THE

Physical Review

A journal of experimental and theoretical physics established by E. L. Nichols in 1893

SECOND SERIES, VOL. 183, No. 1

5 JULY 1969

Spin-Optimized Self-Consistent Field Wave Functions*

Uzi Kaldor†

Department of Chemistry, Stanford University, Stanford, California 94305

and

Frank E. Harris Department of Physics, University of Utch, Salt Lake City, Utah 84112 (Received 19 December 1968)

A method is given for determining a spin-optimized self-consistent-field (SO-SCF) function based on optimized spatial orbitals and a spin function which is optimum in the S, M_S spin space. Results are presented for $\text{Li}(^2S)$, $\text{Li}(^2P)$, $\text{Be}^{+}(^2S)$, $\text{B}^{++}(^2S)$, and $\text{Be}(^1S)$. SO-SCF wave functions and energies are very similar to those obtained with restriction to the normally paired spin function, but the use of the entire spin space greatly improves the description of spin-dependent properties. Whereas single-spin-function calculations of the spin density at the nucleus are for $\text{Li}(^2S)$ and $\text{Li}(^2P)$, respectively, 9% and 100% in error, the corresponding SO-SCF errors are only 2% and 7%.

I. INTRODUCTION

There has been much recent interest in developing practical independent-particle methods capable of yielding good descriptions of spin-dependent properties of open-shell atoms. The methods receiving serious attention are all of the selfconsistent-field (SCF) type in which the spatial orbitals are optimized by applying an energyminimization principle. The simplest and most widely used such procedure, the restricted Hartree-Fock (RHF) method,¹ fails because the orbitals are optimized without regard for spin-dependent perturbations upon the closed-shell electrons. Early attempts to remedy this situation involved the separate optimization of α -spin and β -spin orbitals in a determinantal wave function.² This, the unrestricted Hartree-Fock (UHF) method, in general leads to a wave function which is

not an eigenfunction of $\2 . Improvements beyond the UHF method involve the removal of this defect.

The generation of many-electron spin eigenfunctions was very early discussed by Pauling,³ who gave rules for the handling of the several spin functions spanning an n-spin space of given S^2 and S_z . Serber⁴ and Yamanouchi⁵ considered the matrix representations of the permutation group having spin eigenfunctions as bases, and a comprehensive compilation of the formulas needed to generate spin eigenfunctions from arbitrary orbital products was given wide circulation by Kotani $et al.^6$ The production of a spin eigenfunction from a spin-orbital product was pointed out by Löwdin⁷ to be a projection, and Pauncz, de Heer, and Löwdin⁸ gave rules by which such projections could be characterized. Harriman⁹ gave explicit formulas for density matrices involving projec-

183

1

tors, and several workers¹⁰ have given explicit formulas and discussions of the properties of the "Sanibel" coefficients occurring in the projector algebra. Essentially equivalent is the "spin-free" approach to quantum chemistry described by Matsen.¹¹ More recently Goddard¹² has applied these group-theoretical concepts to derive one-electron SCF-like equations for functions having arbitrary but specified spin coupling.

The necessity to consider all spin functions spanning the spin space of an atom was pointed out some time ago by Pratt.¹³ Actual use of more than one spin function in calculations also has an appreciable history; for example, Harris and Taylor $^{\rm 14}$ found it advisable to include both spin functions in studies of Li and LiH. These calculations were not of the full SCF generality, but they indicated that the optimum spin function did not exactly correspond to singlet coupling of the Li 1s electrons. A similar result was reported by Ritter, Pauncz, and Appel,¹⁵ and recent Hylleraas-type calculations by Larsson¹⁶ on Li (^{2}S) indicate that for a given spatial function, optimization of the spin function usually improves the value calculated for the Fermi contact interaction. Kutzelnigg and Smith¹⁷ have pointed out the increased generality produced by the use of general spin functions in wave functions based on spin-orbital products, and Smith and Larsson¹⁸ have reiterated the role of the full spin space.

