Elastic Scattering of Neutrons from Carbon and Oxygen in the Energy Range 3.0 to 4.7 MeV*

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The elastic scattering of neutrons from oxygen and carbon was studied in the neutron bombarding energy range 3.0 to 4.7 MeV. Angular distributions were obtained by the method of recoiling gas atoms in a proportional counter. Data were taken over the entire energy range in approximately 25-keV intervals, and in smaller intervals near certain resonances. Differential cross sections for C(n,n)C were obtained relative to the total n-p cross section. Similarly, O(n,n)O differential cross sections as well as the $O^{16}(n,\alpha)C^{13}$ total cross section were determined relative to the derived C(n,n)C cross sections. A phase-shift search routine was used to find acceptable least-squares fits to a Legendre-polynomial decomposition of the elastic-scattering results. Energy-level parameters were obtained for five states in O^{17} and are given as follows (in the order: resonance energy relative to the O^{17} ground state in MeV, spin and parity, and dimensionless reduced width): $7.24(\frac{3}{2}^+, 0.15)$, $7.69(\frac{3}{2}^-, 0.07)$, $7.70(\frac{7}{2}^-, <0.03)$, $7.83(\frac{1}{2}^-, 0.04)$, and $8.07(\frac{3}{2}^-, 0.02)$. Level parameters were also determined for two states in C^{13} : $8.27(\frac{3}{2}^+, 0.45)$ and $8.87(\frac{1}{2}^-, 0.03)$.

I. INTRODUCTION

HE primary object of this work was to study the spectroscopy of O¹⁷ in the excitation energy range from 7.0–8.6 MeV by means of the $O^{16}(n,n)O^{16}$ reaction. Total neutron cross-section measurements for O¹⁶ have indicated the presence of broad and narrow overlapping levels in this energy region.^{1,2} The usually difficult problem of determining scattering parameters is simplified, as the entrance channel is composed of a neutron (spin $\frac{1}{2}$) and O¹⁶(spin 0) so that the channel spin s can have only the value $s=\frac{1}{2}$. For a given partial wave of relative orbital angular momentum l, the total angular momentum J of the $O^{16} + n$ system is limited to the values $J = l \pm \frac{1}{2}$ and the parity is given by $(-1)^{l}$. Successful analysis of the differential cross sections in terms of the elastic-scattering phase shifts $\delta(l,J)$ will yield the following level parameters: resonance energy (E_R) , spin (J), parity (π) , and width (Γ) .

In the present investigation $O^{16}(n,n)O^{16}$ differential cross sections were measured over the incident neutron energy range from 3.07-4.73 MeV (excitation energy E_x of O¹⁷ between 7.0 and 8.6 MeV). Measurements were made in intervals of about 25 keV with an incident neutron energy spread ΔE_n , of about 25 keV over the entire energy range, and in the vicinity of certain resonances in 10-keV intervals with $\Delta E_n \approx 18$ keV.

In the course of this work the $O^{16}(n,\alpha)C^{13}$ total cross section and $C^{12}(n,n)C^{12}$ differential cross section (E_x of $\mathrm{C^{13}}$ between 7.78 and 9.32 MeV) were also measured. A phase-shift analysis was performed on the $O^{16}(n,n)O^{16}$ and $C^{12}(n,n)C^{12}$ data and the derived level parameters are presented.

At the time the present work was initiated, previous measurements of the $O^{16}(n,n)O^{16}$ differential cross section in the energy range of this experiment included the work of Baldinger, Huber, and Proctor,³ Hunzinger and Huber,⁴ Sayers,⁵ Phillips,⁶ and Bostrom, Morgan, Prud'homme, and Sattar.7 Results of Refs. 4, 6, and 7 appear in a graphical compilation of neutron cross sections.8 However, only Baldinger et al.,3 extracted phase shifts. They identified three fairly broad levels above $E_x = 7.0 \text{ MeV}$; at $E_n = 3.33 \text{ MeV}$ ($E_x = 7.28 \text{ MeV}$, $\frac{3}{2}$, $\Gamma = 220$ keV), $E_n = 3.80$ MeV (7.73 MeV, $\frac{3}{2}$, 800 keV), and $E_n = 4.40$ MeV (8.29 MeV, $\frac{1}{2}$, 280 keV).

Because the various groups of experimenters that have studied the $O^{16}(n,n)O^{16}$ reaction took data in energy steps too coarse to study levels of intermediate width, there was a clear need for a more detailed and systematic study of this reaction.

More recently, Fowler and Johnson⁹ have measured $O^{16}(n,n)O^{16}$ differential cross sections at a number of incident neutron energies between 2.25 and 3.90 MeV and extracted phase shifts. Their results were in fair agreement with those of Baldinger et al.,3 regarding the broad $\frac{3}{2}^+$ and $\frac{3}{2}^-$ resonances and also identified two other resonances: $E_n = 3.77$ MeV (7.69 MeV, $\frac{5}{2}$, 25 keV) and $E_n = 3.82$ MeV (7.73 MeV, $\frac{3}{2}$, 50 keV).

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B. J. Highes and R. D. Schwartz, reamon cross Sections, Brookhaven National Laboratory Report BNL-325 2nd ed. (Government Printing Office, Washington, D. C., 1958).
 ² J. R. Stehn, M. D. Goldberg, B. A. Magurno, and R. Wiener-Chasman, Neutron Cross Sections, Brookhaven National Labora-

tory Report BNL-325, 2nd ed., Suppl. No. 2 (Office of Technical Services, Department of Commerce, Washington, D. C., 1964), Vol. 1.

⁸ E. Baldinger, P. Huber, and W. G. Proctor, Helv. Phys. Acta 25, 142 (1952). ⁴ W. Hunzinger and P. Huber, Helv. Phys. Acta 35, 351 (1962).

⁵ A. Sayres, Bull. Am. Phys. Soc. 6, 237 (1961). ⁶ D. D. Phillips (unpublished). See Ref. 16.

⁷ N. A. Bostrom, I. L. Morgan, J. T. Prud'homme, and A. R. Sattar, Wright Air Development Center Report WADC-TR-57-446, 1957 (unpublished).

⁸ M. D. Goldberg, V. M. May, and J. R. Stehn, Brookhaven National Laboratory Report BNL-400, 2nd ed. (Office of Tech-nical Services, Department of Commerce, Washington, D. C., 1963), Vol. 1.

⁹ J. L. Fowler and C. H. Johnson (private communication).

II. EXPERIMENTAL PROCEDURE

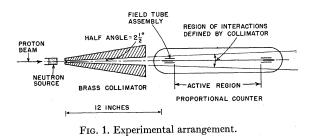
A. Method

The neutron elastic-scattering measurements were made by means of the recoiling atom method,¹⁰ chosen because of the relatively high rate of data accumulation associated with it. The scattering nuclei were present as part of the gas filling of a proportional counter in which the energy of the recoiling nucleus was measured. Application of energy and momentum conservation laws to the elastic scattering process gives the relation,¹¹

$$E_r = 4mM/(m+M)^2 E_{n\frac{1}{2}}(1 - \cos\theta_{\rm c.m.}), \qquad (1)$$

where E_r = kinetic energy of the recoiling nucleus, E_n = incident neutron energy, $\theta_{e.m.}$ = neutron scattering angle in the c.m. system, m = mass of the neutron, and M = mass of the recoil nucleus. From (1) it can be shown that the neutron angular distribution is proportional to the recoil energy distribution. The latter is proportional to the corresponding pulse-height distribution from the detector if W, the average energy loss per electron-ion pair produced in the counter gas, is independent of the recoil energy. This last condition is at least approximately true in practice. For an arbitrary W = W(E) the relationship between the differential cross section, $\sigma(\Omega)$, and the pulse-height distribution, N(X), is given by $\sigma(\Omega) \propto N(X)/W(X)$ where X represents the pulse height corresponding to a given initial recoil energy.

The experimental arrangement is shown in Fig. 1. Neutrons were produced by the reaction $T(p,n)He^3$ (Q=-0.764 MeV). Protons from the Columbia University Van de Graaff electrostatic accelerator were incident on a gaseous T_2 target separated from the vacuum of the drift tube by a 0.05-mil thick nickel window. The collimated neutron beam was coaxial with the proportional counter and the proton beam. Recoil spectra were taken in three series of runs, the counter being filled with carbon dioxide (CO₂) in the first series, propane (C₃H₈) in the second, and methane (CH₄) in the third. In each series, data were obtained at the same neutron energies and for the same experimental geometry. Data and backgrounds were determined by recording alternately spectra from the proportional



¹⁰ H. H. Barschall and M. H. Kanner, Phys. Rev. 58, 590 (1940). ¹¹ S. C. Curran, in *Encyclopedia of Physics*, edited by S. Flügge (Springer-Verlag, Berlin, 1958), Vol. 45, pp. 174 ff.

counter with the brass collimator aperture open and then blocked by a brass insert. Background counts due to neutrons produced in the target cell in the absence of tritium gas were negligible. A "long counter,"¹² using a He³-filled proportional counter as the thermal neutron detector, positioned one meter from the target and 90° to the proton beam, served as a monitor.

