

## $4f^n \rightarrow 4f^{n-1}5d$ Spectra of Rare-Earth Ions in Crystals

EUGENE LOH

Physical Sciences Department, McDonnell Douglas Corporation, Santa Monica, California

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The  $4f^n \rightarrow 4f^{n-1}5d$  absorption spectra of rare-earth ions,  $RE^{2+}$  or  $RE^{3+}$ , in crystals can be fitted with energy gaps in the ground multiplets of the  $4f^{n-1}$  core. Unlike the shielded  $4f^n$  configuration, the  $4f^{n-1}5d$  levels of RE in crystals are therefore very different from those of free ions, and are formed through the interaction between the electron in  $5d$  orbitals with the  $4f^{n-1}$  core.

**T**HE absorption spectra of rare-earth ions,  $RE^{2+}$  or  $RE^{3+}$ , in crystals arise mainly from two types of electronic transitions: weak  $4f \rightarrow 4f$  at low energies, and strong  $4f \rightarrow 5d$  at higher energies. The former are analogous to free-ion levels due to the shielded character of the  $4f^n$  configuration. The  $4f^n \rightarrow 4f^{n-1}5d$  transitions in crystals are not well understood because of uncertainties<sup>1</sup> in the relative magnitude of the crystal field, electrostatic interaction, and spin-orbit coupling in the  $4f^{n-1}5d$  configuration.

In this work we show that the complex patterns in the  $4f^n \rightarrow 4f^{n-1}5d$  absorption spectra of most  $RE^{2+}$  or  $RE^{3+}$  ions in solids can be fitted to energy gaps in the ground multiplets of the  $4f^{n-1}$  core. The  $4f^{n-1}5d$  levels may therefore be understood as formed<sup>2-4</sup> by the electron in the  $5d$  orbitals  $e_g$  or  $t_{2g}$  interacting with the  $4f^{n-1}$  core. As a consequence of this strong crystal-field effect<sup>5</sup> on the  $5d$  electron, the  $4f^{n-1}5d$  configurations of RE in solids are very different from those of free ions.

The  $4f^n \rightarrow 4f^{n-1}5d$  absorption spectra of most  $RE^{2+}$  or  $RE^{3+}$  in crystals exhibit two features. First, they consist of strong bands corresponding to components of  $5d$  orbitals split in the crystal field. Consequently, their spectra are similar [among  $RE^{2+}$  ions,<sup>2,6-8</sup> e.g.,  $Eu^{2+}$  (Ref. 6) and  $Yb^{2+}$  (Ref. 6); or  $RE^{3+}$  ions,<sup>9-12</sup>

e.g.,  $Ce^{3+}$  (Ref. 10),  $Pr^{3+}$  (Ref. 11), and  $Nd^{3+}$  (Ref. 12)] when the ions are embedded in the same type of hosts. Second, the structures of the  $5d$  bands can be fitted to energy differences in the ground multiplets of the  $4f^{n-1}$  core.<sup>13</sup> It is this second feature we wish to demonstrate in the absorption spectra of most  $RE^{2+}$  or  $RE^{3+}$  ions in alkaline-earth fluorides.

Five common  $RE^{2+}$  ions,  $Sm^{2+}$ ,  $Eu^{2+}$ ,  $Dy^{2+}$ ,  $Tm^{2+}$ , and  $Yb^{2+}$ , are shown in Fig. 1. The gross feature in the spectra of some ions has been characterized<sup>6</sup> as the  $e_g$  and  $t_{2g}$  components of the  $5d$  electron in the cubic crystal field. We shall now consider the structure, i.e., strong lines, broad peaks, and shoulders, within each component. We assign most of the strong lines as pure electronic transitions to  $4f^{n-1}5d$  levels and other structures as their vibronics, which are not yet resolved. We assume further that the  $4f^{n-1}5d$  levels are formed through interaction between the  $5d$  ( $e_g$  or  $t_{2g}$ ) electron and the  $4f^{n-1}$  core, and are spaced with the energy gaps in the  $4f^{n-1}$  ground multiplets. In Fig. 1 these levels are marked by bars, which precede most of the structure due to vibronics. For convenience we discuss the spectra of  $RE^{2+}$  in the following order: (1)  $Tm^{2+}$  and  $Dy^{2+}$ , (2)  $Eu^{2+}$  and  $Yb^{2+}$ , and (3)  $Sm^{2+}$ .

