## Low Even Configurations in the First Spectrum of Rhenium (Re I)

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The theory of intermediate coupling with configuration interaction is used to calculate energy levels and g values in the  $5d^56s^2$  and  $5d^66s$  configurations of Re 1. Fifty-one observed levels are fitted with a mean deviation of  $\pm 325$  cm<sup>-1</sup> by using eleven free parameters. The L(L+1) correction is important, as the mean deviation is  $\pm 514$  cm<sup>-1</sup> without it; the parameter for this correction ( $\alpha = 66.8 \pm 13$ ) has a value similar to that in first and second long-period spectra. Standard deviations of the parameters are calculated, and within the limits so defined, Slater's parameters are consistent with expectations based on known behavior in first and second long-period spectra. Calculated g values agree with observed values for thirty-seven levels to about  $\pm 0.03$ . Levels not previously assigned from experiment are given designations; the eigenvectors are used to illustrate qualitative features of these designations.

HE experimental analysis of Re I has been carried out by several investigators over the last twentyfive years. Recently, a new description of this spectrum was made, based on an extended homogeneous line list,<sup>1</sup> and the earlier work was revised and extended with this as a basis.<sup>2,3</sup> The calculations in this paper were carried out primarily to provide assignments for levels of the  $5d^56s^2$  and  $5d^66s$  configurations in this spectrum. These assignments have been published,<sup>2,3</sup> and the present paper gives the complete calculation on which they are based.

As in most spectra of the third long period, an intermediate form of coupling, which includes effects of configuration interaction, is required to describe the observed levels rigorously, but for qualitative descriptions it is customary to assign levels designations based on the LS-coupling notation. In the years that have been devoted to the analysis of ReI, experimental spectroscopists have made such term assignments which contributed greatly to the understanding of this complicated spectrum and helped to bring order out of the forest of levels. But about half the observed levels were not assigned; it was not established, for instance, that only two configurations were present,<sup>4</sup> or that all expected levels in the observable region of the spectrum had been found. The use of theory was indispensable in completing the assignments. The eigenvectors given

in the present paper show the degree of approximation involved in making the assignments.<sup>5</sup>

The eigenvectors have also an application in indicating the accuracy obtainable in atomic theory generally when calculations are made for spectra with many close interacting levels. Since the g values have been observed for thirty-seven levels, a simple criterion is available for estimating this accuracy in many instances. Eigenvectors may also be used in the calculation of intensities of radiative transitions, or in the calculation of hyperfine structure splittings, and the signs of the amplitudes are given, along with the percentage composition, for possible use in these applications.

A brief description of this calculation and two similar ones has been given.<sup>6</sup> It is planned to publish later a detailed description of these other results that were obtained while the Standards Eastern Automatic Computer (SEAC) was available. At present, coding is being set up to carry out similar calculations on the IBM 704.

The matrix elements of electrostatic and spin-orbit interactions for these two configurations are taken mainly from work already published, and checks were made to insure the correctness of these results.7 Subse-

<sup>&</sup>lt;sup>1</sup>W. F. Meggers, J. Research Natl. Bur. Standards 49, 187

levels found in the region lower than  $35000 \text{ cm}^{-1}$  arise from the two configurations  $5d^{8}6s^{2}$  and  $5d^{8}6s$ . Evidence is also noted that two configurations  $3a^{2}03^{2}$  and  $3a^{2}03^{2}$ . Evidence is also noted that supports earlier conclusions that the  $5d^{7}$  configuration is present in the region slightly higher than  $35\ 000$  cm<sup>-1</sup>, so that three configurations probably overlap in this region of the spectrum. Since the  $5d^{7}$  configuration has been neglected in the present calculation, partial results that were obtained for predicted levels higher than  $35\,000$  cm<sup>-1</sup> have been omitted. Three known even levels having a J value of  $2\frac{1}{2}$  have been disregarded for similar reasons (i.e., 44 054, 44 903, and 50 973 cm<sup>-1</sup>).

<sup>&</sup>lt;sup>5</sup> Eigenvectors have been given for the two-electron configura-tions of Th III by G. Racah [Physica 16, 651 (1950)], and for the low even three-electron configuration of Th II by Y. Eisenberg [Physica 18, 177 (1952)]. They have already noted several qualitative features of assignments. Except for one or two levels, their component component of the base of the lower the lower of the several sev their assignments confirmed ones that had been established by experimental methods. Th II and Th III are close to the limit in complexity that can be handled rigorously with desk computers, but are, of course, much simpler than Re I.

<sup>&</sup>lt;sup>9</sup> R. É. Trees, Bull. Am. Phys. Soc. Ser. II, 3, 29 (1958).

<sup>&</sup>lt;sup>7</sup> The matrices of spin-orbit interaction for the  $d^{4s}$  configuration are given by W. R. Bozman and R. E. Trees [J. Research Natl. Bur. Standards 58, 95 (1957)]. Those for the  $d^{5}$  configuration are given by H. Greyber [dissertation, University of Pennsylvania, 1953 (unpublished)]; a few errors were corrected in these results. The matrices of the spin-orbit interaction were checked by calculating the eigenvalues on SEAC, and comparing them with values known from j-j coupling theory. Matrix elements of configuration interaction between  $d^{4s}$  and  $d^{5}$  are given by G. Racah [Phys. Rev. 63, 367 (1943)]; he has used a digital computer to check these matrix elements [G. Racah (private communication)]. No independent check insures consistency of our phases for the matrices of spin-orbit interaction and those of electrostatic

quent steps in the calculation were carried out on SEAC and most of the coding has been described.<sup>8,9</sup>

Numerical matrices were obtained by using preliminary parameters estimated by comparison with parameters obtained in other calculations already carried out for spectra of the third long period (the latter parameters are listed in footnote 22 of reference 8). By using the eigenvectors of the numerical matrices, linear formulas were set up for the eigenvalues of the matrices as functions of the parameters. New parameters were then determined by least squares. In calculations with large matrices it is important to estimate good preliminary parameters so that the linear formulas will be accurate. In spectra where theory and experiment agree very closely, it may be necessary to use the first set of parameters obtained by least squares to obtain more accurate linear formulas, but there was no need to do this for Re 1.<sup>10</sup>

