

ment. Now we define Klein-transformed Green components,  $\Psi_i^{(\alpha)}(x)$ , which are local with respect to all  $\varphi_i^{(\alpha)}$ ,

$$\Psi_i^{(\alpha)}(x) = (-)^{N(i,\alpha)} \varphi_i^{(\alpha)}(x).$$

Here

$$N(i,\alpha) = \sum_{j,\beta} \chi(i,\alpha; j,\beta) n(j,\beta),$$

and  $n(j,\beta)$  is the number operator for the component field  $\varphi_j^{(\beta)}$ . Note that

$$[n(j,\beta), \varphi_k^{(\gamma)}(x)]_- = 0, \text{ unless } j=k, \text{ and } \beta=\gamma.$$

Now we are prepared to use a Reeh-Schlieder argument<sup>3</sup> to prove the desired result. Assume that  $\mathcal{O}$  is a

<sup>3</sup>H. Reeh and S. Schlieder, *Nuovo Cimento* **22**, 1051 (1961). Since a monomial in the  $\Psi_i^{(\alpha)}$  acting on  $\Phi_0$  equals the same monomial in the  $\varphi_i^{(\alpha)}$  acting on  $\Phi_0$  up to a constant factor  $\pm 1$ , the known Reeh-Schlieder property for the  $\varphi_i^{(\alpha)}$  implies the same property for the  $\Psi_i^{(\alpha)}$ .

polynomial of parafields localized in a region  $B$  and that

$$\mathcal{O}\Phi_0 = 0, \quad (2)$$

where  $\Phi_0$  is the vacuum of the parafield representation  $\mathcal{A}$ . Note that  $\Phi_0$  is also the vacuum of  $\mathcal{B}$ . Then multiply Eq. (2) by a polynomial  $\mathcal{O}'$  of Klein-transformed Green component fields  $\Psi_i^{(\alpha)}$  localized in a region  $B'$  space-like to  $B$ . Since  $\Psi_i^{(\alpha)}$  is local with respect to all  $\varphi_j^{(\beta)}$ ,  $\mathcal{O}'$  commutes with  $\mathcal{O}$ . Therefore we have

$$\mathcal{O}'\mathcal{O}\Phi_0 = \mathcal{O}\mathcal{O}'\Phi_0 = 0. \quad (3)$$

Since  $\mathcal{O}'\Phi_0$  is dense in  $\mathcal{B}$  because of the Reeh-Schlieder theorem, Eq. (3) implies that  $\mathcal{O}$  vanishes acting on  $\mathcal{B}$ ,<sup>4</sup> which is the separating property we wanted to prove.

<sup>4</sup>Even for unbounded Bose fields, the Reeh-Schlieder theorem proves that  $\mathcal{O}=0$  on the standard polynomial domain of Wightman.

## Effect of Intermediate Vector Mesons on Neutrino Production of Lepton Pairs\*

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The total cross section for the production of lepton pairs by the scattering of high-energy muon neutrinos and protons is calculated with an assumed weakly interacting vector meson  $W$  which has an anomalous magnetic moment  $\kappa$ . We use a modified Weizsäcker-Williams approximation, with added correction terms which change the result about 50%. In addition to these correction terms, we keep  $l^2$ , the square of the four-momentum of the exchanged photon, in the denominator of the lepton propagators. This results in an additional deviation of 30% from the standard Weizsäcker-Williams approximation. The cross sections are calculated as a function of  $\kappa$  for meson masses of 2.0, 2.5, and 3.0 BeV, and for neutrino energies above the threshold for  $W$  production. Numerical results are presented for  $\kappa=1, 0, -1$ . The cross sections for muon-pair and muon-electron production are equal, and larger by a factor of approximately  $10^4$  than the equivalent cross sections in the four-fermion theory.

### 1. INTRODUCTION

MANY authors have discussed the idea that the  $V-A$  theory of weak interactions is a phenomenological theory, and that the basic theory is one in which a massive vector meson is coupled to parity-non-conserving currents, formed from bilinear combinations of fermion fields.<sup>1,2</sup>

Experiments designed to search for the vector meson, by using high-energy neutrino beams, examine the production of lepton pairs in neutrino-nucleon collisions.

\* This research was supported in part by the National Science Foundation.

<sup>1</sup>T. D. Lee and C. N. Yang, *Phys. Rev.* **119**, 1410 (1960).

<sup>2</sup>G. Feinberg and A. Pais, *Phys. Rev.* **133**, B477 (1964).

If lepton pairs are produced via a four-fermion interaction, then

$$\sigma_{\mu^-l}(AF) \simeq M_p E_\nu G_*^2 \alpha^2,$$

where  $\sigma_{\mu^-l}$  is the cross section for the process

$$\nu_\mu + p \rightarrow \nu_l + p + \mu^- + l,$$

$l$  is a lepton, either  $e^+$  or  $\mu^+$ ,  $M_p$  is the proton mass,  $E_\nu$  is the incoming neutrino laboratory energy.  $G_*$  is the Fermi coupling constant [ $G_* = (1.01 \times 10^{-6})/M_p^2$ ], and  $\alpha$  is the fine-structure constant ( $\alpha = 1/137$ ). For neutrino energies greater than 1 BeV, this estimate yields a cross section approximately equal to  $10^{-42}$  cm<sup>2</sup>. This estimate is also applicable even when vector mesons are

involved, provided that the neutrino energy is below the threshold for real  $W$  production. As soon as real  $W$ 's can be produced, however, the cross section rapidly increases by a factor of about  $10^5$ , since it is then of the order of the cross section for  $W$  production times the branching ratio for  $W \rightarrow$  leptons. For this estimate we obtain  $\sigma_{\mu^-i}(W) \approx G_w \alpha^2 \approx 10^{-37}$  cm<sup>2</sup>. Since the observed  $\sigma_{\mu^-i}$  is well below this number, we conclude that if  $W$ 's exist, their mass must be so high that present-day neutrino beams do not have enough energy to produce them, and this sets a lower limit on the  $W$  mass. The neutrino experiments indicate the lower limit on the meson mass is about 2.0 BeV.<sup>3</sup>

Several authors have calculated the pair-production cross section in the four-fermion theory.<sup>4,5</sup> In this paper, we describe a calculation of the cross section in the vector-meson theory for neutrino energies above the  $W$ -production threshold. The results bear out the crude estimates given above.