(1) In the  $4f^{13} \rightarrow 4f^{12}5d$  spectra of  $Tm^{2+}$ , a set of three bars corresponding to the  $^3H_6$ ,  $^3H_4$ , and  $^3H_5$  levels of the ground multiplet of the  $4f^{12}$  core precedes peaks or humps in both the  $e_g$ , 14 000  $cm^{-1}$  to 28 000  $cm^{-1}$ , and  $t_{2g}$ , 28 000  $cm^{-1}$  to 43 000  $cm^{-1}$  regions. In the  $4f^{10} \rightarrow 4f^9 5d$  absorption spectra of  $CaF_2:Dy^{2+}$ , a set of nine bars corresponding to  $^6H_j$  and  $^6F_j$  levels of the  $4f^9$  core again precedes structures in both the  $e_g$ , 10 000  $cm^{-1}$  to 27 000  $cm^{-1}$ , and  $t_{2g}$ , 27 000  $cm^{-1}$  to 44 000  $cm^{-1}$  regions.

(2) The  $e_g$  band of  $Eu^{2+}$  in  $CaF_2$  and  $SrF_2$  has been resolved directly by Kaplyanskii *et al.*<sup>4</sup> into the pure electronic transitions, shown by bars in Fig. 1, and their associated vibronics at liquid-helium temperature. By analogy, we analyze the  $t_{2g}$  band of  $Eu^{2+}$  in these two hosts by translating the same set of bars from  $e_g$  band to precede the unresolved vibronics in the  $t_{2g}$  band at liquid-nitrogen temperature, Fig. 1. The energy gaps in the ground multiplet of  $4f^6$ , or  $Eu^{3+}$ , are among the

<sup>1</sup> For example, T. S. Piper, J. P. Brown, and D. S. McClure, *J. Chem. Phys.* **46**, 1353 (1967).

<sup>2</sup> D. S. McClure and Z. Kiss, *J. Chem. Phys.* **39**, 3251 (1963).

<sup>3</sup> N. V. Starostin, *Opt. i Spektroskopiya* **23**, 486 (1967) [English transl.: *Opt. Spectry. (USSR)* **23**, 260 (1967)].

<sup>4</sup> A. A. Kaplyanskii and A. K. Przhvuskii, *Opt. i Spektroskopiya* **19**, 597 (1965) [English transl.: *Opt. Spectry. (USSR)* **19**, 331 (1965)].

<sup>5</sup> See, for example, Refs. 2, 3, 4, and 6, and B. P. Zakharchenya, V. P. Makarov, and A. Ya. Ryskin, *Opt. i Spektroskopiya* **17**, 219 (1964) [English transl.: *Opt. Spectry. (USSR)* **17**, 116 (1964)].

<sup>6</sup> A. A. Kaplyanskii and P. P. Feofilov, *Opt. i Spektroskopiya* **13**, 597 (1962) [English transl.: *Opt. Spectry. (USSR)* **13**, 129 (1962)].

<sup>7</sup> J. L. Merz and P. S. Pershan, *Phys. Rev.* **162**, 217 (1967); **162**, 235 (1967); J. L. Merz, Ph.D. thesis, Harvard University, 1966 (unpublished), and references therein, available as Technical Report No. 514, Office of Naval Research, NR-372-012.

<sup>8</sup> V. A. Arkhangelskaya, M. N. Kiselyeva, and V. M. Shraiber, *Opt. i Spektroskopiya* **23**, 509 (1967) [English transl.: *Opt. Spectry. (USSR)* **23**, 275 (1967)]; R. J. Pressley and J. P. Wittke, *IEEE J. Quantum Electron.* **3**, 116 (1967); Z. J. Kiss, *Phys. Rev.* **137**, A1749 (1965); **127**, 718 (1962).

<sup>9</sup> E. Loh, *Phys. Rev.* **147**, 332 (1966).

<sup>10</sup> E. Loh, *Phys. Rev.* **154**, 270 (1967).

<sup>11</sup> E. Loh, *Phys. Rev.* **158**, 273 (1967).

<sup>12</sup> E. Loh (to be published).

<sup>13</sup> G. H. Dieke and H. M. Crosswhite, *Appl. Opt.* **2**, 675 (1963); G. H. Dieke, H. M. Crosswhite, and B. Dunn, *J. Opt. Soc. Am.* **51**, 820 (1961).

