information about binding energies and to make isotopic assignments of resonances. The sum-coincidence spectra yield the results 6.16 ± 0.05 and 5.44 ± 0.09 MeV for the binding energies of W182 (21.2 eV) and W186 (18.8 eV), respectively; these values are not in disagreement with the values 6.29 ± 0.04 and ≈ 4.8 MeV obtained in other

ways.40 The maximum energy of the sum pulses observed in the spectrum for the 15.9-eV resonance is 6.30 ± 0.08 MeV. This energy implies that the resonance must be assigned to either W180 or W182. Similarly, a previously reported weak resonance at 75 eV must be assigned to either W¹⁸⁰ or W¹⁸².

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Atomic Masses from Gallium to Molybdenum*

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The 16-in. double-focusing mass spectrometer at the University of Minnesota has been employed to measure the atomic mass of 42 stable isotopes in the region A = 69 to 100. The standard error associated with these results is approximately 5 parts in 108. Improvements in the instrument are described that result in an increase of useful resolution of a factor of 2 to 3. A set of 64 radioactive masses is calculated from the stable mass data together with β -decay energies and nuclear reaction Q values. The resultant table of masses is used to calculate total nuclear binding energies, separation energies and pairing energies for a number of nuclei in the region near N = 50. The systematics of the separation energies display very smooth characteristics except at the shell closure. Neutron pairing energies show a marked decrease in value following the shell closure.

INTRODUCTION

HE 16-in. double-focusing mass spectrometer at the University of Minnesota has been employed in the past to measure atomic masses in several regions of the periodic table. Measurements of atomic masses for most of the stable isotopes with A < 70 have been reported.¹⁻⁵ In addition, krypton and xenon,⁶ lead and mercury⁷ masses have been measured. Operational difficulties in the mass spectrometer become progressively more apparent in measurements of heavy isotopes where maximum resolution is required. These difficulties necessitated the movement and reconstruction of the instrument.

The improved instrument has been employed to measure a number of stable masses for A > 70. Mass doublets in the region from gallium through molyb-

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denum are reported in this paper and the following paper8 reports results in the region from ruthenium through xenon.

The mass results have been employed to study the nuclear binding energy systematics in the region Z=31to 42 and N=36 to 58. These data include the N=50shell closure and also the possible N = 40 and Z = 40 subshell closure.

THE INSTRUMENT

The instrument employed for all previous measurements has been described in some detail elsewhere.9 The ion optics of the instrument are arranged to yield firstand second-order angle focusing and first-order velocity focusing at the fixed collector slit. Doublets are measured by the peak-matching method.

One of the recurring difficulties in the original instrument was the random modulation of the ion beam resulting from building vibrations and time-varying magnetic fields from nearby ac power lines. The modulation not only limited the maximum useful resolution of the mass spectrometer, but also limited the sweep frequency that could be employed.

To remove these difficulties, the instrument was rebuilt in a ground-floor room which had lower stray magnetic fields. The instrument was mounted on a 2-ton cast-iron surface plate. The use of the surface plate as an

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Mexico.

¹ K. S. Quisenberry, T. T. Scolman, and A. O. Nier, Phys. Rev. 102, 1071 (1956).

² T. T. Scolman, K. S. Quisenberry, and A. O. Nier, Phys. Rev.

<sup>C. F. Giese and J. L. Benson, Phys. Rev. 110, 712 (1958).
K. S. Quisenberry, T. T. Scolman, and A. O. Nier, Phys. Rev. 104, 461 (1956).
K. S. Quisenberry, C. F. Giese, and J. L. Benson, Phys. Rev. 107 (1978).</sup>

^{107, 1664 (1957).}

⁶ H. E. Duckworth, in Proceedings of the International Conference on Nuclidic Masses (University of Toronto Press, Hamilton, 1960),

Chap. 7, p. 446.

7 J. L. Benson, R. A. Damerow, and R. R. Ries, Phys. Rev. 113, 1105 (1959),

⁸ R. A. Damerow, R. R. Ries, and W. H. Johnson, following paper, Phys. Rev. 132, 1673 (1963).
⁹ H. Hintenberger, Nuclear Masses and Their Determination (Pergamon Press, Inc., London, 1957), Session VII, p. 185.

accurately plane reference surface greatly aided in the precise location of the various elements of the spectrometer. The spectrometer and surface plate are supported on four spring mounts, the resonant frequency of the supported mass being about 2 cps. This system then effectively isolates the spectrometer from the higher building vibration frequencies. In order to further minimize the effects of vibration, the spectrometer housing is mounted rigidly to the surface plate. Focusing adjustments are now made by moving the magnet on a track system. The modification also included redesign of several of the adjustable slits so as to give more precise control of the ion beam.

The changes made in the instrument have resulted in improved performance. The maximum usable resolution of the improved instrument is 2 or 3 times that of the original spectrometer. Other changes have considerably decreased the time necessary for focusing the instrument.

MEASUREMENTS

The mass spectrometer has the property that the mass of the ion collected is proportional to the resistance of a circuit element which determines the electric field in the instrument. Thus, one can show that the equation which relates doublet width to resistance is

$$\Delta M = M \Delta R / R \,, \tag{1}$$

where ΔM is the doublet width, M is the mass of one member of the doublet which corresponds to resistance value R, and ΔR is the change in resistance necessary to cause the second ion group to be collected. By measuring known doublets, Eq. (1) may be shown to be correct to a high degree of accuracy.

In the past, this instrument was employed mainly to measure narrow doublets made up of a hydrocarbon comparison ion of known mass and an ion of unknown mass at the same mass number. For the mass region under consideration in this paper, these doublets have $\Delta M/M$ values of approximately 10^{-3} . It has been shown that the instrument is capable of measuring wider doublets with sufficient accuracy to be worthwhile.6 One type of wide doublet is the doublet composed of two adjacent isotopes of the same element. Doublets of this sort are known as isotopic doublets and have been employed previously to supplement the hydrocarbon doublet data.

One of the difficulties that arises when hydrocarbon comparison ions are employed is caused by the rare isotope of carbon, C^{13} . When the hydrocarbon $C_m^{12}H_n^1$ is used as a comparison ion, a satellite ion is also present, $C^{13}C_{m-1}^{12}H_{n-1}^{1}$. In the mass region considered in this paper, a resolution of about 1/50 000 is required to completely resolve the C13 satellite. Resolution of this instrument is defined as $\Delta M/M$ where ΔM corresponds to the width at half-height of an ion peak of mass M. The resolution values ranged from 1/60 000 to 1/200 000 during this period. Thus, the C¹³ satellite was at all times completely resolved.

Table I. Mass doublets.

	Mass	
	differenceb	
Doublet ^a	(mu)	Errore
${}^{ ext{C}_5 ext{H}_9 ext{} ext{Ga}^{69}}_{ ext{C}_5 ext{H}_{11} ext{} ext{Ga}^{71}}$	144.852 7	24
$C_5H_{11}-Ga^{71}$	161.370 2	32
$C_5H_{10}-Ge^{70}$	154.001 3	22
$C_5H_{10}-Ge^{70}$ $C_4H_8O-Ge^{70}H_2$	117.616 1	18
$C_4H_8O-Ge^{72}$	135.438 4	21
$C_4H_0O-Ge^{73}$	141.878 4	21
$CS_2-Ge^{74}H_2$		14
$CS_2 - Ge^{76}$	7.314 0 22.741 6	15
Ge ⁷⁰ H ₂ — Ge ⁷²	17.821 3	17
${ m Ge^{70}H_2 - Ge^{72}} \ { m Ge^{72}H - Ge^{73}}$	6.443 9	13
$Ge^{73}H_2 - Ge^{74}H$	10.105 1	17
Co74H Co76	15.425 0	17
$Ge^{74}H_2^2 - Ge^{76}$ $C_3H_7O_2 - As^{75}$	123.009 8	26
C II C-74		
$C_6H_2-Se^{74}$	93.173 8	38
$C_6H_4 - Se^{76}$	112.099 9	81
$C_6H_5 - Se^{77}$	119.211 9	42
$C_6H_6 - Se^{78}$ $C_6H_8 - Se^{80}$	129.642 6	22
$C_6H_8-Se^{80}$	146.068 5	29
$C_6H_{10} - Se^{82}$	161.545 0	46
$C_{6}H_{8}-HBr^{79}$ $C_{6}H_{10}-HBr^{81}$ $Br^{81}-HBr^{79}$	136.444 3	24
$C_6H_{10}-HBr^{81}$	154.135 3	38
${ m Br^{81} - HBr^{79}}$	990.135 1	125
$\mathrm{C_6H_6}\mathbf{-Kr^{78}}$	126.584 3 ^d	36
${}^{\mathrm{C_6H_6-Kr^{78}}}_{\mathrm{C_6H_8-Kr^{80}}}$	$146.225 7^{d}$	46
$C_6H_{10}-Kr^{82}$	$164.769 \ 8^{d}$	34
C ₆ H ₁₁ — Kr ⁸³	$171.946 8^{d}$	34
C ₆ H ₁₂ -Kr ⁸⁴	182.399 4 ^d	25
$C_6H_{12}-Kr^{84} \ C_6H_{14}-Kr^{86} \ Kr^{83}-Kr^{82}$	198.9367^{d}	27
$Kr^{83} - Kr^{82}$	1000.647 9d	120
Kr84 - Kr83	997.371 6 ^d	120
$C_6H_{13}-Rb^{85}$	189.927 6	39
$C_4H_7O_2-Rb^{87}$	135.417 8	27
C-H Sr84	180.470 8	26
$C_6H_{12} - Sr^{84}$ $C_6H_{14} - Sr^{86}$	200.264 9	36
$C_4H_7O_2-Sr^{87}$	135.722 2	35
C H O - Sr88	146.789 1	47
$C_4H_8O_2 - Sr^{88}$ $Sr^{87} - Sr^{86}$	999.618 1	115
Sr ⁸⁸ — Sr ⁸⁷	996.739 6	116
	133.247 0	34
${}^{\mathrm{C_7H_5-Y^{89}}}_{\mathrm{C_4H_{10}O_2-Zr^{90}}}$		
$C_4H_{10}O_2-Zr^{90}$ $C_7H_7-Zr^{91}$	163.377 1	55 44
$C_7\Pi_7$ — $Z_7^{o_2}$	149.143 1	
$C_7H_8-Zr^{92}$	157.569 4	38
$C_7H_{10} - Zr^{94}$	171.929 4	39 57
$C_{7}H_{12}-Zr^{96}$ $Zr^{91}-Zr^{90}$	185.628 0	57
$Zr^{91}-Zr^{90}$	1000.942 0	116
$Zr^{92} - Zr^{91}$	999.397 2	117
$C_7H_9 - Nb^{93} C_7H_8 - Mo^{92}$	164.046 9	35
$C_7H_8-Mo^{92}$	155.790 0	32
$C_7H_{10} - M_0^{94}$	173.159 6	32
$C_7H_{11}^{10} - Mo^{95}$ $C_7H_{12}^{12} - Mo^{96}$	180.236 5	35
$C_7H_{12}-Mo^{96}$	189.226 9	30
$C_rH_rO_s - M_0^{97}$	122.937 6	23
$C_5H_6O_2 - Mo^{98}$ $C_7H_{16} - Mo^{100}$ $M_2O_2 - Mo^{94}$	131.375 4	28
$C_7H_{16}-Mo^{100}$	217.730 3	42
$Mo^{95} - Mo^{94}$	1000.757 2	122
$Mo^{96} - Mo^{95}$	998.838 5	124
$Mo^{97} - Mo^{96}$	1001.346 3	123
$Mo^{98} - Mo^{97}$	999.386 0	121