The following is a short outline of the calculation: because of its short lifetime ( $<10^{-17}$  sec), the  $W$  is treated in the standard resonance approximation.<sup>6</sup> The electromagnetic interactions with the proton are calculated using the recent experimentally determined form factors.<sup>7</sup> The electromagnetic interaction of the  $W$  is the usual one which includes the effect of an anomalous magnetic moment.<sup>8</sup> Finally, a modified version of the Weizsäcker-Williams approximation (MWW) is employed which includes additional terms that change the result about 50%. In the calculation of the cross section for  $\gamma + \nu_\mu \rightarrow \nu_l + \mu^- + l$ , which is needed for the Weizsäcker-Williams approximation, we do not set  $l^2 = 0$  in the denominators of the lepton propagators ( $l^2$  is the square of the four-momentum of the photon). This results in an additional deviation of 30% from the standard Weizsäcker-Williams approximation.

Since we are using a resonance approximation, it is obviously necessary to have a reasonably accurate value for the total decay rate of  $W \rightarrow$  everything. We use the values  $\Gamma_l/\Gamma = 0.65$  for  $M_W = 2.5$  BeV, and  $\Gamma_l/\Gamma = 0.72$  for  $M_W = 3.0$  BeV, where  $\Gamma_l = \Gamma_{\mu\nu} + \Gamma_{e\nu} = G_w M_W^3 / 3\pi\sqrt{2}$ .<sup>1</sup> For  $M_W = 2.0$  BeV, we use  $\Gamma_l/\Gamma = 1.0$ . A detailed calculation of  $\Gamma$  for  $M_W = 2.5$  and 3.0 BeV has been done by Carhart and Doohar.<sup>9</sup> The value chosen for  $M_W = 2.0$  BeV is justified by noting that phase-space factors depress the rate for  $W \rightarrow 2$  baryons, and form factors depress the rates for  $W \rightarrow$  mesons if we assume that the form factors are dominated by low-lying resonances.<sup>9</sup>

<sup>3</sup> R. Burns, K. Goulianos, E. Hyman, L. Lederman, W. Lee, N. Mistry, J. Rettberg, M. Schwartz, J. Sunderland, and G. Danby, Phys. Rev. Letters 15, 42 (1965).

<sup>4</sup> G. N. Stanciu, Phys. Rev. Letters 13, 288 (1964).

<sup>5</sup> W. Czyz, G. C. Sheppey, and J. D. Walecka, Nuovo Cimento 34, 404 (1964).

<sup>6</sup> Z. Fraenkel, Nuovo Cimento 30, 513 (1963).

<sup>7</sup> J. R. Dunning, Jr., K. W. Chen, A. A. Cone, G. Hartwig, N. F. Ramsey, J. K. Walker, and R. Wilson, Phys. Rev. Letters 13, 631 (1964).

<sup>8</sup> T. D. Lee and C. N. Yang, Phys. Rev. 128, 885 (1962).

<sup>9</sup> R. Carhart and J. Doohar, Phys. Rev. 142, 1214 (1966).

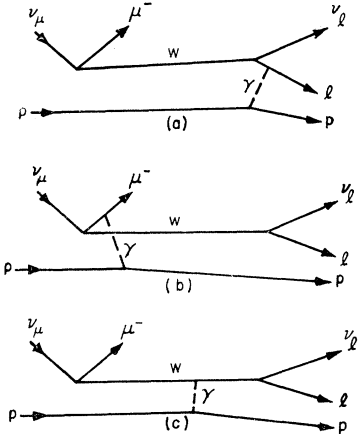


FIG. 1. Feynman diagrams for the process  $\nu_\mu + p \rightarrow \nu_l + p + \mu^- + l$  with an assumed intermediate vector meson  $W$ . Here  $l$  is a lepton, either  $e^+$  or  $\mu^+$ .

Hence, we assume that  $\Gamma$  is almost entirely leptonic for  $M_W = 2.0$  BeV.

Finally, we indicate that it is not immediately obvious that MWW is a good approximation in this problem. The use of MWW is justified in the following section by a detailed examination of the contributions to the cross section from the different regions of phase space.

## 2. DIFFERENTIAL CROSS SECTION FOR $\nu_\mu + p \rightarrow \nu_l + p + \mu^- + l$ ; A MODIFIED WEIZSÄCKER-WILLIAMS APPROXIMATION

The relevant Feynman graphs for the process  $\nu_\mu + p \rightarrow \nu_l + p + \mu^- + l$  are shown in Fig. 1. The cross section can be written as

$$\sigma = \frac{1}{(2\pi)^8} \frac{M_W^2}{|(\mathbf{p}_1 \cdot \mathbf{k}_1)|} \int \frac{d^3 \mathbf{p}_2}{E_2} \int \frac{d^3 \mathbf{p}^+}{\epsilon_+} \int \frac{d^3 \mathbf{p}^-}{\epsilon_-} \int \frac{d^3 \mathbf{k}_2}{\epsilon_2} \times \delta^4(\mathbf{p}^+ + \mathbf{p}^- + \mathbf{p}_2 + \mathbf{k}_2 - \mathbf{p}_1 - \mathbf{k}_1)^{\frac{1}{2}} \sum_{\text{spins}} |\mathfrak{M}|^2, \quad (1)$$

where  $\mathbf{p}_{1\mu}$ ,  $\mathbf{k}_{1\mu}$  are the four-momenta of the incoming proton and neutrino, respectively, and  $\mathbf{p}_{2\mu}$ ,  $\mathbf{k}_{2\mu}$ ,  $\mathbf{p}_\mu^-$ ,  $\mathbf{p}_\mu^+$  are the four-momenta of the outgoing proton, neutrino, muon, and positively charged lepton, respectively.  $\mathfrak{M}$  is the usual Feynman matrix element. In the lowest order in  $g$ , (the weak coupling of the vector meson to fermion currents), and  $\alpha$ , in which the process can occur, we can write  $\mathfrak{M}$  as follows:

$$\mathfrak{M} = (\bar{U}_2 J_\mu U_1)(t^{-2} A_\mu), \quad (2)$$

where  $U_1$  and  $U_2$  are the proton spinors,  $t^{-2} A_\mu$  includes everything else, and  $t^2$  is the square of the momentum transfer to the proton.<sup>10</sup>

The photon-proton vertex can be written

$$\bar{U}_2 J_\mu U_1 = ie\gamma_\mu F_1(l^2) - i\eta(e/2M_p)\sigma_{\mu\nu} t_\nu F_2(l^2), \quad (3)$$

<sup>10</sup> We are using the metric corresponding to  $A \cdot B = A_\mu B_\mu = \mathbf{A} \cdot \mathbf{B} - A_0 B_0$ . We use  $\hbar = c = 1$ .

where we have included the effects of the strong interactions in the form factors  $F_1$  and  $F_2$ .  $\eta$  is the anomalous magnetic moment of the proton ( $\eta=1.79$ ).

