi^{-2}g.
\end{aligned}$$
(11)

It follows from the Schrödinger equation that the correspondence of stationary states is given by

$$\Psi[\psi(\mathbf{x}),\psi^{*}(\mathbf{x}); m^{2},g,u,E] = \xi^{*}\Psi[\xi^{-1}\psi(\xi^{-1}\mathbf{x}),\xi^{-1}\psi^{*}(\xi^{-1}\mathbf{x});\xi^{-2}(m^{2}-2g\ln\xi), \xi^{-2}g,\xi^{-4}u,\xi^{-1}E].$$
(12)

In Eq. (12), s is a constant related to the state functional normalization with the form of the prefactor ξ^* on the right-hand side dictated by the subgroup composition law. For the vacuum state with energy eigenvalue E=0, Eq. (12) implies dilatation invariance; this generalizes the manifest dilatation invariance of the 'bare" vacuum-state functional

$$\Psi[\psi(\mathbf{x}),\psi^*(\mathbf{x}); m^2, 0, u, 0]$$

= exp $\left(-\int \psi^*(\mathbf{x})(-\nabla^2 + m^2)^{1/2}\psi(\mathbf{x})d^3x\right),$ (13)

in which

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$$u = [(-\nabla^2 + m^2)^{1/2}\delta(\mathbf{x})]_{\mathbf{x}=0}$$
 for $g = 0$.

The "dressed" vacuum state satisfies the functional differential equation $H\Psi=0$ with the Hamiltonian given by (8) and (9), and we solve for it by making the dilatation-invariant ansatz

$$\Psi = \Psi [\psi(\mathbf{x}), \psi^*(\mathbf{x}); m^2, g, u, 0]$$

= exp $\left\{ -\int \left[\psi^*(\mathbf{x}) (-\nabla^2 + \hat{m}^2)^{1/2} \psi(\mathbf{x}) + \frac{1}{4} \epsilon g (\psi^*(\mathbf{x}) \psi(\mathbf{x}))^2 \times \left(a - 1 + \ln \frac{\psi^*(\mathbf{x}) \psi(\mathbf{x})}{g}\right) \right] d^3 v \right\}, (14)$

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³G. Birkhoff, *Hydrodynamics* (Princeton University Press, Princeton, N. J., 1960), Chaps. 4 and 5; A. J. A. Morgan, Quart. J. Math. (Oxford) 2, 250 (1952).

where the "dressed" mass \hat{m} must transform under dilatations (11) $\hat{m} \rightarrow \xi^{-1}\hat{m}$, which implies that

$$\hat{m}^2 = m^2 - g \ln g + ag, \qquad (15)$$

with a an absolute constant; the limit $\epsilon \to 0$ is understood to be taken in (14) in such a way that⁶

$$\lim_{\epsilon \to 0} \left\{ \epsilon \left[\delta(\mathbf{x}) \right]_{\mathbf{x}=0} \right\} = 1,$$

implying the transformation character $\epsilon \rightarrow \xi^3 \epsilon$ under dilatations (11). To verify that (14) is the exact⁷ solution to the vacuum-state equation $H\Psi=0$, one simply computes

$$-\frac{\delta^{2}\Psi}{\delta\psi^{*}(\mathbf{x})\delta\psi(\mathbf{x})} = \{ [(-\nabla^{2}+\hat{m}^{2})^{1/2}\delta(\mathbf{x})]_{\mathbf{x}=0} + \epsilon [\delta(\mathbf{x})]_{\mathbf{x}=0} \{ ag\psi^{*}(\mathbf{x})\psi(\mathbf{x}) + g\psi^{*}(\mathbf{x})\psi(\mathbf{x}) \ln [\psi^{*}(\mathbf{x})\psi(\mathbf{x})/g] \} - [(-\nabla^{2}+\hat{m}^{2})^{1/2}\psi^{*}(\mathbf{x})][(-\nabla^{2}+\hat{m}^{2})^{1/2}\psi(\mathbf{x})] + O(\epsilon) \} \Psi \quad (16)$$

⁶ An immediate way to secure this relation is to introduce the wave-number cutoff-limit representation

$$\delta(\mathbf{x}) = \lim_{K \to \infty} \int_{|\mathbf{k}| \leq K} e^{i\mathbf{k} \cdot \mathbf{x}} d^3k / (2\pi)^3,$$

for which $\epsilon = 6\pi^2/K^3$.