a Throughout this paper C, H, O, and S refer to C12, H1, O16, and S32, respectively.

b Mass differences are given in milliunits. All masses and mass differences.

b Mass differences are given in milliunits. All masses and mass differences in this work are measured in a scale in which the atomic mass of Croes is exactly equal to 12 units (symbol u). The symbol mu refers to one milliunit on the unified mass scale. The symbol pu refers to one microunit.

Throughout this paper the errors refer to the last significant figure of the particular result. The errors given in this table are taken from the original experimental data. The resulting error in an atomic mass calculation will be rounded off to the nearest µu.

For completeness, the krypton doublets have been included. They are taken from the M.S. thesis of R. R. Ries, University of Minnesota, 1959 (unpublished).

Metal ions for these measurements were obtained from a variety of sources. In two cases, gases were employed. These were hydrogen bromide and germanium tetrahydride. Pure metals were used for gallium, arsenic, and strontium. In the remaining cases, the following metallic compounds were employed: selenium dioxide, rubidium chloride, yttrium chloride, zirconium tetrachloride, niobium pentachloride, and molybdenum trioxide. Metals and metallic compounds were heated when necessary to obtain sufficient vapor pressure. Sample heating took place in a source furnace essentially the same as that previously described.3 Two baffles have been added between the sample and the ionization region to improve the vapor distribution.

RESULTS

The experimental mass doublet differences are presented in Table I. Both the narrow hydrocarbon-isotope doublets and the wide isotopic doublets are included in this table. The mass scale employed is the unified mass scale adopted by the International Union of Pure and Applied Physics in 1960. In this scale, the atomic mass of C12 is exactly 12 units (symbol u). The errors listed are standard errors and refer to the last quoted figure. These errors are calculated from resistor uncertainties and the standard error of the mean of the original data. The details of the resistor error assignments may be found in an article by Quisenberry et al.4

The doublet values of Table I are combined with the atomic masses of certain secondary mass standards in order to calculate atomic masses. Table II lists the atomic masses of the stable atoms in the mass region $69 \le A \le 100$. The secondary mass standards which were employed to calculate atomic masses from the doublet data are recorded in Table III. In some cases, the mass of a particular atom was overdeterminded by measurement of isotopic doublets in addition to the usual narrow hydrocarbon-isotope doublets. In these cases, a weighted least-squares adjustment was carried out on the data, and the best values found in this process are recorded as the results in Table II. The error associated with an atomic mass is the square root of the sum of the square of the error on the appropriate doublet value and the square of the error in the hydrocarbon reference mass. Whenever the least-squares adjustment was possible, the error on these atomic masses is just the error resulting from the least-squares calculation.

The final column of Table II provides a direct comparison between the present and other previous mass spectroscopic values. For comparison purposes, mass values were calculated from previous doublet results using the standard masses listed in Table III. The first half of this column contains the former Minnesota values determined by Collins et al.10 on a smaller, less

TABLE II. Atomic masses computed from the doublet data of Table I together with a comparison with previous mass spectroscopic values.

	Present r	esulta	1961 Mass	tableb	Other resu	ltsc
Isotope	u	Error	u	Error	u	Error
Ga^{69}	68.925 5	69 3	68.925 682	28	68.925 72d	4
Ga^{71}	70.924 7		70.924 840	50	70.92482^{d}	8
Ge^{70}	69.924 2	47 2	69.924 277	20	$69.924~00^{d}$	6
Ge^{72}	71.922 0	75° 2	71.921 740	50	71.921 60 ^d	5
Ge^{73}	72.923 4		72,923 360	70	$72.923 \ 34^{d}$	3
Ge^{74}	73.921 1	77 2	73.921 150	60	$73.921~00^{d}$	6
$\mathrm{Ge^{76}}$	75.921 4	02 2	75,921 360	90	$75.921 28^{d}$	4
$\mathrm{As^{75}}$	74.921 5	91 3	74.921 580	50	74.921 72d	4
Se^{74}	73.922 4	76 4	73.922 450	60	$73.922 54^{d}$	7
$\mathrm{Se^{76}}$	75.919 1	99 8	75.919 229	48	75.91927^{d}	4
$\mathrm{Se^{77}}$	76.919 9	12 4	76.919 934	48		
$\mathrm{Se^{78}}$	77.917 3	05 2	77.917 348	48		
$\mathrm{Se^{80}}$	79.916 5	29 3	79.916 512	17	79.916 47 ^d	4
Se^{82}	81.916 7	02 5	81.916 660	70	81.916 64 ^d	4
$ m Br^{79}$	78.918 3	28 3	78.918 348	19	78.918 40 ^d	5
Br^{81}	80.916 2		80.916 344	37	80.916 42 ^d	5
$ m Kr^{78}$	77.920 3	64 4	77.920 368	5	77.920 19 ^d	8
$ m Kr^{80}$	79.916 3	72 5	79.916 388	13		
Kr^{82}	81.913 4		81.913 483	8	81.913 46 ^d	6
$ m Kr^{83}$	82.914 1	25 4	82.914 131	8	$82.914~06^{d}$	5
$ m Kr^{84}$	83.911 4		83.911 504	5	83.911 51 ^d	5
Kr^{86}	85.910 6		85.910 617	8	85.910 80 ^d	6
$\mathrm{Rb^{85}}$	84.911 7		84.911 710	60	84.912 03 ^d	6
Rb^{87}	86.909 1		86.909 180	80	86.909 31 ^d	17
Sr ⁸⁴	83.913 4	25 3	83.913 376	11	83.913 399f	17
					83.913 25 ^d	15
Sr^{86}	85.909 2	78 4	85.909 260	80	85.909 156f	23
,					85.909 3 6 ^d	10
					85.909 357g	50
Sr ⁸⁷	86.908 88	82 4	86.908 890	80	86,908 816f	42
					86.908 99 ^d	6
Sr88	87.905 6	34 5	87.905 610	90	87.905 485f	16
					87.905 678g	
					87.906 01 ^d	11
Y^{89}	88.905 8	76 4	88.905 430	90	88.905 862f	15
					88.905 72 ^d	11
$ m Zr^{90}$	89.904 6	96 5	89.904 320	90	89.904 672 ^f	20
					89.904 931g	
					89.904 33 ^d	25
Zr^{91}	90.905 6		90.905 250		90.905 629f	20
Zr^{92}	91.905 0		91.904 590		91.905 093 ^f	22
Zr^{94}	93.906 3		93.906 140		93.906 268f	25
$\mathrm{Zr^{96}}$	95.908 2		95.908 200		95.908 379 ^f	46
$\mathrm{Nb^{93}}$	92.906 3	7 5 4	92.906 020	110	92.906 315f	25
		a = .	04 004 000	420	92.905 66 ^d	8
Mo^{92}	91.906 8		91.906 290		91.906 869f	42
Mo94	93.905 0		93.904 740	130	93.905 166 ^f	52
$\mathrm{Mo^{95}}$	94.905 8		94.905 720	360	94.905 841f	15
$\mathrm{Mo^{96}}$	95.904 6		95.904 550	360	95.904 685f	44
Mo^{97}	96.906 0		96.905 750	400	96.905 952f	30
Mo^{98}	97.905 4		97.905 510	410	97.905 425 ^f 99.907 543 ^f	16 35
$ m Mo^{100}$	99.907 4	64 5	99.907 570	490	99,907 343*	33

^a The atomic masses are computed from the doublet values of Table I and the values of the secondary mass standards listed in Table III.
^b See Ref. 12.
^c In the original references, these results were presented in the old Ols scale. They have been changed to comply to the new Cl² mass scale.
^d See Ref. 10.

precise mass spectrometer. These values have errors approximately 10 times larger than those of the present investigation. No systematic difference seems to be evident for the 28 atomic mass comparisons which can be made, since 14 of the former results are higher and

¹⁰ T. L. Collins, W. H. Johnson, and A. O. Nier, Phys. Rev. 94, 398 (1954).