The next step in the calculation approximates

$$\int \frac{d^3 p^+}{\epsilon_+} \int \frac{d^3 p^-}{\epsilon_-} \int \frac{d^3 k_2}{\epsilon_2} \times \delta^4(p^+ + p^- + k_2 + p_2 - p_1 - k_1) \sum_{\text{spins}} |\mathfrak{M}|^2$$

by the use of a modified version of the Weizsäcker-Williams approximation derived as follows. (For a discussion of the standard covariant Weizsäcker-Williams approximation, see Refs. 11-13.) We can write

$$\sum_{\text{spins}} |\mathfrak{M}|^2 = \frac{J_{\mu\nu} T_{\mu\nu}'}{t^4}, \quad (4)$$

where

$$J_{\mu\nu} = \sum_{\text{proton spins}} (\bar{U}_2 J_\mu U_1) (\bar{U}_1 J_\nu^* U_2),$$

and<sup>14</sup>

$$T_{\mu\nu}' = \sum_{\text{lepton spins}} A_\mu A_\nu^*.$$

Gauge invariance implies that  $t_\mu T_{\mu\nu}' = t_\nu T_{\mu\nu}' = 0$ .

Let us integrate the differential cross section over all the final momenta except that of the proton. We obtain

$$d\sigma = \frac{1}{2} \frac{1}{(2\pi)^8} \frac{M_p^2}{|p_1 \cdot k_1|} \frac{J_{\mu\nu} T_{\mu\nu}'}{t^4} \frac{d^3 p_2}{E_2}, \quad (5)$$

where

$$T_{\mu\nu} = \int \frac{d^3 p^+}{\epsilon_+} \int \frac{d^3 p^-}{\epsilon_-} \int \frac{d^3 k_2}{\epsilon_2} \times \delta^4(p^+ + p^- + k_2 + p_2 - p_1 - k_1) T_{\mu\nu}'.$$

The most general gauge-invariant  $T_{\mu\nu}$  is

$$-T_{\mu\nu} = a \left[ \frac{t^2}{(k_1 \cdot t)} k_{1\mu} k_{1\nu} + (k_1 \cdot t) \delta_{\mu\nu} - k_{1\mu} t_\nu - k_{1\nu} t_\mu \right] + b \left[ t^2 \delta_{\mu\nu} - t_\mu t_\nu \right], \quad (6)$$

where  $a$  and  $b$  are functions of  $k_1 \cdot t$  and  $t^2$ , and have the dimensions of a cross section, i.e., (energy)<sup>-2</sup>.

It can be shown that the functions  $a$  and  $b$  are of the same order of magnitude.<sup>15</sup> We note that  $a(t^2=0) = 2(2\pi)^5 \sigma_p$ , where  $\sigma_p$  is the cross section for the process

<sup>11</sup> V. N. Gribov, V. A. Kolkunov, L. B. Okun, and V. M. Shekhter, Zh. Eksperim. i Teor. Fiz. **41**, 1839 (1961) [English transl.: Soviet Phys.—JETP **14**, 1308 (1962)].

<sup>12</sup> V. Gorgé, M. Locher, and H. Rollnik, Nuovo Cimento **27**, 928 (1963).

<sup>13</sup> V. Gorgé, Nuovo Cimento **35**, 545 (1965).

<sup>14</sup> We use the notation:  $A_\mu^* = (-1)^{\eta(\mu)} A_\mu^\dagger$ , where  $\eta(\mu) = 0$  for  $\mu = 1, 2, 3$ , and  $\eta(4) = 1$ .

<sup>15</sup> E. P. Shabalin, Zh. Eksperim. i Teor. Fiz. **93**, 175 (1962) [English transl.: Soviet Phys.—JETP **16**, 124 (1963)].

$\gamma + \nu_\mu \rightarrow \nu_l + \mu^- + l$ .<sup>11</sup> For  $J_{\mu\nu}$ , we obtain the following:

$$J_{\mu\nu} = (\not{p}_2 \not{p}_1 \not{\nu} + \not{p}_1 \not{p}_2 \not{\nu}) \times \left[ G_E^2 + \frac{t^2}{4M_p^2} G_M^2 \right] \left[ M_p^2 \left( 1 + \frac{t^2}{4M_p^2} \right) \right]^{-1} + \delta_{\mu\nu} \frac{t^2}{2M_p^2} G_M^2 + \text{terms which are proportional to } t_\mu \text{ and } t_\nu. \quad (7)$$

The terms proportional to  $t_\mu$  and  $t_\nu$  will not contribute to  $\sigma$  because of the gauge-invariance requirement.  $G_E$  and  $G_M$  are the electric and magnetic form factors formed from linear combinations of  $F_1$  and  $F_2$ .<sup>16</sup> In this paper we use

$$G_M = \mu G_E = \frac{\mu e}{(1 + 1.23 t^2 / M_p^2)^2},$$

which are obtained from the most recent high-energy electron-proton scattering data.<sup>7</sup>  $\mu$  is the magnetic moment of the proton ( $\mu = 2.79$ ).

