Analytic-Expansion Unrestricted Hartree-Fock Wave Function for the ${}^{5}D$ State of Fe*

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An accurate analytic-expansion unrestricted Hartree-Fock wave function is reported for the ⁵D state of Fe. The computed values of the magnetic dipole hyperfine-interaction constants are compared with the experimental values obtained by Childs and Goodman and are in good agreement.

I. INTRODUCTION

I N the restricted Hartree-Fock (RHF) method,¹ all the orbitals of a given shell are required to have the same radial dependence; thus closed shells cannot make a contribution to magnetic hyperfine interactions or neutron form factors. Several calculations have been made using an unrestricted Hartree-Fock (UHF) method for open-shell systems, which allows different radial functions for the orbitals of a given shell with different spins²; i.e., $R_{nl\alpha} \neq R_{nl\beta}$. In general, these UHF calculations have given results for hyperfine interactions which are in qualitative and often in fair quantitive agreement with experiment.

In order to compare the predictions of the UHF model of core polarization with the hyperfine interaction constants for the ${}^{5}D$ state of Fe experimentally determined by Childs and Goodman,3 we have computed an accurate UHF wave function for this state of Fe. The UHF Fermi contact interactionc onstant is also compared to the value obtained from an earlier, less accurate, UHF calculation of Watson and Freeman⁴ and is found to be in better agreement with experiment. Wood and Pratt⁵ have also performed an UHF calculation for Fe; however, they used Slater's $\rho^{1/3}$ approximation⁶ for the exchange potential in the Fock operator.

In our calculation of the UHF orbitals, we have used the analytic expansion method. With this method, the UHF orbitals, $\varphi_{nlm_lm_s}$, are given as a linear combination of basis functions, $\chi_{p,lmlm_s}$, by

$$\varphi_{nlm_lm_s}(\mathbf{x}) = \sum_{p} C_{nlm_s, p} \chi_{p, lm_lm_s}(\mathbf{x}), \qquad (1)$$

where nlm_lm_s are the usual one-electron quantum

numbers and x includes space and spin co-ordinates. The basis functions are given by

$$\chi_{p,lm_{l}m_{s}} = R_{lp}(r) Y_{lm}(\theta,\phi) \eta_{m_{s}}.$$
(2)

The $Y_{lm}(\theta,\phi)$ are the usual normalized spherical harmonics; the radial functions R_{lp} are normalized nodeless Slater-type orbitals (STO's), namely,

$$R_{lp}(r) = [(2n_{lp})!]^{-1/2} (2\zeta_{lp})^{n_{lp}+1/2} r^{n_{lp}-1} e^{-\zeta_{lp}r}; \quad (3)$$

and the spin function η_{m_s} is

$$\eta_{m_s=+1/2} = \alpha,$$

$$\eta_{m_s=-1/2} = \beta.$$
(4)

The integer n_{lp} is called the principal quantum number of the basis function and ζ_{lp} the orbital exponent. (In the RHF method the additional requirement that $C_{nlm_s=\pm 1/2, p} \equiv C_{nlm_s=-1/2, p}$ is imposed.) Our UHF procedure is a direct extension of

Roothaan's RHF analytic expansion method for openshell systems.^{7,8} The matrix equations for the orbitals, obtained from arbitrary variation of the C's in Eq. (1), are solved without further approximation. This procedure can be applied to any open-shell system, provided that there is no more than one open subshell for each group $l\alpha$ or $l\beta$. Details of the method and computer program will be given in another paper.

II. RESULTS

A. The UHF Wave Function

The UHF calculation of Fe was performed for the configuration

$$(1s\alpha)^{1}1s\beta)^{1}2s\alpha)^{1}2s\beta)^{1}2p\alpha)^{3}2p\beta)^{3}3s\alpha)^{1}3s\beta)^{1}3p\alpha)^{3}3p\beta)^{3}$$

 $3d\alpha)^{5}3d\beta)^{1}4s\alpha)^{1}4s\beta)^{1};$

the wave function is a single Slater determinant. It is an eigenfunction of L^2 , L_Z , and S_Z ; L=3 and $M_S=2$. The value of M_L is determined by the choice of m_I for the $3d\beta$ orbital, $M_L = m_l$; the UHF equations for the radial parts of the orbitals are independent of the choice of m_i . The wave function is not an eigenfunction of S^2 but it is very nearly one with $S \cong 2$. The value of $\langle S^2 \rangle$ will be given later in this paper.

