Beta Decay of Mn^{56} and Co^{56}

C. CHASMAN AND R. A. RISTINEN* Brookhaven National Laboratory, Upton, New York (Received 10 January 1967)

 γ -ray transitions following the decay of Mn⁵⁶ to Fe⁵⁶ have been investigated with particular attention given to weak transitions following hindered β decays. Two such transitions were found, one to the 2.085-MeV state in Fe⁵⁶ (log ft = 8.7 \pm 0.2) and the other to the 3.445-MeV state (log ft = 6.0 \pm 0.2). Precision γ -ray measurements have been made both to corroborate the decay scheme and to provide γ -ray calibration energies. The decay of Co⁵⁶ to Fe⁵⁶ was similarly investigated; several new transitions were found, and precision γ -ray energies are reported.

I. INTRODUCTION

 \sum HE β decays of Mn⁵⁶ and Co⁵⁶ to Fe⁵⁶ have been examined by several investigators who have established the energies, spins, and parities of many of the low-lying levels of $Fe^{56,1-9}$ Many of the levels of Fe⁵⁶ have also been established by charged-particle reactions and inelastic neutron scattering. $8,10$ The spin of the ground state of Mn^{56} is measured¹¹ as 3 and the parity can be taken as positive from both the allowed β decay to the states of positive parity in Fe⁵⁶ and from shell-model considerations. Mn⁵⁶ β decays to the Fe⁵⁶ states at 2.66 $(2+)$, 2.96 $(2+)$, and 3.37 $(2+)$ MeV with allowed $\log ft$ values between 5 and 6 (Fig. 1). However, the log ft to the first excited 2^+ state in Fe⁵⁶ is 7.2 and to the second excited state has been reported' is 7.2 and to the second excited state has been reported² as > 9 and more recently > 10 ,^{3,12} although the 2⁺ and 4+ assignments to these states would indicate an allowed decay. No states in $Fe⁵⁶$ (not even the 3.445-MeV level) had been observed as populated from both the Mn^{56} and Co^{56} sides. As has been pointed out fre-

 3 J. K. Bienlein and H. Dinter, Nucl. Phys. 55, 113 (1964).

⁴ H. Pettersson, O. Bergman, and C. Bergman, Arkiv Fysik 29, 423 (1964). '

⁵ K. W. Dolan, D. K. McDaniels, and D. O. Wells, Phys. Rev. 148, 1151 (1966).

 M. Huguet, H. Forest, and C. Ythier, Compt. Rend. 263, 1342 (1966).

⁸ A. A. Katsanos, J. R. Huizenga, and H. K. Vonack, Phys. Rev. 141, 1053 (1966). Other recent work is summarized in this paper. ⁹ R. L. Auble, W. C. McHarris, and W. H. Kelly, Nucl. Phys.
A91, 465 (1967). This paper was received while the present work was in the final stages of preparation. The energy measurements are in generally good agreement with the reported errors somewhat smaller in this work. Auble, McHarris, and Kelly do not repor
observation of the two weakest transitions in Mn⁵⁶. Transition in Co⁵⁶ reported in this reference not observed in our work are
at 732.2 and 1811 keV. Transitions not observed in this reference

but observed in our work are at 1141.2 and 2213.0 keV.
¹⁰ R.W. Benjamin, P. S. Buchanan, and I. L. Morgan, Nucl.

Phys. 79, 241 (1966).
- "W. J. Childs, L. S. Goodman, and L. J. Kiefer, Phys. Rev.

122, 891 (1961).
 122 , 891 (1961). W. M. Greenberg, and E. B. Arnold, Bull. Am. Phys. Soc. 10, 931 (1965).

quently in the literature the collective nature of the first two excited states¹³ together with the *l*-forbiddenness of the β transitions to these states should result in a retardation of the β transitions. Although it is recognized that the β decays to higher vibrational states are nized that the β decays to higher vibrational states are often increasingly hindered,^{14,15} it is difficult to explain the large differences in transition rates between the decay to the 0.8466- and 2.085-MeV states and especially a $log ft > 10$. A search was made therefore for the 1.238-MeV transition from the second excited state of Fe⁵⁶ as a signature of the β decay to this state. This transition has been dificult to observe in the past for two reasons: (1) the one escape peak $(E_{\gamma}-m_0c^2)$ of the very strong 1.811-MeV transition lies within 65 keV of the transition and dominates the spectrum taken with a XaI crystal and (2) the full energy peak of a strong transition in $Ar⁴¹$, a common background source near reactors, is also within 65 keV. In the course of searching for the 1.238-MeU transition an examination of the decay schemes of Mn^{56} and Co^{56} was made, revealing some new transitions. Accurate energy measurements were made of the γ -ray transitions in Fe⁵⁶.