Using the previous results, we can obtain an expression in terms of  $a$  and  $b$  for the differential cross section,  $\partial^2 \sigma / \partial(\omega^2) \partial(t^2)$ , where  $\omega$  is the center-of-mass energy of the outgoing particles  $\mu^-$ ,  $l$ , and  $\nu_l$ . This is done by using the transformation

$$\frac{d^3 p_2}{E_2} = \frac{\pi d t^2 d \omega^2}{2(k_1 \cdot p_1)} \quad (8)$$

(the angular variable has been integrated over), and substituting Eqs. (6) and (7) into Eq. (5). From this expression we see that  $\partial^2 \sigma / \partial(\omega^2) \partial(t^2)$  falls off rapidly with increasing  $t^2$  and  $\omega^2$ , if we assume that  $a$  varies with  $\omega$  as  $\sigma_p(\omega)$ , and has no more than a quadratic dependence with  $t_\mu$ . For an incoming neutrino energy  $E_\nu$  of 10 BeV, and a meson mass of 2.0 BeV,  $\partial^2 \sigma / \partial(\omega^2) \partial(t^2)$  at  $\omega^2 \approx 9.0$  BeV<sup>2</sup> is about  $\frac{1}{5}$  of its value at  $\omega^2 = 4.4$  BeV<sup>2</sup>, which is the threshold of the resonance region. For  $\omega^2 < 4.4$  BeV<sup>2</sup>, the differential cross section is smaller by a factor of  $10^4$ , since the resonance does not occur in this region. Therefore, the region  $4.4 \text{ BeV}^2 \leq \omega^2 \leq 9.0 \text{ BeV}^2$  yields the most important contribution to the total cross section  $\sigma$ . Since  $\partial^2 \sigma / \partial(\omega^2) \partial(t^2)$  falls off very rapidly with  $t^2$ , due to the proton form factors and the  $t^{-4}$  factor from the exchanged photon, we set  $t^2 \approx (t^2)_{\min}$ , where  $(t^2)_{\min}$  is found from the kinematics<sup>17</sup>:

$$\frac{(t^2)_{\max, \min}}{M_p^2} = \frac{2}{[2(E_\nu/M_p) + 1]} \left\{ \left( \frac{E_\nu}{M_p} \right)^2 - \frac{1}{2} \left( \frac{E_\nu}{M_p} + 1 \right) \frac{\omega^2}{M_p^2} \pm \frac{E_\nu}{M_p} \left[ \left( \frac{E_\nu}{M_p} \right)^2 - \left( \frac{E_\nu}{M_p} + 1 \right) \frac{\omega^2}{M_p^2} + \frac{\omega^4}{M_p^4} \right]^{1/2} \right\}; \quad (9)$$

also

$$\omega_{\max} = [2E_\nu M_p + M_p^2]^{1/2} - M_p. \quad (10)$$

<sup>16</sup> L. N. Hand, D. G. Miller, and R. Wilson, Rev. Mod. Phys. **35**, 335 (1963).

<sup>17</sup> G. F. Chew and F. E. Low, Phys. Rev. **113**, 1640 (1959).

From Eq. (9) we can show that in the region

$$4.4 \text{ BeV}^2 \leq \omega^2 \leq 9.0 \text{ BeV}^2,$$

$$\frac{(\ell^2)_{\min}}{\omega^2} < 0.1, \quad \text{and} \quad \frac{(\ell^2)_{\min}}{M_W^2} \leq 0.1.$$

Although the previous statements were made for  $E_\nu \approx 10$  BeV, they can be shown to hold in essence for any  $E_\nu > W$  threshold.

We can use the following approximations derived from the previous discussion that for  $E_\nu$  above the threshold for  $W$  production:

$$\omega^2 \gg \ell^2, \quad M_p E_\nu \gg \ell^2,$$

since

$$\omega_{\max}^2 \approx 2M_p E_\nu, \quad \text{and} \quad \omega^2 \gg \ell^2.$$

We shall consider only energies  $E_\nu \gg M_p$ . We obtain

$$\frac{\partial^2 \sigma}{\partial(\omega^2) \partial(\ell^2)} = \frac{\alpha}{\pi} \bar{\sigma}_p(\omega^2, \ell^2) F(\omega^2, \ell^2), \quad (11)$$

where

$$\begin{aligned} & t^4 \left( 1 + 1.23 \frac{\ell^2}{M_p^2} \right)^4 F(\omega^2, \ell^2) \\ &= \left( 1 + \mu \frac{\ell^2}{4M_p^2} \right) \ell^2 (\ell^2 + \omega^2)^{-1} \left( 1 + \frac{\ell^2}{4M_p^2} \right)^{-1} \\ & \quad - \frac{1}{4} E_\nu^{-2} \left( 1 + \mu \frac{\ell^2}{4M_p^2} \right) (\ell^2 + \omega^2) \left( 1 + \frac{\ell^2}{4M_p^2} \right)^{-1} \\ & \quad - \frac{1}{2} E_\nu^{-1} M_p^{-1} \left( 1 + \mu \frac{\ell^2}{4M_p^2} \right) \ell^2 \left( 1 + \frac{\ell^2}{4M_p^2} \right)^{-1} \\ & \quad + \frac{1}{8} \mu^2 M_p^{-2} E_\nu^{-2} (\ell^2 + \omega^2) \ell^2, \end{aligned}$$

where

$$\bar{\sigma}_p(\omega^2, \ell^2) = \frac{a(\omega^2, \ell^2)}{2(2\pi)^5}.$$

We note that

$$\bar{\sigma}_p(\omega^2, \ell^2=0) = \sigma_p(\omega^2).$$

We have dropped the  $b$  terms in deriving Eq. (11) because  $b$  is the same order of magnitude as  $a$  and the ratio of the expression multiplying  $b$  to that multiplying  $a$  is less than  $M_p/E_\nu$ .

The standard Weizsäcker-Williams approximation consists in replacing  $\bar{\sigma}_p(\omega^2, \ell^2)$  by its value at the photon pole,  $\sigma_p(\omega^2)$ . Also the approximation  $M_p E_\nu \gg \omega^2$  is usually made.<sup>11</sup> These approximations lead to the equation

$$\begin{aligned} \frac{\partial^2 \sigma}{\partial(\omega^2) \partial(\ell^2)} &= \frac{\alpha}{\pi} \sigma_p(\omega^2) \left( 1 + \frac{\mu \ell^2}{4M_p^2} \right) \ell^{-2} (\ell^2 + \omega^2)^{-1} \\ & \quad \times \left( 1 + \frac{\ell^2}{4M_p^2} \right)^{-1} \left( 1 + 1.23 \frac{\ell^2}{M_p^2} \right)^{-4}. \quad (12) \end{aligned}$$

We find that the approximation  $M_p E_\nu \gg \omega^2$  is not valid in this calculation because the region  $\omega^2 > \omega_t^2$  yields the most important contribution to the total cross section  $\sigma$ , as previously stated.  $\omega_t$  is the threshold for the resonance region:

$$\omega_t = (M_W + m_\mu). \quad (13)$$

Since  $M_W > 2.0$  BeV, relatively large values of  $\omega^2$  contribute to the calculation of  $\sigma$ . Therefore the terms neglected in the derivation of Eq. (12) contribute 50% to the total result in our problem.