See Ref. 10.
 This result is the weighted average of the two narrow Ge⁷² doublets listed in Table I.
 See Ref. 11.
 S.N. R. Isenor, R. C. Barber, and H. E. Duckworth, Can. J. Phys. 38, 819 (1960).

14 lower than the present work. Because of the comparatively large error of the previous results, the comparison of results is not particularly valuable as a test of the reliability of the present data.

Starting with strontium at A = 84, a comparison is made with the work carried out by Demirkhanov et al.11 These results have errors that are about 5 times those quoted for the present results. A comparison of the 14 stable isotopic masses from Y89 to Mo100 reveals moderate agreement between the present results and the results of Demirkhanov. This order of agreement does not hold in the case of several of the strontium masses (Sr⁸⁶ and Sr88) where the disagreement between the two results to more than 5 times their combined errors. Smaller disagreements occur at Zr⁹², Zr⁹⁶, and Nb⁹³. The reason for these discrepancies is unknown.

The present results are also compared with the values found in the Mass Table due to König et al. 12 The results of König represent "best" values from a least-squares adjustment of much of the available nuclear reaction and disintegration data, together with selected mass spectroscopic information. It should be mentioned that in the mass region from A = 69 to A = 93, the mass results of Collins et al. 10 served as the mass spectroscopic input data for the least-squares adjustment of the Mass Table. The comparison of the present results with this table is quite good from A = 73 to A = 88. A particularly encouraging feature of this table is the fact that, with three exceptions, the table always adjusted the older Minnesota results of Collins in the right direction to agree with the more precise, present results. Until recently, precise mass information from A = 88 to A = 100 was not available. This is reflected in the poor agreement of the Mass Table results with the present work in this mass region. The discrepancies are particularly obvious for the zirconium and molybdenum

Table III. Secondary mass standards.

	Present va	lue	Refer
Standard	u	Error	ence
C^{12}	12.000 000 0		а
$\mathbf{H}^{_1}$	1.007 824 7	2	ь
n^{e}	$1.008\ 665\ 4$	4	d
H^2	2.014 102 2	1	d
N^{14}	14.003 073 1	4	е
O^{16}	15.994 914 2	5	b
S^{32}	31.972 073 8	11	d
Cl^{35}	34.968 853 1	19	f

TABLE IV. Least-squares adjusted values for overdetermined stable isotopes and isotopic differences.

Isotope or isotopic difference	Measured re u l	sultª Error		djusted ^b sult Error	Differ- ence ^c µu
Ge ⁷⁰	69.924 246	2	69.924	247 2	+1
Ge ⁷²	71.922 073	3	71.922		+1
Ge ⁷³	72.923 458	3	72.923		-1
Ge ⁷⁴	73.921 180	3 3 2 2 2 2 2 3	73.921	177 2	$-\frac{1}{3}$
Ge ⁷⁶	75.921 402	3	75.921	102 2	-3
$Ge^{72} - Ge^{70}$	1.997 828	2	1.997		0
$Ge^{73} - Ge^{72}$	1.001 381	2	1.001	382 2	+1
$Ge^{74} - Ge^{73}$	0.997 720	2	0.997	720 2	71
$Ge^{76} - Ge^{74}$	2.000 224	2	2.000	224 2	0
Br ⁷⁹	78.918 328	2	78.918	229 2	0
Br ⁸¹	80.916 287	4	80.916	287 4	0
Br ⁸¹ —Br ⁷⁹	1.997 960	12	1.997		-1
Kr ⁸²	81.913 477		81.913		
Kr ⁸³		4			0
	82.914 125	4	82.914		0
Kr ⁸⁴ Kr ⁸³ – Kr ⁸²	83.911 497	3	83.911		0
	1.000 648	12	1.000	048 5	0
Kr84 - Kr83	0.997 372	12	0.997	3/2 5	0
Sr86	85.909 280	4	85.909		-2
Sr ⁸⁷	86.908 879	4	86.908		+3
Sr88	87.905 637	.5	87.905		-3
$Sr^{87} - Sr^{86}$	0.999 618	12	0.999	603 5	-15
$Sr^{88} - Sr^{87}$	0.996 740	12	0.996		+12
Zr^{90}	89.904 698	6	89.904		-2
Zr^{91}	90.905 630	5	90.905		+1
$ m Zr^{92}$	91.905 028	4	91.905		0
$Zr^{91}-Zr^{90}$	1.000 942	12	1.000		 7
$Zr^{92} - Zr^{91}$	0.999 397	12	0.999		0
$\mathrm{Mo^{94}}$	93.905 087	4	93.905		-1
$\mathrm{Mo^{95}}$	94.905 835	4	94.905	835 4	0
$\mathrm{Mo^{96}}$	95.904 669	4	95.904		+1
$\mathrm{Mo^{97}}$	96,906 014	3	96.906	014 3	0
$\mathrm{Mo^{98}}$	97.905 401	3	97.905		0
${ m Mo^{95}-Mo^{94}}$	1.000 757	12	1.000	749 5	-8
${ m Mo^{96}-Mo^{95}}$	0.998 838	12	0.998		-4
$Mo^{97} - Mo^{96}$	1.001 346	12	1.001		-1
$Mo^{98} - Mo^{97}$	0.999 386	12	0.999	387 4	+1

^a Calculated from the doublet data of Table I and the standard masses listed in Table III. These results serve as input data for the least-squares adjustment.

isotopes, where the mass table values also have very large errors.

Hydrocarbon mass unit doublets of the form $C_mH_n-C_mH_{n-1}$ were measured concurrently with the doublets of Table I in order to make a continuous check on the validity of the dispersion relation, Eq. (1). The average of these mass unit measurements, taken during the measurement of any one element, was used to calculate a value for the hydrogen atomic mass from Eq. (1); and this result was compared with the accepted value for hydrogen listed in Table III. The ratio of the accepted value of the hydrogen mass to the measured value of the hydrogen mass, as calculated from the mass unit doublet, is called the dispersion constant. Thus, when the mass unit measurements lead to results which agree on the average with the accepted mass of hydrogen, the dispersion constant is just equal to unity, and Eq. (1) is used as it stands. However, any deviation from the accepted value of hydrogen defines a dispersion constant different from unity. All doublets which are measured concurrently with this particular mass unit

 $^{^{\}rm a}$ This is the definition of the C $^{\rm 12}$ mass scale. $^{\rm b}$ See Ref. 5. $^{\rm o}$ The symbol n stands for the neutron mass. $^{\rm d}$ See Ref. 12. $^{\rm c}$ See Ref. 2. $^{\rm f}$ See Ref. 3.

¹¹ R. A. Demirkhanov, V. V. Dorokhov, and M. I. Dzkuya, Zh. Eksperim. i Teor. Fiz. 40, 1572 (1961) [translation: Soviet Phys.—JETP 13, 1104 (1961)].

¹² L. A. König, J. H. E. Mattauch, and A. H. Wapstra, Nucl. Phys. 31, 18 (1962).

h These results and errors are the least-squares adjusted values.
 Adjusted data minus originally measured results (in microunits).

TABLE V. Comparison of mass spectrometer and nuclear reaction mass differences for stable isotopes.

Mass	Present re		Nuclear re result		Reaction	Refer-
difference	u	Error	u	Error	employed	enceb
Ge ⁷³ – Ge ⁷²	1.001 382	2 2	1.001 687	172	$Ge^{73}(\gamma,n)Ge^{72}$	c
Se ⁷⁷ - Se ⁷⁶	1.000 713	9	1.000 703	10	$\mathrm{Se}^{76}(n,\gamma)\mathrm{Se}^{77}$	d
Se ⁷⁸ - Se ⁷⁷	0.997 394	1 5	0.997 411	15	$\mathrm{Se}^{77}(n,\gamma)\mathrm{Se}^{78}$	d
Kr84 – Kr83	0.997 372	2 5	0.997 317	21	$\mathrm{Kr}^{83}(n,\gamma)\mathrm{Kr}^{84}$	5-2-10, 6
Rb87 Sr87	0.000 30	1 5	0.000 293	3	$Rb^{87}(\beta^{-})Sr^{87}$	60-3-61
			0.000 261	32	$Rb^{87}(p,n)Sr^{87}$	e
Sr87 - Sr86	0.999 603	3 5	0.999 629	19	$Sr^{86}(n,\gamma)Sr^{87}$	d
			0.999 540	215	$Sr^{86}(d,p)Sr^{87}$	60-3-62
Sr88 - Sr87	0.996 752	2 6	0.996 737	15	$\mathrm{Sr}^{87}(n,\gamma)\mathrm{Sr}^{88}$	d
$Zr^{91} - Zr^{90}$	1,000 935	7	1.000 888	32	$Zr^{90}(d,p)Zr^{91}$	f
$Zr^{92} - Zr^{91}$	0.999 397	7 6	0.999 332	32	$Zr^{91}(d,p)Zr^{92}$	g
			0.999 368	43	$Zr^{91}(n,\gamma)Zr^{92}$	d
$Mo^{96} - Mo^{95}$	0.998 834	1 5	0.998 831	11	$Mo^{95}(n,\gamma)Mo^{96}$	d
Mo ⁹⁷ - Mo ⁹⁶	1.001 345	6 4	1.001 436	322	$Mo^{96}(d,p)Mo^{97}$	h
			1.001 043	322	$Mo^{97}(\gamma, n)Mo^{96}$	60-6-49
Mo98 - Mo97	0.999 387	7 4	0.999 772	107	Mo97(d,p) Mo98	h