In the energy region under investigation the following conditions prevail: (1) Neutron capture by O and C is negligible compared to elastic scattering^{1,2}; (2) elastic scattering is the only other open channel in the C¹²+*n* interaction, and (3) the O¹⁶(n,α)C¹³ (Q = -2.215 MeV) cross section is negligible compared to elastic scattering for $E_n < 4.1$ MeV.¹³ Furthermore, for incident neutron energies greater than 3.1 MeV the group of pulses from the O¹⁶(n,α)C¹³ reaction has greater pulse height than pulses due to the most energetic recoils in the CO₂ filling.

B. Apparatus

1. Proportional Counter

Theoretical and practical discussions on proportional counters are well covered in the literature.^{11,14} The counter used in this experiment had cylindrical geometry. A 4-mil stainless steel anode wire was positioned axially with respect to a 2.75-in. i.d., 0.125-in. wall stainless steel cylindrical cathode. The $9\frac{7}{8}$ -in. long active volume was defined by a pair of $\frac{1}{2}$ -in. o.d. field tubes concentric with the center wire, which served to keep the electric field radial over the entire active volume. An important feature of the construction was that all electrical feed-throughs were located off-axis so as not to be in the path of the axially collimated neutron beam.

Commercially available high-purity gases were used for the counter fillings with no further purification. The suppliers listed the following purities: CO₂, 99.9998%; C₃H₈, 99.99 mole%; CH₄, 99.68 mole%. A measure of the resolution was given by noting the shape of the high-energy cutoff of the recoil pulse-height distributions. The resolution was found to be about 3% for proton recoils of 3.07-MeV energy, 10% for 0.87-MeV carbon recoils in both CH₄ and C₃H₈, and 11 and 13%, respectively, for 0.95-MeV carbon ions and 0.74-MeV oxygen ions in CO₂. The average energy expended per electron ion pair in stopping 0.87-MeV carbon ions in CH₄ was observed to be about 30% greater than the same quantity for 3.07-MeV protons, which is in rough agreement with previous observations.¹⁵

 ¹² A. O. Hanson and J. L. McKibben, Phys. Rev. 72, 673 (1947).
 ¹³ R. B. Walton, J. D. Clement, and F. Boreli, Phys. Rev. 107, 1065 (1957).

¹⁴ B. B. Rossi and H. H. Staub, *Ionization Chambers and Counters* (McGraw-Hill Book Company, Inc., New York, 1949), Chapt. 4, 7.

¹⁵ A. T. G. Ferguson, in *Fast Neutron Physics*, edited by J. B. Marion and J. L. Fowler (Interscience Publishers Inc., New York, 1960), Part I, p. 179.

2. Neutron Energy Calibration

Precautions were taken to determine accurately the neutron energy for reasons to be mentioned below. Proton energies were inferred by the standard procedure of monitoring the field of the 90° momentum-analyzing magnet by means of the nuclear magnetic resonance frequency f of hydrogen calibrated at the $Li^{7}(p,n)Be^{7}$ threshold at $E_p = 1.8811$ MeV. The calibration constant k, in the expression,

$$E_{p} = k f^{2} (1 - k f^{2} / 2mc^{2})$$

where mc^2 = rest energy of the proton, which includes the first-order relativistic correction, was assumed to be constant over the energy range studied. For the analyzing magnet used, past experience has shown that this calculated E_p would be at most greater than the true E_p by about 5 keV at the uppermost energy used.

Tables¹⁶ of neutron energies as a function of proton energy and neutron emission angle for the $T(p,n)He^{3}$ reaction were used to obtain the average neutron energy E_n after having taken into account proton energy losses in traversing the nickel foil and the tritium, and the variation of neutron energy in the solid angle of the collimated neutron beam. The thickness of the nickel window was ascertained by noting its effect in shifting the $T(p,n)He^3$ threshold energy from the published value¹⁷ of 1.019 MeV. The measured shift was 135 ± 5 keV as compared with a calculated ionization loss of 129 keV for a 0.05-mil nickel thickness.¹⁸ The neutron energies that will be quoted in this paper may be considered accurate to ± 10 keV on an absolute scale. Energy differences are accurate to ± 5 keV. The mean energy spread of the neutrons was obtained by folding into the energy loss of protons in traversing the tritium a width of 12 keV representing the combined effects of all other energy-spreading factors present. The latter was determined empirically from the observed widths of narrow resonances in $O^{16}(n,\alpha)C^{13}$ when a very thin tritium target was employed.

3. Electronics

Pulses from the proportional counter were first fed into a low-noise pre-amplifier, then into a pulse shaper and finally into one-half of a 512-channel pulse-height analyzer. The entire system was linear to $\pm 1\%$. A running dead-time correction was accomplished by normalizing to the total number of monitor counts that arrived at times when the multichannel analyzer was not "busy."

C. Handling of Data

The reduction of the data was treated in the following manner. Dead-time correction and background subtraction were performed for each run and the results were machine-plotted. Figures 2 and 3 illustrate such results and show pulse-height distributions for CO₂ and $C_{3}H_{8}$, respectively, at $E_{n} = 3.286$ MeV. The number of proton recoils of energy greater than the upper bias of Fig. 3 were counted in a scaler. The proton recoil con-

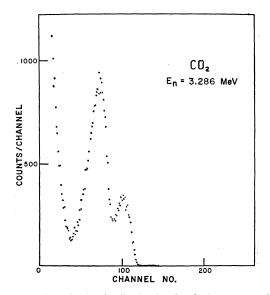


FIG. 2. The pulse-height distribution for the interaction of neutrons with CO₂ observed at 3.286 MeV.

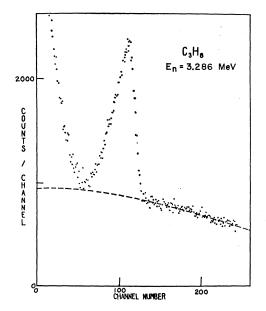


FIG. 3. The pulse-height distribution for the interaction of neutrons with $C_{2}H_{8}$ observed at 3.286 MeV. The higher-energy part has been biased-out electronically. The assumed protonrecoil contribution is shown by the dashed curve.

¹⁶ J. B. Marion and B. Allen, Shell Development Company, Houston, Texas, 1955 (unpublished).
¹⁷ J. E. Brolley, Jr., and J. L. Fowler, in *Fast Neutron Physics*, edited by J. B. Marion and J. L. Fowler (Interscience Publishers Inc., New York, 1960), Part I, p. 73.
¹⁸ J. B. Marion, *1960 Nuclear Data Tables* (U. S. Government Printing Office, Washington, D. C., 1960), Part 3, p. 16.

tribution to the C_3H_8 data was subtracted leaving the pulse-height distribution of the carbon recoils. In Fig. 3. the proton recoil contribution is shown as the dashed curve. Its shape is not quite flat because of wall effect. The $C^{12}(n,n)C^{12}$ scattering is backward-peaked at this energy and its endpoint appears near channel 120. The portion of the dashed curve in the region of the carbon recoils was estimated in a consistent manner on all the C₃H₈ plots so as to join smoothly with the higher channel distribution and take wall and end effects into account. The CH₄ recoil distributions were similarly treated. A check on this procedure was to compare the resulting carbon-recoil distributions with the best published^{19,20} $C^{12}(n,n)C^{12}$ angular distributions at energies where they were available. Figure 4(a) shows the resulting carbon recoil distribution at $E_n = 3.286$ MeV.

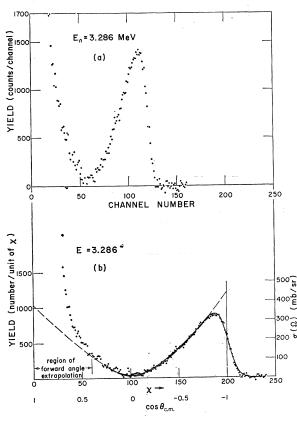


FIG. 4(a): The carbon-recoil contribution to the pulse-height distribution for the interaction of neutrons with C3H8 at 3.286 MeV. (b): The redistribution of the carbon-recoil spectrum of Fig. 4(a) onto a $x=100(1-\cos\theta_{\rm c.m.})$ scale. The dashed curve represents this distribution corrected for counter resolution, extrapolated into the forward-angle region, and put on an absolute cross section scale to give a plot of the C(n,n)C differential cross section at 3.286 MeV. The solid curve results from distorting the dashed curve with an 8% FWHM resolution function.