The replacement of  $\bar{\sigma}_p(\omega^2, \ell^2)$  by its value at the photon pole is also a poor approximation in our problem. The reason for this is that for a large region of phase space,  $\ell^2 > m_\mu^2$ , and  $\ell^2 > m_l^2$ . Since  $\ell^2$  occurs in the denominators of the lepton propagators, the position of the poles of the propagators are much different for  $\ell^2=0$  and  $\ell^2 \neq 0$ . The terms involving the lepton propagators which contribute to  $\sigma$  can differ by a factor of 2 for  $\ell^2=0$  and  $\ell^2 \neq 0$ . This is the most important difference between  $\bar{\sigma}_p(\omega^2, \ell^2)$  and  $\sigma_p(\omega^2)$ . Contributions to the ratio  $[\sigma_p(\omega^2) - \bar{\sigma}_p(\omega^2, \ell^2)]/\sigma_p(\omega^2)$ , other than those which arise by including the correct  $\ell^2$  dependence in the denominators of the lepton propagators, are either of the order of  $\ell^2/\omega^2$  or  $\ell^2/M_W^2$ ; therefore these contributions may be neglected. The result of this analysis is that  $\bar{\sigma}_p(\omega^2, \ell^2)$  can be approximated by calculating  $\sigma_p(\omega^2)$  and putting the correct  $\ell^2$  dependence in the denominators of the lepton propagators.

If we calculate  $\sigma$  for  $\ell^2=0$  in the lepton propagators, we obtain results that are approximately 30% larger than those with  $\ell^2$  in the propagators.

### 3. TOTAL CROSS SECTION FOR

$$\gamma + \nu_\mu \rightarrow \nu_l + \mu^- + l^+$$

The Feynman graphs pertinent to the calculation of the total cross section of the process  $\gamma + \nu_\mu \rightarrow \nu_l + \mu^- + l^+$

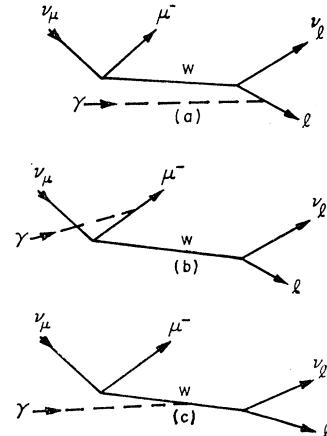


FIG. 2. Feynman diagrams for the process  $\gamma + \nu_\mu \rightarrow \nu_l + \mu^- + l$  with an assumed intermediary vector meson  $W$ . Here  $l$  is a lepton, either  $e^+$  or  $\mu^+$ .

are shown in Fig. 2.

$$\sigma_p = \frac{1}{(2\pi)^5} \frac{1}{2} \frac{1}{|(k_1 \cdot t)|} \int \frac{d^3 p^+}{\epsilon_+} \int \frac{d^3 p^-}{\epsilon_-} \int \frac{d^3 k_2}{\epsilon_2} \times \delta^4(p^+ + p^- + k_2 - t - k_1) \frac{1}{2} \sum_{\text{spins}} |\mathfrak{N}_p|^2. \quad (14)$$

$t_\mu$  is now the four-momentum of the incoming photon. For  $\mathfrak{N}_p$ , we use the Feynman rules. The lepton-photon vertex is the usual one.

For a vector-meson-lepton vertex, we have

$$\mathfrak{N}_{p, \nu_l, W} = ig \phi_\mu^m \bar{\nu}_k (1 - \gamma_5) \gamma_\mu u_{s, p}. \quad (15)$$

This corresponds to the reaction  $l \rightarrow \nu_l + W$ .  $\phi_\mu^m$  is the polarization vector of the  $W$ .  $\bar{\nu}_k = \nu_k^\dagger \gamma_4$  is the spinor for a neutrino with momentum  $\mathbf{k}$ .  $u_{s, p}$  is the spinor for a lepton with spin  $s$  and momentum  $\mathbf{p}$ .  $g^2 = G_F M_W^2 / \sqrt{2}$ .

$$\begin{aligned} \mathfrak{N}_p = & ig^2 e \frac{\bar{u}_p \gamma_\alpha (1 + \gamma_5) \nu_{k_1} \bar{\nu}_{k_2} (1 - \gamma_5) \gamma_\alpha (t - p^+) \gamma \cdot \epsilon^r v_{p^+ c}}{[q_1^2 + M_W^2](-2p^+ \cdot t + t^2)} + ig^2 e \frac{\bar{u}_p \gamma \cdot \epsilon^r (p^- - t) \gamma_\alpha (1 + \gamma_5) \nu_{k_1} \bar{\nu}_{k_2} (1 - \gamma_5) \gamma_\alpha v_{p^+ c}}{(-2p^- \cdot t + t^2)[q_2^2 + M_W^2]} \\ & - ig^2 e \frac{\bar{u}_p \gamma_\alpha (1 + \gamma_5) \nu_{k_1} \bar{\nu}_{k_2} (1 - \gamma_5) \gamma_\alpha v_{p^+ c} [(q_1 + q_2) \cdot \epsilon^r]}{[q_1^2 + M_W^2][q_2^2 + M_W^2]} - i(1 + \kappa) g^2 e \frac{\bar{u}_p \gamma \cdot \epsilon^r (1 + \gamma_5) \nu_{k_1} \bar{\nu}_{k_2} (1 - \gamma_5) t v_{p^+ c}}{[q_1^2 + M_W^2][q_2^2 + M_W^2]} \\ & + i(1 + \kappa) g^2 e \frac{\bar{u}_p t (1 + \gamma_5) \nu_{k_1} \bar{\nu}_{k_2} (1 - \gamma_5) \gamma \cdot \epsilon^r v_{p^+ c}}{[q_1^2 + M_W^2][q_2^2 + M_W^2]}, \quad (17) \end{aligned}$$

where  $A = \gamma_\mu A_\mu$  and  $v_{p^+ c}$  is the spinor for the positively charged lepton.  $\omega$  is now the center-of-mass energy of the incoming photon and neutrino. We have kept  $t^2$  in the denominators of the lepton propagators, as indicated in the preceding discussion.

Since we are doing the calculation in the resonance region  $\omega > \omega_t$ , we may neglect the lepton masses in the numerator terms of  $\sum_{\text{spins}} |\mathfrak{N}_p|^2$ .