^aThese results were calculated from the appropriate Q value and the masses listed in Table III. ^bThe series of numbers in this column refer to the page numbers of the Nuclear Data Sheets of Ref. 16.

doublet are then corrected by the appropriate amount. Dispersion constants were determined for each element during this investigation and they ranged from $0.999\ 967\pm 2$ to $1.000\ 015\pm 3$. The average value of this constant for the entire block of measurements was $0.999997\pm 2.$

The effect of the least-squares adjustment on the data where isotopic masses were overdetermined can be seen in Table IV. The second column lists the results which are calculated directly from the doublet data and which then serve as the input data for the adjustment. The third column contains the results of the least-squares process, while the last column records the difference between the adjusted value and the input data. It can be seen that the atomic masses change at most by 2 or 3 microunits, and these changes are always covered by the original errors. The adjusted errors on the atomic mass values generally remain the same or become one microunit smaller. The isotopic mass differences are generally adjusted by a larger amount, as in the case of strontium where the two isotopic mass differences are changed by 12 and 15 microunits. With the exception of the $\mathrm{Sr}^{87}\mathrm{-Sr}^{86}$ doublet, these adjustments are also covered by the original errors.

A comparison of mass differences determined both from mass spectroscopic measurements and from nuclear reaction energies is contained in Table V. This table records mass differences only when nuclear reactions connecting 2 stable atoms are available. The conversion factor between Q values and mass units is taken from Cohen¹³ and is given by 1 unit=931.476 ± 0.004 MeV.

The number of comparisons in this mass region which involve only 2 stable atoms is not very large, but the over-all agreement is quite good. The agreement is particularly good for recent precise nuclear reaction results. Particular mention should be made of the excellent agreement between the present results and the (n,γ) reaction results determined by Kinsey and Bartholomew.¹⁴ With the exception of the Sr⁸⁷-Sr⁸⁶ mass difference, all of these results agree with the present work within the combined errors. This comparison also verifies the gamma spectrum assignments made by Kinsey and Bartholemew when determining the groundstate Q values.

Nuclear reaction and disintegration energies are combined with the stable mass data of Table II in order to calculate the atomic masses of radioactive atoms. The results of such calculations are recorded in Table VI, which lists a final adopted mass for 65 radioactive atoms in the mass region from A = 67 to A = 100. A schematic view of this mass region, exhibiting the various paths by which the radioactive masses may be calculated from the stable atomic masses, is presented in Fig. 1. Several zinc masses from the 1961 Mass Table¹¹ were utilized to calculate some radioactive gallium masses.

As is evident from Fig. 1, some radioactive masses may be calculated from the stable masses in only one way, while others may be determined by several reaction paths. When only one such path is available, the final adopted value is just the calculated result, so this entry is included only in the final column. In the 25 cases where the radioactive mass can be calculated in more than one way, the result of each calculation is recorded in the second column. The final adopted value for such a radioactive mass is then the weighted average of the several individual results.

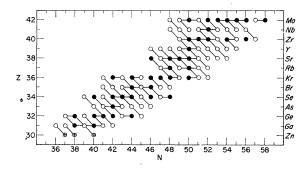


Fig. 1. Reaction scheme showing the stable nuclei whose atomic masses have been measured in this investigation. Solid circles represent stable nuclei, open circles represent radioactive nuclei, and connecting lines indicate available nuclear reaction and B-decay mass differences. The three crossed circles represent zinc isotopes not measured in this work.

See Ref. 17.
d See Ref. 14.

A. J. Elwyn, H. H. Landon, S. Oleksa, and G. N. Glasoe, Phys. Rev. 2, 1200 (1958).
R. L. Preston, H. J. Martin, Jr., and M. B. Sampson, Phys. Rev. 121, 41 (1961). 1301). 1. Martin, Jr., M. B. Sampson, and R. L. Preston, Phys. Rev. 125, (1962).

^{942 (1962).}h N. S. Wall, Phys. Rev. 96, 664 (1954).

¹³ E. R. Cohen, Bull. Am. Phys. Soc. 7, 305 (1962). ¹⁴ B. B. Kinsey and G. A. Bartholomew, Can. J. Phys. 31, 1051 (1953).

Table VI. Atomic masses of radioactive isotopes computed from measured isotopic masses of Table II and available nuclear reaction and beta-decay energies.