^{*} Based on work performed under the auspices of the U.S. Atomic Energy Commission.

Atomic Energy Commission. ¹ D. R. Hartree, The Calculation of Atomic Structures (John Wiley & Sons, Inc., New York, 1957). ² See, for example, N. Bessis et al., Phys. Rev. 135, A588 (1964); R. E. Watson and A. J. Freeman, *ibid*. 123, 2027 (1961); D. A. Goodings, *ibid*. 123, 1706 (1961). For a review of the UHF method and extensive references, see A. J. Freeman and R. E. Watson, in Magnetism, edited by G. T. Rado and H. Suhl (Academic Press, Inc., New York, 1965), Vol. IIA. ⁸ W. J. Childs and L. S. Goodman, preceding paper, Phys. Rev. 148, 74 (1966).

R. E. Watson and A. J. Freeman, Phys. Rev. 123, 2027 (1961).

 ⁵ J. H. Wood and G. W. Pratt, Jr., Phys. Rev. 107, 995 (1957).
 ⁶ J. C. Slater, Phys. Rev. 81, 385 (1951).

⁷ C. C. J. Roothaan, Rev. Mod. Phys. 32, 179 (1960).

⁸ C. C. J. Roothaan and P. S. Bagus, *Methods in Computational Physics* (Academic Press Inc., New York, 1963), Vol. II.

	Exponent		1sa	1sß	2sa	2sβ	3sa	3sβ	4sα	4 <i>s</i> β	
$\begin{array}{c} \chi_1(1s) \\ \chi_2(1s) \\ \chi_3(2s) \\ \chi_4(2s) \\ \chi_5(3s) \\ \chi_6(3s) \\ \chi_7(3s) \\ \chi_8(4s) \\ \chi_9(4s) \end{array}$	38.73300 25.87820 21.33300 11.01230 9.50024 6.99020 4.63797 3.92345 1.95370	C_1 C_2 C_3 C_4 C_5 C_6 C_7 C_8 C_9	$\begin{array}{c} 0.02056\\ 0.95050\\ 0.04069\\ -0.00497\\ 0.00654\\ -0.00554\\ 0.00277\\ -0.00101\\ 0.00022\\ 0.00015\end{array}$	$\begin{array}{c} 0.02057\\ 0.95050\\ 0.04071\\ -0.00501\\ 0.00653\\ -0.00554\\ 0.00277\\ -0.00101\\ 0.00022\\ 0.00015\end{array}$	$\begin{array}{c} 0.00644\\ -0.31094\\ -0.18016\\ 1.03806\\ 0.14489\\ 0.00437\\ 0.00415\\ -0.00023\\ 0.00006\\ 0.00002\end{array}$	$\begin{array}{c} 0.00629 \\ -0.31144 \\ -0.18129 \\ 1.04258 \\ 0.14567 \\ -0.00182 \\ 0.00475 \\ -0.00079 \\ 0.00015 \end{array}$	$\begin{array}{c} -0.00290\\ 0.11682\\ 0.07296\\ -0.42236\\ 0.36701\\ 0.77597\\ 0.16592\\ 0.00769\\ -0.0303\end{array}$	$\begin{array}{c} -0.00274\\ 0.11535\\ 0.07273\\ -0.41937\\ -0.30949\\ 0.38109\\ 0.75070\\ 0.18267\\ 0.00796\\ -0.00415\end{array}$	$\begin{array}{r} 0.0003\\ -0.0261\\ -0.0179\\ 0.1001\\ 0.0713\\ -0.0741\\ -0.2498\\ 0.0163\\ 0.5226\\ 0.5395\end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	6766972968
$\chi_{10}(4s) \chi_{11}(4s)$	0.80019	$C_{10} \\ C_{11}$	0.00013	0.00013	0.00002	0.00009	0.00166	0.00175	0.0375	0.1100)2
	Exponent		2 <i>p</i> α	2 <i>þ</i> β	3pa	3 <i>p</i> β		Exponent		$3d\alpha$ $3d\mu$	β
$\begin{array}{c} \chi_1(2p) \\ \chi_2(2p) \\ \chi_3(3p) \\ \chi_4(3p) \\ \chi_5(3p) \\ \chi_6(3p) \end{array}$	$\begin{array}{c} 17.02370\\ 10.09200\\ 6.47648\\ 5.46805\\ 3.45397\\ 2.55178\end{array}$	$C_1 \\ C_2 \\ C_3 \\ C_4 \\ C_5 \\ C_6$	$\begin{array}{c} 0.14292\\ 0.85741\\ 0.03107\\ 0.00174\\ -0.00439\\ 0.00238\end{array}$	$\begin{array}{c} 0.14313\\ 0.86163\\ 0.02122\\ 0.00589\\ -0.00639\\ 0.00270\end{array}$	$\begin{array}{r} -0.04165 \\ -0.35146 \\ -0.12939 \\ 0.69843 \\ 0.48392 \\ 0.06146 \end{array}$	$\begin{array}{r} -0.04019 \\ -0.34681 \\ -0.10041 \\ 0.65699 \\ 0.48070 \\ 0.08245 \end{array}$	$\begin{array}{c} \chi_1(3d) \\ \chi_2(3d) \\ \chi_3(3d) \\ \chi_4(3d) \\ \chi_5(3d) \end{array}$	$\begin{array}{c} 11.50000\\ 4.09542\\ 6.00405\\ 2.61691\\ 1.44880 \end{array}$	$\begin{array}{ccc} C_1 & 0.0 \\ C_2 & 0.2 \\ C_3 & 0.2 \\ C_4 & 0.4 \\ C_5 & 0.1 \end{array}$	3066 0.0293 6060 0.2364 7315 0.2601 6291 0.4252 3241 0.2360	31 44 13 28 06