The resonance is treated by replacing  $M_W$  by  $M_W - \frac{1}{2}i\Gamma$ .<sup>6</sup>  $\Gamma$  is the total decay width of the vector meson. We have  $M_W \gg \Gamma$  and  $M_W \Gamma \gg \Gamma^2$ .<sup>1</sup> The resonance occurs at  $q_2^2 = -M_W^2$ . Therefore the terms with  $q_2^2 + M_W^2$  in the denominator will be the most important ones. When  $\sum_{\text{spins}} |\mathfrak{N}_p|^2$  is computed, and we integrate to obtain  $\bar{\sigma}_p$ , the squares of the resonant terms are the most significant by a factor of  $10^4$  [Graphs (b) and (c) in Fig. 2].<sup>18</sup> For  $\bar{\sigma}_p$ , in the region  $\omega > \omega_t$ , we obtain

$$\bar{\sigma}_p = \sigma_0^p [\kappa^2 \beta_2^p + \kappa \beta_1^p + \beta_0^p], \quad (18)$$

where

$$\sigma_0^p = 6\sqrt{2}\alpha G_F \left(\frac{m_\mu}{M_W}\right)^2 \frac{\Gamma_l}{\Gamma}.$$

$$\beta_2^p = R_2(\omega) \ln\left(\frac{x}{\lambda}\right)^2 + S_2(\omega) [z_0^2 - 1]^{1/2},$$

The propagator of the vector meson is given by

$$-i[\delta_{\mu\nu} + q_\mu q_\nu / M_W^2] (q^2 + M_W^2)^{-1},$$

where  $q_\mu$  is the four-momentum of the vector meson.

The photon-meson vertex, corresponding to  $W \rightarrow W + \gamma$ , is given by<sup>8</sup>

$$\begin{aligned} \mathfrak{N}_{\gamma W} = & \phi_\alpha^m \phi_\beta^n \epsilon_\mu^r V_{(\mu)\alpha\beta}, \quad (16) \\ V_{(\mu)\alpha\beta} = & ie[\delta_{\alpha\beta}(q_1 + q_2)_\mu - \delta_{\alpha\mu}(-\kappa q_2 + q_1 + \kappa q_1)_\beta \\ & - \delta_{\beta\mu}(-\kappa q_1 + q_2 + \kappa q_2)_\alpha]. \end{aligned}$$

$q_{1\mu}$ ,  $\phi_\mu^m$  are the momentum and the polarization vectors of the incoming meson.  $q_{2\mu}$ ,  $\phi_\mu^n$  are the momentum and polarization vectors of the outgoing vector meson.  $\epsilon_\mu^r$  is the polarization vector of the photon.  $\kappa$  is the anomalous magnetic moment of the vector meson.

Using the Feynman rules, we obtain

where

$$x = \frac{\omega}{m_\mu}, \quad \lambda = \frac{M_W}{m_\mu}, \quad z_0 = \frac{x^2 + 1 - \lambda^2}{2x}.$$

$$R_2(\omega) = \frac{\lambda^2}{6} + \frac{1}{6} \frac{\lambda^4}{x^2} + \frac{2}{3} \frac{\lambda^2}{x^2} + \frac{1}{6} \frac{\lambda^4}{x^4},$$

$$S_2(\omega) = \frac{\lambda^2}{8x} + \frac{1}{3} \frac{\lambda^4}{x^3} + \frac{1}{24} \frac{\lambda^6}{x^5} + \frac{25}{24} \frac{1}{x},$$

$$\beta_1^p = R_1(\omega) \ln\left(\frac{x}{\lambda}\right)^2 + S_1(\omega) [z_0^2 - 1]^{1/2},$$

$$R_1(\omega) = \frac{\lambda^2}{3} + \frac{\lambda^4}{3x^2} + \frac{\lambda^2}{3x^2} + \frac{2}{3} \frac{\lambda^4}{x^4},$$

$$S_1(\omega) = \frac{31}{12} \frac{\lambda^2}{x} + \frac{\lambda^4}{x^3} + \frac{1}{12} \frac{\lambda^6}{x^5} + \frac{49}{12} \frac{1}{x},$$

$$\beta_0^p = R_0(\omega) \ln\left(\frac{x}{\lambda}\right)^2 + S_0(\omega) [z_0^2 - 1]^{1/2} + T_0(\omega)$$

$$\times \ln \left\{ \frac{z_0(1 - y^2/x^2) + y^2/x + [z_0^2 - 1]^{1/2}(1 + y^2/x^2)}{z_0(1 - y^2/x^2) + y^2/x - [z_0^2 - 1]^{1/2}(1 + y^2/x^2)} \right\},$$

<sup>18</sup> The full details of the calculations of  $\sigma_p$  are given in J. Doohar, thesis, Stevens Institute of Technology, 1965 (unpublished).

where  $y^2 = t^2/m_\mu^2$ . The term involving  $T_0(\omega)$  includes the effect of keeping  $t^2$  in the denominator of the muon

TABLE I. The numerical values of the total cross section for the process  $\nu_\mu + \bar{p} \rightarrow \nu_l + \bar{p} + \mu^- + l$ ;  $l$  is a lepton, either  $e^+$  or  $\mu^+$ . An intermediate vector meson is assumed with a mass of 2.0, 2.5, and 3.0 BeV.