The differences between calculated and observed energy levels when the L(L+1) correction is omitted are given in the last column, "no  $\alpha$ ," of Table I. The mean deviation is  $\pm 514$ . The errors are mostly positive for small J values, and negative for the large J values. This systematic behavior shows conclusively that the correction has an over-all importance for all levels of this spectrum. In Ta II, over-all importance could not be demonstrated because only a few levels with large J value (more rigorously, L value) were observed.<sup>8</sup> The results are from linear formulas. Inaccuracy of linear formulas usually leads to an overestimate for the mean deviation, but in this case the effect is negligible. It can be estimated by omitting the errors for the two lowest levels with<sup>10</sup>  $J = 3\frac{1}{2}$  and regarding the adjustment by least squares as being made to fit forty-five observations rather than forty-seven. The mean deviation would then be reduced from  $\pm 514$  to  $\pm 508$  cm<sup>-1</sup>.

The calculated energy levels and g values obtained using the L(L+1) correction are given in the rest of

this code available and discussing its use with me. <sup>10</sup> No allowance was made for inaccuracy of the linear formulas, as is sometimes done by omitting pairs of levels that are close, or by using an average value. The errors in the linear formulas were less than 100 cm<sup>-1</sup> for all levels but the ones at 19 758 and 27 141 cm<sup>-1</sup>, where the errors were about 200 cm<sup>-1</sup>, and the levels at 14 217 and 15 058 cm<sup>-1</sup> where the errors were about 450 cm<sup>-1</sup>. The fact that the mean error of the final calculations is  $\pm 371$ cm<sup>-1</sup> is justification for ignoring these inaccuracies. Four levels were omitted from the calculation by least squares to minimize the computer time and the ground level was one of these omissions.

TABLE I. C	alculated energ	y le	vels	(in cn	$n^{-1}$ ) and g	valı	ies in the
$5d^{6}6s + 5d^{5}6s^{2}$	configurations	of	the	first	spectrum	of	rhenium
(Re 1).							

			Level	Calc		g value	Calc	No α Calc
	Labe	I Obs.	Calc.	-Obs.	Obs.	Calc.	-Obs.	-Obs
ł	1 2 3 4 5 6	15 166 17 238 19 758 27 385 30 132	15 265 17 406 19 605 26 764 30 476 34 366	99 168 153 621 344 	2.368 2.521 0.983	2.165 2.785 0.901 1.831 0.375 1.939 Sum =	-0.203 0.264 -0.082  -0.021	692 263 96 
13	1 2 3 4 5 6 7 8 9 10	13 826 16 328 20 482 22 423 26 132 27 828 30 527 31 461 	13 858 16 619 20 397 22 427 26 812 27 848 30 707 31 888 34 081 35 289	32 291 -85 4 680 20 180 427 	1.485 1.706 1.451 0.781 0.650 0.888  	1.453 1.760 1.404 0.802 0.590 0.958 0.913 1.373 1.054 1.058 Sum =	-0.032 0.054 -0.047 0.021 -0.060 0.070   0.006	312 389 657 621 595 810 663
21/2	1 2 3 4 5 6 7 8 9 10 11 12 13 14	0 11 584 14 621 15 770 19 458 23 155 24 425 26 661 28 030 29 800 31 186 32 435 33 282	681 11 465 14 726 15 863 19 059 23 073 24 657 26 367 27 643 30 219 30 601 32 391 33 279 34 613	$\begin{array}{c} 681 \\ -119 \\ 93 \\ -399 \\ -82 \\ 232 \\ -294 \\ -387 \\ 419 \\ -585 \\ -44 \\ -3 \\ \cdots \end{array}$	1.950 1.278 1.151 1.309 1.361 1.189 1.067 1.32 1.12 1.17 1.17 	1.944 1.251 1.087 1.421 1.320 1.182 1.045 1.428 1.039 1.034 1.131 1.170 0.842 1.256 Sum =	$\begin{array}{c} -0.006\\ -0.027\\ -0.064\\ 0.112\\ -0.041\\ -0.007\\ -0.022\\ 0.11\\ -0.08\\ -0.14\\ -0.04\\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ $	44( 152 43 452 122 592 -173 872 -795 466 151
31/2	1 2 3 4 5 6 7 8 9 10 11 12	14 217 15 058 17 331 21 775 24 724 27 141 28 542 28 810  31 983 33 824 	14 484 14 895 17 282 21 750 24 746 27 099 28 045 29 260 31 000 31 405 33 978 35 163	$\begin{array}{c} 267 \\ -163 \\ -49 \\ -25 \\ 22 \\ -47 \\ 450 \\ \cdots \\ -578 \\ 154 \\ \cdots \end{array}$	1.567 1.153 1.255 1.135 1.03 1.34  0.93 	1.565 1.153 1.260 1.134 1.038 1.152 1.027 0.972 1.205 1.215 1.027 1.378 Sum =	-0.002 0.000 0.005 -0.001 0.01 -0.19  0.04   -0.138	$\begin{array}{c} 614\\ -638\\ -579\\ -110\\ -136\\ -170\\ -913\\ \cdots\\ -672\\ 10\\ \end{array}$
41	1 2 3 4 5 6 7 8 9	11 754 16 619 22 160 27 161 27 514 30 645 31 399 34 194 	11 882 16 403 22 232 26 880 27 967 30 139 31 702 33 574 36 472	$ \begin{array}{r}     128 \\     -216 \\     72 \\     -280 \\     453 \\     -506 \\     303 \\     -620 \\     \cdots \end{array} $	1.545 1.175 1.198  1.13  	1.534 1.171 1.190 1.117 1.157 1.173 1.042 1.138 1.101 Sum =	-0.011 -0.004 -0.008  0.03    0.007	$ \begin{array}{r} 16\\ -59\\ -31\\ -49\\ 40\\ -59\\ -39\\ -76\\ \cdots \end{array} $
51	1 2 3 4 5 6	16 307 23 956 27 244 30 560 33 318 	15 982 23 559 27 447 30 869 33 711 36 565	-325 -397 203 309 393 	1.242 0.995 1.18 1.07 	1.247 0.988 1.173 1.136 1.127 1.023 Sum =	0.005 -0.007 -0.01 0.07  0.057	-72 -1214 -50 -520 214
6 <del>1</del>	1 2 3	26 348 27 130	25 796 27 537 38 724	-552 407 	1.100 1.208	1.080 1.217 1.088 Sum =	-0.020 0.009  -0.011	···· ····
	M Mea	$A (d^{6}s) A (d^{5}s^{2}) B (d^{6}s) B (d^{5}s^{2}) C (d^{6}s) C (d^{6}s) C (d^{5}s^{2}) t (d^{5}s^{2}) t (d^{5}s^{2}) G_{2} H_{2} \alpha ean error n deviatio$	n	36 949 20 033 470.5 498.5 1842.6 1881.6 2347.1 2425.2 2954.9 477.6 66.8 ±3 ±3	$\pm 1080$ $\pm 1148$ $\pm 51$ $\pm 31$ $\pm 95$ $\pm 104$ $\pm 80$ $\pm 104$ $\pm 26$ $\pm 13$ 71 25		$\begin{array}{c} 36 \ 955 \\ 19 \ 041 \\ 499.7 \\ 465.4 \\ 2032.5 \\ 2125.0 \\ 2277.5 \\ 2419.6 \\ 2740.1 \\ 541.9 \\ \cdots \\ \pm 580 \\ \pm 514 \end{array}$	
	ca							