TABLE I. Basis functions and expansion coefficients for the UHF orbitals of Fe.

The UHF wave function was computed using the basis set obtained by Clementi⁹ for his analytic expansion RHF calculation on the 5D state of Fe. Clementi optimized the values of the exponents of the STO's used in order to minimize the total RHF energy. Automatic exponent optimization procedures which make it possible to obtain very good values of the exponents are provided by the RHF computer program that Clementi used.8 The exponent variation was careful enough and the basis set sufficiently large, 11 s-type, 6 p-type, and 5 d-type STO's, to permit a fairly good representation of the exact RHF orbitals.¹⁰ We believe that this basis set, which very nearly spans the manifold of RHF orbitals, should also nearly span the UHF manifold. Almost all of the differences between the RHF and UHF orbitals can be obtained by allowing the $C_{nlm_s,p}$ to be different.

In Table I we give the symmetry type, principal quantum number, and orbital exponent of the STO's used and the UHF coefficients, $C_{nlm_s,p}$, obtained. The principal quantum number and symmetry type of the STO are given in parentheses after the symbol χ .

The orbitals are orthonormal. Since the basis functions used are normalized but not orthogonal, the condition of orthonormality for the $C_{nlm_{s},p}$ becomes

$$\delta_{nlm_s}^{\dagger} \mathbf{S}_l \mathbf{c}_{n'lm_s} = \delta_{nn'}, \qquad (5)$$

where $\mathbf{c}_{nlm_{\bullet},p}$ is a vector formed from the coefficients $C_{nlm_{\bullet},p}$ and \mathbf{S}_{l} is the overlap matrix of the STO's for symmetry l,

$$S_{lpq} = \int R_{lp}(\mathbf{r}) R_{lq}(\mathbf{r}) \mathbf{r}^2 d\mathbf{r}.$$
 (6)

⁹ E. Clementi, IBM J. Res. Develop. Suppl. 9 (1965).

In Table II we give several properties of the UHF wave function and compare them with values obtained from Clementi's RHF wave function.9,11 We include the total electronic energy E and the virial coefficient V/T, the ratio of the potential to the kinetic energy. The nearness of V/T to the value -2 is a measure of how well the exponents of the STO's have been optimized.¹⁰ The deviation of V/T from -2 is small and about the same for the RHF and UHF functions. This supports our belief that the choice of a large, carefully optimized RHF basis set is satisfactory for a UHF calculation. The value of $\langle S^2 \rangle$ is given to indicate the deviation of the UHF function from an eigenfunction of S^2 ; the RHF function is an eigenfunction of S^2 , (S)(S+1)=6. For each orbital we give the orbital energy ϵ and the expectation values of r and r^2 ; for s orbitals we give the absolute value of the orbital at the origin $|\varphi(0)|$, and for d orbitals we give the expectation value of $1/r^3$.