Next, the carbon recoil distribution was converted to an angular distribution, the energy dependence of $\langle W \rangle$ being taken into account. $\langle W \rangle$ was taken to be $\langle W(E) \rangle$ $=W(E_0)+(E-E_0)d\langle W\rangle/dE$. E_0 was some reference energy. Values of $d\langle W \rangle/dE$ for the various recoil ions were obtained from plots of recoil distribution endpoints versus neutron energy. The absolute scale for the differential cross section was obtained by normalizing the estimated number of p+n collisions that occurred within the sensitive volume of the counter to the known n-p total cross section.²¹ Also, the angular distributions were corrected for the finite resolution of the counter. This was especially important near the end point of the recoil distribution (i.e., near $\theta = 180^{\circ}$). The large errors in the forward-angle data due to uncompensated low-energy background were avoided by extrapolating the angular distribution curve in this region so that the integrated cross section was equal to the $C^{12}+n$ total cross-section results of Fossan *et al.*²² and Bockelman et al.23 Figure 4(b) illustrates these steps for $C^{12}(n,n)C^{12}$ at $E_n = 3.286$ MeV.

Extraction of the $O^{16}(n,n)O^{16}$ angular distributions from the CO₂ recoil data was treated in a somewhat similar way. First, the CO₂ recoil pulse-height distribution was redistributed so that the carbon-recoil part was on a $\cos\theta$ scale. Normalization was made by matching the spectrum of carbon recoils of greater energy than the oxygen recoil endpoint to the corresponding part of the previously determined $\sigma(\Omega)[C^{12}(n,n)C^{12}]$ curve, taking into account the distortion due to the finite counter resolution. The oxygen recoil distribution, which remained after the carbon recoil contribution was subtracted from the CO₂ recoil spectrum, was then converted to a differential cross-section curve. The $O^{16}(n,\alpha)C^{13}$ total cross section was obtained from the integrated number of pulses larger than those due to the highest energy carbon recoils. Correction to the (Ω) results was made for counter resolution. A forward angle extrapolation was performed such that $\int \sigma(\Omega) [O^{16}(n,n)O^{16}] d\Omega =$ the total elastic-scattering cross section. The latter was taken to be the Wisconsin total neutron cross-section results^{22,23} minus σ_{α} [O¹⁶(n, α)C¹³]. It was necessary to increase the energy scale of the work of Fossan et al.,22 by 10 keV to conform with the energies we found for the narrow states in the vicinity of E_n^{\perp} between 4.5 and 4.7 MeV. Figures 5(a) and (b) illustrate the procedure for extracting $\sigma(\Omega) [O^{16}(n,n)O^{16}]$ at $E_n = 3.286$ MeV.

The $O^{16}(n,n)O^{16}$ and $C^{12}(n,n)C^{12}$ differential crosssection curves were Fourier analyzed as a sum of

¹⁹ J. E. Wills, Jr., J. K. Bair, H. D. Cohn, and H. B. Willard, Phys. Rev. **109**, 891 (1958). ²⁰ R. W. Meier, P. Scherrer, and G. Trumpy. Helv. Phys. Acta

R. W. Meier, P. Scherrer, and G. Trumpy, Helv. Phys. Acta 27, 577 (1954).

²¹ J. L. Gammel, in Fast Neutron Physics, edited by J. B. Marion and J. L. Fowler (Interscience Publishers Inc., New York, 1963), Part II, p. 2185.

²² D. B. Fossan, R. L. Walter, W. E. Wilson, and H. H. Bar-

schall, Phys. Rev. 123, 209 (1961). ²⁸ C. K. Bockelman, D. W. Miller, R. K. Adair, and H. H. Barschall, Phys. Rev. 84, 69 (1951).

600

200

800

60

40

20

50

(m.b/sr)

۲ (D)

/ield∕unit of

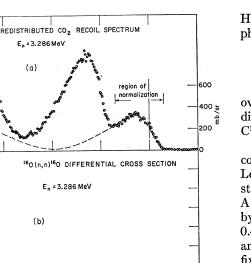


FIG. 5(a): The redistributed spectrum for the interaction of neutrons with CO₂ at 3.286 MeV. The carbon-recoil part is on a $x=100(1-\cos\theta_{\rm c.m.})$ scale. The dashed curve represents the result of distorting the ideal C(n,n)C angular distribution with a 10% FWHM resolution function. (b): The O¹⁶(n,n)O¹⁶ angular distribution on an absolute cross-section scale before (points) and after (dashed curve) forward-angle extrapolation and correction for resolution.

x٠

150

200

Legendre polynomials through the sixth order:

$$\sigma(\Omega) = \sum_{L=0}^{6} A_L P_L(\cos\theta_{\rm c.m.}).$$
 (2)

We used Price's method²⁴ to obtain the expansion coefficients. This Fourier analysis served a double purpose. First, it is a concise form in which to express the angular distributions. Second, it is a useful intermediate step in carrying out the phase-shift analysis (see Sec. V). The cutoff of the series (2) at L=6 was necessitated by the effective experimental angular resolution. It corresponded to ignoring contributions from partial waves of $l\geq 4$, which was justifiable on physical grounds since neutron penetrabilities for $l\geq 4$ waves are quite small at these energies

III. ERRORS

The various steps in the procedure for reduction of data described in Sec. II introduced a variety of interacting systematic uncertainties. The net effects of all the uncertainties fell into three main categories: Statistical errors, normalization errors, and all others. How these uncertainties are taken into account in the phase-shift analysis is discussed in Sec. V.

A. Statistical Errors

Statistical errors contributed an average uncertainty over back angles of from 5 to 9 mb/sr in the $O^{16}(n,n)O^{16}$ differential cross section and from 8 to 13 mb/sr in the $C^{12}(n,n)C^{12}$ differential cross section.

As applied to the present analysis, Price's method could not yield quantitatively the uncertainties in the Legendre polynomial expansion coefficients arising from statistical fluctuations in the pulse-height distributions. A fair indication of these errors was, however, found by performing a least-squares fit to the data in the region $0.4 < \cos\theta < -0.8$, with each channel given unit weight and with the total cross section value (i.e., A_0) held fixed. Then the statistical uncertainty in a coefficient, defined as the deviation in that coefficient alone needed to double the value of X^2 with respect to its minimum value, was calculated. The errors obtained are shown in Table I.

B. Normalization Errors

Normalization errors were of two kinds. The first group included those that directly affected $\sigma(\Omega)$ at back angles but left the total elastic-scattering cross section σ_n (elastic) unchanged. Examples are the normalization of $\sigma(\Omega)[C^{12}(n,n)C^{12}]$ to $\sigma_T(n-p)$ through the integrated proton-recoil counts in the C3H8 runs, and the normalization of $\sigma(\Omega)[O^{16}(n,n)O^{16}]$ to the backangle C+n scattering. As a result of this kind of uncertainty, $\sigma(\Omega)$ would be in error at back angles by a constant fraction; and the computed $\sigma(\Omega)$ at forward angles would be in error in the opposite direction in a more complicated way. It turned out that the errors in the A coefficients could be related to all the A coefficients through a linear matrix with a multiplicative constant, ϵ , denoting the size of the error. For C+n scattering $|\epsilon|$ was about 8% and for O+n scattering $|\epsilon|$ was about 17%.

The second kind of normalization error is associated with uncertainties in the normalization to σ_n (elastic) values and affects only the computed $\sigma(\Omega)$ at forward angles. Again uncertainties were propagated into the *A* coefficients in a way that could be expressed as a linear matrix with a multiplicative constant, ϵ . In the case of C¹²(*n*,*n*)C¹², $|\epsilon| \sim 3\%$ (the quoted statistical error in the published total neutron cross section,²²

TABLE I. Statistical errors in A coefficients.

		R	ange of	error (±	=) in ml	o/sr	
Reaction	A_0	A_1	A_2	A_3	A_4	A_5	A_6
$O^{16}(n,n)O^{16}$	5-9	13-22	14-23	17-28	18-31	19-32	23-39
$C^{12}(n,n)C^{12}$	8–13	20-32	23-35	28-42	30-45	32-48	37-58

²⁴ P. C. Price, Phil. Mag. 45, 237 (1954).

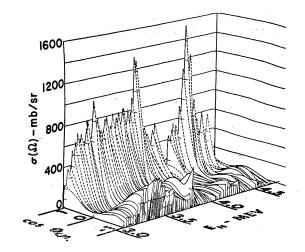
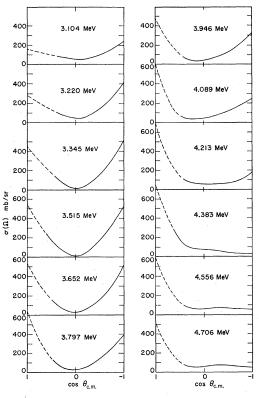


FIG. 6. $O^{16}(n,n)O^{16}$ differential cross sections as a function of c.m. scattering angle and neutron energy.

except for E_n between 4.20 and 4.25 MeV, where $|\epsilon| \sim 6\%$. In the O+*n* scattering case the situation was somewhat more complex owing to the existence of several important relatively narrow resonances and the presence of the O¹⁶(*n*, α)C¹³ reaction channel. Here



C12 (n,n)C12 DIFFERENTIAL CROSS SECTION

FIG. 7. Typical differential cross-section curves for $C^{12}(n,n)C^{12}$ at various energies over the energy range studied. Correction for counter resolution has been included. The dashed line portions represent the forward-angle extrapolations.

we can write $|\epsilon| = |\epsilon(\sigma_n, \text{ statistical})| + |\epsilon(\text{reaction})| + |\Delta\sigma_n(10 \text{ keV})/\sigma_T|$, where the first and second terms are self-explanatory. The third term represents the effect of an assumed 10-keV uncertainty in our energy calibration relative to that of the total neutron cross-section work.^{22,23} These terms had roughly the following values: $|\epsilon(\sigma_n, \text{ statistical})| \sim 3\%$; $|\epsilon(\text{reaction})| \simeq 0.17\sigma_{\alpha} \times [O^{16}(n,\alpha)C^{13}]/\sigma_T \sim 3\%$ maximum, but was usually negligible; and $|\Delta\sigma_n(10 \text{ keV})/\sigma_T|$ which could be as high as 30% on the steep slopes of certain resonances, but usually was < 10%. Because of the forward-angle errors in the C¹²(n,n)C¹² results propagated into the O¹⁶(n,n)O¹⁶ results.