interaction, but this is a question of the correct interpretation of Racah's formulas rather than one of numerical error. After setting up matrices of  $d^4s+d^5$ , the matrices for  $d^6s+d^5s^2$  were obtained formally by use of negative values for the spin-orbit botamet formally by use of negative values for the spin-order parameters. The orders of the final matrices, followed by the J values in parentheses, are 13(1/2), 19(3/2), 24(5/2), 20(7/2), 14(9/2), 7(11/2), and 3(13/2). <sup>8</sup> Trees, Cahill, and Rabinowitz, J. Research Natl. Bur. Standards 55, 335 (1955).

<sup>&</sup>lt;sup>9</sup> A code used to evaluate standard deviations of the parameters has not been described. Definitions of these deviations, which are given in Table I, must be obtained from books on statistics, since we do not know of any atomic spectra for which they have been calculated previously. The probable errors are two-thirds of the standard deviations. I wish to thank J. M. Cameron for making

Table I. The mean deviation between calculated and observed energy levels is  $\pm 325$  cm<sup>-1</sup>. The errors still show appreciable systematic behavior, being greater the higher the energy. Excluding the ground level, none of the levels lower than 25 000 cm<sup>-1</sup> have errors in excess of 400 cm<sup>-1</sup>, but errors of about half the higher levels are greater than this. This is probably because we neglected the  $5d^7$  configuration.<sup>4</sup> It is reasonable to expect that the mean deviation would be about half as great if the  $5d^7$  configuration were included, so that the agreement would be similar to that obtainable in first and second long-period spectra. It is likely that three high even levels omitted from the calculation<sup>4</sup> originate in the three low even configurations. Our calculation with two configurations does not explain these three levels well. In one instance (for the level observed at 50 973) a discrepancy of 2800  $cm^{-1}$  would have to be accepted, which is outside the probable error, as noted in the next paragraph. This is an indication that there is direct experimental evidence for the position of the  $5d^7$  configuration, so that a calculation with three configurations would not have to be made solely on the basis of the relatively small perturbations produced in the levels lower than 35 000 cm<sup>-1</sup>. (We did not attempt this calculation because the matrices would have been too large for SEAC.)

The error in predicting positions of levels well isolated from the observed region of the spectrum is often not realized, this being the same as the failure to distinguish between interpolation and extrapolation.<sup>11</sup> By using the full statistical matrix,<sup>9</sup> which we omit for the sake of brevity, probable errors can be estimated for any predicted level, but the qualitative behavior can be seen from the standard deviations. The errors in different parameters are correlated so that the effects of the standard deviations, which are additive, are compensated in the well observed regions of the spectrum, but the sign of the correlation will change and be additive for isolated levels. If the errors were random, the mean error of  $\pm 371$  cm<sup>-1</sup> might apply for the accuracy of the prediction of, say, the ninth level with a J value of  $3\frac{1}{2}$  (calculated position 31 000 cm<sup>-1</sup>). But in predicting positions of levels by extrapolation, as done for the ground level in this calculation,<sup>10</sup> the standard deviations of about  $1000 \text{ cm}^{-1}$  in the parameters A would be a better estimate (these two parameters enter into all the linear formulas for eigenvalues with a total coefficient of unity). For this reason, a single isolated level can often be well fitted when it is included in the calculation by least squares without making much change in the mean error, so that this is not a sensitive way to establish the identity of isolated unassigned levels. However, when the level is included, the standard deviations of certain of the parameters will be greatly reduced. Checks on the internal consistency of the parameters, such as those discussed in the next paragraph, then become more critical and provide a better way to establish the identity.

Because B, C, and  $\zeta$  were adjusted independently in the  $5d^56s^2$  and  $5d^66s$  configurations, it is possible to check the internal consistency of the calculation from these three pairs of similarly defined parameters. In each pair, the parameter evaluated in the  $5d^56s^2$  configuration is expected to be larger by 5 to 15%, from analogy with results obtained in first and second longperiod spectra. Since errors in different parameters are correlated, this is more than a check on B, C, and  $\zeta$ individually. When the L(L+1) correction is included, all three pairs of parameters behave as expected. The standard deviations are rather large, however, and there would not necessarily have been an inconsistency if the parameter in the  $5d^56s^2$  configuration had been slightly smaller in one of the three pairs. This happens for the parameter B in the calculation with the L(L+1)correction omitted, but the standard deviations are larger in this calculation, approximately in the ratio of the mean errors, so that this is not improbable from the statistics. However, the parameter  $C(d^5s^2)$  changes by 240 cm<sup>-1</sup> in the two calculations, and a change in excess of 150 cm<sup>-1</sup> is unlikely [i.e., the standard deviation in C when the L(L+1) correction is omitted. This discrepancy is present because it is a particularly poor approximation to regard the errors as random when the L(L+1) correction is neglected. For both  $d^{6}s$  and  $d^{5}s^{2}$ , the parameter C will be systematically too large when the correction is omitted.<sup>12</sup>