All units used in Table II are atomic units. The unit of energy is the Hartree, 1 Hartree= 27.211 eV; the unit of length, the Bohr (a_0) , 1 Bohr= 0.52917×10^{-8} cm; and the unit of angular momentum \hbar .

B. The Hyperfine Interaction Constants

The Fermi contact interaction is given by

$$H_c = (8\pi/3)g_{e\mu}g_{I\mu}N\sum_i \delta(\mathbf{r}_i)\mathbf{s}_i \cdot \mathbf{I}, \qquad (7)$$

where $g_e = 2$ and g_I are the electronic and nuclear g factors, respectively; μ_e and μ_N are the electronic and nuclear magnetons (note that $\mu_e = |e| \hbar/2mc$), **I** is the nuclear spin, and \mathbf{s}_i the spin of the *i*th electron. If LS coupling is assumed and if matrix elements are taken

¹⁰ A detailed discussion of the accuracy to which analytic expansion orbitals may represent the exact HF orbitals and techniques of determining this accuracy are given in P. S. Bagus, Argonne National Laboratory Technical Report No. ANL-6959 (unpublished).

¹¹ In order to obtain properties of the RHF wave function not given in Ref. 9, we have recomputed the function with our computer programs using Clementi's basis set. Properties given in Table II but not available from Ref. 9 were obtained from our recalculation.

only between the degenerate states of the electrostatic Hamiltonian (different values of M_L and M_S for a given term), then the contact contribution to the hyperfine splitting may be written as^{12,13}

$$E_{c} = c'\mathbf{I} \cdot \mathbf{S} = c'(g_{J} - 1)\mathbf{I} \cdot \mathbf{J} = c\mathbf{I} \cdot \mathbf{J}, \qquad (8)$$

where

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$$g_{J} = 1 + \frac{(J)(J+1) + (S)(S+1) - (L)(L+1)}{2(J)(J+1)}, \qquad (9)$$

$$c' = \frac{8\pi}{3} \frac{g_{e}\mu_{e}g_{I}\mu_{N}}{S} \langle LS; M_{L}, M_{S} = S |$$

$$\times \sum_{i} \delta(\mathbf{r}_{i}) s_{zi} | LS; M_{L}, M_{S} = S \rangle. \qquad (10)$$

 (\mathbf{z})

If a UHF wave function is used to evaluate the

TABLE II. Properties of the UHF wave function for Fe and a comparison with RHF values. (All values in atomic units; 1 Hartree=27.211 eV, 1 Bohr=0.52917 Å.)

	UHI	F	RHF ^a
$\begin{array}{c} E \\ V/T \\ \langle S^2 \rangle \end{array}$	-1262.4 -2.0 6.0	-1262.4425 - 2.000009 6	
$ert egin{array}{c} \epsilon \ ert arphi(0) \ \langle r angle \ \langle r^2 angle \end{array}$	$\begin{array}{r} 1 s \alpha \\ -261.37445 \\ 73.39271 \\ 0.05911 \\ 0.004686 \end{array}$	1sβ - 261.37484 73.39303 0.05911 0.004686	1s -261.37311 73.39286 0.05911 0.004686
$egin{array}{c} \epsilon \ arphi(0) \ \langle r angle \ \langle r^2 angle \end{array}$	$\begin{array}{c} 2s\alpha \\ -31.99928 \\ 22.21865 \\ 0.2688 \\ 0.08526 \end{array}$	2sβ - 31.87643 22.27583 0.2681 0.08471	2s -31.93623 22.24735 0.2685 0.08498
$ert egin{array}{c} \epsilon \ arphi(0) \ ert \ \langle r angle \ \langle r^2 angle \end{array}$	3sα -4.34862 8.28171 0.8176 0.7700	3sβ 3.99403 8.19521 0.8209 0.7767	3s -4.16972 8.23853 0.8192 0.7733
$ert egin{array}{c} \epsilon \ ert arphi(0) \ ert \ \langle r angle \ \langle r^2 angle \end{array}$	$\begin{array}{r} 4s\alpha \\ -0.27719 \\ 1.89744 \\ 3.1258 \\ 11.3694 \end{array}$	$\begin{array}{r} 4s\beta \\ -0.24221 \\ 1.63679 \\ 3.3927 \\ 13.4276 \end{array}$	4s 0.25832 1.76340 3.2577 12.3672
$\epsilon \ \langle r angle \ \langle r^2 angle$	$2plpha \ -27.47531 \ 0.2367 \ 0.06875$	$2p\beta$ -27.35667 0.2356 0.06802	2p -27.41431 0.2361 0.06838
$\left. egin{array}{c} \epsilon \ \langle r angle \ \langle r^2 angle \end{array} ight.$	$3p\alpha$ -2.95830 0.8617 0.8735	3 <i>þβ</i> 2.52954 0.8704 0.8929	3p - 2.74242 0.8658 0.8827
$\left. egin{array}{c} \epsilon \ \langle r^{-3} angle \ \langle r angle \ \langle r^2 angle \end{array} ight.$	$3dlpha - 0.68803 \ 5.0662 \ 1.0580 \ 1.4449$	$\begin{array}{c} 3deta\ -0.46872\ 4.5522\ 1.1596\ 1.7908 \end{array}$	$3d \\ -0.64710 \\ 4.9780 \\ 1.0734 \\ 1.4948$