Actual applications by SCF-type methods have lagged somewhat behind the ideas indicated in the preceding paragraph. Löwdin⁷ made the concrete proposal that the spatial orbitals be optimized after spin projection of a Slater determinant built from pure-spin orbitals. This method, sometimes denoted spin-extended Hartree-Fock (SEHF), corresponds to the use of an *a priori* determined spin eigenfunction, and has the practical virtue that the spatial orbitals can be required to satisfy certain orthogonality relations.¹⁹ SEHF calculations of first-row atoms have been carried out by Goddard²⁰ (who refers to this method as "GF"), Kaldor,²¹ Sando,²² and Lunell.²³

A potential defect of the SEHF method is that its spin function does not correspond to close coupling of α and β spins into pairs, and it may often be better to optimize spatial orbitals using the spin eigenfunction which maximally couples pairs. We call this procedure "maximally paired Hartree-Fock" (MPHF); it is called the "G1" method by Goddard. The limited-basis-set studies of Ritter, Pauncz, and Appel¹⁵ on Li, larger-basis-set studies of Goddard²⁴ and Lunell,²³ and numerical studies by Hardcastle, Gammel and Keown,²⁵ all indicate that MPHF energies can be appreciably better than SEHF energies. However, the arbitrary choice of the maximally paired spin function interferes with good estimation of spin-dependent properties. For example,

MPHF functions have zero spin density at the nuclei of atoms having no s open shells, such as Li (^{2}P) , N (^{4}S) , or Mn (^{6}S) , in contradiction with experiment.

The obvious next step is to choose an optimum spin eigenfunction as well as to optimize the spatial orbitals. This method, which we call spinoptimized SCF (SO-SCF), will be seen to yield considerable qualitative improvement in the description of spin-dependent quantities. In this paper we report SO-SCF studies of several threeand four-electron systems.²³

II. THE SO-SCF FUNCTION

One way to write a wave function based on an orbital product and a general spin eigenfunction is the form

$$\Psi = \mathfrak{A} \Xi \Theta , \qquad (1)$$

where α is the antisymmetrizer, Ξ is a product of spatial one-electron functions,

$$\Xi = \chi_1(1)\chi_2(2)\cdots\chi_n(n) \tag{2}$$

and Θ is a linear combination of all independent spin functions spanning the space of appropriate S and $M_{\rm S}$:

$$\Theta = \sum_{k} t_{k} \theta_{k} \quad . \tag{3}$$

The best function Ψ is obtained by simultaneous optimization of the orbitals in Ξ and the coefficients t_k in Θ . This is the function we call the spin-optimized self-consistent field (SO-SCF) function.

The SO-SCF functions presented in this paper can alternatively be described as an antisymmetrized spin-projected Hartree product of spin orbitals

$$\Psi = \mathfrak{a} \circ \varphi_1(1) \varphi_2(2) \cdots \varphi_n(n) , \qquad (4)$$

where O projects a spin state characterized by S and M_S , and the spin orbitals φ_i are of the mixed-spin form

$$\varphi_{i} = \chi_{i} (a_{i} \alpha + b_{i} \beta) .$$
 (5)

Both the χ_i and the ratios a_i/b_i are to be determined variationally.

However, a given wave function often corresponds to an infinite number of different sets of a_i/b_i ratios, so that these ratios cannot be determined uniquely. For a two-electron singlet state, for instance, both ratios are completely immaterial and cancel out upon projection,²⁶ as there is only one spin function describing this state, viz. $\alpha\beta - \beta\alpha$. A three-electron doublet state has two independent spin functions, which may be chosen as

$$\theta_1 = 2^{-1/2} (\alpha \beta \alpha - \beta \alpha \alpha)$$

$$\theta_2 = 6^{-1/2} (2 \alpha \alpha \beta - \alpha \beta \alpha - \beta \alpha \alpha)$$
(6)

so that although there are three a_i/b_i ratios, there is only one actual spin parameter to be determined variationally, namely the ratio t_2/t_1 in Eq. (3) for Θ . These illustrations indicate why the SO-SCF function is represented more conveniently by Eqs. (1)-(3) than by Eqs. (4) and (5). It may also be noted that there exist SO-SCF functions (such as that for ${}^4S N$) which cannot assume the form given in Eqs. (4) and (5).