Very roughly, calculated g values agree with the observed values to  $\pm 0.03$ , or to about 1% of the range of the Landé g values for the levels in pure LS-coupling. The mean error for the calculation of energy levels is about 1% of the observed range of levels, so a crude correspondence can be established between the accuracy in the two calculations. The somewhat arbitrary figure of  $\pm 0.03$  is regarded as an upper limit, but based on the consideration that larger errors are always found in pairs (or triples) of relatively close levels, and in these pairs the observed and calculated g sums check to this accuracy, except for three instances where the g value in a nearby level is not observed. Since the sensitivity of calculated g values to the parameters varies greatly for different levels, it is difficult to evaluate the significance of either very good or very bad agreement in

<sup>&</sup>lt;sup>11</sup> Extrapolations made by D. S. Bowman [Phys. Rev. **59**, 386 (1941)] and R. E. Trees [Phys. Rev. **83**, 756 (1951)] predict essentially the same positions for levels of the  $3d^34s$  configuration in Mn II. Their mean error is about  $\pm 250$  cm<sup>-1</sup>, but the error in the predictions increases regularly, the higher the level, and has a magnitude of about 3000 cm<sup>-1</sup> for the highest levels, as shown by a recent extension of the experimental analysis made by L. Iglesias [J. Opt. Soc. Am. **46**, 449 (1956)]. This may exceed the error normally expected in extrapolation, because eight observations were fitted with five parameters (in Trees' calculation) and statistical assumptions would not apply well with so few observations.

 $<sup>^{12}\,\</sup>rm R.$  E. Trees, Phys. Rev. 84, 1089 (1951), Table I; 85, 382 (1952), references 5 and 9.

Total	$b^4P$	$30132\\b^4D$	$27 \ 385 \\ a^2S$	19 758 a <sup>4</sup> D	17 238 a <sup>6</sup> D	$15 \ 166 \ a^4 P$	Level Desig.	LS position
97.0	5.1	1.2	1.1	2.2	71.4-	16.0-	( <sup>5</sup> D) <sup>6</sup> D	18 936
94.4	0.6	2.2	8.2	26.5 -	9.4	47.5-	`´4P	19 380
80 5	2.8	0.4-	0.2	56.2	11.9	18.0-	$^{4}D$	20 868
54.4	8.6-	38.6-	1.7	3.8		1.7 -	$(^{3}D)^{4}D$	32 916
67.6	174-	6.0	36.3	2.9	0.5-	4.5	<sup>2</sup> S	33 590
61.9	10.8-	38.1	5.0-	5.6	1.2	1.2-	$({}^{5}D){}^{4}D$	33 593
47.3	37.0	1.6	010	1.7	0.3	6.7	$(\dot{4}^{3}P)^{4}P$	34 347
23.9	12.8-	01-	4.0-		4.9-	2.1-	$(2^3P)^4P$	39 609
33.9	14-	2.3-	28.0 -	0.6-		1.6-	$(4^{1}S)^{2}S$	38 266
8.3	1.8	0.2	5.9	0.0	0.1	0.3	$(0^{1}S)^{2}S$	70 075
67	1.0	5.2-	1.5				$(4^3P)^2P$	42 142
10.4	04-	42	53-	0.3	0.1	0.1-	$(2^3P)^2P$	49 153
4.5	1.2-		2.7	0.2-	0.1-	0.3	$\tilde{2P}$	48 952

TABLE II. Composition of the states with J = 1/2.

TABLE III. Composition of the states with J=3/2.

And the second se		Contract of the second s	The second se	the rest of the second second second	the second s	A THE R P WAS A PROPERTY OF A		the state of the s				
LS position	Level Desig.	$13 826 \\ a^4 P$	16 328 a <sup>6</sup> D	20 482 a <sup>4</sup> D	$22\ 423\ a^2D$	26 132 a <sup>4</sup> F	27828 $b^4D$	$30\ 527\\b^4F$	${31\ 461\ b^4 P}$	$b^2D$	c⁴D	Total
18 232	$({}^{5}D){}^{6}D$	6.5	79.9	1.7	1.7	0.1-	2.5	2.2	3.1-	0.4	0.3	98.4
19 380	$^{4}P$	40.2	1.9-	40.1 -	5.7 - 1	0.3	3.2	2.2	2.0 -	0.4	0.1-	96.1
20 869	$^{4}D$	34.6	6.3-	38.2	2.0 -	0.3 -	1.4 -	0.1	2.6-	0.3-	6.9	92.7
26 361	$5^2D$	2.7	0.7 -	3.4 -	41.2	12.7	4.3 -	2.0 -	0.1 -	1.1	2.8	71.8
52 911	$1^2D$	1.0	0.2 -	1.6-	11.5		3.6-	0.3	0.3	3.2	3.4	25.1
27 525	4F	1.9	0.8 -	1.1	20.1	46.4	1.7 -	13.6	0.9	3.4-	7.5-	97.4
30 016	$(4^{3}F)^{4}F$		0.5 -	0.2	3.8	17.3	24.5	25.6	0.9	3.4	5.7	81.9
44 102	$(2^3F)^4F$					1.9-	1.0 -	5.0		5.1 -	0.1-	13.1
31 027	$(4^{3}P)^{4}P$	5.1 -	1.8 -	3.2	0.9	0.7 -	0.8	11.8	20.9 -	6.9	3.5	55.6
41 755	$(2^3P)^4P$	1.0	4.6			0.2	0.3 -	2.9-	16.6	7.8-	1.8	35.2
32 536	`(⁵D)́4D	2.2	1.2 -	5.2	0.1	3.7-	11.9	1.5	41.3	1.5	0.6	69.2
33 112	$(^{3}D)^{4}D$	3.0	0.6	2.4	0.8 -	0.3	13.2 -	14.1-		27.8	26.0 -	88.2
35 269	$(\dot{4}^{3}P)^{2}P$		0.3	0.1 -	0.7-	2.5	7.8-	5.0 -	0.2 -	3.2-	11.3	31.1
53 677	$(2^{3}P)^{2}P$		0.2-	0.6	0.3	1.8-	9.0	3.5	0.4	3.8	10.4 -	30.0
37 256	$3^2D$	1.0-	0.3	1.8 -	0.5 -	3.3	0.7-	5.3	6.6	26.9	13.3	59.7
39 748	$(4^{1}D)^{2}D$		0.2-	0.1	1.4		8.9	3.3-	3.6-	2.8 -	1.9-	22.2
59 240	$(2^1D)^2D$				0.2 -		4.2 -		0.3	0.3		5.0
41 976	$(^{3}D)^{2}D$	0.5	0.2-		8.9	8.3-	0.4	1.5-		1.7-	2.9-	24.4
48 953	`_2P			0.3-		0.2	0.7-	0.1-		0.1-	1.4	2.8

TABLE IV. Composition of the states with J = 5/2.

