

## Inelastic Model of the $n$ - $p$ Mass Difference\*

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A simple model of the  $n$ - $p$  mass difference is exhibited, in field-theoretic terms, in order to illustrate one way in which the inelastic feedback mechanism may be expected to operate.

RECENTLY, two new derivations of the  $n$ - $p$  mass difference have been given<sup>1,2</sup> which attempt to include inelastic Compton effects by exhibiting a so-called feedback term with the effect of reversing the over-all sign of earlier estimates. The purpose of this note is to give a simple, field-theoretic example of how the feedback mechanism may be expected to operate. We consider the difference,  $\Delta\Sigma(\omega)$ , of the nucleon self-energies obtained by iterating the difference,  $\Delta\Sigma^{(0)}(\omega)$ , of the Feynman-Speisman (FS)-type<sup>3</sup> electromagnetic self-energies between an increasing number of virtual pions. This difference of sums over the subset of proper self-energy rainbow (or ladder) graphs, illustrated in Fig. 1, is supposed to approximate the  $n$ - $p$  mass difference, and provide the mechanism whereby inelastic states of one photon plus many pions can react on the strong-interaction determination of the nucleon masses. These graphs are chosen only because they can be generated by the iteration of a linear integral equation, which we do not attempt to solve.<sup>4</sup> Rather, we observe that, if a solution exists for which  $\text{Im}\Delta\Sigma$  is not pathological near threshold and vanishes rapidly for large  $\omega$ , then the feedback mechanism can be sufficient to reverse the sign of the initial  $\Delta\Sigma^{(0)}$  estimate. When evaluated with the aid of reasonable charge and moment form factors, the latter leads to the incorrect result for  $\Delta m = m_p - m_n$ , of amount  $\Delta m^{(0)} = -Z\Delta\Sigma^{(0)}(m) \sim +\frac{1}{2}$  MeV, where  $Z$  denotes the nucleon wave-function renormalization constant.

We represent by  $\delta\Sigma = x_p\delta\Sigma_p + x_n\delta\Sigma_n$  the sum of such self-energy graphs for the nucleon, obtained by the iteration of the simpler, electromagnetic  $\delta\Sigma^{(0)} = x_p\delta\Sigma_p^{(0)} + x_n\delta\Sigma_n^{(0)}$ . Here,  $x_{p,n}$  denote isotopic projection operators for proton and neutron, respectively, while  $\delta\Sigma_{p,n}$  and  $\delta\Sigma_{p,n}^{(0)}$  are the corresponding self-energy functions.  $\delta\Sigma^{(0)}$  is illustrated by the first term on the right side of Fig. 1; it is of order  $e^2$  and is understood to be given in terms of realistic charge and moment distributions

of the electromagnetic vertex. The model may thus be defined by the equation<sup>5</sup>

$$\delta\Sigma(-i\gamma \cdot p) = \delta\Sigma^{(0)}(-i\gamma \cdot p) + \frac{ig^2}{(2\pi)^4} \sum_{i=1}^3 \int \frac{d^4k}{k^2 + \mu^2} \gamma_5 \tau_i \frac{1}{m + i\gamma \cdot (p+k)} \times \delta\Sigma[-i\gamma \cdot (p+k)] \frac{1}{m + i\gamma \cdot (p+k)} \gamma_5 \tau_i, \quad (1)$$

where we neglect the variation of pion and nucleon masses in Eq. (1) because  $\delta\Sigma$  is already of order  $e^2$ . Both  $e^2$  and  $g^2$  are taken as renormalized, with  $g^2/4\pi = 15$ . Inserting isotopic representations, one easily finds, for  $\Delta\Sigma = \delta\Sigma_p - \delta\Sigma_n$ ,

$$\Delta\Sigma(-i\gamma \cdot p) = \Delta\Sigma^{(0)}(-i\gamma \cdot p) - \frac{ig^2}{(2\pi)^4} \int \frac{d^4k}{k^2 + \mu^2} \gamma_5 \times \frac{1}{m + i\gamma \cdot (p+k)} \Delta\Sigma[-i\gamma \cdot (p+k)] \frac{1}{m + i\gamma \cdot (p+k)} \gamma_5, \quad (2)$$

where  $\Delta\Sigma^{(0)} = \delta\Sigma_p^{(0)} - \delta\Sigma_n^{(0)}$ .