III. METHOD OF CALCULATION

The SO-SCF wave function is determined by a procedure similar to that used by one of us^{21} for the SEHF function. It can be shown that the operator *H*-*EI* has vanishing matrix elements between the SO-SCF function and all states singly excited therefrom (within the space spanned by

the basis orbitals). This extension of Brillouin's theorem is the basis for the following operations.

First, an arbitrary guess is made for Ψ , choosing Ξ and Θ in the forms given by Eqs. (2) and (3). Then all possible linearly independent singly excited states

$$\Psi_i^a = \mathbf{a} \Xi_i^a \Theta \tag{7}$$

are form, where $\Xi_i^{\ a}$ indicates the replacement of χ_i in Ξ by χ_a , with χ_a orthogonal to χ_i . The same Θ is used for Ψ and all $\Psi_i^{\ a}$. A configuration-interaction (CI) calculation is made to find the optimum wave function of the form $\Psi + \sum_{i,a} \chi_i^{\ a}$, following which the effects of configuration mixing are incorporated (to first order) in Ψ by modifying the orbitals:

$$\chi_i' = \chi_i + \sum_a C_{ia} \chi_a . \tag{8}$$

This process is repeated until convergence is achieved.

The spin function Θ is then optimized by per-

TABLE I.	SO-SCF	wave	functions	for	² S	states	of	three-electron a	systems.
----------	--------	------	-----------	-----	----------------	--------	----	------------------	----------

Basis orbital ^a χ_{1s}		<i>χ</i> _{1s'}	x_{2s}	
	Li (^2S) : $\Theta = 0.999992$	$8\theta_1 - 0.0037864\theta_2^{b}$		
1s(3.0000)	1.143 59	0.60995	0.02804	
3s(5.4000)	-0.07779	0.07839	0.00060	
3s(2.9972)	0.007 11	0.27138	0.02058	
3s(1.3465)	0.011 55	-0.00230	0.28475	
3s(0.8451)	-0.007 64	0.00243	0.64913	
4s(5.3300)	-0.10462	0.12827	-0.00230	
4s(0.7257)	0.003 22	-0.00122	0.11476	
$\chi(0)^{C}$	3,35256	1.788 13	0.08221	
	$Be^+(^2S): \Theta = 0.99998$	$51\theta_1 - 0.0054581\theta_2$		
1s(4.000)	1.14392	0.67073	0.04074	
4 <i>s</i> (6.870)	-0.076 21	0.101 20	-0.00230	
3 <i>s</i> (6.870)	-0.10780	0.11123	0.00049	
3s(4.040)	0.01209	0.19032	0.02525	
3s(2.002)	0.00382	-0.00089	0.38774	
4s(1.327)	0.00032	-0.00058	0.62580	
χ(0)	5.163 09	3.02734	0.183 88	
	$B^{++}(^{2}S): \Theta = 0.99998$	$23\theta_1 - 0.0059523\theta_2$		
1s(5.000)	1.141 19	0.71021	0.04770	
4s(7.860)	-0.02254	0.05939	- 0.001 29	
3 <i>s</i> (7.860)	-0.161 03	0.176 96	0.00099	
3s(4.840)	0.02115	0.117 14	0.02654	
3s(2.702)	-0.00232	-0.00188	0.39035	
3s(1.875)	0.00312	-0.001 13	0.60952	
χ(0)	7.19842	4.47989	0.30087	

^aOrbital exponents are given in parentheses.