LS position	Level Desig.	0 a <sup>6</sup> S	$11 584 a^4 P$	14 621 a4G	$15\ 770\ a^6D$	19 458 a4D	$23\ 155\ a^2D$	$24 \ 425 \\ a^{2}F$	26 661 b4P	$28 \ 030 \\ b^{4}F$	$29 \ 800 \\ a^4 F$	31 186 b <sup>4</sup> D	${32 \ 435 \atop b^2 F}$	33 282 b4G	Total
2586	<sup>6</sup> S	87.4	6.7	1.1	0.7-	1.8-	0.8-	0.4	0.7	0.1					99.7
$17\ 058$	$({}^{5}D){}^{6}D$		1.4	23.9	60.0	0.9	5.1		3.1	3.0	1.0-	0.6	0.3 -		99.2
18 315	${}^{4}G$		13.1	40.2-	14.4	4.2	12.6—	2.0	0.6	0.8	0.6 -	1.3	4.4	3.5	97.7
19 380	$^{4}P$	10.8 -	18.0	5.7	4.3 -	16.4 -	10.9 —	7.1	13.9	2.1	1.1-	0.4 -	1.5 -	0.4	92.6
20 869	4D	0.6	19.2	5.8	9.7—	37.0		5.6 -	3.3	0.3-	0.2	0.7-	0.1 -	1.0 -	83.5
$25\ 241$	$(4^{3}P)^{4}P$	0.3	3.0-	4.7 —	0.7—	1.8	12.1	1.6—	23.9	3.8		0.3	0.1	2.3	54.6
45584	$(2^{3}P)^{4}P$		0.8	2.8	2.0	0.2 -	5.3 -	0.2	21.9-	2.5 -	0.3	0.1	2.1	0.4 -	38.5
26 361	$5^2D$	0.5 -	10.2	0.2 -	0.3 -	15.7 -	14.7	2.4 -	0.6 -		1.1	6.3	10.3	0.1	62.5
52 911	$1^2D$	0.3 -	4.8		0.2 -	4.4 -	4.6	0.1-	0.1 -		0.8	0.2	0.9	0.1	16.4
27 188	$3^2F$	0.1	11.9 -	9.9	1.7 -	1.3	7.7 —	9.8	1.6	1.8	3.3	19.6	19.1	1.3	89.1
27 525	$^{4}F$		3.6	0.2 -	0.5 -	4.8	8.4	20.5	5.0-	11.5	34.2	0.1-	3.4-	0.3	92.5
30 112	$(4^{3}F)^{4}F$			0.1-	1.2 -	0.4	1.0	4.5	7.2 -	24.8	19.6—	0.2 -	1.9	15.7 —	76.5
43 680	$(2^{3}F)^{4}F$				0.4	0.1 -			2.3	4.0-	4.6	2.2	2.8 -	0.3	16.7
30 776	$({}^{5}D){}^{4}D$		1.5	0.5	1.7 -	3.8	0.4 -	5.2 -	7.2 -	0.9	6.1—	33.0	6.4 -	8.2	74.9
31 401	$5^2F$		0.4 -		0.1	1.1-	8.3-	25.2-	0.2	23.6	6.4		0.7	4.5 -	70.5
32 905	$({}^{3}G){}^{4}G$		0.4 -	1.6	0.9-	0.1 -	1.3		2.3 -	4.2	7.3—	11.2 -	0.3	45.0	74.6
33 438	$(^{3}D)^{4}D$		1.5	2.4	0.3	2.3	0.4 -	0.9-	1.8 -	1.1-	0.8	12.0 -	35.1		58.5
37 256	$3^2D$		1.2 -	0.2 -	0.6	2.2-	1.9 -	0.4 -	0.1 -	6.0	3.7	1.3 -	2.4 -	0.9	21.0
39 028	$(4^{3}F)^{2}F$			~ .		0.1	0.8	8.6	0.7	5.7 -	1.3-	0.6	0.4	2.2 -	20.5
52 558	$(2^{3}F)^{2}F$		0.1 -	0.1	0.1 -		0.3 -			1.7		0.2		1.5	4.0
39 748	$(4^{1}D)^{2}D$			0.1-	0.2 -		1.2	1.0	1.0	0.2	6.0 -	0.9	0.1	5.4—	16.1
59 240	$(2^{1}D)^{2}D$		0.2	0.1		~ <b>^</b>		0.4 -	0.7 —	0.7 -	1.2	0.2-	1.3	1.0	5.8
42 896	$({}^{1}F){}^{2}F$		0.1 -	0.2	~ 4	0.2		2.8		0.2 -	0.3 -		1.5	6.0	11.4
44 628	$(^{s}D)^{2}D$		1.5		0.1-	1.1-	2.0	1.2 -	1.6-	0.8 -	0.3	8.7	5.1		22.2