Neutrino energy (BeV)	$M_W$ (BeV)	$\Gamma_l/\Gamma$	$\beta_2$	$\beta_1$	$\beta_0$	$\sigma$ ( $\kappa=1$ ) (cm <sup>2</sup> )	$\sigma$ ( $\kappa=0$ ) (cm <sup>2</sup> )	$\sigma$ ( $\kappa=-1$ ) (cm <sup>2</sup> )
4.37						Threshold for $W$ production		
5.63	2.0	1.0	0.268	0.949	3.75	$8.94 \times 10^{-39}$	$6.75 \times 10^{-39}$	$5.52 \times 10^{-39}$
9.38			4.63	15.9	52.2	$1.31 \times 10^{-37}$	$9.40 \times 10^{-38}$	$7.37 \times 10^{-38}$
13.1			12.9	41.1	128.7	$3.29 \times 10^{-37}$	$2.32 \times 10^{-37}$	$1.80 \times 10^{-37}$
16.9			23.6	70.8	215.7	$5.58 \times 10^{-37}$	$3.88 \times 10^{-37}$	$3.03 \times 10^{-37}$
20.6			35.7	102.0	304.6	$7.96 \times 10^{-37}$	$5.48 \times 10^{-37}$	$4.29 \times 10^{-37}$
5.99						Threshold for $W$ production		
7.50	2.5	0.65	0.180	0.659	2.97	$2.85 \times 10^{-39}$	$2.22 \times 10^{-39}$	$1.86 \times 10^{-39}$
12.2			4.27	15.4	57.4	$5.76 \times 10^{-38}$	$4.29 \times 10^{-38}$	$3.46 \times 10^{-38}$
16.9			13.0	43.8	153.0	$1.57 \times 10^{-37}$	$1.14 \times 10^{-37}$	$9.14 \times 10^{-38}$
21.6			24.9	79.2	267.3	$2.78 \times 10^{-37}$	$2.00 \times 10^{-37}$	$1.59 \times 10^{-37}$
8.25						Threshold for $W$ production		
10.3	3.0	0.72	0.155	0.585	2.97	$2.13 \times 10^{-39}$	$1.71 \times 10^{-39}$	$1.46 \times 10^{-39}$
14.1			2.00	7.65	34.0	$2.51 \times 10^{-38}$	$1.96 \times 10^{-38}$	$1.63 \times 10^{-38}$
17.8			6.18	22.7	94.3	$7.08 \times 10^{-38}$	$5.42 \times 10^{-38}$	$4.47 \times 10^{-38}$
21.6			12.4	43.8	173.3	$1.32 \times 10^{-37}$	$9.96 \times 10^{-38}$	$8.14 \times 10^{-38}$

propagator [Graph (b) in Fig. 2].

$$R_0(\omega) = \frac{\lambda^2}{6} + \frac{1\lambda^4}{2x^2} + \frac{4\lambda^6}{3x^4} + \frac{4\lambda^8}{3x^6} + \frac{2\lambda^2}{3x^2} + \frac{5\lambda^4}{6x^4} + \frac{10\lambda^6}{3x^6},$$

$$S_0(\omega) = \frac{1\lambda^2}{8x} + \frac{1\lambda^4}{3x^3} + \frac{65\lambda^6}{24x^5} + \frac{73}{24x},$$

$$T_0(\omega) = \frac{2\lambda^4}{3x^2} + \frac{4\lambda^6}{3x^4} + \frac{4\lambda^8}{3x^6} + \frac{4\lambda^2}{3x^2} + \frac{16\lambda^4}{3x^4} + \frac{\lambda^6}{x^6} + 8.$$

In the derivation of these expressions we have used  $M_W \gg \Gamma$  and  $M_W \Gamma \gg \Gamma^2$ .

For the evaluation of a typical resonant integral, see the Appendix.

#### 4. TOTAL CROSS SECTION FOR

$$\nu_\mu + \bar{p} \rightarrow \nu_l + \bar{p} + l + \bar{p}$$

To obtain  $\sigma$ , we integrate  $\partial^2 \sigma / \partial(t^2) \partial(\omega^2)$ , as given by Eq. (11), numerically over the region  $\omega_{\min}^2 \leq \omega^2 \leq \omega_{\max}^2$ ,  $(t^2)_{\min} \leq t^2 \leq (t^2)_{\max}$ . The final cross section takes the form

$$\sigma = \sigma_0 (\kappa^2 \beta_2 + \kappa \beta_1 + \beta_0). \tag{19}$$

$\beta_2, \beta_1$ , and  $\beta_0$  are functions of  $E_\nu$  and  $M_W$ . They are given in Table I.

$$\sigma_0 = 8.14 \left( \frac{M_p}{M_W} \right)^2 \frac{\Gamma_l}{\Gamma} \times 10^{-39} \text{ cm}^2.$$

The values used for  $\Gamma_l/\Gamma$  have been discussed in the Introduction. The numerical results are summarized in Table I.

#### 5. DISCUSSION OF RESULTS

Deriving the final expression for MWW, we have neglected terms which are smaller by factors of  $t^2/\omega^2$ ,

$t^2/M_W^2$ ,  $t^2/M_p E_\nu$ , or  $M_p/E_\nu$  than those terms which we have included.  $t^2/\omega^2$ ,  $t^2/M_W^2$ , and  $t^2/M_p E_\nu$  are all less than 0.1 in the regions of  $t^2$  and  $\omega^2$  which contribute to  $\sigma$  for  $E_\nu \geq W$  threshold. For  $E_\nu \geq 10$  BeV, we have  $M_p/E_\nu < 0.1$ . Therefore, for  $E_\nu \geq 10$  BeV we believe that 10% is a reasonable upper limit on the error in  $\sigma$ . For  $5 \text{ BeV} \leq E_\nu \leq 10 \text{ BeV}$  the error may be slightly larger.

By including the correct  $t^2$  dependence in the denominators of the lepton propagators, we can extend the validity of MWW for our problem over a wide region of the range of the variables  $t^2$  and  $\omega^2$ . The only requirement is that the most important contributions of phase space to the cross section occur in the region  $t^2/\omega^2 \ll 1$ ,  $t^2/M_W^2 \ll 1$ . In its usual form, the Weizsäcker-Williams approximation (WW) is good only if the positions of the poles of the lepton propagators are not much different for  $t^2=0$  and  $t^2 \neq 0$ .<sup>19</sup> This will be true only if  $t^2 \ll m_l^2$ , where  $l$  corresponds to the particular lepton propagator being considered. This condition is extremely restrictive and is not satisfied in our problem. However, the conditions  $t^2/\omega^2 \ll 1$  and  $t^2/M_W^2 \ll 1$ , are satisfied. In general, for a process involving exchanged photon and lepton propagators, the requirements for the validity of MWW are easiest to meet than those for WW.