 $b_{\theta_1} = (\alpha\beta\alpha - \beta\alpha\alpha)/\sqrt{2}; \ \theta_2 = (2\alpha\alpha\beta - \alpha\beta\alpha - \beta\alpha\alpha)/\sqrt{6}.$

^cWave function at nucleus.

183

forming a CI calculation including the functions Ψ_k ,

$$\Psi_{k} = \mathbf{a} \Xi \,\theta_{k} \,, \tag{9}$$

where Ξ is the spatial function calculated in the previous step and θ_k are the spin functions spanning the space of appropriate S and M_{c} . The resulting CI coefficients are used to construct a new approximation to Θ according to Eq. (3). The spatial orbital variation and the spin function optimization are repeated until the changes fall below a prescribed limit. In practice we do not wait for orbital convergence to carry out the spin function optimization: the fastest way is to determine Θ after each Ξ iteration. It should be noted that whereas the Ξ optimization is an iterative procedure, the best Θ for a particular Ξ is found in one step, which is in general faster than a step in the orbital variation. It therefore does not take significantly longer to optimize both Ξ and Θ than to optimize Ξ for a predetermined Θ . However, unlike the SEHF case, any transformation among the occupied orbitals will in general alter the wave function (this is generally true for any form of Θ but the SEHF function²⁷). No orthogonalization is therefore possible, and the improvement in the wave function is paid

for by having to calculate more complicated matrix elements.

IV. RESULTS AND DISCUSSION

SO-SCF calculations were carried out for the ^{2}S ground state of Li, Be⁺, and B⁺⁺. The basis sets were those used by $Goddard^{24}$ in his MPHF calculations. No exponent reoptimization was necessary, as the SO-SCF orbitals differ only slightly from the MPHF orbitals. Our SO-SCF wave functions are given in Table I. In addition, we determined the SEHF, MPHF, and SO-SCF wave functions for the lowest ${}^{2}P$ excited state of Li (Table II). The nine-orbital basis set of Weiss²⁸ was used. The basis orbital exponents were optimized for the SO-SCF function by the method of steepest descent with parabolic interpolation,²¹ and these exponents were also used for the MPHF function. Weiss's original exponents were found to be optimal for the SEHF function, Finally, we determined MPHF and SO-SCF functions for the ¹S ground state of Be, using a five-orbital set with exponents optimized for the SO-SCF function. These wave functions are shown in Table III.

The energy, spin density at the nucleus, and orbital overlaps given by the SO-SCF functions for the three-electron systems are collected in

Basis orbital ^a	x _{1s}	x _{1s'}	Basis orbital	x _{2p}
	SEHF func	tion: $\Theta = (2\alpha\beta\alpha - \alpha\alpha\beta)$	$(3 - \beta \alpha \alpha) / \sqrt{6}$	
1s(3.00)	0.88913	0.89341	2p(1.500)	J.13291
3s(9.60)	-0.00082	-0.000 98	4p(2.120)	0.04436
3s(3.38)	0.10454	0.11034	4p(1.275)	0.32486
3s(2.52)	$0.058\ 21$	0.04414	4p(0.785)	0.50649
/			4p(0.566)	0.16075
χ(0) ^b	2.606 60	2.61914	-	0.0
	MPHF :	function: $\Theta = (\alpha \beta \alpha - \beta \beta)$	$(3\alpha\alpha)/\sqrt{2}$	
1 <i>s</i> (2.9440)	1.17057	0,627 94	2p (1.5053)	0.13056
3s(9.5460)	-0.00101	-0.00905	4p(2.1161)	0.05527
3s(4.3055)	-0.23753	0.23917	4p(1.2041)	0.38840
3s(2.9228)	0.04148	0.21648	4p(0.7316)	0.51010
			4p(0.5316)	0.085 56
X(0)	3.33604	1.78958	-	0.0
	SO-SCF function	on: $\Theta = 0.9999940\theta_1 +$	$0.00344504\theta_2^{\rm C}$	
1s(2.9440)	1.17054	0.627 97	2p(1.5053)	0.13063
3s(9.5460)	-0.00101	-0.00904	4p(2.1161)	0.055 29
3s(4,3055)	-0.23738	0.23902	4p(1.2041)	0.38843
3s(2.9228)	0.04135	0.21662	4p(0.7316)	0.51007
- 0 (00)			4p(0.5316)	0.08550
χ(0)	3.335 93	1.78967		0.0