	Calculated	massa	Nuclear reaction		Final add			Calculated	massa	Nuclear reaction		Final ado	
Isotope		Error	employed	Reference ^b	u	Error	Isotope		Error	employed	Referenceb	. u	Erro
Ga ⁶⁷			$Zn^{67}(p,n)Ga^{67}$	d	66.928 220	12		83.914 561	s 86	$Rb^{85}(\gamma,n)Rb^{84}$	1		
Ga68			$Zn^{68}(p,n)Ga^{68}$	d	67.927 996	11	Rb86	85.911 186	7	Rb86 (β-)Sr86	60-3-50	85.911 186	
Ga70	69.925 936	22	$Ga^{69}(n,\gamma)Ga^{70}$	d			l	85.911 173	43	$Rb^{87}(\gamma, n)Rb^{86}$	60-3-61		
	69.926 018	11	Ga70(B-)Ge70	d			Rb88	87.911 217	107	Rb ⁸⁸ (β ⁻)Sr ⁸⁸	60-3-71	87.911 260	9
	69.926 048	17	$Zn^{70}(p,n)Ga^{70}$	d	69.926 012	20	l	87.911 435	215	$Rb^{87}(d,p)Rb^{88}$	60-3-71		
	69,925 956	65	$Ga^{71}(\gamma,n)Ga^{70}$	е			Rb89			Rb89(8-)Sr89	60-3-82	88.911 654	5
Ga72	71.926 359	11	Ga ⁷² (β ⁻)Ge ⁷²	e			Sr85	84,912 959	322	Sr86 (γ.n)Sr85	60-3-52	84,912 981	3.
	71.926 372	9	$Ga^{71}(n,\gamma)Ga^{72}$	t	71.926 367	7		84,912 981	33	Rb85(p,n)Sr85	60-3-42		
	71.925 8838	19	$Ga^{71}(n,\gamma)Ga^{72}$	•			Sr89	88.907 447	6	Sr89 (B-) Y89	60-3-83	88.907 446	
Ga73			Ga ⁷³ (β ⁻)Ge ⁷³	59-1-37	72.925 123	43		88,907 339	75	$Sr^{88}(d,p)Sr^{89}$	60-3-83		
Ge69	68.927 956	5	Ga69(p,n)Ge69	h			Sr90			Sr90(B-)Y90	60-4-33	89,907 755	3:
	68.927 971	13	$Ge^{69}(\beta^{+})Ga^{69}$	d	68.927 958	5	Sr91			Sr91 (B-) Y91	60-5-74	90.910 150	10
	68.928 5718		$Ge^{70}(\gamma, n)Ge^{69}$	i			Sr92			Sr92(B-)Y92	60-5-84	91.910 949	
Ge71	70.924 966	32	$Ga^{71}(p,n)Ge^{71}$	59-2-44			Y85			Y ⁸⁵ (β ⁺)Sr ⁸⁵	j	84,916 728	
	70.924 951	5	Ge71(e)Ga71	59-2-43	70.924 951	5	Y86			Y86 (β+)Sr86	i	85.914 968	
Ge75	74.922 858	22	Ge ⁷⁵ (β ⁻)As ⁷⁵		74,922 857	22	¥87			$Sr^{87}(p,n)Y^{87}$	60-3-63	86,910 693	
	74.922 720	215	$Ge^{76}(\gamma,n)Ge^{75}$	ė			Y88	87.909 526	12	$Y^{88}(\beta^{+})Sr^{88}$	m	00.020	
Ge77			Ge ⁷⁷ (β ⁻)As ⁷⁷	59-3-44	76.923 598	55	*	87.909 329		Y ⁸⁸ (β ⁺)Sr ⁸⁸	n		
As73			$Ge^{73}(p,n)As^{73}$	0 7 0 11	72.923 855	32		87,909 654	86	$Y^{89}(\gamma,n)Y^{88}$	1	87,909 522	8
As74	73,923 936	11	As ⁷⁴ (β ⁻)Se ⁷⁴	59-4-71				87.909 9018		$\mathbf{Y}^{89}(\gamma,n)\mathbf{Y}^{88}$. 0	0.1.505 022	
	73.923 938	31	$As^{75}(\gamma,n)As^{74}$	e .	73.923 932	8		87.909 519	8	$Sr^{88}(p,n)Y^{88}$	p		
	73.923 928	11	$As^{74}(\beta^{+})Ge^{74}$	e	.0.,20 ,02	-	V90	89.907 133	12	$Y^{90}(\beta^{-})Zr^{90}$	60-4-34		
As^{76}	75.922 387	13	As ⁷⁶ (β ⁻)Se ⁷⁶	59-5-39	75,922 390	12	1	89.907 189	10	$\mathbf{Y}^{89}(n,\gamma)\mathbf{Y}^{90}$	q	89,907 171	32
113	75.922 420	43	$As^{75}(n,\gamma)As^{76}$	59-5-41	13.722 070	12		89.907 420	54	$Y^{89}(d,p)Y^{90}$	r	09.907 171	32
As^{77}	75.722 420	40	$As^{77}(\beta^{-})Se^{77}$	59-3-46	76,920 646	11	Y91	89.907 420	34	$Y^{91}(\beta^{-})Zr^{91}$	60-5-75	90.907 284	1:
As78			As ⁷⁸ (β ⁻)Se ⁷⁸	59-5-48, 50	77.921 889	215	Y92			$Y^{92}(\beta^{-})Zr^{92}$	60-5-85	91.908 893	
As79			As ⁷⁹ (β ⁻)Se ⁷⁹	59-2-49	78,920 969	107	Y93			$Y^{93}(\beta^{-})Zr^{93}$	60-5-98	92.909 547	24
Se ⁷³			Se ⁷⁸ (β+)As ⁷⁸	59-1-41	72.926 807	34	Y94			$Y^{94}(\beta^{-})Zr^{94}$	60-5-11	93.911 685	
Se ⁷⁵			$As^{75}(p,n)Se^{75}$	59-1-48	74.922 520	4	Zr89	88.908 923	8	Zr ⁸⁹ (β ⁺)Y ⁸⁹	60-3-11	93.911 003	21.
Se ⁷⁹			$Se^{79}(\beta^{-})Br^{79}$	59-2-50	78.918 500	6	ZIos		97	$Zr^{90}(\gamma,n)Zr^{89}$	60-4-36	88,908 923	. 10
Se ⁸¹			Se ⁸¹ (β ⁻)Br ⁸¹	59-1-64	80.917 790	54		88.908 678		*** *	60-3-85	00.900 923	
Br ⁷⁶			Br ⁷⁶ (β ⁺)Se ⁷⁶	59-5-43	75.924 127	13	7.00	88.908 923	8	$Y^{89}(p,n)Zr^{89}$	60-5-99	92.906 444	10
Br ⁷⁷			$Se^{77}(p,n)Br^{77}$	59-3-49	76.924 127	6	Zr ⁹³	92.906 443	4	Zr ⁹³ (β-)Nb ⁹³	60-5-99	92.900 444	10
Br78	77.921 156	11		59-5-52	77.921 156	11		92.906 550	43	$Zr^{92}(d,p)Zr^{93}$		04 000 027	13
DI	77.921 136		$Se^{78}(p,n)Br^{78}$ $Br^{79}(\gamma,n)Br^{78}$	8	77.921 130	11	Zr ⁹⁵	94.908 097	54	$Zr^{94}(d,p)Zr^{95}$	60-5-127	94.908 037	13
Br80		150 12	$Br^{80}(\beta^-)Kr^{80}$	59-1-56				94.908 034	12	$Zr^{95}(\beta^{-})Nb^{95}$	60-5-127	06.010.011	33
BLon	79.918 519	12 7	(1- /	59-1-56	79.918 544	10	Zr97			Zr ⁹⁷ (β ⁻)Nb ⁹⁷	60-6-47	96.910 944	
	79.918 556		Br ⁸⁰ (β ⁺)Se ⁸⁰	39-1-30 e	19,910 344	10	Nb89			$Nb^{89}(\beta^{+})Zr^{89}$	60-3-86	88.913 088	98 12
	79.918 496	38 14	$Br^{81}(\gamma,n)Br^{80}$				Nb90			$Nb^{90}(\beta^{+})Zr^{90}$	60-4-37	89.911 255	
T) . 00	79.918 535	14	$Br^{79}(n,\gamma)Br^{80}$		01 016 706	4	Nb91			$Mo^{91}(\beta^+)Nb^{91}$	60-5-78	90.906 795	82
Br82			Br ⁸² (β ⁻)Kr ⁸²	59-1-70	81,916 796		Nb92	91.906 871		$Zr^{92}(p,n)Nb^{92}$		04.007.404	
Br83			Br ⁸³ (β ⁻)Kr ⁸³	59-1-80	82.915 199	22		91.907 186	41	$Nb^{93}(\gamma,n)Nb^{92}$	60-5-100	91.907 184	47
Kr77			Kr ⁷⁷ (β ⁺)Br ⁷⁷		76.924 468	22		91.907 325	161	$Nb^{92}(\beta^{+})Zr^{92}$	60-5-81		
Kr79	78.920 221		Kr ⁷⁸ (d,p)Kr ⁷⁹	59-2-53	78.920 068	14	Nb94	93.907 322	32	$Nb^{98}(n,\gamma)Nb^{94}$	60-5-114		
	78.920 067	6	Kr ⁷⁹ (β ⁺)Br ⁷⁹	59-2-52	00 044 405	407		93.907 253	107	$Nb^{93}(d,p)Nb^{94}$	60-5-114	93.907 314	27
Kr81			$Kr^{80}(d,p)Kr^{81}$	58-1-39	80.916 605	107		93.907 308	54	$Nb^{94}(\beta^{-})Mo^{94}$	60-5-113		
Kr85	84.912 514	9	Kr ⁸⁵ (β ⁻)Rb ⁸⁵	60-3-40	84.912 516	6	Nb95			$Nb^{95}(\beta^{-})Mo^{95}$	60-5-128	94.906 832	6
	84.912 519	9	Kr84(d,p)Kr85	j			Nb^{97}			${\rm Nb^{97}}(\beta^{-}){\rm Mo^{97}}$	60-6-48	96.908 088	3.
Kr87	86.913 499		$Kr^{87}(\beta^{-})Rb^{87}$	60-3-60	86.913 363	52	Mo90			$Mo^{90}(\beta^+)Nb^{90}$	60-4-40	89.913 982	108
	86.913 354	54	$Kr^{86}(d,p)Kr^{87}$	j			Mo^{91}			$Mo^{92}(\gamma,n)Mo^{91}$	60-5-88, 81	90.911 583	75
Kr88			$Kr^{88}(\beta^{-})Rb^{88}$	60-3-70	87.914 266	235	Mo^{98}	92.906 898	43	$Nb^{93}(p,n)Mo^{93}$	60-5-101	92,906 953	70
Rb84	83.914 392		$Rb^{84}(\beta^-)Sr^{84}$	5-2-11			1	92.907 041	54	${ m Mo^{92}}(d,p){ m Mo^{93}}$	60-5-101		
	83.914 363	11	$Rb^{84}(\beta^{+})Kr^{84}$	5-2-11	83.914 374	10	Mo99			${ m Mo^{100}}(\gamma,n){ m Mo^{99}}$	61-1-66	98.907 495	322
	83.914 411	86	$Rb^{85}(\gamma,n)Rb^{84}$	k			l						

In 15 of the 25 cases where several reactions were available to determine one radioactive mass, the results agree very well; see, for example, As74. When the several independent reaction and disintegration energies combined with different stable atomic masses lead to these similar results, both the energy values and the mass

^a For comparison purposes, this column lists the results of the calculations whenever more than one nuclear reaction was available for a particular isotope. No entry is made in this column when only one reaction energy is used. In these cases, the final adopted value is also the calculated value.

^b Whenever possible, references are taken from large compilations of nuclear data. References to the original literature are made if that is the only source or if the experimental results lead to calculated masses which differ by large amounts. The sequence of numbers in this column refer to the year, the set, and the page numbers of the Nuclear Data Sheets of Ref. 16.

^a The final adopted value is the weighted average of the various calculated results from different reactions. When only one reaction energy is available, the adopted value is simply the calculated result. The error associated with the adopted values is the larger of the internal error and the external error as defined in the Nuclear Data Sheets.

^d F. Everling, L. A. König, J. H. E. Mattauch, and A. H. Wapstra, Nucl. Phys. 25, 177 (1961).

<sup>L. A. König, J. H. E. Mattauch, and A. H. Wapstra, Nucl. Phys. 28, 1 (1961).
See Ref. 15.
This result is not included in the weighted average.
See Ref. 12.
D. M. Van Patter and Ward Whaling, Rev. Mod. Phys. 29, 757 (1957).
A. H. Wapstra (private communication).
See Ref. 17.
See Ref. 18.
See Ref. 20.
See Ref. 20.
See Ref. 21.
Ge Ref. 21.
G. A. Bartholomew, P. J. Campion, J. W. Knowles, and G. Manning, Nucl. Phys. 10, 590 (1959).
N. S. Wall, Phys. Rev. 96, 670 (1954).</sup>

values are confirmed. In the remaining 10 cases, different reaction paths lead to results that do not agree within the combined error. Several of these cases will now be discussed individually.

The energies of the gamma rays in the complex spectrum resulting from the gallium (n,γ) reaction were measured by Campion et al.,15 but the unambiguous assignment of the ground-state gamma energy of Ga⁷² was not possible. A listing of the unassigned gamma rays was given by these authors. An estimate of 7.1 MeV for the last neutron binding energy in Ga⁷² was given in the Nuclear Data Sheets.16 This would suggest that the unassigned 6.971-MeV gamma corresponds to the groundstate reaction. However, this Q value leads to a Ga⁷² mass which differs by about 475 keV from the result calculated from the beta decay of Ga⁷² to Ge⁷². The level schemes for these 2 isotopes seem to be well known and thus we have chosen the beta disintegration energy in order to calculate the mass of Ga⁷². Assuming that result as correct, we have then assigned the 6.516-MeV gamma as the ground-state gamma for the $Ga^{71}(n,\gamma)Ga^{72}$ reaction.

The (γ,n) reactions for Rb⁸⁴ are available. This reaction is difficult to interpret because of the large nuclear spin change involved. The value by Tobin et al. 17 determined by studying the reaction to the isomeric state of Rb⁸⁴ is preferred to the value by Geller et al.¹⁸ because of its agreement with the two well-known beta decays of Rb84.