FIG. 1. Energy-level diagram of Fe⁵⁶ from the decay of Mn^{56} . Energies are in keV. Log \tilde{t} values obtained from the γ -ray intensities are given alongside the different β branches.

¹³ J. Bellicard and P. Barreau, Nucl. Phys. 36, 476 (1962). ¹⁴ L. S. Kisslinger and R. A. Sorensen, Rev. Mod. Phys. 35, 853 (1963)[~]

159 915

f Work performed under the auspices of the U. S.Atomic Energy Commission.

[~] Presently on leave of absence at the University of Colorado, Denver, Colorado.

P. Dagley, M. A. Grace, J. M. Gregory, and J. S. Hill, Proc. Roy. Soc. (London) 250A, 550 (1959). '

² P. Kienle and R. E. Segel, Phys. Rev. 114, 1554 (1959).

⁷ R. Schöneberg, M. Schumacher, and A. Flammersfeld, Z. Physik **192**, 305 (1966).

¹⁵ M. Sakai, Nucl. Phys. **33**, 96 (1962).

SOURCE MOVER

FIG.. 2. The source holder is advanced by the lead screw by a 6-rpm motor driven from a ratemeter. In this fashion the counting rate can be maintained constant to within 5%

II. METHOD

 Mn^{56} was made by neutron capture in monoisotopic Mn⁵⁵ in the Brookhaven Graphite Reactor. Samples of $MnO₂$ and of Mn (99.9% purity) were used of about 10—20 mg irradiated in polyethylene capsules. Gammaray spectra were then taken with a 4-cm' lithium-drifted germanium detector and a conventional pulse-height analysis system. Since the intensity of the transition sought was very small the counting rate was increased to about 8000 counts/sec at the cost of pulse-height resolution (about 5 keV full width at half-maximum, FWHM). The amplifier clipping and integrating times were also adjusted to compromise for faster counting rather than high resolution.

Because the source has a comparatively short half-life (2.6 h) and the electronics were counting at a moderately high rate, the resolution was seriously degraded by ratedependent pulse-height changes. To overcome this difficulty a source mover advanced by a rate meter was employed. In this arrangement the source capsule is placed on an arm extending perpendicularly from the axis of a 30-in. threaded shaft (13 threads/in.) . This shaft is driven by a 6-rpm motor controlled by a rate meter so that if the rate drops below a preset value the motor is turned on and then turned off again when the rate exceeds a diferent preset value. The input to the rate meter is the output of a single-channel analyzer

TABLE I. Mn⁵⁶ γ -ray energies (MeV) and intensities.

This work		Bienlein and Dinter ^a		Reidy and Wiedenbeckb	
Energy	Intensity	Energy	Intensity	Energy	
0.84664 ± 0.0001	100	$0.845 + 0.005$	100	$0.84679 + 0.00010$	
$+0.001$ 1.8112	27	$1.80 + 0.001$	34	$1.81098 + 0.00063$	
± 0.002 1.238	0.15				
$+0.001$ 2.1124	14	± 0.01 2.11	18.0	±0.0016 2.110	
$+0.001$ 2.523	1.0	2.52 ± 0.02	0.87		
± 0.002 2.600	0.04				
$+0.001$ 2.6576	0.63	$+0.01$ 2.67	0.69		
2.9581 $+0.001$	0.29	3.00 $+0.02$	0.55		
±0.001 3.3693	0.14	± 0.02 3.38	0.15		

^a Reference 3.
^b J. J. Reidy and M. L. Wiedenbeck, Nucl. Phys. **70,** 518 (1965).