C. Miscellaneous Errors

Sources of various other uncertainties together with their maximum contributions appear in Table II.

IV. CROSS-SECTION RESULTS

The $\sigma(\Omega)[O^{16}(n,n)O^{16}]$ results are displayed in Fig. 6 as a function of energy and the derived Fourier coefficients, A_L , of Eq. (2) are listed in Table III.

Table IV lists the A coefficients for $\sigma(\Omega)[C^{12}(n,n)C^{12}]$. In Fig. 7 are shown $\sigma(\Omega)[C^{12}(n,n)C^{12}]$ results at a number of energies to illustrate how the angular distributions vary over the range of energies studied. The solid-line portions of the curves in Figs. 6 and 7 represent visually drawn curves through the data and include correction for counter resolution, while the dashedline portions represent the forward-angle extrapolations.

The measured $O^{16}(n,\alpha)C^{13}$ total cross sections are plotted as the circles and crosses of Fig. 8 and have not been corrected for wall or end effects. Such a correction would be no more than 4%, and is small compared

TABLE II. Miscellaneous sources of error.

Sources uncertair		Maximum expected contribution
 Impurities Multiple scatt Pile up of puls Wall and end 	es	1% 1% 1% 1%
5. Proton-recoil e		Mainly, ~ 20 mb/sr in the coefficient A_1 in C+n scattering. Also, a carry-over of ~ 4 mb/sr in A_2 in O+n scattering.
6. Finite resolution	on of counter	Negligible for A_0 , increasing with the order of A coefficient to about 20% for A_6 .
7. Energy spread	l of neutron	Averaged each A coefficient over the energy spread.
8. Quadratic app	roximation angle extra-	Individual uncertainties in any A coefficient depended in part on values of the other A coef- ficients. The effect was negli- gible for the low-order coeffi- cients, but could be appreci- able for high-order coefficients.

/sr)		
A_4	A_5	A_{6}
	/sr) A4	

TABLE III. A Coefficients for neutron elastic scattering from oxygen.																																																																																													ί.	L	ı	ı	ı	1	n	r	1	3	г	e	(ŗ	ρ	1	v	7	x	02	0	1	L	r	r	3	1	r	f	1	ŗ	Q	ı	r	'n	i	r	3	e	t	t1	ιt	a	a	с	ŝ	5	;	С	с	ic	ti	t	s
---	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	--	----	---	---	---	---	---	---	---	---	---	---	---	---	---	---	---	---	---	---	----	---	---	---	---	---	---	---	---	---	---	---	---	---	---	----	---	---	---	---	---	----	----	---	---	---	---	---	---	---	---	----	----	---	---

Energy spread of neutron beam -25 keV 3.166 115 54 66 21 -34 -1 7 3.167 126 33 119 71 -16 -6 7 3.245 200 76 303 140 -54 -1 3.256 -210 74 327 155 -54 -1 3.167 151 110 222 132 -24 -2 117 3278 28 178 3343 130 -58 -3 -3 3.169 171 126 268 148 -16 -1 2 3.220 173 302 122 -223 -22					I ABLE J	ш. <i>А</i> С	oemcien	is for neu	cron elastic s		ig from	oxygen.				
Energy spread of neutron beam -25 keV 3.166 115 54 66 21 - 34 - 17 3.167 125 76 139 77 - 16 7 3225 210 74 327 155 - 54 3.167 151 71 126 268 148 - 16 - 1 2 3.245 200 76 303 140 - 54 - 1 3.250 127 143 776 139 77 - 28 - 2 17 3.250 127 123 27 155 - 36 - 34 3.220 157 102 220 143 - 43 - 14 18 3.245 201 74 327 155 - 36 - 46 - 3 3.245 201 74 327 155 - 36 - 46 - 3 3.245 201 74 327 155 - 36 - 46 - 3 3.245 201 74 327 155 - 36 - 46 - 3 3.245 201 74 327 155 - 36 - 46 - 3 3.245 201 75 30 124 - 67 - 5 3.245 201 79 310 140 - 44 - 14 18 3.246 221 72 330 124 - 67 - 5 3.346 221 92 330 126 - 34 - 14 4 3.341 223 190 368 136 - 41 - 10 3.346 224 105 376 94 - 9 - 12 4 3.346 224 105 376 94 - 9 - 12 4 3.346 224 105 376 94 - 9 - 12 4 3.346 224 1105 376 94 - 9 - 12 4 3.346 224 113 331 18 46 35 - 15 3.346 224 1105 376 94 - 9 - 12 4 3.346 324 117 331 84 152 - 15 3.346 224 113 331 71 224 70 98 140 - 7 - 7 - 16 3.346 324 152 206 71 13 28 140 - 13 3.346 234 117 318 184 52 - 164 - 31 3.347 324 117 331 84 17 - 185 3.346 224 113 327 162 206 71 2 - 14 - 1 3.346 326 197 305 102 - 24 - 9 3.513 223 117 247 50 5 - 3 - 6 3.346 236 197 305 102 - 24 - 9 3.513 223 198 226 244 128 22 - 71 - 4 3.346 236 197 305 102 - 24 - 9 3.513 223 198 226 244 128 22 - 9 3.546 224 197 305 102 - 145 3.547 203 224 2197 126 144 28 - 18 - 15 3.647 249 229 217 126 146 - 9 3.548 224 157 125 34 47 - 3 - 6 3.349 234 214 127 233 106 12 - 16 - 5 3.547 241 222 217 128 34 69 1 - 1426 3.548 228 246 24 128 22 - 9 3.548 248 256 171 325 44 - 136 3.549 238 246 244 128 22 - 9 3.668 238 246 244 128 22 - 9 3.668 238 240 244 128 22 - 9 3.668 238 240 244 128 22 - 9 3.664 239 209 106 0 - 24 - 28 3.664 239 209 106 0 - 24 - 28 3.674 239 204 126 105 106 - 136 3.649 238 240 127 106 144 - 1711 4.458 100 189 160 161 - 3 - 9 -4.458 100 189 160 161 - 3 3.777 284 277 383 329 171 - 9383 3.8				1					E			4		/sr)		
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	(MeV)	A_0	A_1	A_2	A_3	A4	A_{5}	A 6	(MeV)	A_0	A_1	A_2	A_3	A_4	A_5	A 6
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$		Ener	gy spre	ad of ne	utron b	eam=25	5 keV				gy sprea		utron b	eam = 18	8 keV	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$			54 83		21 71	-34 -16						303 327				2 19
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.127	135	76	137	77	-28	-3	14	3.267	216	103	347	156	-46	-4	10
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$														-45 - 58		7 21
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.196	171	126	268	148	-16	-1	2	3.282	223	68	330	114	-67	-5	8
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$										229 233						64
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.266	214	76	339	130	-38	-12	11	3.312	236	.95	378	136	-29	-15	-0^{6}
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$																$-0 \\ 4$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.349	241	105	376	94	-9	-12	4		242		335	84	23	-15	-15
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$						-12 - 7	$-7 \\ -4$			259			84			$-11 \\ -23$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$										244						$-23 \\ -18 \\ -15$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$						26	-1			237		268				4
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.488	234	154	256		2	-11			236	197		102			$-13 \\ -21$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		232	197		94					236	181	270	66	9		-19
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-28			236			56			$-\frac{8}{1}$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.585	236	221	253	106	12	16	-5	3.521	237	176	240	36	-22	-28	-9
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$																$-20 \\ -2$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	3.654	237	213	197	107	-4	-24	-9	3.664	239	209	187	106	16	-35	-2
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.675 3.702	233 224					$-26 \\ -32$			239 237				11 30		$-12 \\ -6$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.724	232	207	183	177	41	-45	9	3.696	236	251	231	168	38	-44	-13
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-72 - 53			235						$-\frac{4}{7}$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			226			41	78			236	198		158	35	- 56	-5
$\begin{array}{cccccccccccccccccccccccccccccccccccc$						5	13			255	229		250		-83	-2
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			189			4		-7		284	277				-93	$-11 \\ -10$
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	3.925	177	149	76	73	-24	-14	5	3.786	353	275	612	395			-60
$\begin{array}{cccccccccccccccccccccccccccccccccccc$						-2	-9 -8			$319 \\ 245$					45	$-44 \\ -50$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.995	166	194	111	69	-14	-6	-0	3.816	211	170	229	132	-32	39	-27
$\begin{array}{cccccccccccccccccccccccccccccccccccc$						$-\frac{2}{-2}$										$-25 \\ -19$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	4.067	104	113	122	93	25	-2	-17	4.513	101	114	122	138	31	-15	-12
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$					17	-13										$-5 \\ -8$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$											81			43	-9	-14
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	4.188	150	86	186	47	-17	-6	-1		127	105		152	69	-32	$-39 \\ -22$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							$-\frac{2}{5}$									$-34 \\ -33$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	4.263	113	59	78	30	17	-1	-15	4.592	98	74	37	115	53	-24	-20
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								-13 -30				51 58			-17 -29	$-31 \\ -33$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	4.337	219	312	341	179	60	1	-25	4.622	102	41	30	89	39	-38	-6
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	4.383				357 305		-10								$-30 \\ -34$	$-3 \\ 1$
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	4.409	171	208	202	204	28	-28	-3	4.655	104	108	104	154	65	-15	-29
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	4.458	135	194	201	195	59	-19	-22	4.005	90	104	100	139	00	-4	-14
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.482	96		126	124		-19	-4								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.531	99	43	64	68	0	-13	-11								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.557						-23 -22									
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.609	116	87	66	118	72	-27	-22								
4.682 84 36 -5 45 14 -17 -11	$4.034 \\ 4.658$						-28 -11									
	4.682	84	36	-5	45	. 14	-17	-11								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$				-11												
									<u> </u>							