The positron decay energy of Y88 to Sr88 has been a controversial value for some time. Two distinct groups of results were found for this disintegration energy, and they are best represented by Ramaswamy et al.19 who quotes the value 3442±31 keV, and by Rhode et al.20 who determined the value 3625±10 keV. Two conflicting $\mathbf{Y}^{89}(\gamma,n)\mathbf{Y}^{88}$ reaction results were not previously helpful in resolving the discrepancy because of the large error in the mass values for Y89. The recent and precise $Sr^{88}(p,n)Y^{88}$ reaction measured by Shafroth, ²¹ however, gives a result which agrees closely with the disintegration energy of Rhode. In addition, the present mass determination of Y⁸⁹ permits a more precise calculation of Y⁸⁸ from the two (γ,n) reactions. The mass result calculated from the Q value of Geller et al. 18 agrees reasonably well with the results due to Rhode and

Shafroth. The (γ,n) result of Chidley et al.²² is far outside the error of these values, and still further from the result of Ramaswamy's measurement. The final adopted value chosen here for the mass of Y88 is the weighted average of the three results due to Rhode, Shafroth, and Geller.

In the remaining cases where the agreement between the several possible calculations was not good, evidence did not seem to favor one value over another, so the final adopted value in these cases is just the weighted average.

NUCLEAR SYSTEMATICS

The atomic masses of stable and radioactive isotopes listed in Table II and Table IV may be employed to calculate the nuclear binding energies for various combinations of nucleons. In order to calculate the total nuclear binding energy (TNBE), the proper atomic mass value is substituted into Eq. (2)

TNBE
$$(Z,N) = ZM_H + NM_n - zM_N^A - E_b(Z,N)/c^2$$
, (2)

where $M_{\rm H}$ is the hydrogen mass, M_n is the neutron mass, zM_N^A is the atomic mass of the atom with Z protons and N neutrons. The term $E_b(Z,N)$ is an estimate of the total electronic binding energy. The functional form of this relation was taken from Foldy.23 The value ranged from 51 μ u for gallium to 104 μ u for molybdenum, with a stated accuracy of 10%. Consideration of the small binding energy of the electron in hydrogen was neglected in view of the much larger uncertainty in the value of $E_b(Z,N)$.

Values of the total nuclear binding energy are useful in the calculation of various binding energy quantities of interest in the study of nuclear systematics. One quantity which indicates the general tendency of nuclear binding is the average binding energy, TNBE(Z,N)/A. The binding energy of a given system of nucleons in a particular nucleus may be found by forming the appropriate difference between two total binding energy terms. As an example, the neutron separation energy and the proton separation energy are found from Eq. (3) and Eq. (4), respectively.

$$S_n(Z,N) = \text{TNBE}(Z,N) - \text{TNBE}(Z,N-1),$$
 (3)

$$S_p(Z,N) = \text{TNBE}(Z,N) - \text{TNBE}(Z-1,N)$$
. (4

Table VII lists the total nuclear binding energy, the average nuclear binding energy, the neutron separation energy, and the proton separation energy in nuclei for which these quantities may be calculated from the present mass values.

In Fig. 2, the average nuclear binding energy for stable isotopes is plotted as a function of the mass number. The even-A points for each element are connected by a solid line and the odd-A points for all the elements are connected by a dashed line. The character-

¹⁵ P. J. Campion and G. A. Bartholomew, Can. J. Phys. 35, 1361 (1957).

 ¹⁶ Nuclear Data Sheets, compiled by K. Way et al. (Printing and Publishing Office, National Academy of Sciences-National Research Council, Washington 25, D. C., 1958-1961).
 ¹⁷ R. Tobin, J. McElhinney, and L. Cohen, Phys. Rev. 110, 1388

^{(1958).} 18 K. N. Geller, J. Halpern, and E. G. Muirhead, Phys. Rev. 118,

^{1302 (1960).} 19 M. H. Ramaswamy and P. S. Jastram, Nucl. Phys. 19, 243

<sup>(1960).

20</sup> J. I. Rhode, O. E. Johnson, and W. G. Smith, Phys. Rev. 129, 815 (1963).

21 S. M. Shafroth, Nucl. Phys. 28, 649 (1961).

²² B. G. Chidley, L. Katz, and S. Kowalski, Can J. Phys. 36, 407 (1958).
²³ L. L. Foldy, Phys. Rev. **83**, 397 (1951).

Table VII. Total nuclear binding energy, average binding energy per nucleon, and nucleon separation energies for those nuclei between A=67 and A=100 where sufficient data is available.

Isotope	. Z	N	TNBEa mu			S _n	c Error	S mu	p ^d Error	Isotope	. Z	N	TNBE ^a			S _n		S _p	d Error
Zn66	30	36	620,600°							Kr86		50	804.278	9.3521					
Zn^{67}	30	37	628.165							Kr87	36 36	51	810.190	9.3321	3 7	10.572 5.912	10 52		
Zn^{68}	30	38	639.114°							Kr88	36	52	817.952	9.2949		7.762			
Zn^{69}	30	39	645.992°							Rb84	37	47	782.337	9.3135	3			7.571	11
Zn^{70}	30	40	655.962e							Rb ⁸⁵	37	48	793.583	9.3363		11.246	11	7.523	5
$ m Zn^{71} \ Zn^{72}$	30 30	$\begin{array}{c} 41 \\ 42 \end{array}$	662.005e							Rb86	37	49	802.856	9.3355	3	9.273	8	9.150	9
Ga^{67}	31	36	670.901° 626.249	9.3470	3			5.649	16	$ m Rb^{87}$ $ m Rb^{88}$	37 37	50 51	813.524 820.112	9.3509 9.3195		10.668 6.588	8 96	9.246 9.922	
Ga^{68}	31	37	635.138	9.3403		8.889	16	6.973	16	Rb89	37	52	828.384	9.3077	7	8.272	110	10.432	
Ga^{69}	31	38	646.230	9.3657		11.092	11	7.116	9	Sr84	38	46	782.440	9.3148	3	0.212	110	10.102	211
Ga^{70}	31	39	654.453	9.3493		8.223	20	8.461	35	Sr85	38	47	791.549	9.3123	5	9.109	33	9.212	34
Ga^{71}	31	40	664.430	9.3582		9.977	20	8.468	16	Sr86	38	48	803.917	9.3479		12.368	33	10.334	5
$ m Ga^{72}$ $ m Ga^{73}$	31 31	$\frac{41}{42}$	671.429 681.338	9.3254 9.3334		6.999 9.909	8	9.424 10.437	220	Sr ⁸⁷	38	49	812.979	9.3446	3	9.062	5	10.123	8
Ge ⁶⁹	32	37	642.997	9.3334		9.909	44	7.859	224 12	Sr ⁸⁸ Sr ⁸⁹	38 38	50 51	824.892 831.746	9.3738 9.3455	3	11.913 6.854	6 10	11.368 11.634	5 96
$\widetilde{\mathrm{Ge^{70}}}$	32	38	655.374	9.3625		12.377	5	9.144	3	Sr90	38	52	840.102	9.3345	4	8.356	33	11.718	63
Ge^{71}	32	39	663.335	9.3427	3	7.961	5	8.882	21	Sr ⁹¹	38	53	846.373	9.3008	3	6.271	36	11.,10	00
Ge^{72}	32	40	674.876	9.3733		11.541	5	10.446	4	Sr92	38	54	854.239	9.2852	9	7.866	83		
$ m Ge^{73}$ $ m Ge^{74}$	32	41	682.160	9.3447	3	7.284	3	10.731	7	Y85	39	46	786.957	9.2583	5	40 405		4.517	37
Ge ⁷⁵	32 32	42 43	693.105 700.091	9.3663 9.3345		10.945 6.986	3 22	11.767	43	Y^{86} Y^{87}	39 39	47 48	797.382 810.322	9.2719 9.3140		10.425 12.940	65 220	5.833	
Ge ⁷⁶	32	44	710.212	9.3449		10.121	22			V88	39	49	820.159	9.3140	3	9.837	215	6.405 7.180	215 9
Ge^{77}	32	$\frac{1}{45}$	716.680	9.3075		6.468	55			Y89	39	50	832.470	9.3536		12.311	9	7.578	6
$ m As^{73}$	33	40	680.917	9.3276				6.041	32	\mathbf{Y}^{90}	39	51	839.841	9.3316	4	7.371	32	8.095	33
As^{74}	33	41		9.3176		8.589	33	7.346	8	\mathbf{Y}^{91}	39	52	848.393	9.3230		8.552	34	8.291	34
As75	33	42	700.512	9.3402		11.006	9	7.407	3	Y^{92}	39	53	855.450	9.2984	4	7.057	34	9.077	36
${ m As^{76}} \over { m As^{77}}$	33 33	43 44	708.378 718.788	9.3208 9.3349		7.866 10.410	12 16	8.287 8.576	25 11	Y^{93} Y^{94}	39 39	54 55	863.461 869.988	9.2845 9.2552	4 23	8.011 6.527	40 216	9.222	84
As^{78}	33	45	726.210	9.3104		7.422	215	9.530	222	Zr^{89}	40	49	828.578	9.2332	3	0.327	210	8.419	13
$\mathrm{As^{79}}$	33	46	735.796	9.3139		9.586		2.000		Z_{r}^{190}	40	50	841.470	9.3497		12.892	11	9.000	
$\mathrm{Se^{73}}$	34	39	677.120	9.2756						Zr^{91}	40	51	849.201	9.3319	3	7.731	7	9.360	32
Se^{74}	34	40	690.117	9.3259		12.997	34	9.200	32	Zr^{92}	40	52	858.469	9.3312	3	9.268	6	10.076	13
Se ⁷⁵ Se ⁷⁶	34 34	41 42	698.738 710.725	9.3165 9.3516		8.621 11.987	6 9	9.232 10.213	9 9	$ m Zr^{93}$ $ m Zr^{94}$	40 40	53 54	865.718	9.3088	3	7.249	11	10.268	34
Se ⁷⁷	34	43	718.678	9.3335		7.953	9	10.213	13	Zr ⁹⁵	40	55	874.510 881.456	9.3033 9.2785	3	8.792 6.946	11 14	11.049 11.468	24 215
Se ⁷⁸	34	44	729.949	9.3583		11.271	5	11.161	11	Zr96	40	56	889.890	9.2697	3	8.434	14	11.700	213
Se^{79}	34	45	737.420	9.3344	. 3	7.471	6	11.210	215	Zr ⁹⁷	40	57	895.880	9.2359	$\overset{\circ}{4}$	5.990	34		
Se^{80}	34	46	748.056	9.3507		10.636	7	12.260	107	Nb^{89}	41	48	823.566	9.2536					
Se^{81}	34	47	755.461	9.3267	7	7.405	54			Nb^{90}	41	49	834.064	9.2674		10.498	99	5.486	16
$\mathrm{Se^{82}}$ $\mathrm{Br^{76}}$	34 35	48 41	765.214 704.951	9.3319 9.2757	3	9.753	54	6.213	14	${ m Nb^{91}} \ { m Nb^{92}}$	41 41	50 51	847.190 855.466	9.3098 9.2985	9 6	13.126 8.276	83 95	5.720 6.265	82 47
Br^{77}	35	42	716.367	9.3035		11.416	14	5.642	10	Np ₃₃	41	52	864.940	9.3004	3	9.474	47	6.471	5
$\mathrm{Br^{78}}$	35	43	725.253	9.2981	- 3	8.886	13	6.575	12	Nb^{94}	41	53	872.667	9.2837	4	7.727	27	6.949	29
$\mathrm{Br^{79}}$	35	44	736.746	9.3259		11.493	11	6.797	3	$\mathrm{Nb^{95}}$	41	54	881.814	9.2823	3	9.147	28	7.304	7
Br^{80}	35	45	745.196	9.3150		8.450	10	7.776	12	Nb^{96}	41	55	889.260	9.2631	4	7.446	33	7.804	35
$ m Br^{81}$ $ m Br^{82}$	35 35	$\frac{46}{47}$	756.118 764.274	9.3348		10.922	11	8.062	5	Nb ⁹⁷	41	56	897.889	9.2566	3	8.629	33	7.999	10
Br^{83}	35	47	774.537	9.3204 9.3318		8.156 10.263	6 22	8.813 9.323	54 23	$Mo^{90} Mo^{91}$	42 42	48 49	830.491 841.555	9.2277 9.2479	12 9	11.064	131	6.925 7.491	146 76
Kr^{77}	36	41	712.431	9.2524		10.200	44	7.480	26 26	Mo ⁹²	42	50	854.996	9.2479		13.441	75	7.806	82
$ m Kr^{78}$	36	42	725.200	9.2974	3	12.769	22	8.833	7	$\mathrm{Mo^{93}}$	42	51	863.516	9.2851	8	8.520	70	8.050	84
Kr^{79}	36	43	734.161	9.2932	3	8.961	14	8.908	18	$\mathrm{Mo^{94}}$	42	52	874.048	9.2984		10.532	70	9.108	5
Kr ⁸⁰	36	44	746.523	9.3315		12.362	15	9.777	5	Mo ⁹⁵	42	53	881.964	9.2838	3	7.916	5	9.297	27
$ m Kr^{81}$ $ m Kr^{82}$	36 36	45 46	754.955 766.749	9.3204 9.3506		8.432 11.794	107 107	9.759 10.631	107 5	$ m Mo^{96} m Mo^{97}$	42 42	54 55	891.796 899.116	9.2895 9.2692	3	9.832 7.320	5 4	9.982	7
Kr ⁸³	36	47	774.766	9.3345	3	8.017	5	10.031	5	Mo ⁹⁸	42	56	908.395	9.2692	3	9.279	4	9.856 10.506	32
Kr^{84}	36	48	786.060	9.3579		11.294	$\overset{\circ}{4}$	11.523	22	$\mathrm{Mo^{99}}$	42	57	914.966	9.2421	33	6.571	320	10.000	
Kr^{85}	36	49	793.706	9.3377		7.646	7			$\mathrm{Mo^{100}}$	42	58	923,662	9.2366	3	8.696			