'

whose window is adjusted to maintain the counting rate desired. A block diagram of the source mover is shown in Fig. 2. Sy this method the counting rate could be kept constant to about 5% until the source was too weak to be of value. An average source would last about 24 h. In some of the later experiments the analyzer was digitally stabilized from external pulsers at two points: one at the high-energy end and one at the lowenergy end of the multichannel analyzer. Since the electronic line shape is a function of the counting rate, digital stabilization is only satisfactory for high resolution when the counting rate is unchanging. Stability was maintained better than 1/10 000 per day.

Measurements were made with more than fifteen different samples of Mn and $MnO₂$. To check for contaminents irradiation times were changed from 30 sec to 150 sec and the spectra obtained were stored before using from 15 min to $2\frac{1}{2}$ h with no significant change in the spectra obtained. Measurements were also made under different background conditions. The line at 1.238 MeV appeared in every run. Figure 3 is the sum of three runs in the energy region of interest. For comparison the intensity of the 1.811-MeV transition is also shown. The intensity of the 1.238-MeV line was compared to that of the 1.811-MeV transition which was measured as a 27% transition. Since the 1.238-MeV transition could also appear if Mn^{56} could decay to the 3.445- or 3.123-MeV level in $Fe⁵⁶$ it was decided to search for a 2.600-MeV and a 1.038-MeV γ -ray transition. The 2.600-MeV transition was found and its intensity is measured as 0.04% of the 0.846-MeV line. A search for the 1.038-MeV transition was negative and an upper limit of 0.02% decays to the 3.123-MeV state is established. Figures 4 and 5 show the γ -ray spectrum from the Mn^{56} decay covering the total decay energy. Table I lists γ -ray intensities and energies from the Mn⁵⁶ decay. Gamma-ray intensities have an uncertainty of about 15% .

FIG. 3. The γ -ray spectrum from the decay of Mn⁵⁶ in the region of 1250 keV, showing the 1238-keV transition and the oneescape peak of the 1811-keV transition. The insert shows the
1811-keV full-energy peak. Note suppressed zero. These data represent the sum of three runs.

FIG. 4. The γ -ray spectrum from the decay of Mn⁵⁶ from 0 to 2.2 MeV.

Because the 2.600-MeV transition was found it was necessary to measure the branching ratio of the 2.600and 1.360-MeV transitions to find the feeding of the 2.085-MeV level from the 3.445-MeV level. To do this a $Co⁵⁶$ source was prepared since the 3.445-MeV level is strongly populated in the Co⁵⁶ positron decay.

The Co⁵⁶ source was made at the Brookhaven Cyclotron by irradiating natural iron (99.9% purity) with 9.5-MeV protons. Co⁵⁷ and Co⁵⁸ are the only long-lived contaminents expected. Co⁵⁷ has a decay energy of about 0.850 MeV with the highest-energy γ -ray transition¹⁶ of 0.7064 MeV and is consequently not troublesome. Co⁵⁸ comes from the (p,n) reaction on Fe⁵⁸ which is only $1/3\%$ abundant. The three lines which are known in the beta decay of $Co⁵⁸$ are at 0.81, 0.86, and 1.67 MeV, and do not interfere with the measurements made.

Besides measuring the ratio of intensities of the 2.599- and 1.360-MeV transitions the whole decay of $Co⁵⁶$ was examined (Fig. 6). Most of the transitions found are in essential agreement with those of Pettersson et al.⁴ and Dolan et al.⁵ Table II is a summary of these results comparing the energies and intensities found in the three experiments. Intensities have an uncertainty of about 15% . Gamma-ray transitions of 788.6, 977.51, 1141.2, 1175.5, 1964.6, 2113.0, and 2213.0 keV were not reported seen by Dolan et al. Transitions of 1141.2, 2213.0, and 3548.1 keV were not observed by Pettersson *et al.* (See also Ref. 17.) Figures 7 and 8 show γ -ray spectra from the decay of Co⁵⁶. An upper limit on the γ -ray intensity of the 734-keV transition observed only in the conversion electron spectrum by Pettersson was measured as a 0.08% relative to the intensity of the 846-keV transition.