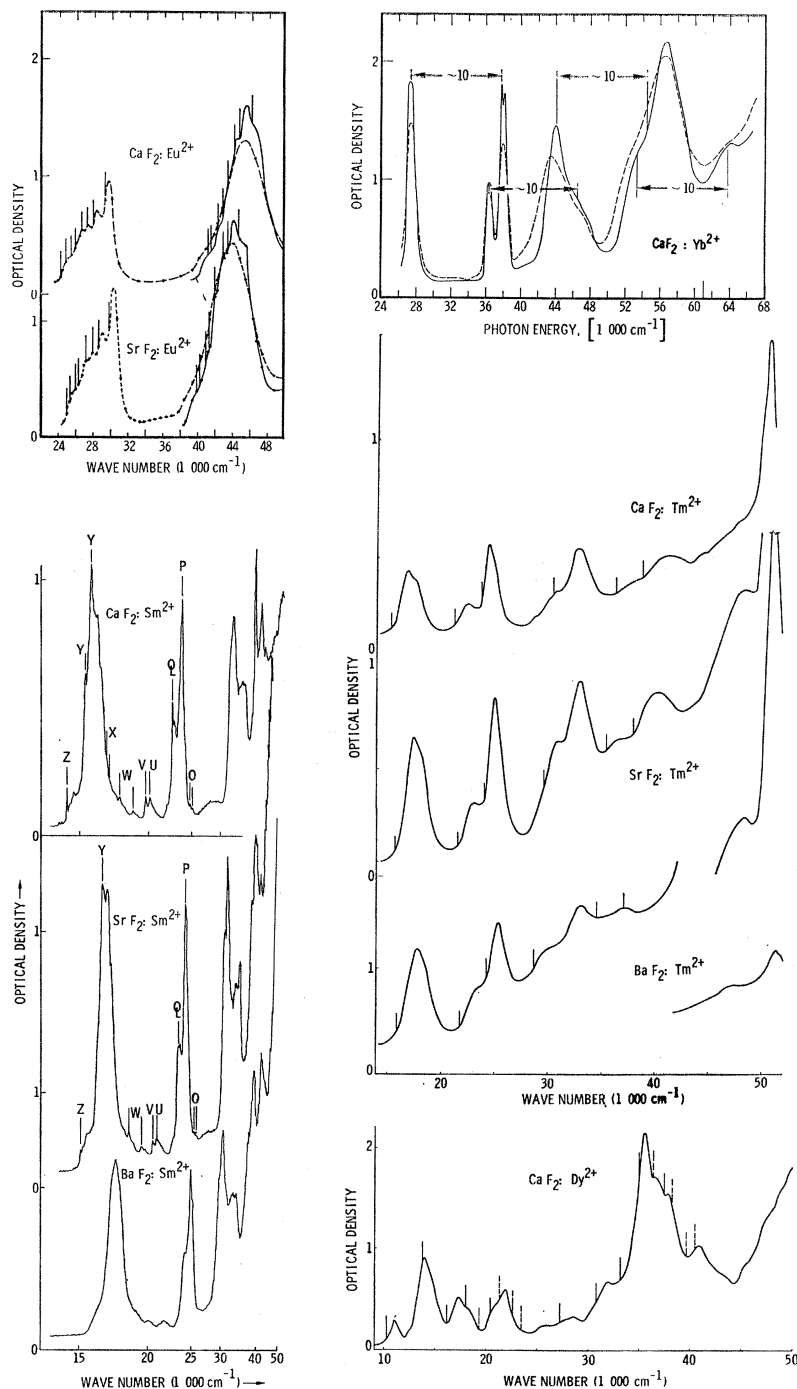


FIG. 1.  $4f^n \rightarrow 4f^{n-1}5d$  absorption spectra of some divalent rare-earth ions  $\text{Sm}^{2+}$ ,  $\text{Eu}^{2+}$ ,  $\text{Dy}^{2+}$ ,  $\text{Tm}^{2+}$ , and  $\text{Yb}^{2+}$  in alkaline-earth fluorides. ( $\text{Sm}^{2+}$  and solid curves of  $\text{Eu}^{2+}$  and  $\text{Yb}^{2+}$  at liquid-nitrogen temperature,  $\text{Dy}^{2+}$  and  $\text{Tm}^{2+}$  at room temperature.)