 TABLE IV. A Coefficients for neutron elastic scattering from carbon.

$\begin{array}{c c c c c c c c c c c c c c c c c c c $					A_L (mb/	(er)		
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	E (MeV)	A_0	A_1				A_5	A_{6}
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		Ener	du enro.	d of m	utron h	0 m - 2 f	koV	-
$ \begin{array}{ccccccccccccccccccccccccccccccccccc$	3 068							1
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			-26		-18	-13		7
$ \begin{array}{ccccccccccccccccccccccccccccccccccc$					-23			5
$ \begin{array}{ccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.220		-36	210	-33	-24	4	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$					-11		2	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			-29 -14					
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.307	175	-15	283	-25	-42		
$\begin{array}{cccccccccccccccccccccccccccccccccccc$					-8			
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			-0 -14		-29 -30			
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			9					21
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$			4	329	2	-9		
$\begin{array}{cccccccccccccccccccccccccccccccccccc$					23			
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-9	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.847		47	288	68	5	-3	-4
$\begin{array}{cccccccccccccccccccccccccccccccccccc$						5		-2
$\begin{array}{cccccccccccccccccccccccccccccccccccc$						13	-6	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	3.946	156	33	245	53	16	-10	-11
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-10	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								-15 -13
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-9	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-11 -21	-0 -10
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.310	165	258	204	110	47	-15	-0
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$			253					
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-17	-26
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.408	148	217	182	109	35	-10	-18
$\begin{array}{cccccccccccccccccccccccccccccccccccc$							-13	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								-9
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.506	135	221	236	169	73	-8	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$								-15 -10
$\begin{array}{cccccccccccccccccccccccccccccccccccc$				167	126		-21	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	4.606	120		156	112	36	-21	-11
$\begin{array}{cccccccccccccccccccccccccccccccccccc$				94 156	109			
4.706 114 133 140 113 32 -16 -14			145	157	117	41	-14	-19
4.729 112 123 132 101 29 -12 -12	4.706		133					
	4.729	112	123	132	101	49	12	- 12

with the over-all uncertainty in σ_{α} [O¹⁶(n,α)C¹³] of about 20%.

V. EXTRACTION OF SCATTERING PARAMETERS

A. Phase-Shift Analysis

The elastic-scattering differential cross section for the interaction between a neutron $(\text{spin } \frac{1}{2})$ and a spin-0 nucleus for the case of unpolarized neutrons and a spininsensitive detector is given by²⁵

$$\sigma(\Omega) = \frac{1}{2k^2} \{ \sum_{l=0}^{\infty} \left[(l+1) \sin \delta_l^+ \exp(i\delta_l^+) + l \sin \delta_l^- \exp(i\delta_l^-) \right] P_l(\cos\theta) |^2 + \sin^2\theta \sum_{l=0}^{\infty} \left[\sin \delta_l^+ \exp(i\delta_l^+) - \sin \delta_l^- \exp(i\delta_l^-) \right] P_l'(\cos\theta) |^2 \}, \quad (3)$$

where $\delta_l^{\pm} = \delta(l, J = l \pm \frac{1}{2})$ denotes the (l, J) phase shift, $P_l(\cos\theta) = l$ th-order Legendre polynomial, $P_l'(\cos\theta)$ $= dP_l(\cos\theta)/d(\cos\theta)$, and k = wave number in the center-of-mass system. If elastic scattering is the only open channel, all the phase shifts are real. If nonelastic channels are also important, the phase shifts are complex quantities, and the problem of determining them becomes considerably more difficult.

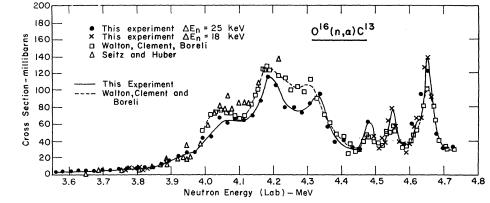
Pisent²⁶ has recently discussed the problems attending phase-shift analysis of angular distributions in elastic scattering. First, there are the "inherent" ambiguities,²⁷ for which $\sigma(\Omega)$ is unchanged: (a) If the signs of all the phase shifts are simultaneously reversed, and (b) if $\delta(l,l+\frac{1}{2})$ and $\delta(l,l-\frac{1}{2})$ are interchanged for all l simultaneously. In addition, there may be "accidental" ambiguities, i.e., two or more "independent" sets of phase shifts not related by the transformations mentioned above, yet satisfying $\sigma(\Omega)$ within experimental error at a given energy.

The following requirements, however, are usually enough to relieve all the ambiguities. (1) The phase shifts should vary with energy in a "physical" way. That is, (a) they should be continuous functions of energy; (b) there should be no unusual variation of too many phase shifts together to account for resonances; and (c) over a single resonance, one phase shift should increase with increasing energy with a total increase of approximately 180°, the other phase shifts remaining fairly constant. (2) Nonresonant phase shifts should be given approximately by the "potential" scattering phase shifts; i.e., phase shifts calculated for scattering from an appropriate single-particle potential. (3) The phase shifts should join with those obtained at lower energies, if available.

²⁵ C. L. Critchfeld and D. C. Dodder, Phys. Rev. **76**, 602 (1949).

⁽¹⁹⁴⁹⁾.
 ²⁶ G. Pisent, Helv. Phys. Acta 36, 248 (1963).
 ²⁷ J. M. Blatt and L. C. Biedenharn, Rev. Mod. Phys. 24, 258 (1952).

FIG. 8. The $O^{16}(n,\alpha)C^{13}$ total cross section plotted as a function of neutron energy. Comparison with the results of Walton, Clement, and Boreli (Ref. 13) and Seitz and Huber (Ref. 40) is shown.



In principle, if n is the number of phase shifts involved, an *n*-dimensional search for all independent phase-shift solutions would be desirable at each energy. One would then apply the various criteria of reasonable physical behavior in order to single out the best solution. However, such a complete search program was not practical in this instance because it would have required an excessive amount of computer time.

In the present work, the extraction of phase shifts from the angular distributions was handled in the following way. First, we note that the right-hand side of (3) can be expressed as the sum of Legendre



$$\sigma(\Omega) = 1/2k^2 \sum_L B_L P_L(\cos\theta).$$
(4)

In (4), the *B* coefficients contain the entire dependence of $\sigma(\Omega)$ on the phase shifts, whereas, the k^{-2} and $P_L(\cos\theta)$ factors contain the energy and angular dependences, respectively. Values of the *B* coefficients up to $L_{\text{max}} = 6$ were obtained from the A coefficients of Sec. II and were plotted versus E_n . Smooth curves were drawn visually through the points such that possible resonances

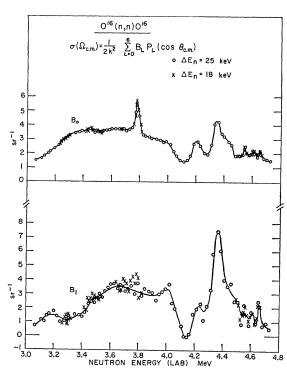


FIG. 9. The coefficients B_0 and B_1 in the expansion

$$\sigma(\Omega)[O^{16}(n,n)O^{16}] = (2k^2)^{-1} \sum_{L=0}^{6} B_L P_L(\cos\theta_{\rm c.m.})$$

plotted as a function of neutron energy. The solid curves serve to guide the eye and were used for the interpolation of the values of these coefficients in the phase-shift analysis.