Calibration of the γ -ray energies from both the Mn⁵⁶ and Co⁵⁶ decays was done using the standard γ -ray energies given in Table III and calibrating the electronics with a high-precision mercury pulser. Our method has assumed that two-escape peaks and oneescape peaks fall $2m_0c^2$ and m_0c^2 , respectively, below the full-energy peak linearly in the number of hole-electron pairs per keV energy loss. Because of the very stable electronics and in particular the digital stabilization it was possible to use very large pulse-height suppression in a biased amplifier followed by a large amount of gain. This resulted in multichannel-analyzer dispersions of about 300 eV/channel. Since the high-energy end of the $Co⁵⁶$ spectrum is considerably above convenient calibration γ rays with the listed sources the two-escape peak of the 3253.7-keV transition was standardized against the *n-p* capture γ ray measured as 2223.39 \pm 0.100 keV by comparison of its two escape peaks with the Co⁶⁰ standards given in Table III. This value is in good agreement with that of 2223.29 ± 0.07 keV given by Greenwood and Black.¹⁷ This γ ray thus established the calibration for the high-energy Co^{56} γ rays without need for a large extrapolation with the pulser.

FIG. 5. The γ -ray spectrum from the decay of Mn⁵⁶ from 1.7 to 3.5 MeV.

¹⁷ R. C. Greenwood and W. W. Black, Phys. Letters 21, 702 $(1966).$

¹⁶ O. C. Kistner and A. W. Sunyar, Phys. Rev. 139, B295 (1965).

FIG. 6. Energy-level diagram of Fe^{56} from the decay of Co^{56} . Energies are in keV.

III. DISCUSSION

A. Co⁵⁶

The energies and intensities of the Co⁵⁶ measurements are in good agreement with the results of Pettersson

FIG. 7. γ -ray spectrum from the decay of Co⁵⁶ from 0 to 2.6 MeV. Spectra used for calibration of γ -ray energies were taken at more than twice this dispersion.

et al. and Dolan et al. as shown in Table II. Unlike Ref. 6, no transitions were found to correspond to the doublets found by Hinrichsen et al. in their particlegamma coincidence reaction studies nor evidence for levels at 4459.4, 3600, and 2658 keV.¹⁸ Attention here is given only to the transitions found in this experiment that were not observed by both Pettersson et al. and Dolan et al.

The transition at 788.6 keV seen by Pettersson et al. but not Dolan et al. is confirmed but remains unassigned. A transition of 3548.1 keV is assigned to the decay of a level at 4394.7 keV and is the only evidence for such a state. This transition was also seen and similarly as-

FIG. 8. The γ -ray spectrum from the decay of Co⁵⁶ from 2.5 to 3.6 MeV. Spectra used for calibration of γ -ray energies were taken at more than twice this dispersion.

¹⁸ P. F. Hinrichsen, M. H. Shapiro, and D. M. Van Patter, Bull. Am. Phys. Soc. 10, 427 (1965); M. H. Shapiro, P. F. H. Hinrichsen, R. Middleton, and R. K. Mohindra, Phys. Letters 19, 573 (1965).