TABLE II. Wave functions for Li (^{2}P) .

^aOrbital exponents are given in parentheses.

^bWave function at nucleus.

 $c_{\theta_1} = (\alpha\beta\alpha - \beta\alpha\alpha)/\sqrt{2}; \ \theta_2 = (2\alpha\alpha\beta - \alpha\beta\alpha - \beta\alpha\alpha)/\sqrt{6}.$

Basis orbital ^a	x1s	<i>x</i> 1 <i>s</i> ′	χ_{2s}	x _{2s'}
	MPHF function: $\Theta =$	$\theta_1 = (\alpha\beta\alpha\beta + \beta\alpha\beta\alpha -$	αββα – βααβ)/2	
1 <i>s</i> (5.6858)	-0.04904	0,469 05	0.00844	0.04160
1s(2.9479)	1.088 03	0.65136	0.004 95	-0.02368
3 <i>s</i> (3.6747)	-0.04566	-0.16056	0.018 59	-0.11770
2 <i>s</i> (1.3767)	-0.04740	0.04889	-0.25646	0.92741
2 <i>s</i> (0.8441)	0.04230	-0.01185	1.20511	0.17591
χ(0) ^b	2.7319	5,447 9	0.0787	0.4199
	SO-SCF function:	$\Theta = 0.9998615\theta_1 -$	$\cdot 0.0166399 heta_2^{\mathbf{C}}$	
1 <i>s</i> (5.6858)	-0.05033	0.468 60	0.01052	0.03683
1s(2.9479)	1.09546	0.64941	0.01972	-0.05545
3 <i>s</i> (3.6747)	-0.04410	-0.16137	0.01711	-0.11573
2 <i>s</i> (1.3767)	-0.058 96	0.05532	- 0.256 63	0.93794
2s(0.8441)	0.02999	-0.006 00	1.20286	0.17636
χ(0)	2.7432	5.4389	0.1368	0.4199

TABLE III. Wave functions for Be $({}^{1}S)$.

^aOrbital exponents are given in parentheses.

 $\mathbf{b}_{\theta_2} = \left(2\alpha\alpha\beta\beta + 2\beta\beta\alpha\alpha - \alpha\beta\alpha\beta - \beta\alpha\beta\alpha - \alpha\beta\beta\alpha - \beta\alpha\alpha\beta\right)/\sqrt{12}.$

^cWave function at nucleus.

183

TABLE IV. SO-SCF results for three-electron systems.

	Energy	Spin density	Orbital overlaps			
	(hartree)	at nucleus	$\langle \chi_{1s} \chi_{1s'} \rangle$	$\langle \chi_{1s} \chi_{2s} \rangle$	$\langle \chi_{1s'} \chi_{2s} \rangle$	
Li(² S)	-7.447 565	0.2265	0.92832	0,136 15	0.227 60	
$Li(^2P)$	-7.380087	-0.0169	0.92766	• • •	• • •	
Be $+(^{2}S)$	-14.291620	0.9937	0.94915	0.20042	0.29992	
B ⁺⁺ (² <i>S</i>)	- 23.389919	2.5166	0,960 65	0.237 14	0,333 65	

TABLE V. Comparison of calculated and experimental results for Li.