narrowest of all RE, while that of  $4f^{13}$ , or  $\text{Yb}^{3+}$ , is the largest. The  $\text{Yb}^{2+}$  spectrum in Fig. 1 clearly demonstrates this contrast by showing the huge  $\text{Yb}^{3+}(^2F_{7/2}) - \text{Yb}^{2+}(^2F_{7/2}) \approx 10\,000\text{ cm}^{-1}$  separation between  $e_g$  bands<sup>12</sup>  $37\,800\text{ cm}^{-1} - 27\,400\text{ cm}^{-1} = 10\,400\text{ cm}^{-1}$ , and  $46\,800\text{ cm}^{-1} - 36\,400\text{ cm}^{-1} = 10\,400\text{ cm}^{-1}$ , and  $t_{2g}$  bands<sup>12</sup>  $54\,500$

$\text{cm}^{-1} - 44\,100\text{ cm}^{-1} = 10\,400\text{ cm}^{-1}$ , and  $63\,500\text{ cm}^{-1} - 53\,100\text{ cm}^{-1} = 10\,400\text{ cm}^{-1}$ .

(3) We assign most of the strong lines in the  $e_g$ ,  $\sim 14\,000\text{ cm}^{-1}$  to  $\sim 30\,000\text{ cm}^{-1}$  region of the  $\text{Sm}^{2+}$  spectra in  $\text{CaF}_2$  and  $\text{SrF}_2$ , Fig. 1, as pure electronic transitions and mark them with vertical bars. Following