It is useful to employ for  $\Delta\Sigma(\omega)$  the form of the very general representations<sup>6</sup> valid for  $\Sigma(\omega)$ ,

$$\Delta\Sigma(\omega) = \int_{m+\mu}^{\infty} dn \left[ \frac{\Delta\rho_+(n)}{n-\omega} + \frac{\Delta\rho_-(n)}{n+\omega} \right], \quad (3)$$

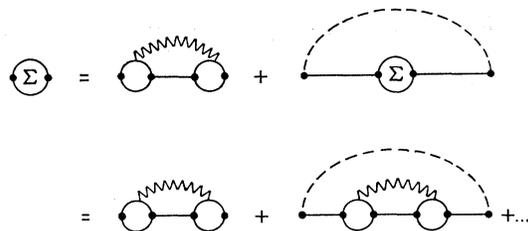


FIG. 1. Pictorial representation of an approximate integral equation for  $\delta\Sigma$ , and its iterative solution.

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<sup>1</sup> H. Pagels, Phys. Rev. **144**, 1261 (1966).

<sup>2</sup> H. M. Fried and T. N. Truong, Phys. Rev. Letters **16**, 559 (1966); **16**, 884(E) (1966); Phys. Rev. (to be published). The relation of this method to that of the conventional calculation is described in the latter paper.

<sup>3</sup> R. P. Feynman and G. Speisman, Phys. Rev. **94**, 500 (1954). References 1 and 2 contain or lead to an exhaustive list of the relevant literature.

<sup>4</sup> For simple  $\Delta\Sigma^{(0)}$ , this equation has been studied by many authors; e.g., D. Falk, Phys. Rev. **115**, 1069 (1959).

<sup>5</sup> More precisely, one should begin with the unrenormalized  $m_0, g_0$ , and include self-energy corrections to the pion and nucleon propagators and to the  $ps$  vertex; extract the appropriate renormalization constants which convert  $g_0$  to  $g$ ; and then approximate the renormalized propagators by their respective pole terms, and each renormalized  $ps$  vertex by  $\gamma_5 \tau_i$ .

<sup>6</sup> See, for example, M. Ida, Phys. Rev. **136**, B1767 (1964).

where  $\Delta\rho_{\pm}(n) = \pm(n \mp m)^2 \Delta\tau_{\pm}(n)$ . Insertion of Eq. (3) into Eq. (2), followed by the computation of an elementary Feynman integral, produces

$$\Delta\Sigma(\omega) = \Delta\Sigma^{(0)}(\omega) + \frac{g^2}{8\pi^2} \int_{m+\mu}^{\infty} dn \times \left[ \frac{\Delta\rho_+(n)}{n-m} f_+(n, \omega, \mu) + \frac{\Delta\rho_-(n)}{n+m} f_-(n, \omega, \mu) \right], \quad (4)$$

where

$$f_{\pm}(n, \omega, \mu) = \int_0^1 dx \times \left\{ \frac{xm[m - (1-x)\omega]}{\mu^2(1-x) + xm^2 - x(1-x)\omega^2} - \frac{1}{2} \frac{[n \mp (1-x)\omega]}{(n \mp m)} \times \ln \left[ \frac{\mu^2(1-x) + xn^2 - x(1-x)\omega^2}{\mu^2(1-x) + xm^2 - x(1-x)\omega^2} \right] \right\}. \quad (5)$$