	Energy	Spin density	(Orbital overlaps	
	(hartree)	at nucleus ^a	$\langle \chi_{1s} \chi_{1s'} \rangle$	$\langle \chi_{1s} \chi_{2s} \rangle$	$\langle \chi_{1s'} \chi_{2s} \rangle$
		$Li(^2S)$			
$\mathbf{RHF}^{\mathbf{b}}$	-7.432726	0.1666 (28%)	1.000 00	0.0	0.0
\mathbf{SEHF}^{21}	-7.432813	0.2412 (4%)	0.999 90	0.0	0.0
$MPHF^{24}$	-7.447560	0.2100 (9%)	0.928 22	0.14005	0.233 09
$SO-SCF^{c}$	-7.447565	0.2265 (2%)	0.92832	0.13615	0.227 60
Expt.	-7.4780 ^d	0.2313 ^e	• • •	• • •	• • •
		$Li(^2P)$			
\mathbf{RHF}^{28}	-7.365 068	0.0 (100%)	1.000 00	•••	• • •
SEHF ^C	-7.365 091	-0.0218 (20%)	0,99996	• • •	• • •
MPHF ^C	-7.380082	0.0 (100%)	0.92766	• • •	• • •
$SO-SCF^{c}$	-7.380 087	-0.0169 (7%)	0.92765	•••	•••
 Expt.	-7.410 16 ²⁸	-0.0181 ^f	•••	•••	•••

^aError in percent of experimental value is given in parentheses. ^bE. Clementi, IBM J. Res. Develop. Suppl. <u>9</u>, 2

^cThis work.

^dC. E. Moore, National Bureau of Standards Circular No. 467 (U.S. Government Printing Office, Washington, D.C., 1949).

^eP. Kusch and H. Taub, Phys. Rev. <u>75</u>, 1477 (1949). ^fK. C. Brog, T. G. Eck and H. Wieder, Phys. Rev. 153, 91 (1967).

^{(1965).}

	Energy						
	(hartree)	$\langle \chi_{1s} \chi_{1s'} \rangle$	$\langle \chi_{2s} \chi_{2s'} \rangle$	$\langle \chi_{1s} \chi_{2s} \rangle$	$\langle \chi_{1s} \chi_{2s'} \rangle$	$\langle \chi_{1s'} \chi_{2s} \rangle$	$\langle \chi_{1s'} \chi_{2s'} \rangle$
RHF ^a	- 14.573 02	1.000 00	1.000 00	0.0	0.0	0.0	0.0
${ m SEHF}^{20}$	-14.58721	0.95313	0.99547	0.0	-0.01155	0.03685	0.0
${ m MPHF}^{ m b}$	-14.589485	0.95106	0.87392	0.18070	0,36293	0.11534	0.24637
$\mathrm{SO}\operatorname{-SCF}^{\mathrm{b}}$	-14.589498	0.95354	0.87965	0.17594	0.31382	0.14049	0.22630
Expt. c	- 14.667 4	•••	•••		•••	•••	

TABLE VI. Comparison of calculated and experimental results for $Be({}^{1}S)$.

^aE. Clementi, IBM J. Res. Develop. Suppl. <u>9</u>, 2 (1965). ^bThis work.

^cE. Clementi, J. Chem. Phys. <u>38</u>, 2248 (1963).

Table IV, and the Li calculations are compared in Table V with the RHF, SEHF, and MPHF results and with experimental data. It is obvious that the SO-SCF functions are almost identical with the corresponding MPHF functions, and that the contribution of the "open-shell" spin function is very small. These observations confirm those of earlier studies.^{15,16} The energy improvement of the SO-SCF function over the MPHF function is insignificant, but the spin density at the nucleus is greatly improved. Whereas the MPHF spin density for Li (^2S) is 9% in error, the inclusion of both possible spin functions reduces the error to 2%. The effect is even more dramatic in the case of Li $({}^{2}P)$. The MPHF function yields an identically zero spin density at the nucleus for this state, as there are no unpaired s electrons. The very small contribution of the second spin function to the SO-SCF function brings the spin density to within 7% of the experimental value. These facts suggest that both spin functions are

needed to obtain a good general description of the spin properties of lithium.