^{*} Total nuclear binding energy in milliunits. The mass equivalent of the electronic binding energy has been subtracted from the total atomic binding energy. Errors are not listed because, for most purposes, the differences in 2 TNBE values are employed. An estimate of the error associated with TNBE may be found by multiplying the error assigned to TNBE/A by the corresponding A.
b Average nuclear binding energy per nucleon in milliunits.

istic parabolic shape in the even-A curves for each element appear in this region as has been observed in other regions. The obvious change in the curve connecting odd-A points near A = 89 is a result of the shell closure at N=50. There is an increase in the odd-A

curve in the region from A = 79 to A = 89. This rise does not appear at the shell closure at $N=82^{24}$ or $N=126.2^{25}$

(1957). ²⁵ V. B. Bhanot, W. H. Johnson, and A. O. Nier, Phys. Rev. **120**,

<sup>Neutron separation energy in milliunits.
d Proton separation energy in milliunits.
The zinc masses were not measured in this investigation. In order to study the nuclear systematics of Ga and Ge it was necessary to calculate TNBE for the listed zinc isotopes. The atomic masses of zinc were taken from the 1961 Mass Table of Ref. 12.</sup>

²⁴ W. H. Johnson, Jr., and A. O. Nier, Phys. Rev. 105, 1014

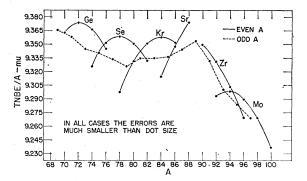


Fig. 2. Average binding energy per nucleon for stable isotopes.

A more detailed study of binding-energy systematics may be made by considering the neutron separation energy and the proton separation energy. Figures 3 and 4 plot the neutron separation energy as a function of the neutron number for even N and odd N, respectively. In each case, successive points for a particular element are connected by a straight line. The sharp discontinuity beyond N=50 is shown in these graphs with greater precision than previously available. The generally smooth character of the curves on both sides of N=50is perhaps the most significant result inferred from these plots. The N=42 data points for bromine and the N=40data point for gallium seem to contradict this smooth behavior. In the case of bromine, the assignment of a mass to the radioactive Br77 and Br76 may be in error. In the case of gallium, these is no obvious error in either mass used to calculate this separation energy. This value may indicate an anomaly at N = 40. The S_n values for germanium and selenium at N=40, however, show no particular anomaly.

Proton separation energies S_p have been plotted in Fig. 5 for even Z. A line connects data points of constant N. These curves indicate, as has been pointed out previously, that the closure of a neutron shell seems to have no effect on the proton separation energies. The variation of S_p for a given N value as a function of Z is also smooth. There is a persistent change in slope at Z=40

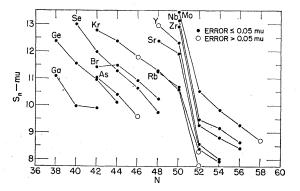


Fig. 3. Neutron separation energies for nuclei with an even number of neutrons.

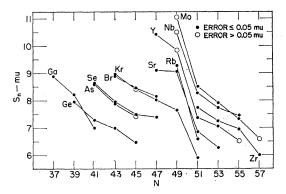


Fig. 4. Neutron separation energies for nuclei with an odd number of neutrons.

for the four curves that have data points at both Z=38 and Z=42. As in the neutron data, this may be an indication of a slight change in nuclear structure near nucleon number 40.

The study of the systematics of the binding energy of the last pair of nucleons in a nucleus is worthwhile because stable masses with small experimental errors are employed in most cases in the calculation. Table VIII lists the value of the binding energy of the last pair of neutrons S_{2n} for even N nuclei. These data are plotted in Fig. 6 as a function of neutron number N. Values from the same element are connected by a line. Once again the shell closure at N=50 is clearly visible. The smooth behavior of these curves for values on either side of N=50 is clearly evident.

The values of the binding energy of the last pair of protons for even Z nuclei are listed in Table IX. These data are plotted in Fig. 7. Once again the smooth variation of these data with changes in Z is evident. These results also indicate a change in slope at Z=40. Because there are only two other values of N for which there are 3 data points for a given N value, the consistency of slope for values of Z other than 40 is difficult to demonstrate.

A number of pairing energies for neutron and proton pairs may also be calculated from the mass data. The pairing energy for the last pair of neutrons added to a

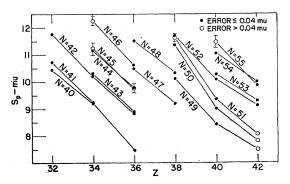


Fig. 5. Proton separation energies for nuclei with an even number of protons.

Table VIII. Binding energy and pairing energy of the last pair of neutrons for the listed isotopes.

Table IX. Binding energy and pairing energy of the last pair of protons for the listed isotopes.