 $\ensuremath{\,^{\circ}}$ The RHF values are obtained from Clementi's calculation on Fe; see Refs. 9 and 11.

TABLE III. Values of the contact hyperfine-interaction constants χ and c for the ⁵D state of Fe⁵⁷ (χ in atomic units).

	x	(Mc/sec)	Difference of <i>c</i> with experiment
Childs and Goodman Experimental results ^a		-5.1	
Present UHF Calculation	-0.768	-4.4 ^b	14%
Watson and Freeman UHF Calculation °	-0.59	-3.4 ^b	33%

^a See Ref. 3. ^b Obtained from χ for $g_I = +0.1806$. ^c See Ref. 4.

matrix element in Eq. (10), we obtain

$$c' = \frac{2}{3} g_e \mu_e g_I \mu_N \chi \,, \tag{11}$$

$$\chi = (2\pi/S) \sum_{s \text{ orbitals}} \left(\left| \varphi_{i\alpha}(0) \right|^2 - \left| \varphi_{i\beta}(0) \right|^2 \right). \quad (12)$$

It should be recalled that the UHF wave function used to evaluate the matrix element in Eq. (10) is not an eigenfunction of S^2 . An eigenfunction of S^2 may be obtained by projection from the UHF wave function. However, the value of χ calculated with the projected wave function is likely to be significantly different from the χ calculated directly with the UHF function; and it is not certain that the use of the projected χ is to bepreferred to the use of the UHF χ .¹⁴ Heine¹⁵ has suggested an alternative procedure for forming an eigenfunction of S^2 which should produce only a small change in the value of χ . In this procedure, the UHF function is augmented only with determinants which are small. Because of the computational difficulties in obtaining these eigenfunctions of S^2 , we only give the value of χ for the UHF function.

In Table III we give the value of χ obtained from our values of $|\varphi(0)|$, listed in Table II, together with the χ obtained by Watson and Freeman⁴ from their earlier UHF calculation. Using the value¹⁶ of $g_I = +0.1806$ for Fe^{57} , we also give in Table III the values of c obtained from our calculation and Watson and Freeman's calculation together with the value obtained from experimental data by Childs and Goodman.³ Our value of the contact hfs constant is in good agreement with experiment and in rather better agreement than that obtained from Watson and Freeman.

The remaining contribution to the magnetic dipole hyperfine interaction is

$$H_{l} = g_{e}g_{I}\mu_{e}\mu_{I} \left(\sum_{i} \frac{l_{i} \cdot \mathbf{I}}{r_{i}^{3}} + \frac{3(\mathbf{s}_{i} \cdot \mathbf{r})(\mathbf{r}_{i} \cdot \mathbf{I}) - (\mathbf{r}_{i} \cdot \mathbf{r}_{i})\mathbf{s}_{i} \cdot \mathbf{I}}{r_{i}^{5}} \right). \quad (13)$$

¹⁴ For a review of these problems see Sec. III.5 of the article by A. J. Freeman and R. E. Watson in *Magnetism*, cited in Ref. 2.
¹⁵ V. Heine, Phys. Rev. 107, 1002 (1957).
¹⁶ G. W. Ludwig and H. H. Woodbury, Phys. Rev. 117, 1286 (1960).