The results for Be $({}^{1}S)$ are qualitatively consistent with the observations already made for lithium. As shown in Table VI, the SO-SCF and MPHF energies are very nearly equal, and the second spin function appears with a very small coefficient in the SO-SCF function. In Be, as in Li, the increased flexibility provided by the second spin function results in a small, but significant decrease in the splitting of the spatial orbitals within each shell. We anticipate that these observations will probably apply also to heavier atoms, where UHF²⁹ and SEHF²¹ calculations do not in general give good agreement with experiment.

ACKNOWLEDGMENT

It is a pleasure to thank Dr. Vedene H. Smith for a number of helpful discussions.

*Supported in part by the National Science Foundation, Grant GP-5555.

[†]Present address: Department of Chemistry, Tel Aviv University, Tel Aviv, Israel.

¹D. R. Hartree, <u>The Calculation of Atomic Structure</u> (John Wiley & Sons, Inc., New York, 1957); C. C. J. Roothaan and P. S. Bagus, Methods Computational Phys. 2, 47 (1963).

²J. A. Pople and R. K. Nesbet, J. Chem. Phys. <u>22</u>, 571 (1954).

³L. Pauling, J. Chem. Phys. <u>1</u>, 280 (1933).

⁴R. Serber, Phys. Rev. <u>45</u>, 461 (1934); J. Chem. Phys. 2, 697 (1934).

⁵T. Yamanouchi, Proc. Phys.-Math. Soc. Japan <u>20</u>, 547 (1938).

⁶M. Kotani, A. Amemiya, E. Ishiguro, and T. Kimura, <u>Table of Molecular Integrals</u> (Maruzen Co., Ltd., ⁸R. Pauncz, J. de Heer, and P. O. Löwdin, J. Chem. Phys. 36, 2247 (1962).

- ⁹J. E. Harriman, J. Chem. Phys. <u>40</u>, 2827 (1964). See also A. Hardisson and J. E. Harriman, *ibid.* <u>46</u>, 3639 (1967).
- ¹⁰R. Pauncz, J. Chem. Phys. <u>37</u>, 2739 (1962);
- F. Sasaki and K. Ohno, J. Math. Phys. 4, 1140 (1963);
- N. Karayianis and C. A. Morrison, ibid. 6, 876 (1965);
- V. H. Smith and F. E. Harris, *ibid.*, to be published;
- V. H. Smith, J. Chem. Phys. 41, 277 (1964); F. E.
- Harris, ibid. 46, 2769 (1967).

¹¹F. A. Matsen, Advan. Quantum Chem. <u>1</u>, 59 (1964).

¹²W. A. Goddard, Phys. Rev. <u>157</u>, 73 (1967).

¹³G. W. Pratt, Phys. Rev. <u>102</u>, 1303 (1956).

¹⁴F. E. Harris and H. S. Taylor, Physica <u>30</u>, 185 (1964); H. S. Taylor, Ph.D. dissertation, University

of California, Berkeley, 1960 (unpublished).

¹⁵Z. W. Ritter, R. Pauncz, and K. Appel, J. Chem. Phys. 35, 571 (1961).

Tokyo, 1955).

⁷P. O. Löwdin, Phys. Rev. <u>97</u>, 1509 (1955).

¹⁷W. Kutzelnigg and V. H. Smith, J. Chem. Phys. <u>41</u>, 896 (1964).

¹⁸V. H. Smith and S. Larsson in "Proceedings of the Conference on Density Matrices, Queens University,

Kingston, Ontario, Canada, 1967" (unpublished).

¹⁹P. O. Löwdin, J. Appl. Phys. <u>33</u>, 251 (1962).