These relations provide a linear equation for  $\Delta\rho_{\pm}$ , in terms of the absorptive part of  $\Delta\Sigma^{(0)}(\omega)$  and the cuts of  $f_{\pm}$ . If a solution exists, we are interested in  $\Delta\Sigma(m)$ ,

$$\int_m^{\infty} dn [(n-m)\Delta\tau_+(n) - (n+m)\Delta\tau_-(n)] = \Delta\Sigma^{(0)}(m) + \frac{g^2}{8\pi^2} \int_m^{\infty} dn [(n-m)\Delta\tau_+(n) f_+(n, m, 0) - (n+m)\Delta\tau_-(n) f_-(n, m, 0)], \quad (6)$$

where the  $\Delta\rho_{\pm}$  have been replaced by the  $\Delta\tau_{\pm}$  and, for simplicity, all dependence on  $\mu$  has been dropped: this latter approximation does not lead to an infrared logarithm in  $\Delta\Sigma(m)$ , but introduces errors proportional only to  $(\mu/m)^2$ . Writing

$$f_{\pm}(n, m, 0) = 1 - l_{\pm}(n), \quad (7)$$

where

$$l_{\pm}(n) = \frac{1}{2} \int_0^1 dx \left[ \frac{n \mp (1-x)m}{n \mp m} \right] \ln \left( 1 + \frac{n^2 - m^2}{xm^2} \right), \quad (8)$$

and defining the average value  $l$  of Eq. (8),

$$l \equiv \int_m^{\infty} dn [(n-m)\Delta\tau_+(n)l_+(n) - (n+m)\Delta\tau_-(n)l_-(n)] / \int_m^{\infty} dn [(n-m)\Delta\tau_+(n) - (n+m)\Delta\tau_-(n)], \quad (9)$$

we may use the definition  $\Delta m = -Z\Delta\Sigma(m)$  to obtain

$$\Delta\Sigma(m) = \Delta\Sigma^{(0)}(m) + \frac{g^2}{8\pi^2} (1-l)\Delta\Sigma(m), \quad (10)$$

or

$$\Delta m = \Delta m^{(0)} \left( 1 - \frac{g^2}{8\pi^2} (1-l) \right)^{-1}.$$

For negative  $l$ , or for sufficiently small positive  $l$  ( $l < 0.57$ ), Eq. (10) exhibits the desired change in sign. The value of  $l$  depends on the detailed solution to the model, but it is apparent that the effect depends upon reasonable threshold behavior together with strong damping, at higher energies, of  $\Delta\tau_{\pm}$ .<sup>7</sup> For example, if a solution exists with  $\Delta\tau_+ \approx \Delta\tau_- = \Delta\tau$ , where  $\Delta\tau$  increases from threshold in a smooth way and is essentially cut off for  $n > 2m$ , then we may approximate  $\Delta\tau$  by

$$\Delta\tau(n) \approx 2\Delta\tau_0 \left\{ \left( \frac{n}{m} - 1 \right) \theta \left( \frac{3}{2} - \frac{n}{m} \right) \theta \left( \frac{n}{m} - 1 \right) + \left( 2 - \frac{n}{m} \right) \theta \left( 2 - \frac{n}{m} \right) \theta \left( \frac{n}{m} - \frac{3}{2} \right) \right\},$$

where  $\Delta\tau_0$  is some constant. This, together with the relation

$$\frac{1}{2} \int_0^1 dx (1-x) \ln \left( 1 + \frac{n^2 - m^2}{xm^2} \right) \approx \frac{n}{m} - 1,$$

approximately valid in the region  $2m > n > m$ , leads to the value  $l = \frac{1}{2}$ , satisfying our criterion. The magnitude of Eq. (10) should not be taken seriously; the model itself has only the virtue of providing a simple example of the feedback mechanism.

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<sup>7</sup> Similar damping is essential to the success of the feedback mechanisms of Refs. 1 and 2. Even if a solution with these properties does not exist here, it might still be possible to find a sufficiently small  $l$ ; otherwise, a damping factor, corresponding, e.g., to the use of more realistic  $ps$  vertex functions, must be included.