LS position	Level Desig.	14 217 a <sup>6</sup> D	15 058 a4G	17 331 a <sup>4</sup> D	$21\ 775 \\ a^2 F$	24 724 b4F	$27\ 141\ b^2 F$	28 542 a <sup>2</sup> G	$\begin{array}{c} 28\ 810\\ a^4H \end{array}$	$a^4F$	31 983 b <sup>4</sup> D	33 824 b4G	c⁴D	Total
15 415	( <sup>5</sup> D) <sup>6</sup> D	93.3	0.4-	0.2		2.1	0.3-	0.3-	1.3			0.1-	1.5-	99.5
18 315	`4 <i>G</i>	0.4	48.7	32.3 -	1.6	2.6	3.5	1.8 -	1.3 -	3.5	0.3	2.8	0.4 -	99.2
20 869	$^4D$	0.1	24.5	45.5		5.3-	0.3	1.4	6.0-	0.3	4.8	0.2	10.9 —	99.3
27 188	$3^2F$	0.1	12.3 -	0.4	17.5	1.5	25.8	10.2 -	3.6-	0.2 -	18.9	0.3	0.9-	91.7
27 525	${}^4F$		6.6	1.3	24.7	1.0	12.3 -		5.8	33.4 –	2.4	0.9	8.6	97.0
28 312	$({}^{5}D){}^{4}D$	0.4 -	3.5	10.8	1.5 -	5.0	19.7	11.8—	6.7	1.3-	29.0-	0.1 -	1.1	90.9
29 645	$(4^{3}F)^{4}F$	2.7 -	0.2	0.7	1.6	10.6	7.1—	3.5-	14.4	8.4	2.9	17.3—	11.6—	81.0
43 690	$(2^{3}F)^{4}F$	1.2		0.5 -		4.4-	1.6	0.5	3.2 -	9.4—	0.7 —	0.9	0.2 -	22.6
29 941	$5^2G$		0.1-	0.2	10.3	9.9	7.9	51.1	0.4	0.8	1.9 -	0.3		82.9
30 868	$(^{3}H)^{4}H$			0.1-	0.2	14.2 -	5.6		45.9	1.6	4.4	6.2		78.2
31 401	$5^2F$				27.4 —	5.0	2.4	4.8	1.9	2.1 -	24.8	0.2	0.2	68.8
32 084	$({}^{3}G){}^{4}G$	0.2-	1.8—	2.9	0.1-	8.0	9.7 —	1.0 -	0.1-	9.3	0.2-	60.0		93.3
33 894	$(^{3}D)^{4}D$	1.3	1.0	3.7		0.1 -	0.1	0.1 -	0.8-	17.8	3.5	3.0-	61.6	93.0
37 975	$(4^{3}F)^{2}F$		0.3 -	0.3 -	7.4	6.9—	0.7 —		1.4 -	3.6	0.4 -	0.7 —	2.1 -	23.8
52 699	$(2^{3}F)^{2}F$	0.1 -		0.3	0.1	3.7	1.8	1.9—		1.2 -	0.1	1.2	0.1-	10.5
41 183	$(^{3}G)^{2}G$				1.7	0.2 -	0.6	5.5	1.8	3.4	0.6-	0.8-	0.1 -	14.7
41 680	$3^2G$		0.2 -	0.1	1.9	1.0 -		1.3	0.6	1.2	4.9-	0.1		11.3
42 896	$({}^{1}F){}^{2}F$		0.2 -	0.2	3.6		0.5	1.9-	0.8-	2.0		0.2	0.4	9.8
43 842	$(4^{1}G)^{2}G$			0.2		12.3		2.7	2.1 -			0.2 -		17.5
50 492	$(2^{1}G)^{2}G$				0.2	6.0-	0.2	0.1-	1.8	0.4	0.1 -	4.6		13.4

TABLE V. Composition of the states with J = 7/2.

TABLE VI. Composition of the states with J=9/2.

LS position	Level Desig.	11 754 a <sup>6</sup> D	16 619 a4G	22 160 a <sup>2</sup> G	$\begin{array}{c} 27 \ 161 \\ a^4 H \end{array}$	$27 514 a^{4}F$	$30\ 645\\b^4F$	31 399 a <sup>2</sup> H	34 194 b4G	$b^2G$	Total
13 303	(5D)6D	90.7		3.6	1.9	0.2	2.7		0.3		99.4
18 315	4 <u>G</u>		84.6-	2.8 -	5.9	1.0-	0.3	0.4 -	4.5-	0.2	99.7
27 525	${}^{4}\overline{F}$	0.2	5.3-	18.4	0.2	47.7	13.8 -	8.8	0.4		94.8
28 800	$(4^{3}F)^{4}F$	5.1 -		9.8	11.2	3.4	23.5	0.2	10.4	0.2	63.8
43 949	$(2^3F)^4F$	3.3		4.0 -	5.4 -		11.9 -			1.3-	25.9
29 886	`2H		3.1 -	4.2	9.2 -	25.8 -	0.3	46.6	8.0	0.2 -	97.4
29 941	$5^2G$		1.4-	31.4	20.0 -	2.9 -	1.7 -	13.9-	4.8 -		76.1
30 164	$(^{3}H)^{4}H$		0.1 -	2.4 -	24.5 -	10.7	38.7	1.6	11.4 -	8.1-	97.5
31 027	$(^{3}G)^{4}G$	0.2 -	4.9	7.1	16.0	2.7 -	0.1 -	11.4	52.6-	0.3-	95.3
39 070	$({}^3G){}^2G$			2.7	2.9 -	0.1 -	3.3	3.1-	0.8	43.3	56.2
39 264	$({}^{3}H)^{2}H$			1.2 -	0.1	1.6	2.0 -		0.2 -	0.9-	6.0
41 680	$3^2G$		0.4 -	0.7 -	0.1 -	0.6		2.4	0.1 -	33.2	37.5
43 842	$(4^1G)^2G$	0.2 -		9.3	0.5	2.7 -	0.7	8.5-		6.2 -	28.1
50 492	$(2^1G)^2G$	0.1		2.5-	2.0 -	0.6	0.9-	3.1	6.5-	5.9	21.6

TABLE VII. Composition of the states with J = 11/2.