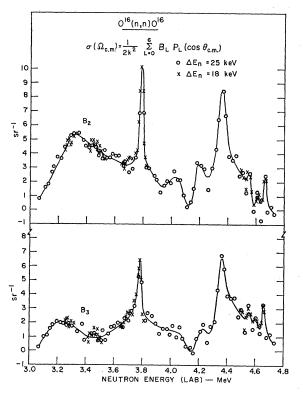


FIG. 10. The coefficients B_2 and B_3 in the expansion

$$\sigma(\Omega)[O^{16}(n,n)O^{16}] = (2k^2)^{-1} \sum_{L=0}^{6} B_L P_L(\cos\theta_{\rm c.m.})$$

plotted as a function of neutron energy. The solid curves serve to guide the eye and were used for the interpolation of the values of these coefficients in the phase-shift analysis.

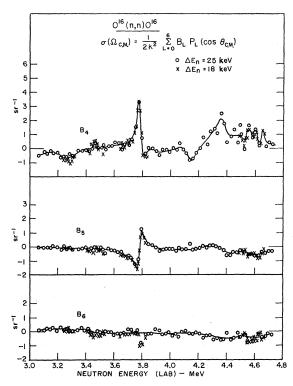


FIG. 11. The coefficients B_4 , B_5 , and B_6 in the expansion, $\sigma(\Omega)[O^{16}(n,n)O^{16}] = (2k^2)^{-1} \sum_{L=0}^{6} B_L P_L(\cos\theta_{c.m.}),$

plotted as a function of neutron energy. The solid curves serve to guide the eye and were used for the interpolation of the values of these coefficients in the phase-shift analysis.

of width $\Gamma \ll$ neutron energy spread were ignored. Figures 9-11 and Fig. 12 are plots of the B coefficients for $O^{16}(n,n)O^{16}$ and $C^{12}(n,n)C^{12}$, respectively, with the curves used for them superposed. Next, the B values were read from the curves at a selected energy in a lowenergy region where all the B's were varying relatively smoothly with energy. Trial sets of phase shifts were chosen from previously published results in the case^{19,20} of $C^{12}(n,n)C^{12}$. For $O^{16}(n,n)O^{16}$ the calculated singleparticle potential phase shifts of Kolesov et al.,²⁸ were used. The B coefficients and the trial set of phase shifts were fed into a computer search code that (a) varied singly or in pairs any number of the phase shifts, each over 180° (a full cycle, since the phase shifts enter as $\exp[i2\delta(l,J)]$, (b) calculated for each set of phase shifts, $\delta'(l,J)$, so generated, the quantity χ^2 , defined as

$$\chi^2 = \sum_{L=0}^{6} 2/2L + 1(B_L' - B_L)^2,$$

where $B_L = L$ th experimentally determined coefficient and $B_L' = L$ th coefficient constructed from the $\delta'(l,J)$; (c) kept the $\delta'(l,J)$ sets that gave the 21 lowest values of χ^2 . This last procedure involved a gradient descent method in which all the phase shifts of a set were allowed to vary simultaneously.

The 21 resulting good-fit sets of phase shifts were grouped into "independent" sets and new searches were performed at the same energy starting with these sets to generate, if possible, other independent sets. For example, six independent sets were found for $O^{16}(n,n)O^{16}$ at the first energy selected, $E_n=3.30$ MeV, and two for $C^{12}(n,n)C^{12}$ at $E_n=3.10$ MeV. Each resulting independent solution together with the appropriate *B* coefficient then served as input to similar searches at a second energy chosen reasonably close to the starting energy.

This procedure was repeated over the energy range as far as it could successfully be applied. For $C^{12}(n,n)C^{12}$ this included the entire energy range studied; for $O^{16}(n,n)O^{16}$ it was $3.07 < E_n < 4.20$ MeV. In each case, imposition of the abovementioned requirements of "physical" behavior on the phase shifts narrowed down the number of possible solutions to just one.

Some indication of the uncertainties in the phase shifts was obtained by going back and using the search routine at a given energy, and taking as the initial set of phase shifts their most likely values. $\sigma(\Omega)$ was found in many instances to be much less sensitive to certain simultaneous variations of the phase shifts than to individual variation of the phase shifts about their "best" values.

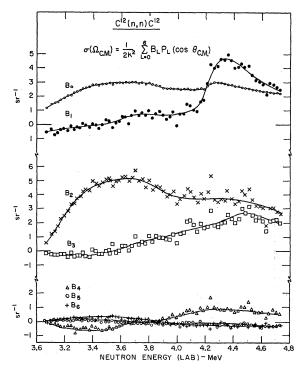


FIG. 12. The B-coefficients in the expansion,

$$\sigma(\Omega)[C^{12}(n,n)C^{12}] = (2k^2)^{-1} \sum_{L=0}^{6} B_L P_L(\cos\theta_{c.m.})$$

plotted as a function of neutron energy. The solid curves serve to guide the eye and were used for the interpolation of the values of these coefficients in the phase-shift analysis.

²⁸ V. E. Kolesov, V. P. Korotkikh, and V. G. Malashkina, Izvest. Akad. Nauk, SSSR Ser. Fiz. **27**, 903 (1963).

In addition, the 25-keV wide resonance in O¹⁶(n,n)O¹⁶ at $E_n = 3.78$ MeV, which was too narrow to be analyzed by the method described above, was studied in the following way. Differential cross sections were artifically generated at ten energies over the resonance for the three cases: each of $\delta(2,\frac{5}{2})$, $\delta(3,\frac{5}{2})$, or $\delta(3,\frac{7}{2})$ alone changing over the resonance according to the dispersion formula,

143

$$\tan^{-1}\delta(l,J) = \Gamma/2/E_R - E$$

with the other two set equal to zero. It was not necessary to consider l>3 as these would result in unreasonably large reduced widths for this level. Interpolated values were used for the remaining phase shifts. The angular distributions so generated were compared with the observed differential cross sections. Figure 13 illustrates this comparison at four of the energies used. Note that no energy averaging was included in the calcluted curves. These comparisons clearly indicated an assignment of $J^{\pi} = \frac{\pi}{2}$.

The resulting phase shifts for $O^{16}(n,n)O^{16}$ are shown in Fig. 14 and those for $C^{12}(n,n)C^{12}$ in Fig. 15. In these figures, the points represent the best values of the phase shifts at energies where they were determined. The error bars are but a rough indication of uncertainties at selected energies. Smooth curves have been drawn to indicate how the phase shifts probably vary with energy.

B. Resonance Parameters

Because of the experimental analytical uncertainties in this work, it was decided simply to apply the onelevel approximation to the appropriate phase shifts at resonances whether or not this approximation could otherwise be justified. This corresponds to using Breit-Wigner forms for the resonances. Therefore, uncertainties in the resulting values for the reduced width γ^2 might be as great as 30%. Consistent with this approxi-

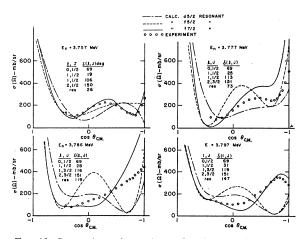


FIG. 13. Comparison of experimental and artifically generated angular distributions for $O^{16}(n,n)O^{16}$ in the vicinity of the resonance at $E_n=3.78$ MeV. The points represent the experimental results for a neutron energy spread of 18 keV and include correction for counter resolution. The various curves were calculated under the assumptions that the resonance is $d_{5/2}$, $f_{5/2}$, or $f_{7/2}$ and has a width of 22 keV.

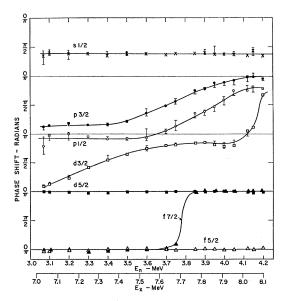


FIG. 14. The phase shifts for $O^{16}(n,n)O^{16}$ plotted as a function of neutron energy and excitation energy.

mation, channels other than elastic scattering were ignored.

The phase shift $\delta(l,J)$ in the vicinity of an isolated (l,J) resonance corresponding to the unbound state λ , may be written as

$$\delta(l,J) = \beta_{\lambda} + \varphi(l,J),$$

where λ denotes all the quantum number of the state, $\varphi(l,J)$ is the nonresonant part of the phase shift and

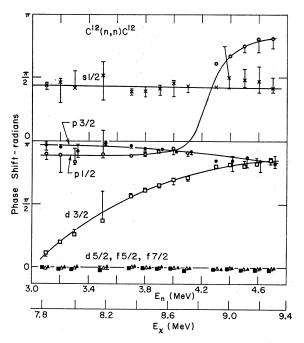


FIG. 15. The phase shifts for $C^{12}(n,n)C^{12}$ plotted as a function of neutron energy and excitation energy.