This work		Pettersson et al. ^a			Dolan <i>et al.</i> ^b	
Energy	Intensity	Energy	Intensity	Energy	Intensity	
788.6 ± 1.5 846.64 ± 0.10 977.5 ± 0.5 $1037.87 + 0.20$	0.36 100 1.5 14.0	787.91 846.82 977.46 1038.02	1.04 ± 0.21 100 $1.73 + 0.35$ 14.1 ± 1.5	846.5 ± 0.1 1038.1 ± 0.1	100 12.4 ± 0.5	
$1141.2 + 1.0$ $1175.5 + 0.5$ $1238.4 + 0.2$ $1360.2 + 0.15$ 1771.4 ± 0.15	0.17 1.6 64 4.0 14.0	1175.25 1238.56 1360.58 1771.90	2.1 ± 0.6 66.8 ± 4.0 4.0 ± 0.8 16.2 ± 1.4	$1238.6 + 0.1$ 1359.9 ± 0.2 1770.8 ± 0.2	$71.2 + 2.6$ $3.8 + 0.3$ 15.6 ± 1.3	
1964.6 ± 1.2 2015.4 ± 0.3 2034.9 ± 0.3 2113.0 ± 0.1 2213.0 ± 1.5	0.68 2.6 6.6 0.56 0.60	1964.8 2015.60 2035.50	$0.75 + 0.27$ 4.1 ± 1.2 9.2 ± 1.7	2015.6 ± 0.3 $2034.7 + 0.2$	$3.8 + 0.7$ $7.8 + 1.0$	
$2598.70 + 0.15$ 3010.0 ± 0.4 $3201.9 + 0.3$ $3253.7 + 0.2$ 3273.2 ± 0.3 3452.1 ± 0.3 3548.1 ± 0.3	14.0 0.60 2.9 7.2 1.6 0.72 0.2	2599.7 3010.4 3202.3 3254.0 3272.8 3452.5	17.4 ± 1.5 1.3 ± 0.4 3.2 ± 0.5 8.5 ± 0.6 1.5 ± 0.4 $0.95 + 0.15$	2598.9 ± 0.2 3009.5 ± 0.2 $3202.3 + 0.3$ 3254.0 ± 0.3 3273.6 ± 0.2 3452.6 ± 0.3 3548.3 ± 0.3	$16.0 + 2.7$ $1.9 + 0.8$ $2.9 + 1.1$ $5.8 + 2.7$ $1.2 + 0.5$ $0.7 + 0.3$ $0.2 + 0.1$	

TABLE II. $Co⁴⁵ \gamma$ -ray energies (keV) and intensities.

^a Reference 4.

^b Reference 5.

signed by Dolan et al. With the exception of this state all levels above and including 3445.3-keV decay to the levels at 2085 and 846 keV. However the branching ration of these transitions are such that it is impossible to definitely say whether the spins and parities of these states are the same. States of 4298.3 and 4100.3 keV have, in addition, a decay mode to the 3122.8-keV level but does not allow spin or parity assignments either. The new transitions of 1141.2 and 2113.0 keV sum to within 0.5 keV of the energy difference between the 4100.3- and 846.6-keV levels. This suggests a level of 2938 keV. Such a level is known from the Mn^{56} decay to decay by a transition to the ground state. However, the intensity of this γ ray precluded its observation in the Co⁵⁶ decay. The γ ray of 1175.5 keV is not a Co⁶⁰ contaminent because the 1332.-keV transition is not observed. This γ ray is satisfactorily placed depopulating the level at 4298.5 keV.

It is very difficult to assign spins and parities of the higher states on the basis of branching ratios. These states are presumably $3^+, 4^+, 5^+$ because of the allowed β decay from the 4⁺ Co⁵⁶ ground state.

⁴ G. T. Ewan and A. J. Tavendale, Canadian Atomic Energy Commission Report No. AECL 2079, 1964 (unpublished).

$B. Mn^{56}$

The β decay of Mn⁵⁶ to the 2085-keV state in Fe⁵⁶ has been found. The transition rate measured by comparing the intensity of the 1238-keV transition to that of 1811-keV transition corrected for feeding from the 3445-keV state yields $\log ft = 8.7 \pm 0.2$ for the decay to the 2.085-MeV state. This is in distinct contrast to the recently reported^{3,12} value of $\log ft > 10$. A β transition to the 3.445-MeV state in Fe⁵⁶ has been found with an allowed $\log ft = 6.0 \pm 0.1$. The lower $\log ft$ value is anticipated because at this excitation energy the transition need not be l forbidden. The absence of a 1038-keV transition implies a $\log ft > 7.9 \pm 0.2$ for populating the 3112-keV state. The states at 3445, 2085, and 2958 keV appear to be the only ones fed in the β decay of both Mn^{56} and Co⁵⁶. The log ft values for the Mn⁵⁶ and Co⁵⁶ decays to the 2085-keV state are very similar although both very large for allowed decays. It is reasonable to interpret this as indicative of the *l* forbiddenness of the decays together with the collective nature of the 2085keV state. The lower logft values for the 3445-keV level from the Mn⁵⁶ side might be interpreted as the transition of a $f_{7/2}$ proton to a $p_{3/2}$ neutron. It is not surprising then that the β transition from the Co⁵⁶ ground state is retarded to this level.

ACKNOWLEDGMENT

The authors would like to thank Dr. R. Wiener-Chasman for many helpful suggestions and discussions.