LS position	Level Desig.	16 307 a4G	$23 956 a^{2}I$	27 244 a <sup>4</sup> H	${30\ 560\ b^4G}$	$\begin{array}{c} 33 \ 318 \\ a^2 H \end{array}$	$b^2I$	$b^2H$
18 315	4 <i>G</i>	82.3-	5.7-	3.6	0.1	8.2	0.1-	
25 928	2I	2.0-	70.4	2.8 -	7.6 -	13.5	3.7	
29 304	$(^{3}H)^{4}H$	0.2 -	0.7-	61.8-	22.0	8.8	6.3	0.1-
29 736	$(^{3}G)^{4}G$	5.2	3.1	30.8	33.9	22.3	3.9	0.7
29 886	$^{2}H$	10.3-	19.3	0.1	14.3	44.9-	9.5	1.5
37 543	$(^{3}H)^{2}H$		0.4 -		14.4-	1.8	33.8	49.5
37 843	$(1I)^2I$		0.4-	0.9-	7.7	0.4 -	42.5-	48.1

individual instances. On an absolute basis, the calculated g values are in exceptionally poor agreement with experiment for the three lowest levels with a J value of one half. But the eigenvectors show that the range of Landé g values is several times as great for these three levels as it is in other close levels, so that the disagreement is not exceptional on a percentage basis.

The percentage compositions of the eigenvectors are given in Tables II to VIII inclusive. In the left-hand column, "LS-position," are given the positions of the levels as they would be in pure LS coupling with configuration interaction and second-order spin-orbit

TABLE VIII. Composition of the states with J = 13/2.

LS position	Level Desig.	$26 348 a^{2}I$	$27\ 130\ a^{4}H$	$b^2I$
25 927	$^{2}I$	96.6	2.7	0.7
28 287	$(^{3}H)^{4}H$	1.9-	91.1	7.0-
37 843	$(1I)^2I$	1.5-	6.2	92.3

		Assign	nent
J	Level	This paper	AEL
11/2	30 560 33 318	$b^4G a^2H$	$a^2 H \ b^4 G$
7/2	24 724 28 542 27 141 31 983	$b^4F$ $a^2G$ $b^2F$ $b^4D$	$a^2G\ a^4F\ b^4D\ b^2F$
5/2	28 030 29 800 31 186	$b^4F$ $a^4F$ $b^4D$	$a^4F$ $b^4D$ $b^4F$
3/2	30 527 31 461	$b^4F$ $b^4P$	$b^4P$ $b^4F$

TABLE IX. Alternate classifications given in AEL.

effects omitted. Excepting some of the duplicated terms, these are simply the diagonal elements of the matrices. In the second column the LS-coupling names of the vector components are given in the scheme with Racah's seniority number diagonal. The configuration  $5d^{6}6s$  or  $5d^{5}6s^{2}$  is implied by the presence or absence of a parent term, respectively. These names correspond, of course, to the diagonal matrix elements given in the left-hand column already described. The top row, "level," lists the experimentally observed levels. The second row, "designation," gives the assignments of the levels in the abbreviated LS-coupling notation customary in experimental work and used in AEL. The full assignment requires, in addition, a configuration and, possibly, a parent term; this is obtained from the second column opposite the boldface component for that level.

Some assignments may be arbitrary choices between equally good alternatives, and it is also possible that we have overlooked possibilities that would be slightly better than the ones given. No assignments will insure that every level has a high purity for the particular LS component specified, or that the latter will correspond to the dominant component of the vector, or indicate the particular vector that contains most of that component. As an example, the assignment of 27 828  $cm^{-1}$ as  $5d^6({}^5D)6s b^4D_3$  is one where all three criteria are invalid. By using the sum of the purities as a criterion, it would be slightly better to make a cyclic interchange of the three assignments  $b^4D_{\frac{2}{3}}$ ,  $b^4P_{\frac{2}{3}}$ , and  $b^4F_{\frac{2}{3}}$ ; the total purity for the three levels would be increased from 80.0%, as given, to 81.5% but the assignment of 30 527 cm<sup>-1</sup> as  $5d^6(a^3P)6s b^4P_{\frac{3}{2}}$  is then one where all three criteria are invalid. This change should not be made however, because it is desirable to consider the LS character of the levels, apart from the particular configuration and parent. When this is done, the purities of the  $b^4 F_{\frac{3}{2}}$  and  $b^4 D_{\frac{3}{2}}$  are both greater by about 13%, while the corresponding effects in the alternate assignment would be about 2%. From this viewpoint, it follows that strong configuration interactions that are present in this spectrum tend to stabilize the assignments to a greater extent than indicated by the purities. The assignment of the level<sup>13</sup> at 27 385 cm<sup>-1 13</sup> as  $5d^56s^2 a^2S_{\frac{1}{2}}$  is definitely established by the presence of strong configuration interaction, but the configuration should not be taken too seriously, since the level contains a nearly equal amount of  $5d^6(a^1S)6s^2S_{\frac{1}{2}}$ .

Complicated criteria could be set up so that the vectors would define the assignments uniquely, but this would not be worthwhile unless the probable accuracy with which the vectors represent the true composition of the levels were also calculated and taken into account. An analysis of this accuracy has not been made, but a comparison of these vectors with those obtained from the preliminary parameters indicates that there would not have been much difference in the assignments if the latter had been used. Inaccuracy would be most important in levels higher than 25 000 cm<sup>-1</sup>, since here there are eight pairs of levels where either the observed or calculated separation is less than 1000 cm<sup>-1</sup>, and in five of these pairs the observed and calculated separations differ by a factor of two or more. The neglect of the  $5d^7$  configuration may be a factor which is also important for these levels.<sup>14</sup> It was felt, therefore, that assignments for these higher levels should be regarded as indications of what would be reasonable, and that some interchanges suggested by experimental criteria should be accepted. To prevent confusion, and also indicate generally the degree of uncertainty that we consider might arise from inaccuracy in the theory, we list in Table IX the differences between the assignments in this paper and the assignments finally accepted and given in AEL.<sup>3,15</sup>

It is expected that in some instances pairs of vectors can be changed greatly by making small changes in the parameters, even when the corresponding levels are separated by more than  $1000 \text{ cm}^{-1}$ . This is the basis for considering that the interchange of assignments for the

<sup>&</sup>lt;sup>13</sup> In reference 2 this level is classified differently on the basis of a single pattern for the g value, but this pattern has since been rejected as unreliable. Large isotope and hyperfine structure splittings in this spectrum often caused trouble in interpreting the experimental data. For instance, in the early literature the level at 27 385 cm<sup>-1</sup> was regarded as two levels with a separation of 0.8 cm<sup>-1</sup>. As noted by Meggers in reference 1, many of the hyperfine structure patterns were fully resolved with a grating and it was not always possible to average these patterns out in making the line list. By increasing the tolerance to allow for this, we have found a few lines listed in reference 1 that are hyperfine structure components of classified lines given in reference 2.