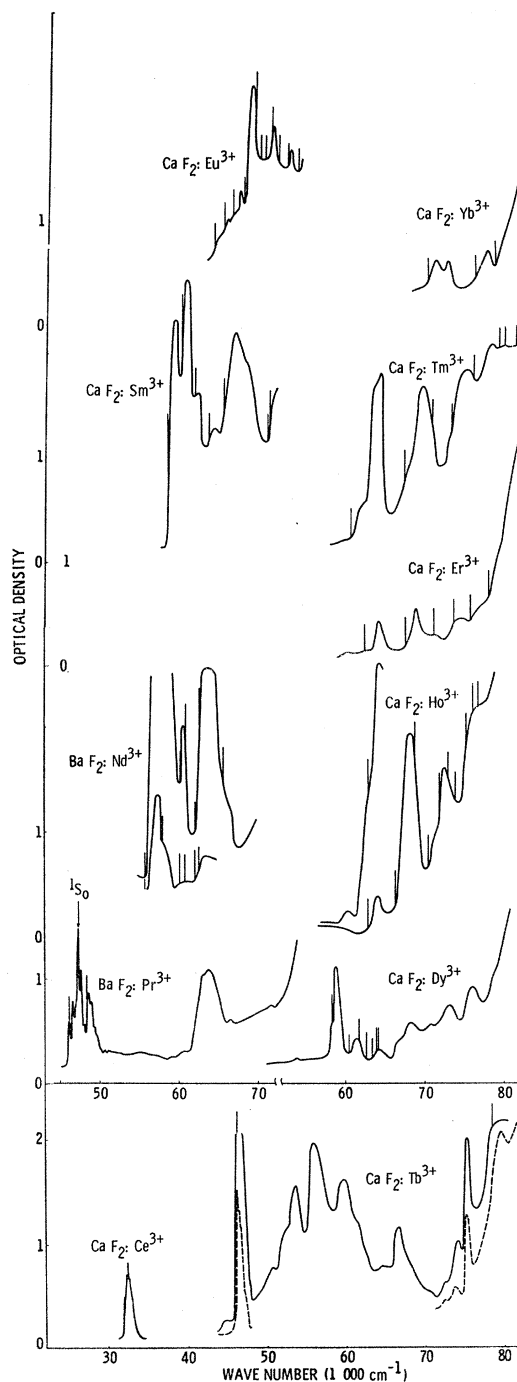


FIG. 2.  $4f^n \rightarrow 4f^{n-1}5d$  absorption spectra of trivalent rare-earth ions  $\text{Ce}^{3+}$ ,  $\text{Pr}^{3+}$ ,  $\text{Nd}^{3+}$ ,  $\text{Sm}^{3+}$ ,  $\text{Eu}^{3+}$ ,  $\text{Tb}^{3+}$ ,  $\text{Dy}^{3+}$ ,  $\text{Ho}^{3+}$ ,  $\text{Er}^{3+}$ ,  $\text{Tm}^{3+}$ , and  $\text{Yb}^{3+}$  in alkaline-earth fluorides. (All  $\text{RE}^{3+}$  are in  $\text{CaF}_2$  except  $\text{Pr}^{3+}$  and  $\text{Nd}^{3+}$  in  $\text{BaF}_2$ . All are measured at liquid-nitrogen temperature, except  $\text{Ho}^{3+}$  and  $\text{Er}^{3+}$  at room temperature. Concentration and sample thickness are 0.15%  $\text{Ce}^{3+} \times 0.29$  mm, 0.005%  $\text{Pr}^{3+} \times 0.95$  mm, 0.005%  $\text{Nd}^{3+} \times 1$  mm and 0.01%  $\text{Nd}^{3+} \times 13$  mm, 0.005%  $\text{Sm}^{3+} \times 10$  mm, >0.006%  $\text{Eu}^{3+} \times 10$  mm, 0.2%  $\text{Tb}^{3+} \times \sim 1$  mm and 0.02%  $\text{Tb}^{3+} \times \sim 1$  mm, 0.01%  $\text{Dy}^{3+} \times \sim 1$  mm, 1%  $\text{Ho}^{3+} \times \sim 1$  mm and 0.1%  $\text{Ho}^{3+} \times \sim 1$  mm, 0.01%  $\text{Er}^{3+} \times \sim 1$  mm, 0.1%  $\text{Tm}^{3+} \times \sim 1$  mm and >0.005%  $\text{Yb}^{3+} \times \sim 10$  mm.)

the work<sup>4</sup> on the  $e_g$  band of  $\text{Eu}^{2+}$ , we compare our assignment in the spectra<sup>14</sup> of  $\text{Sm}^{2+}$  with the ground multiplets<sup>13</sup> of  $4f^5$  or  $\text{Sm}^{3+}$  in Table I. The agreement in corresponding energy gaps is satisfactory. We do not venture to make similar comparison in the  $t_{2g}$  region of  $\text{Sm}^{2+}$  from  $\sim 30\,000$   $\text{cm}^{-1}$  to  $\sim 40\,000$   $\text{cm}^{-1}$  because of the presence of strong defect absorption,<sup>7,15</sup> e.g., at  $\sim 33\,000$   $\text{cm}^{-1}$  in the  $\text{Tm}^{3+}$  spectra and probably at  $\sim 36\,000$   $\text{cm}^{-1}$  in the  $\text{Dy}^{2+}$  spectrum of Fig. 1.

Based on the above assignment of  $4f^{n-1}5d$  levels of  $\text{RE}^{2+}$  and other considerations, such as the dependence of the absorption spectra on lattice parameters, we are able to assign the regions of  $e_g$ ,  $t_{2g}$ , and their separation, i.e., the crystal-field strength  $10Dq$ , more reliably. The data due to gross crystal-field effects on  $5d$  electrons of some  $\text{RE}^{2+}$  will be summarized in a separate publication.

All available  $\text{RE}^{3+}$  ions are shown in Fig. 2. Only  $\text{Gd}^{3+}$  and  $\text{Lu}^{3+}$  are absent because of their excessive  $4f \rightarrow 5d$  energies. The  $4f^{n-1}5d$  levels are again marked by vertical bars. Following this procedure used for  $\text{RE}^{2+}$ , these bars are placed in the spectra to precede most structures due to vibronics. For  $\text{Pr}^{3+}$  the  $1S_0$  level<sup>16,11</sup> at  $\sim 47\,200$   $\text{cm}^{-1}$  accidentally coincides with the  $4f(^2F_{7/2})5d(e_g)$  level in  $\text{CaF}_2$ . We use the spectrum of  $\text{BaF}_2 \cdot \text{Pr}^{3+}$  to avoid this overlap and hence unambiguously demonstrate the  $4f(^2F_{5/2})5d(e_g)$  and  $4f(^2F_{7/2})5d(e_g)$  levels as indicated by bars. For  $\text{Tb}^{3+}$  a huge separation between  $4f^7(^8S)5d(e_g)$  and  $4f^7(^6P_{7/2})5d(e_g)$  corresponding to the energy gap of  $\sim 32\,000$   $\text{cm}^{-1}$  between  $^8S$  and  $^6P_{7/2}$  levels<sup>13</sup> of  $\text{Gd}^{3+}$  is expected. Accordingly, we assign some of the structures between  $72\,000$  and  $80\,000$   $\text{cm}^{-1}$  in the  $\text{Tb}^{3+}$  spectra of Fig. 2 as unresolved vibronics of the  $4f^7(^6P_{7/2})5d(e_g)$  level as suggested by the bar at  $78\,400$   $\text{cm}^{-1}$ . We abandon our previous<sup>9</sup> speculation on the strong absorption peak at  $75\,200$   $\text{cm}^{-1}$  as due to  $4f \rightarrow 5d$  absorption of  $\text{Tb}^{4+}$ , since the energy of  $\text{Tb}^{4+}$  absorption should be higher than that of the isoelectronic  $\text{Gd}^{3+}$  because of higher effective nuclear charge in  $\text{Tb}^{4+}$  ion.

