Rare Decay Modes of K^* [†]

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Assuming ρ -meson dominance, the branching ratio $\Gamma(K^* \to K\pi\pi)/\Gamma(K^* \to K\gamma)$ has been calculated to be 1/63. Estimating the $K*K\rho$ coupling from unitary symmetry and the decay of the π^0 gives $\Gamma(K^* \to K\gamma)/$ $\Gamma(K^* \to K\pi) \cong 0.15\%$, a factor of 5 lower than previous estimates.

A SSUMING ρ -meson dominance in the sense of dispersion theory, the branching ratio $\Gamma(K^* \to K \pi \pi)$
 $\Gamma(K^* \to K \gamma)$ can be estimated independent of arbitrary coupling constants. The ρ is treated as almost stable; i.e. , the propagator is taken to be that of a stable particle. The calculation is similar to that of Gell-Mann, Sharp, and Wagner¹ for the ω decay.

We take the following interactions:

$$
(f_{K^*K\rho}/M)\epsilon_{\mu\nu\lambda\sigma}k_{\mu}\epsilon_{\nu}q_{\lambda}\eta_{\sigma}
$$
 at the $K^*K\rho^0$ vertex,

 $f_{\rho\pi\pi}\eta_{\mu}(p_1-p_2)_{\mu}$ at the $\rho\pi\pi$ vertex,

 $\gamma_{\gamma\rho}$ at the $\rho\gamma$ vertex.

(See Figs. 1 and 2 for notation.)

$$
K^*
$$

\n
$$
K^*
$$

\n
$$
K^*
$$

\nFIG. 1. Lower order diagram for
\n
$$
K^* \rightarrow K \pi \pi
$$
, showing notation.

Then

Then
\n
$$
\Gamma(K^* \to K\gamma) = \frac{1}{24} \frac{f^2 K^* K \rho}{4\pi M^2} \frac{\gamma^2 \gamma \rho}{m_\rho^4} \frac{(m^2 K^* - m^2 K)^3}{m^3 K^*},
$$
\n
$$
\Gamma(K^* \to K\pi\pi) = \frac{1}{\pi} \frac{f^2 K^* K \rho}{4\pi M^2} \frac{f^2 \rho \pi \pi}{4\pi} \times \frac{|\mathbf{p}_1 \times \mathbf{p}_K|^2 dE_1 dE_K}{|\mathbf{p}_1 \times \mathbf{p}_K|^2 dE_1 dE_K}
$$
\n
$$
\times \int \frac{|\mathbf{p}_1 \times \mathbf{p}_K|^2 dE_1 dE_K}{[\mathbf{m}_p^2 - m_{K^*}^2 + 2m_{K^*} E_K - m^2 \kappa]^2},
$$
 for

where the statistical factors due to isospin conservation have already been inserted [i.e., $\Gamma(K^* \to K \pi \pi)$ include all charge configurations in the final state]. Since there is only about 100-MeV kinetic energy released in the $K\pi\pi$ decay mode, the three-body phase space can be treated nonrelativistically. The result is

$$
\Gamma(K^* \to K\pi\pi) = \frac{1}{4} \frac{1}{\mu + 1} \left[\frac{\mu - 1}{\mu + 1} \right]^{3/2} m_{K^*} m_{\pi}^2
$$

$$
\times \frac{\frac{f^2 K^* K \rho}{4\pi M^2} \frac{f^2 \rho \pi \pi}{4\pi} \frac{Q^4}{\left[m_p^2 - (m_{K^*} - m_K)^2 \right]^2},
$$

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'M. Gell-Mann, D. Sharp, and W. G. Wagner, Phys. Rev. Letters 8, 261 (1962).

where

$$
Q = m_{K^*} - m_K - 2m_{\pi}
$$

$$
\mu = 1 + m_K/m_{\pi},
$$

and we have put $E_K = m_K$ in the ρ propagator. [An expansion of the propagator to first order in $(E_K - m_K)/m_K$ changes the result by about 10%.

Now if the ρ meson also dominates the nucleon isovector form factor, we have $\gamma_{\gamma\rho} = \epsilon m_{\rho}^2/f_{\rho\pi\pi}^2$. Also, $f_{\rho\pi\pi}$ is determined from the observed ρ decay width to be $f_{\text{part}}^2/4\pi \cong 2$. Thus, we obtain

$$
\Gamma(K^* \to K\pi\pi) \cong 1.9 \times 10^{-4} \frac{f^2 \kappa \ast \kappa_\rho}{4\pi M^2} m_\pi^3,
$$

\n
$$
\Gamma(K^* \to K\gamma) \cong 1.2 \times 10^{-2} \frac{f^2 \kappa \ast \kappa_\rho}{4\pi M^2} m_\pi^3,
$$

\n
$$
\frac{\Gamma(K^* \to K\pi\pi)}{\Gamma(K^* \to K\gamma)} \cong \frac{1}{63}.
$$

Using the result of unitary symmetry $f^2_{K^*K\rho}/4\pi M^2$ $=\frac{3}{4}f^2_{\omega\rho\pi}/4\pi M^2$ and estimating $f^2_{\omega\rho\pi}/4\pi M^2=0.02/m_\pi^2$.

Fig. 2. Lowest order diagram for
$$
K^* \to K\gamma
$$
.

from the decay of the π^0 , we find

$$
\Gamma(K^* \to K\gamma) \cong 2.5 \times 10^{-2} \text{ MeV}
$$

$$
\Gamma(K^* \to K\gamma)/\Gamma(K^* \to K\pi) \cong 0.15\%,
$$

which is lower than previous estimates by about a factor of $5^{3,4}$

Fujii,⁴ using the Fermi statistical model, finds $\Gamma(K^* \to K \pi \pi) / \Gamma(K^* \to K \gamma) \approx 5$. However, in this result, the angular momentum barrier due to the K^* spin was not taken into account.

I should like to thank Professor J. J. Sakurai for suggesting this calculation.

² J. J. Sakurai, in *Proceedings of the "Enrico Fermi" Internation* School of Physics (Villa Monastero, Varenna, Como, Italy). See
also, M. Gell-Mann and F. Zachariasen, Phys. Rev. 124, 953 (1961) ; Y. Nambu and J. J. Sakurai, Phys. Rev. Letters $8, 79$

^{(1962).} ' M. A. B. Beg, P. C. De Celles, and R. B. Marr, Phys. Rev. 124, 622 (1961). ⁴ A. Fujii, Phys. Rev. 124, 1240 (1961).