²⁰W. A. Goddard, Phys. Rev. <u>157</u>, 93 (1967); J. Chem. Phys. 48, 450, 1008 (1968).

²¹U. Kaldor, J. Chem. Phys. <u>48</u>, 835 (1968); <u>49</u>, 6 (1968); Phys. Rev. <u>176</u>, 19 (1968).

²²K. M. Sando, Ph. D. dissertation, University of Wisconsin, 1968 (unpublished).

²³S. Lunell, Phys. Rev. <u>173</u>, 85 (1968), has carried out MPHF and optimized-spin-function studies on $Li(^2S)$

which differ from the present work in that only the 2s spatial orbital was optimized. His results are qualitatively similar to those reported here.

²⁴W. A. Goddard, Phys. Rev. 169, 120 (1968).

²⁵D. L. Hardcastle, J. L. Gammel, and R. Keown, J. Chem. Phys. 49, 1358 (1968).

²⁶A similar phenomenon was observed by C. F. Bunge, Phys. Rev. <u>154</u>, 70 (1967), using orbitals of the form $\chi_i^+ \alpha + \chi_i^- \beta$ for two-electron systems.

²⁷F. E. Harris, Advan. Quantum Chem. 3, 61 (1967).

²⁸A. W. Weiss, Astrophys. J. <u>138</u>, 1262 (1963).

²⁹R. K. Nesbet, in <u>Quantum Theory of Atoms, Molecules</u> and the Solid State, edited by P. O. Löwdin (Academic Press Inc., New York, 1966), p. 157.

PHYSICAL REVIEW

VOLUME 183, NUMBER 1

5 JULY 1969

Annihilation of Positrons in Argon I. Experimental

P. H. R. Orth* and G. Jones

Department of Physics, University of British Columbia, Vancouver 8, British Columbia, Canada (Received 24 October 1968)

The annihilation of positrons in argon has been investigated as a function both of argon density and applied electric field. The direct annihilation rate decreases with increasing field as observed in earlier work. These electric field results are compared with simple theoretical models of the positron-argon atom interaction in the following paper. A small nonlinear dependence on density of the direct annihilation rate became apparent at densities larger than about 10 amagats. The linear portion of the direct rate was characterized by $Z_{eff} = 27.3 \pm 1.3$. From the density dependence of the orthopositronium lifetime, the free orthopositronium annihilation rate (λ_1) and linear quenching rate in argon (λ_q) were found to be

$$\lambda_1 = (7.53 \pm 0.18) \times 10^6 \text{ sec}^{-1},$$

 $\lambda_{\alpha} = (0.24 \pm 0.02) \times 10^{6} \text{ sec}^{-1} \text{ amagat}^{-1}.$

The role that impurities play in these measurements is also discussed.

I. INTRODUCTION

Recent investigations of the lifetime of positrons in argon have shown that the free positron annihilation rate cannot be described by a single exponential.¹⁻⁴ Time spectra of the annihilation γ rays show clear evidence of a shoulder followed by an exponential decay presumed to characterize annihilation of positrons at thermal velocities. It has been shown that the shoulder is removed, and the lifetime of the exponential increased, when a moderate static electric field is applied.⁵ Typically, a field of about 80 V cm⁻¹amagat⁻¹ is sufficient to increase the direct lifetime by a factor of 2. These results provide the only available experimental test of the validity of models that describe the positron-argon interaction at low relative velocities. A further series of these measurements has been made with improved instrumentation offering greater experimental precision; the measurements were performed to a greater degree of statistical accuracy. The results of these measurements are used to test the validity of several empirical potentials describing the effective positron-argon interaction in paper II which follows.

While engaged in this program, several other features of the decay of positrons in argon were measured in order to facilitate comparison with the results of other workers. These measure-

¹⁶S. Larsson, Phys. Rev. <u>169</u>, 49 (1968).