The  $4f \rightarrow 5d$  absorption spectra of  $\text{RE}^{3+}$  in alkaline-earth fluorides are more complicated than that of  $\text{RE}^{2+}$  because of the presence of charge compensators<sup>10,11</sup> accompanying trivalent ions. The charge compensators, e.g., interstitial fluorine ion, may lower the symmetry at  $\text{RE}^{3+}$  sites and can introduce cluster absorptions<sup>10,11</sup> due to clusters of  $\text{RE}^{3+}$ -charge-compensator complexes. The deduction of data on  $e_g$ ,  $t_{2g}$ , and  $10Dq$  of single-ion  $\text{RE}^{3+}$  in alkaline-earth fluorides from their spectra is therefore not as reliable as for  $\text{RE}^{2+}$  in the divalent host.

In conclusion, the  $4f^n \rightarrow 4f^{n-1}5d$  absorption spectra of  $\text{RE}^{2+}$  and  $\text{RE}^{3+}$  in crystals can be fitted to energy gaps in the ground multiplets of the  $4f^{n-1}$  core. The  $4f^{n-1}5d$

<sup>14</sup> D. L. Wood and W. Kaiser, Phys. Rev. **126**, 2079 (1962); W. Kaiser, C. G. B. Garrett, and D. L. Wood, Phys. Rev. **123**, 766 (1961).

<sup>15</sup> D. L. Wood (private communication).

<sup>16</sup> E. Loh, Phys. Rev. **140**, A1463 (1965).

TABLE I. Pure electronic transitions in  $e_g$  component of  $4f^6 \rightarrow 4f^5 5d$  absorption of  $\text{Sm}^{2+}$  in  $\text{CaF}_2$  and  $\text{SrF}_2$  (wave number  $\nu$  in  $1000 \text{ cm}^{-1}$ ).

$\text{CaF}_2:\text{Sm}^{2+}$		$\text{SrF}_2:\text{Sm}^{2+}$		$\text{Sm}^{3+}$ ground multiplets
$\nu$	$\Delta\nu$	$\nu$	$\Delta\nu$	$\Delta\nu$
14.5	0	15.2	0	0 Z or ${}^6\text{H}_{5/2}$
15.47	0.97	16.45	1.25	1.1 Y or ${}^6\text{H}_{7/2}$
15.8	1.3			
16.76	2.26			2.3 X or ${}^6\text{H}_{9/2}$
16.96	2.46			
17.68	3.18	18.3	3.1	3.6 W or ${}^6\text{H}_{11/2}$
18.68	4.18			
19.7	5.2	20.39	5.19	5 V or ${}^6\text{H}_{13/2}$
20.12	5.62	20.77	5.57	5.75 U or ${}^6\text{F}_{1/2}$
21.85	7.35	22.55	7.35	7 R or ${}^6\text{F}_{5/2}$
22.47	7.97	23.15	7.95	7.9 Q or ${}^6\text{F}_{7/2}$
23.64	9.14	24.12	8.92	9.05 P or ${}^6\text{F}_{9/2}$
24.7	10.2	25.33	10.13	10.4 O or ${}^6\text{F}_{11/2}$
25.1	10.6			
		25.62	10.42	

configurations of RE in crystals are therefore understood as formed through interaction of the  $5d$  orbital with the  $4f^{n-1}$  core.

The author wishes to thank D. L. Wood of Bell Telephone Laboratories for sending original tracings of  $\text{Sm}^{2+}$ , C. H. Anderson and R. J. Pressley of RCA Laboratories for supplying  $\text{Tm}^{2+}$ - and  $\text{Dy}^{2+}$ -doped

crystals, and Z. J. Kiss of RCA Laboratories for mentioning Ref. 1. He is also indebted to Hughes Research Laboratories for use of their vacuum-ultraviolet spectrometer, to the late G. Dorosheski for constructing some graphs, to R. R. Carlen and L. Podoksik for spectral measurements, and to J. R. Henderson for reading the manuscript.