¹² R. E. Trees, Phys. Rev. 92, 308 (1953).

¹³ A. Abragam, J. Horowitz, and M. H. L. Pryce, Proc. Roy. Soc. (London) **A230**, 169 (1955).

^{(1960).}

With the same assumptions as before about LS coupling and the matrix elements taken and using the notation and expressions given by Trees,¹² we obtain for the ⁵D state of Fe

$$E_l = r_J a_l \mathbf{I} \cdot \mathbf{J} \tag{14}$$

$$a_l = g_e g_I \mu_e \mu_I \langle 3d\beta | 1/r^3 | 3d\beta \rangle, \qquad (15)$$

where r_J depends only on *J*, *L*, and *S*. Using the value of $\langle 3d\beta | 1/r^3 | 3d\beta \rangle$ given in Table II, we obtain $a_l = 78.4$ Mc/sec. The value obtained from experimental data by Childs and Goodman³ is $a_l = 74.8$ Mc/sec; the error of the UHF value is 5%.

Because the UHF wave function is not an eigenfunction of S^2 , equivalent expressions for E_l will give different results when evaluated with an UHF function.¹⁷ The results will involve different combinations of $\langle 3d\alpha | 1/r^3 | 3d\alpha \rangle$ and $\langle 3d\beta | 1/r^3 | 3d\beta \rangle$ in the expression for E_l . However, since these integrals have nearly the

¹⁷ D. A. Goodings, Phys. Rev. 123, 1706 (1961).

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same value (see Table II), the numerical results for E_1 will be nearly the same, and we will not pursue the matter further.

III. CONCLUSION

Good agreement between the magnetic dipole hyperfine-interaction constants obtained from an UHF calculation and from experimental data⁴ has been found. In particular, the agreement for the Fermi contact interaction constant (14%) supports the model of exchange core-polarization used in the UHF method.

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Excited-State Mixing in the Optical Pumping of Alkali-Metal Vapors

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The effect of excited-state mixing on the optical pumping of alkali-metal vapors is considered for the case where the energy separation between the ${}^{2}P_{1/2}$ and ${}^{2}P_{3/2}$ first excited states is relatively large, allowing the isolation of mixing effects within a particular J level. Two models for the mixing process are proposed, one a random reorientation of the total electronic angular momentum \mathbf{J} , and the other a reorientation of \mathbf{J} subject to the selection rule $\Delta m_J = 0, \pm 2$. The probabilities for mixing transitions among J-state sublevels have been calculated for an alkali atom of nuclear spin $\frac{3}{2}$, and can be used to calculate the optical-pumping transition probabilities for a vapor subject to any degree of excited-state mixing. Several possible experiments suggested by these calculations are discussed.

I. INTRODUCTION

THE primary purpose of this paper is to analyze the role that excited-state interactions play in the optical pumping of alkali-metal vapors. In an experiment typical of those to be considered,¹ a beam of circularly polarized $D_1({}^2S_{1/2} \leftrightarrow {}^2P_{1/2})$ photons is passed through a cell containing the vapor to be studied. Since the prevailing selection rules are $\Delta m_F = \pm 1$ for absorption, and $\Delta m_F = \pm 1$, 0 for re-emission, the vapor is gradually pumped from ground-state sublevels of low to

high (F, m_F) , with all atoms ultimately being forced into the highest sublevel if there is no ground-state relaxation. In practice, of course, relaxation exists, being caused by collisions of the alkali atoms with buffer gas molecules, impurity atoms, or the walls of the cell. Collisional interactions can also disrupt the optical pumping process by transferring or mixing atoms among excited-state sublevels prior to de-excitation. The equilibrium population distribution in the ground state is thus determined both by the strengths and characteristics of the relaxation mechanisms involved, and by the various perturbations that the alkali atoms may suffer while in the excited state.

In order to calculate the equilibrium electronic and nuclear spin polarizations, and to describe the transient

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¹Comprehensive surveys of optical pumping techniques can be found in : G. V. Skrotskii and T. G. Izyumova, Usp. Fiz. Nauk 73, 423 (1961) [English transl., Soviet Phys.—Uspekhi 4, 177 (1961)]; and R. Benumof, Am. J. Phys. 33, 151 (1965).