We see from Table I that for neutrino energies varying from 5 to 20 BeV, the cross section varies from  $10^{-39}$  to  $10^{-37}$  cm<sup>2</sup>. Calculations using the four-fermion theory show that in a similar energy range, the cross sections vary from  $10^{-44}$  to  $10^{-43}$  cm<sup>2</sup>.<sup>4,5</sup> The results are in accord with the crude estimates discussed in the Introduction. We also note that the rate for  $\mu^-, \mu^+$  production is equal to that for  $\mu^-, e^+$  production in the vector-meson theory. In the four-fermion theory, however,  $\mu^-, e^+$  production

<sup>19</sup> V. Gorgé, M. Locher, and H. Rollnik, in passing, mention the effect of lepton propagators on the regions of phase space which contribute to the cross section of a process such as  $e^- + p \rightarrow e^- + p + \gamma$  (Ref. 12). See Ref. 13 by Gorgé for a longer discussion.

is larger by a factor of 2 to 3 for neutrino energies from 5 to 20 BeV.<sup>4,5</sup> This can be understood by noting that  $\sigma_{\mu^- \mu^+ (W)} \approx \sigma_{W \text{ production}} \times (\Gamma_{\mu\nu}/\Gamma)$ , and that  $\sigma_{\mu^- e^+ (W)} \approx \sigma_{W \text{ production}} \times (\Gamma_{e\nu}/\Gamma)$ . For a large-mass boson, we have  $\Gamma_{\mu\nu} = \Gamma_{e\nu}$ . Hence the cross sections are equal in the vector-meson theory. This result corresponds to the fact that we neglected graph (a) in Fig. 2, and neglected the lepton masses in numerator terms. The only place where lepton masses are significant is in the denominators of lepton propagators. In graph (b) only the muon propagator is involved. In graph (c) no lepton propagators are involved.

Finally, we note that even for very high neutrino energies, ( $E_\nu \approx 20$  BeV), the cross section falls off as  $M_W$  is increased. This occurs because the proton form factors rapidly fall off as the momentum transfer increases. Even though  $l^2 \approx (l^2)_{\min}$ ,  $(l^2)_{\min}$  will increase with  $M_W$  for fixed  $E_\nu$ , because it increases with  $\omega$ , and, hence  $\omega_t$ .

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#### APPENDIX: A TYPICAL FULLY RESONANT INTEGRAL

We consider the following integral typical of a term arising from the contribution of the square of graph (b) in Fig. 2 to  $\bar{\sigma}_p$

$$I = \int \frac{d^3 p^-}{\epsilon_-} \frac{1}{(-2p^- \cdot t + l^2)} \frac{1}{(q_z^2 + M_W^2)^2 + \Gamma^2 M_W^2}. \quad (\text{A1})$$

We have done the integration over  $d^3 p^+$  and  $d^3 k_2$  covariantly using the properties of the energy and momentum delta functions.<sup>20</sup> Letting  $x = \omega/m_\mu$ ,  $z = \epsilon_-/m_\mu$ ,  $\lambda = M_W/m_\mu$ ,  $\Gamma_0 = \Gamma/m_\mu$ , and  $y^2 = l^2/m_\mu^2$ , we may write, after the angular integration has been done,

$$I = \frac{2\pi}{m_\mu^3 \omega} \int_1^{x/2} dz (|2xz - 2xz_0|^2 + \Gamma_0^2 \lambda^2)^{-1} \times \ln \left\{ \frac{z(1 - y^2/x^2) + y^2/x + (z^2 - 1)^{1/2}(1 + y^2/x^2)}{z(1 - y^2/x^2) + y^2/x - (z^2 - 1)^{1/2}(1 + y^2/x^2)} \right\}, \quad (\text{A2})$$

where

$$z_0 = (x^2 + 1 - \lambda^2)/2x.$$

<sup>20</sup> J. D. Jackson, in *Brandeis Summer Institute of Theoretical Physics, 1962* (W. A. Benjamin, Inc., New York, 1963).

We have kept  $y^2$  in the logarithm but neglected over-all factors proportional to  $y^2/x^2$ .

For  $x > x_0 = 1 + \lambda$  we can have  $q_z^2 = -M_W^2$  in the region of integration. Using this fact, we shall show that

$$(|2xz - 2xz_0|^2 + \Gamma_0^2 \lambda^2)^{-1} = (\pi/2x\Gamma_0\lambda)\delta(z - z_0), \quad (\text{A3})$$

for  $x > 1 + \lambda + 10\Gamma_0$ , (i.e.,  $x$  sufficiently far from the resonance threshold).

Since a relatively large region contributes to  $\sigma$  in the integration of  $\partial^2 \sigma / \partial(\omega^2) \partial(l^2)$  over  $\omega$ , an error in the region  $M_W + m_\mu \leq \omega \leq M_W + m_\mu + 10\Gamma$  will be negligible since  $M_W \gg \Gamma$ . Therefore, once Eq. (A3) is derived we may use it throughout the calculation.

To prove Eq. (A3) we consider the following integral:

$$I_1 = \int_1^{x/2} \frac{z dz}{|2xz - 2xz_0|^2 + \Gamma_0^2 \lambda^2}. \quad (\text{A4})$$

This integral is straightforward, and for  $x > x_0$  we obtain

$$I_1 = \frac{z_0}{2x\Gamma_0\lambda} \tan^{-1} \left\{ \frac{2xz - 2xz_0}{\Gamma_0\lambda} \right\} \Big|_1^{x/2}. \quad (\text{A5})$$

We used  $M_W \gg \Gamma$ .

For  $x > x_0 + 10\Gamma_0$  we may expand the arctangent as follows:

$$\tan^{-1} \left\{ \frac{2xz - 2xz_0}{\Gamma_0\lambda} \right\} \Big|_1^{x/2} \approx \pi - \Gamma_0\lambda \left[ \frac{1}{\lambda^2 - 1} + \frac{1}{x^2 - 2x - (\lambda^2 - 1)} \right]. \quad (\text{A6})$$

For  $x > x_0 + 10\Gamma_0$  we may drop the second term in Eq. (A6). We obtain

$$I_1 = (\pi/2\Gamma_0\lambda)(z_0/x). \quad (\text{A7})$$

This is the result obtained by using Eq. (3). Equation (A3) can be shown to hold for<sup>18</sup>:

$$I_n = \int_1^{x/2} \frac{z^n dz}{|2xz - 2xz_0|^2 + \Gamma_0^2 \lambda^2}, \quad n=0, 1, \dots \quad (\text{A8})$$

Using Eq. (A3) we obtain for  $I$

$$I = \frac{\pi^2}{m_\mu^2 \omega^2} \left( \frac{1}{\Gamma_0\lambda} \right) \times \ln \left\{ \frac{z_0(1 - y^2/x^2) + y^2/x + [z_0^2 - 1]^{1/2}(1 + y^2/x^2)}{z_0(1 - y^2/x^2) + y^2/x - [z_0^2 - 1]^{1/2}(1 + y^2/x^2)} \right\}. \quad (\text{A9})$$

The remaining resonant integrals contributing to  $\bar{\sigma}_p$  can be done in the same way.