TABLE V. Leve	l parameters for	states of O ¹⁷ .
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$E_n(\text{MeV})$	$E_{R^{\mathbf{a}}}(\mathbf{MeV})$	$E_{\lambda}(\text{MeV})$	J^{π}	$\Gamma_{e.m.}(\text{keV})$	Γ_n/Γ	$\gamma_{\lambda n^2}({ m MeV})$	$ heta_{\lambda n}{}^2$	$\varphi_{\lambda}(\mathrm{deg})$
3.29 ± 0.02 3.77 ± 0.02	7.24 7.69	7.02 7.67	$\frac{3+2}{2}$	400 ± 30 360 ± 30	>0.99 >0.99	0.25 0.12	0.15 0.07	-20 + 17
3.78 ± 0.01	7.70	7.60	27- 2-	<23	>0.99	< 0.05	< 0.03	0
3.92 ± 0.02 4.175 ± 0.01	7.83 8.07	$7.82 \\ 8.05$	1- 2- 3+	$245 \pm 30 \\ 75 \pm 20$	>0.95 >0.90	0.07 0.03	$\begin{array}{c} 0.04 \\ 0.02 \end{array}$	$-12 \\ -23$

* $E_R = (16/17)E_n + 4.142$ MeV.

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 β_{λ} is the resonant part of the phase shift and is given by

$$B_{\lambda} = \tan^{-1}(\frac{1}{2}\Gamma_{\lambda}/E_R - E).$$

 Γ_{λ} is the width of the level and E_R is the resonance energy, defined as the energy corresponding to $\beta_{\lambda} = \pi/2$.

Estimates of the various $\varphi(l,J)$ were made by inspection of the $\delta(l,J)$ versus *E* curves, guided by a reasonable extrapolation of the single-particle phase shifts of Kolesov *et al.*²⁸ The width Γ was taken as

or

$$E(\beta=135^{\circ})-E(\beta=45^{\circ})$$

$$2|E(\beta=90^{\circ}\pm45^{\circ})-E(\beta=90^{\circ})|,$$

if the latter was more practical.

First-order expressions relating the experimental parameters to R-matrix parameters²⁹ are given by (c.m. system is implied).

$$\begin{split} &\Gamma_{\lambda} = 2 P_{I} \gamma_{\lambda}^{2}, \\ &\theta_{\lambda} = \gamma_{\lambda}^{2} / \left(\hbar^{2} / m a_{c}^{2} \right), \\ &E_{R} = E_{\lambda} + \Delta_{\lambda}, \\ &\Delta_{\lambda} = -S_{I} \gamma_{\lambda}^{2}, \end{split}$$

 P_l is the penetrability, γ_{λ}^2 the reduced width, θ_{λ}^2 the dimensionless reduced width, a_c the channel radius, m the reduced mass, E_{λ} the characteristic energy, Δ_{λ} the level shift, and S_l the shift factor. θ_{λ}^2 may be compared to θ_c^2 , the single-particle reduced width for a given channel. $\theta_c^2 = \frac{3}{2}$ times the quantity usually referred to

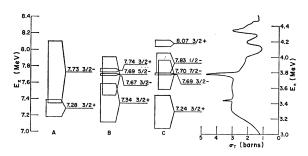


FIG. 16. A graphical presentation of level parameters assigned to states of O^{17} , according to (A) Baldinger, Huber, and Proctor (Ref. 3), (B) Fowler and Johnson (Ref. 9), and (C) this work. The total neutron cross section of oxygen is shown on the same energy scale.

²⁸ A. M. Lane and R. G. Thomas, Rev. Mod. Phys. 30, 257 (1958).

as the Wigner limit on the reduced width. A reasonable but rough estimate of the magnitude of θ_c^{2} is 0.5.³⁰

VI. CONCLUSION

Application of the phase-shift analysis program to the $O^{16}(n,n)O^{16}$ differential cross sections, as described in Sec. V, was successful in assigning level parameters to five levels of O^{17} in the excitation energy range between 7.2 and 8.1 MeV. These are listed in Table V. Above an excitation energy of about 8.1 MeV either the levels were too narrow, or the level structure was too complex to be handled by the phase-shift analysis program. All the levels considered in this study are $T=\frac{1}{2}$ since the lowest $T=\frac{3}{2}$ level in O^{17} would probably correspond to the N¹⁷ ground state and have an energy of about 11.2 MeV.³¹

Figure 16 is a comparison of O¹⁷ level parameters found in this and in previous studies. The broad, overlapping $\frac{3}{2}^+$, 7.24 MeV and $\frac{3}{2}^-$, 7.69-MeV levels found in this work are in fair agreement with the results of Baldinger et al.,3 and Fowler and Johnson.9 The assignment of $\frac{7}{2}$ to the 7.70-MeV level is in disagreement with the interpretation of Fowler and Johnson, who found that a combination of a $\frac{5}{2}$ and a $\frac{3}{2}$ resonance best suited their data in this energy region. It may be mentioned, however, that the differential cross-section results of this work and the work of Fowler and Johnson are in good agreement. Other assignments to this level have been $J = \frac{5}{2}$ by Fossan *et al.*,²² and $J \ge \frac{5}{2}$ by Walton, Clement, and Boreli.¹³ The basis of the $\frac{7}{2}$ assignment from the present work was the comparison of the various computed angular distributions with those measured in the vicinity of the resonance, as typified by Fig. 13. Furthermore, the behavior of the coefficient B_5 in Fig. 11 near $E_n = 3.78$ MeV is well explained by the interference between the $\frac{7}{2}$ and $\frac{3}{2}$ phase shifts. In order to account similarly for such behavior in B_5 , a $\frac{5}{2}$ resonance would have to be accompanied by a considerable negative amount of $J \geq \frac{5}{2}$, positive parity phase shift, which is not required at other energies.

The $\frac{1}{2}$ -, $E_x = 7.83$ -MeV level reported here may perhaps be identified with the $J = \frac{1}{2}$, 110-keV wide level at 7.94 MeV observed in C¹³(α ,n)O¹⁶ work¹³ and later

³⁰ F. C. Barker, Nucl. Phys. 28, 96 (1961).

³¹ T. Lauritsen and F. Ajzenberg-Selove, in *American Institute* of *Physics Handbook* (McGraw-Hill Book Company, Inc., New York, 1963), 2nd ed., Chap. 8, p. 79.

E_n^{a} (MeV)	$E_{R}^{a}(MeV)$	E_{λ} (MeV)	J^{π}	$\Gamma_{c.m.}(keV)$	Γ_n/Γ	$\gamma_{\lambda n}({ m MeV})$	$ heta_{\lambda n}^2$	$\varphi_{\lambda}(\mathrm{deg})$
3.60 ± 0.05 4.25 ± 0.02	8.27 8.87	7.96 8.86	$\frac{32}{12}$ + $\frac{12}{2}$ -	$1050 \pm 100 \\ 180 \pm 50$	>0.99 >0.99	0.68 ^b 0.06	0.35 ^b 0.03	-22^{0}

TABLE VI. Level parameters for states of C¹³.

^a E_n is the laboratory neutron energy at resonance; i.e., $\beta = 90^\circ$. $E_R = (12/13)E_n + 4.947$ MeV. ^b If the more exact expression for γ^2 is used, which takes into account the first-order energy dependence of the shift factor, these values are increased by about 30%. They represent about half the single-particle values.

assigned negative parity by means of the $C^{13}(\alpha,\alpha)C^{13}$ interaction.³² It is possible that the true values for the width and energy of this level may be distorted in $C^{13} + \alpha$ studies because of the rapid increase of the Coulomb barrier penetrability with energy in the $C^{13} + \alpha$ channel. In any case our $\frac{3}{2}$ + assignment to the 8.07-MeV level agrees with the combined results of $C^{13} + \alpha$ work.^{13,32,33}

Although the $O^{16}(n,n)O^{16}$ phase shifts could not be obtained in this work above $E_n \sim 4.2 \text{ MeV}(E_x=8.1$ MeV), some comments may be made regarding the 4.3- to 4.5-MeV region, where total neutron crosssection work indicates the existence of one or more resonances of width \geq 50 keV. Baldinger *et al.*³ found a $\frac{1}{2}$, $\Gamma_{e.m.} = 250$ -keV level at $E_n = 4.40$ MeV. Fossan et al.²² found a $J=\frac{3}{2}$, $\Gamma_{e.m.}=60$ keV, resonance at $E_n = 4.32$ MeV from their total neutron cross-section work on O¹⁶. Studies^{13,33} of C¹³(α,n)O¹⁶ indicated the presence of a $J = \frac{3}{2}$, $\Gamma_{c.m.} = 60$ keV, resonance at an energy corresponding to E_n of about 4.30 MeV. Resonances were observed in the $C^{13}(\alpha, \alpha)C^{13}$ work of Barnes et al.,³² at energies corresponding to lab neutron energies for $O^{16}+n$ of 4.32, 4.45, and 4.50 MeV and given spinparity assignments of $\frac{3}{2}$, probably $\frac{1}{2}$, and $\frac{5}{2}$, respectively. We note that the observed peaking of the coefficient B_3 near 4.36 MeV in Fig. 10 would seem to require either two overlapping $\frac{3}{2}$ resonances of opposite parity or a resonance of $J \geq \frac{5}{2}$ interfering with a nonnegligible phase shift of opposite parity.