Isotope	Z	N	S_{2n} mu	a Error	P, mu	e Error	Isotope	Z	N	S_{2p}	Error	P_{i}	p ^b
										mu		mu	Error
Ga ⁶⁹ Ga ⁷¹	31 31	38 40	19.981	12	2.203	25	Ge ⁶⁹	32	37	14.832	12	0.886	25
Ga ⁷³	31	40 42	18.200 16.908	4	1.754	40	Ge^{70}	32	38	16.260	9	2.028	11
Ge ⁷²	32	42 40	19.502	43	2.910	45	$_{ m Ge^{72}}^{ m Ge^{72}}$	32	39	17.343	29	0.421	50
Ge ⁷⁴	32 32	40 42	19.502	3	3.580	10		32	40	18.914	16	1.978	18
Ge ⁷⁶	32	42 44	17.107	3 2	3.661	5 44	Ge ⁷³ Ge ⁷⁴	32	41	20.155	220	1.307	220
As ⁷⁵	33	42	19.595	32	$\frac{3.135}{2.417}$	36	Se ⁷³	32 34	42	22.204	220	1.330	235
As^{77}	33	42 44	18.276		$\frac{2.417}{2.544}$	30 27	Se ⁷⁴		39	13.785	34	2.450	
As ⁷⁹	33	46	17.008	11 108	$\frac{2.344}{2.164}$	445	Se ⁷⁵	34 34	40	15.241	4	3.159	64
Se ⁷⁶	33 34	42	20.608	108	3.366		Se ⁷⁶	$\frac{34}{34}$	41	16.578	4	1.886	17
Se ⁷⁸	34	42 44	19.224	8		12 12	Se ⁷⁷	34 34	42	17.620	8	2.806	10
Se ⁸⁰	34 34	44 46	18.107		3.318		Se ⁷⁸		43	18.587	22	2.013	33
Se ⁸²	34 34	48	17.158	4	3.165	13 108	Se ⁷⁹	34	44	19.737	3	2.585	22
Br ⁷⁹	3 4 35	48 44	20.379	6 7	2.348		Se ¹³ Kr ⁷⁷	34	45	20.740	55	1.680	433
Br ⁸¹	35 35	44 46	19.372	5	$\frac{2.607}{2.472}$	23		36 36	41	13.693	22	1.267	34
Br ⁸³	35 35	48	18.419	22		21 24	$ m Kr^{78}$ $ m Kr^{79}$		42	14.475	9	3.191	15
Kr ⁸⁰	36	46 44	21.323		2.107	2 4 29	Kr80	36 36	43	15.483	15	2.333	26
Kr ⁸²	36	46	20.226	6	$\frac{3.401}{3.362}$	29 214	Kr ⁸¹		44	16.574	5	2.980	7
Kr ⁸⁴	36	48 48	20.226 19.311	6 4	3.302		Kr^{61} Kr^{82}	36	45	17.535	107	1.983	109
Kr ⁸⁶	36	50	18.218		2.926	8	Kr ⁸² Kr ⁸³	36	46	18.693	4	2.569	.9
Kr ⁸⁸	36	50 52	13.674	$\begin{array}{c} 4 \\ 235 \end{array}$	1.850	13	Kr ⁸⁴	36	47	19.305	$5\frac{4}{2}$	1.679	55
Rb^{87}	30 37	52 50				257	Sr ⁸⁴	36	48	20.846	5	2.200	44
Rb ⁸⁹	37 37	50 52	19.941	5	1.395	15		38	46	15.691	4		20
Sr ⁸⁶	37 38	32 48	14.860 21.477	54	1.684	200	Sr ⁸⁵ Sr ⁸⁶	38	47	16.783	33	1.641	39
Sr ⁸⁸		48 50		5	3.259	66		38	48	17.857	4 7	2.811	9
Sr ⁹⁰	38	50	20.975	6	2.851	9	Sr ⁸⁷	38	49	19.273	7	0.973	16
Sr ⁹⁰	38	52	15.210	32	1.502	37	Sr88	38	50	20.614	5	2.122	8
Y ⁸⁷	38 39	54 48	14.137	87	1.595	93	Sr ⁸⁹	38	51	21.556	53	1.712	199
Y 89		48 50	23.365	220	2.515	243	Sr ⁹⁰	38	52	22.150	237	1.286	259
Y^{93}	39	50	22.148	215	2.474	216	Zr^{89}	40	49	15.599	11	1.239	19
\mathbf{Y}^{91}	39	52	15.923	13	1.181	65	Zr^{90}	40	50	16.578	7	1.422	10
Zr^{92}	39	54	15.068	27	0.954	69	Zr^{91}	40	51	17.455	10	1.265	65
Zr^{92} Zr^{94}	40	52	16.999	7	1.537	11	Zr^{92}	40	52	18.367	32	1.785	40
	40	54	16.041	6	1.543	21	Zr^{93}	40	53	19.345	19	1.191	67
Zr ⁹⁶	40	56	15.380	7	1.488	27	Zr^{94}	40	54	20.271	81	1.827	94
Nb ⁹¹	41	50	23.624	128	2.628	130	$\mathrm{Mo^{91}}$	42	49	12.977	76	2.005	79
Nb^{93}	41	52	17.750	82	1.198	125	Mo^{92}	42	50	13.526	_6	2.086	164
Nb^{95}	41	54	16.874	7	1.420	54	Mo^{93}	42	51	14.315	70	1.785	117
Nb^{97}	41	56	16.075	10	1.183	65	$\mathrm{Mo^{94}}$	42	52	15.579	.5	2.637	9 55
Mo^{92}	42	50	24.505	108	2.377	185	Mo^{95}	42	53	16.246	11	2.348	55
Mo ⁹⁴	42	52	19.052	5	2.012	140	Mo^{96}	42	54	17.286	5	2.678	13
Mo^{96}	42	54	17.748	4	1.916	8	Mo^{97}	42	55	17.660	13	2.052	65
Mo^{98}	42	56	16.599	$\frac{4}{2}$	1.959	6	Mo^{98}	42	56	18.505	6	2.507	17
$\mathrm{Mo^{100}}$	42	58	15.267	5	2.125	644	${ m Mo^{99}}$	42	5 7	19.086	322		

a Binding energy of the last pair of neutrons in milliunits. b Pairing energy of the last pair of neutrons in milliunits.

nucleus of Z protons and N neutrons is given by Eq. (5). $P_n(Z,N) = \text{TNBE}(Z,N) - 2\text{TNBE}(Z,N-1) \\ + \text{TNBE}(Z,N-2), \quad (5)$

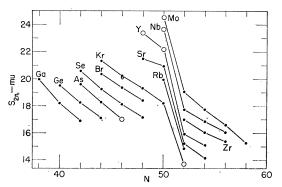


Fig. 6. Binding energy of the last two neutrons.

where N is even. The proton pairing energy P_p is defined in a similar manner. For light nuclei Mayer and Jensen²⁶ have concluded that a correlation exists be-

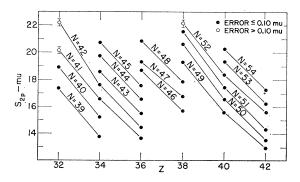


Fig. 7. Binding energies of the last two protons.

²⁶ M. G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure* (John Wiley & Sons, Inc., New York. 1955).

 $^{^{\}rm a}$ Binding energy of the last pair of protons in milliunits. $^{\rm b}$ Pairing energy of the last pair of protons in milliunits.

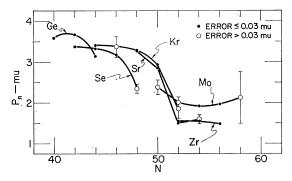


Fig. 8. Neutron pairing energy.

tween the pairing energy and the j value of the odd nucleon of the pair, with larger pairing energies correlated to higher j values. In regions of high j values, one finds that it is energetically possible to have the pair occupy a high j-value state rather than to pair in the lower spin state of the preceding odd nucleon. This mechanism is used to explain the absence of the highest j values from the ground-state spins of odd nuclei.

Neutron and proton pairing energies that may be calculated from the present masses are listed in Tables VIII and IX, respectively. The P_n values are plotted in Fig. 8 as a function of N. Values from the same element are connected by a curved line. An attempt has been made to correlate the magnitude of these pairing energies with (a) the j value of the previous odd neutron and (b) with the j value which the pair is assumed to have according to the filling scheme of Mayer and Jensen.²⁶ Neither comparison produces very convincing results. The correlation between the value of j and the pairing energy is in some cases what Mayer and Jensen have suggested; in others, the opposite. The one positive statement that may be made is that the value of P_n decreases, rather strikingly, following the shell closure at N=50. The general tendency of the curves for each element is smooth, and in most instances continuously decreasing with increasing N to N=50. The character of the curves changes abruptly following N=50. Beyond

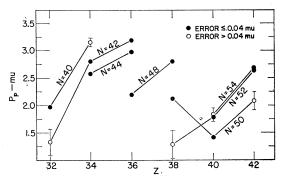


Fig. 9. Proton pairing energy.

N=50, the P_n values are small, and for each element, are essentially constant for the region plotted. Note that there is nothing anomalous about the one value at N=40. Values of P_n for zinc from Quisenberry et al. 1 at N=40 and N=38 further strengthen this conclusion.

Figure 9 is an illustration of the proton pairing energy P_p as a function of Z. In this illustration, points with the same N value are connected by lines. An attempt to correlate the P_p with the j value for the pair is again not particularly fruitful. The P_p values at Z=40, with $j=\frac{1}{2}$ for this pair, appear to be lower than practically all other values in this region. A j value of $\frac{1}{2}$ occurs only at Z=40; values for other pairs in this region are all larger. Thus, in this instance, low P_p is correlated with low j. There are other instances, however, where this correlation is reversed. It is of interest to note that the value of P_p at Z=40 is so small. This may indicate a structure change near Z=40 that is not indicated at N=40.

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