<sup>&</sup>lt;sup>14</sup> Only the lowest pair of levels with  $J = \frac{7}{2}$  are as close as this below 25 000 cm<sup>-1</sup>, and this could account for the systematic behavior of the errors, rather than neglect of the  $5d^7$  configuration. This explanation indicates that the accuracy of calculations made in spectra where LS coupling is applicable, is generally better than that for calculations made in spectra with a strong breakdown of coupling. Calculations that have been made in intermediate coupling often lead to closer agreement as noted by G. Racah [Physica 16, 651 (1950)], but they have been made for spectra with relatively few allowed levels, so this agreement is not conclusive evidence.

<sup>&</sup>lt;sup>15</sup> Some changes were suggested by Charlotte E. Moore so that the g values, line intensities, or term groupings would agree better with what was expected in LS coupling.

two levels  $a^2H_{11/2}$  and  $b^4G_{11/2}$  given in AEL should be made, as indicated by the g value observed for one of them, even though the observed and calculated separations of the two levels are large and agree well (i.e., 2758 cm<sup>-1</sup> and 2842 cm<sup>-1</sup>, respectively). Our qualitative criteria for admitting the possibility of a critical dependence on the parameter values in this instance, are that the pure LS levels are nearly degenerate, as can be seen from the left-hand column in Table VII, and there is no direct interaction between the pure levels to stabilize the composition.<sup>16</sup>

<sup>16</sup> This critical dependence of eigenvectors on the parameters is E. U. Condon and G. H. Shortley, *Theory of Atomic Spectra* (Cambridge University Press, Cambridge, 1951), p. 39.

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## Photodetachment of $O_2^{-+}$

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The cross section for photodetachment of  $O_2^-$  has been measured for the range of photon energies 0.5 to 3.0 ev. No onset energy is discovered in this range but analysis of the data gives an extrapolated threshold at  $0.15\pm0.05$  ev. The curve is found to rise gradually with increasing slope, reaching a value of  $2.4\times10^{-18}$  cm<sup>2</sup> at 3.0 ev photon energy. Comparison of the data with the predictions of threshold law theory and the results of previous experiments results in an interpretation in terms of assumed potential curves for O<sub>2</sub><sup>-</sup>.

## I. INTRODUCTION

GLANCE at the history of molecular-negativeion research reveals a great many experiments which have been performed giving valuable information on their rates of production and destruction as well as their behavior under various circumstances. Most of these experiments have been necessarily of the swarm type involving mixtures of electrons, atoms, molecules, and ions with various distributions of energy. Unfortunately, the observations have seldom been unequivocally interpretable in terms of the microscopic character of the ions. The weakness of the forces which maintain negative ions causes them to be very delicate, which fact mitigates against their assembly in sufficiently high densities to permit spectroscopic research of the conventional types. Further, molecular ions have complicated structures, and this has hampered theoretical investigation of their nature. Thus, the mechanics of formation, the energy of binding, and the energy-level schemes have still not been established for any molecular negative ion.

Many different kinds of diatomic negative ions are now known to exist and among the most diligently studied of these has been the ion of oxygen,  $O_2^{-}$ . It is readily formed in gas discharges containing oxygen. The presence of small amounts of the gas in devices such as counters tubes can seriously affect their behavior by capture of electrons from the ionized plasma.<sup>1</sup> Also, the

possibility that  $O_2^-$  may form in the D region of the ionosphere makes it of interest to several fields of study.<sup>2</sup>

The configuration of electrons in the ground state of  $O_2^-$  is believed to consist of the orbitals of  $O_2$  plus an extra electron in the antibonding  $(\pi_q 2p)$  orbital.<sup>3,4</sup> Formation of the ion may take place by direct radiative attachment although the cross section for this process is expected to be very much smaller than is required to explain the copious production observed in gas discharges. Bloch and Bradbury<sup>5</sup> proposed that the electron capture leaves the ion in an excited vibrational level which is then stabilized by collision. This explanation also requires a cross section for the stabilizing collision considerably larger than is supposed likely.<sup>6</sup> It has also been suggested<sup>7</sup> that the ions may be formed in electron- or ion-exchange collisions of O2 with O-.

Published values of the binding energy of the extra electron of  $O_2^-$  range from 0.07 ev<sup>5</sup> to 0.9 ev.<sup>8</sup> The low binding energy allows the ion to be destroyed easily by collisional detachment in discharges at high E/p.<sup>9</sup> It is also possible<sup>7</sup> that they may be destroyed in discharges by charge exchange in collisions with other types of molecules. Further, they may be destroyed by the

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<sup>&</sup>lt;sup>1</sup> B. B. Rossi and H. H. Staub, Ionization Chambers and Counters (McGraw-Hill Book Company, Inc., New York, 1949), first edition, p. 29 ff.

<sup>&</sup>lt;sup>2</sup> Smith, Burch, and Branscomb, Ann. Géophys. 14, 225 (1958), <sup>3</sup> H. S. W. Massey, *Negative Ions* (Cambridge University Press, Cambridge, 1950), second edition, p. 28.

<sup>&</sup>lt;sup>4</sup> D. R. Bates and H. S. W. Massey, Trans. Roy. Soc. (London) <sup>6</sup> F. Bloch and N. B. Bradbury, Phys. Rev. 48, 689 (1935).
<sup>6</sup> H. S. W. Massey, reference 3, p. 72 ff.
<sup>7</sup> D. S. Burch and R. Geballe, Phys. Rev. 106, 183 (1957); 106,

<sup>188 (1957).</sup> <sup>8</sup> I. A. Kazarnovski, Doklady Akad. Nauk S.S.S.R. 59, 67

<sup>(1948)</sup> 

<sup>&</sup>lt;sup>9</sup> L. B. Loeb, Phys. Rev. 48, 684 (1935).