If the principle of charge symmetry holds, then there should be a correspondence between states of O¹⁷ and its mirror nuclues F¹⁷. Level parameters of states of F¹⁷ have been obtained by Salisbury and Richards³⁴ from 2.6 to 7.4 MeV of excitation and by Dangle et al.³⁵ from 7.3 to 9.1 MeV. Reference 34 shows excellent correspondences between levels in O¹⁷ and F¹⁷ up to about 6.5 MeV. Table VII lists possible analog states in the energy region considered in the present work.

Only two levels in C¹³ were observed between 7.78 and 9.32 MeV excitation energy. They are listed in Table VI. The broad $\frac{3}{2}$, 8.3-MeV level has previously been identified by Wills et al.,19 and by Meier, Scherrer, and Trumpy.²⁰ Its dimensionless reduced width, θ^2 has

a value of about one-half the single-particle value.³⁶ We have found the $J=\frac{1}{2}$, 8.85-MeV level, previously observed in total cross-section work,^{22,37} to have negative parity. These states correspond to the $\frac{3}{2}$, 8.1 MeV and $\frac{1}{2}$ 8.90-MeV levels, respectively, in N¹³, which have been investigated by Shute et al.36 Theoretical studies by Kurath³⁸ and Barker³⁰ on positive parity states in C¹³ predict a state corresponding to the $\frac{3}{2}$, 8.3-MeV level. Kurath and Lawson³⁹ have studied the negative parity states of mass 13 nuclei and predict a $\frac{1}{2}$ level which, for a reasonable value of their parameter a/K of 5.3, can be identified with the above-mentioned $\frac{1}{2}^{-}$ levels in C¹³ and N¹³.

The $O^{16}(n,\alpha)C^{13}$ total cross section measured in this experiment is compared to the measurements of Walton. Clement, and Boreli¹³ and Seitz and Huber⁴⁰ in Fig. 8. Agreement is reasonably good. At worst, the results of this work lie below those of Walton, Clement, and Boreli by about 30%, which is within the combined experimental errors. To about the same extent the present results are in agreement with the cross section¹³ of the inverse reaction, $C^{13}(\alpha, n)O^{16}$, through the principle of detailed balance.

Theoretical calculations in O¹⁷ have so far been limited to identifying the single-particle states,⁴¹ that is, $\frac{5}{2}$, g.s.; $\frac{1}{2}$, 0.871 MeV; $\frac{3}{2}$, 5.07 MeV; the indication by Christy and Fowler⁴² that the $\frac{1}{2}$, 3.06-MeV state has the configuration $(1p)^{-3}(2s,1d)^4$, and the suggestion by Harvey⁴³ that the $\frac{1}{2}$, 5.94-MeV level is the lowest level in a $K = \frac{1}{2}^{-}$ band.

TABLE VII. Corresponding energy levels of O¹⁷ and F¹⁷.

	O17			F17 a	
J^{π}	$E_{x^{\mathbf{b}}}$ (MeV)	$ heta^2$	J^{π}	$E_{x^{\mathbf{b}}}$ (MeV)	θ^2
$\frac{\frac{3}{2}+c}{\frac{3}{2}-d}$ $\frac{\frac{3}{2}-d}{\frac{7}{2}-}$	7.24 7.69 7.70	$0.15 \\ 0.07 \\ < 0.03$	$\frac{3}{2}+$ $\frac{3}{2}-$ $\frac{7}{2}-$	7.21 8.22 7.55	0.19 0.09 0.013

^a Values taken from Refs. 34 and 35. Here, $θ^2$ is the quantity $γ^2/(3\hbar^2/2\mu a)$ for elastic scattering multiplied by $\frac{3}{2}$, ^b E_x has been taken to be equal to E_R as defined in the text. ^c Correspondence made in Ref. 34. ^d Correspondence made in Ref. 35.

³⁶ G. G. Shute, D. Robson, V. R. McKenna, and A. T. Berztiss, ⁵⁰ G. G. Snute, D. Robson, V. R. McKenna, and A. T. Berztiss, Nucl. Phys. 37, 535 (1962).
⁵⁷ K. Tsukada and T. Fuse, J. Phys. Soc. Japan 15, 1994 (1960).
⁵⁸ D. Kurath, Phys. Rev. 101, 216 (1956).
⁵⁹ D. Kurath and R. D. Lawson, Nucl. Phys. 23, 5 (1961).
⁴⁰ J. Seitz and P. Huber, Helv. Phys. Acta 28, 227 (1955).
⁴¹ A. M. Lane, Rev. Mod. Phys. 32, 519 (1960).
⁴² R. F. Christy and W. A. Fowler, Phys. Rev. 96, 851 (1954).
⁴³ M. Harvey, Phys. Letters. 3, 209 (1963).

²² B. K. Barnes, R. L. Steele, T. A. Belote, and J. R. Risser, Bull. Am. Phys. Soc. 8, 125 (1963).

³³ J. P. Schiffer, A. A. Kraus, Jr., and J. R. Risser, Phys. Rev. 105, 1811 (1957). S. R. Salisbury and H. T. Richards, Phys. Rev. 126, 2147

^{(1962).} ³⁵ R. L. Dangle, L. D. Oppliger, and G. Hardie, Phys. Rev.

In view of the present existence of a large body of spectroscopic information in the O¹⁷ and F¹⁷ systems extending from the ground state to excitation energies of about 8 MeV, there is a need at this point for more detailed theoretical study of the A=17, $T=\frac{1}{2}$ system. In particular, resonances of moderate reduced widths

observed in the $O^{16}+n$ interaction would probably be associated with states of O^{17} having configurations that couple fairly strongly to the continuum. Comparison of theory with experiment then would shed light on both the form of the interaction and the configurations of the intermediate states involved.

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Dependence of Proton Optical-Model Parameters upon Experimental Uncertainties*

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The dependence of proton optical-model parameters upon experimental uncertainties has been studied for four types of errors. These were errors in over-all normalization, errors affecting the shape of the angular distribution, errors in incident energy, and errors in scattering angle. By varying the data, it is found that the optical-model parameters are affected differently by the four types of errors; this suggests that a single, over-all assigned error is inadequate for error analysis in optical-model studies. The minimum X^2 criterion is critically examined as a criterion for obtaining optimum optical-model parameters. The role of the experimental errors in the theoretical analysis is studied. Because the errors associated with the individual cross sections that form an angular distribution are not independent, we conclude that these data do not make up a collection of several independent random samples, but that, collectively, they resemble a single random sample. This conclusion is statistically important in the study of reaction mechanisms using results of optical-model analysis.

I. INTRODUCTION

 \mathbf{R} ECENT systematic optical-model analysis of proton elastic scattering from elements heavier than Al showed that good fits to the experimental data could be obtained.¹ Excellent fits to experimental data for E_p between 9 and 22 MeV were obtained when most of the parameters of the model were allowed to vary; however, the resulting best-fit parameters exhibited considerable variation from one angular distribution to the next. When large fluctuations in the parameters as a function of mass number and energy are necessary to describe the scattering, the usefulness of an optical-model description is much reduced.

To make the optical-model description more useful in the proton analysis¹ an average set of geometrical parameters was obtained to determine whether the data could be reproduced with smoothly varying real and imaginary well depths. A distinctive feature of this analysis was that the data were very adequately reproduced by the predictions of this restricted model, and the variations in the well-depth parameters were substantially reduced.

The next goal has been to ascertain if there is a systematic variation of these parameters as a function

of mass number which might yield information about the mechanisms involved in the reaction. An example is the suggestion of Lane² that the real optical potential should exhibit an isobaric spin dependence. For nucleon elastic scattering this dependence results in a nuclearsymmetry term in the optical potential of the form

$$\pm V_1(N-Z)/A, \qquad (1)$$

where the + and - signs refer to neutrons and protons, respectively. Several groups^{1,3–9} have attempted to obtain a value for V_1 by studying proton elastic scattering; results from some of these studies are presented in Table I. Examination of Table I shows that all of the values are reasonably consistent with $V_1 \simeq 25$ MeV except that of Durisch and Gould.⁶ The tin data they analyzed were reported with a precision of $\pm 3\%$,¹⁰ a

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⁹ G. R. Satchler, R. M. Drisko, and R. H. Bassel, Phys. Rev. 136, B637 (1964).

¹⁰ J. E. Durisch, R. R. Johnson, and N. M. Hintz, Phys. Rev. 137, B904 (1965).