⁷ To be precise, we have

$$\lim_{\epsilon \to 0} (H_{\epsilon} \Psi) = 0, \text{ where } H_{\epsilon} \equiv H - \int \Lambda_{\epsilon}(\mathbf{x}) d^3 x,$$

with

 $\Lambda_{\epsilon}(\mathbf{x}) \equiv -\frac{1}{2}\epsilon_{g}\psi^{*}(\mathbf{x})$

$$\times \{ (-\nabla^2 + \hat{m}^2)^{1/2}, \psi^*(\mathbf{x})\psi(\mathbf{x})[a - \frac{1}{2} + \ln\psi^*(\mathbf{x})\psi(\mathbf{x})/g] \} \psi(\mathbf{x}) \\ - \frac{1}{4} \epsilon^2 g^2 [\psi^*(\mathbf{x})\psi(\mathbf{x})]^3 [a - \frac{1}{2} + \ln\psi^*(\mathbf{x})\psi(\mathbf{x})/g]^2,$$

i.e., the terms represented by $O(\epsilon)$ in (16). Such a modification of the Hamiltonian definition $H \to \lim_{\epsilon \to 0} H_{\epsilon}$ is necessary because $\Lambda_{\epsilon}(\mathbf{x})$ is not uniformly convergent to zero for all $\psi^{*}(\mathbf{x}), \psi(\mathbf{x})$ as $\epsilon \to 0$.

and puts $u = [(-\nabla^2 + \hat{m}^2)^{1/2} \delta(\mathbf{x})]_{\mathbf{x}=0}$. The constant a in (14) is fixed by the functional-integral normalization condition⁵ $\langle \Psi | \Psi \rangle = 1$, and thus a is related to mass renormalization with \hat{m}^2 required to be finite and non-negative.

It is easy to abstract the analysis presented here for the specific model theory based on (1), and similar dilatation-covariance considerations can be applied to any other local relativistic field theory. For the generic self-interacting complex scalar field theory with the Hamiltonian operator

$$H = \int \left[\pi^*(\mathbf{x})\pi(\mathbf{x}) + \nabla \psi^*(\mathbf{x}) \cdot \nabla \psi(\mathbf{x}) + U(\psi^*(\mathbf{x})\psi(\mathbf{x})) - u \right] d^3x, \quad (17)$$

the vacuum-state functional satisfying $H\Psi=0$ is given by⁸

$$\Psi = \exp{-\int \left[\psi^*(\mathbf{x})(-\nabla^2 + \hat{m}^2)^{1/2}\psi(\mathbf{x}) + \epsilon F(\psi^*(\mathbf{x})\psi(\mathbf{x}))\right] d^3x}, \quad (18)$$

where

$$F(\rho) \equiv \int_0^{\rho} \left[U(\tau) - \hat{m}^2 \tau \right] \ln(\tau/\rho) d\tau , \qquad (19)$$

with the "dressed" mass \hat{m} prescribed in terms of physical constants in $U(\psi^*(\mathbf{x})\psi(\mathbf{x}))$ by dilatation covariance.

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⁸ In order to make this formal solution rigorous for $F(\rho)$ such that $\lim_{\rho\to\infty} F(\rho)/\rho\neq 0$, one must add to the Hamiltonian appropriate terms proportional to ϵ and ϵ^{2} . If the self-interaction energy density is such that $\lim_{\rho\to\infty} F(\rho)/\rho=0$, the vacuum-state solution (18) is a specialized form of the author's generic result in an unpublished report.